

Obituaries

Prepared by the Historical Astronomy Division

ARTHUR EDWIN COVINGTON, 1913–2001

Arthur Edwin Covington, Canada's first radio astronomer and founder of the daily 10.7-cm solar flux patrol, died peacefully in his home in Kingston, Ontario after a lengthy illness on 17 March 2001. He was eighty-eight years old. His wife Charlotte and their four children, Nancy, Eric, Alan, and Janet survive him. Covington was born in Regina and educated in Vancouver. Deeply absorbed with radio science and astronomy from his youth, Covington graduated in mathematics and physics from the University of British Columbia (UBC) in 1938. He stayed at UBC to complete a Master's thesis on lens design for electron microscopes. But in 1940 he moved to the University of California, Berkeley, to begin graduate studies in nuclear physics. He met and married fellow physics student Charlotte Anne Riche in Berkeley. In 1942, the Covingtons moved to Ottawa to join Canada's wartime research effort in radar at the Radio Branch of the National Research Council's Laboratories (NRCL).

Well aware of Jansky's and Reber's discoveries of "cosmic radio noise" at metric wavelengths, at the end of the war Covington proposed to use converted radar equipment operating at a wavelength of 10.7 cm to probe the galactic center using the Sun's decimetric emission as a calibrator for deriving the cosmic-noise spectrum. His first attempts to measure the integrated flux over the solar disk in July 1946 did not produce consistent flux levels from day to day. The sun was quite active with large sunspots, so Covington was uncertain whether his instruments, local interference, or the Sun caused the daily variations, and he struggled to interpret these inconsistencies. Charlotte Covington, his wife, pointed out that a forthcoming partial solar eclipse would be visible from Ottawa on 23 November 1946. Covington seized this chance to measure the solar flux variations as the Moon occulted a large sunspot. In that single observation, he provided the first decisive proof that dark sunspots are associated with discrete hot sources of decimetric radiation. Over the next year of uninterrupted daily measurements of the integrated flux, Covington established that the variability of the 10.7-cm flux closely matched that of the sunspot numbers. Both measures were modulated by the comings and goings of active regions and the solar rotation.

Covington's groundbreaking studies of what came to be called the "slowly varying component of solar radio emission" quickly took advantage of other wartime developments at NRCL. A unique combination of a 46 m-long slotted wave-guide and two parabolic cylinders, erected in 1951 south of Ottawa, was the first device built in Canada to detect radio emissions from discrete astronomical sources. It was one of the earliest compound microwave interferometers; its narrow fan beam enabled Covington and N. W. Broten to measure solar limb brightening and the temperatures above isolated sunspots. With W. J. Medd, Covington began a series of experiments in the early 1950's to improve the relative and absolute accuracies of the daily observations of integrated Solar radio flux. In the process, they laid the



Arthur E. Covington, 1913–2001
Photo courtesy of Eric Covington

foundation for future confidence in the daily 10.7-cm flux as an objective index of solar activity.

Over the next decade, these early successes paved the way for the NRC to establish the Algonquin Radio Observatory for galactic and solar astronomy at a remote radio-quiet site beside Lake Traverse, Ontario. Covington chose to continue with solar research for the rest of his career. He had, from the beginning, grasped the practical importance of solar-terrestrial research using the 10.7-cm flux as a proxy for ionizing solar radiation. The ever-growing list of applications for the daily 10.7-cm flux is a tribute to Covington's persistence in raising the quality of the measurements to the highest standards. The monitoring program was extended long past his retirement in 1978. The program, now in its 54th year, continues at the Dominion Radio Astrophysical Observatory near Penticton.

Covington's colleagues at NRCL honored him, shortly after he retired, by erecting a unique sundial at a site overlooking the building where he pursued his solar researches for over 25 years. It is an 0.9-m paraboloidal reflector supported on a framework of 10-cm wave-guide, with a 10-cm dipole feed at the focus as a gnomon.

Covington never forgot that his youthful passion for astronomy was nurtured among amateur enthusiasts in the Royal Astronomical Society of Canada. He contributed generously and with evident pleasure in many ways to the Ottawa Centre throughout his career.

Beneath Covington's reserved, orderly exterior there laid a quirky sense of humor and a generous spirit. He was fascinated by arcane topics as well, such as religious mysticism, especially if they had a solar connection. He was an avid collector of books, new and old, that reflected his wide-ranging interests, particularly in the history of radio science and the interrelationship of astronomy with other disciplines. He and Charlotte set up the Riche-Covington Trust at Queen's University, Kingston, to house his collection on the development of radio science in Canada. He leaves these tangible legacies, but his many friends will fondly remember him as much for his devotion to his family and to solar science.

Victor Gaizauskas

Herzberg Institute of Astrophysics (Editor's Note: A similar obituary appeared in the *Journal of the Royal Astronomical Society of Canada*.)

MERTON EDWARD DAVIES, 1917–2001

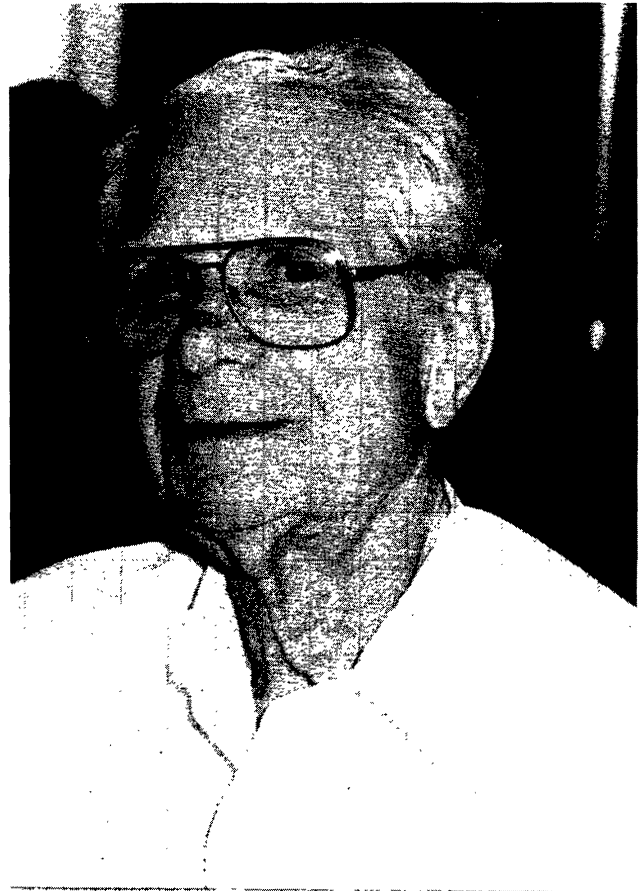
Merton E. Davies was a great friend to all who knew him. The diversity of that large group of fortunate people reflected the wide range of his professional and personal interests.

Born in St. Paul, Minnesota on 13 September 1917, Davies was the youngest of the three children of Albert Daniel and Lucille McCabe Davies. Soon after his birth, the family moved to Palo Alto, California, where Davies received an AB in mathematics from Stanford University in 1938. He taught mathematics for two years at the University of Nevada, but was then absorbed as a mathematics group leader in aviation development at Douglas Aircraft as that corporation geared up for World War II.

Mert, as he was known to everyone, joined the RAND Corporation in 1947 and remained on the RAND staff for the rest of his life. He began his career in military reconnaissance. In the early 1950s, he was a key contributor to the development of the CORONA system, which became the world's first reconnaissance satellite. The first pictures taken of Earth from orbit were reconnaissance photographs registered on a roll of film that was returned to Earth on 18 August 1960 in a CORONA capsule.

Mert was one of the ten founders of the National Reconnaissance Office, an agency of the Department of Defense. He received the George W. Goddard Award in 1966 for his "distinguished contributions to the development of photoreconnaissance." He became a valued consultant to the United States Arms Control and Disarmament Agency.

This top-secret work for the military was a prologue to Mert's highly productive career as a major player in the Golden Age of planetary exploration. He understood immediately that his pioneering work on space photography could be applied to his keen interest in the moon and planets. By 1958, just two years after the successful launch of the Soviet Sputnik, Mert had already written a report about methods for obtaining pictures of the moon from a spin-stabilized space-



Merton E. Davies, 1917–2001
Photo courtesy of M. Randel Davies

craft. When the opportunities for deep space missions arrived, he was ready to use his expertise in photogrammetry to start the systematic mapping of the planets and their satellites. This led to his participation in a long series of NASA missions, making him the only person on Earth who had made virtual visits to every planet in the solar system save Pluto, for which there were no missions in his lifetime.

Mert's contributions to these missions were instrumental to their astounding successes. Before launch, he participated in the design of the camera systems and the development of imaging strategies. When the data came in, it was Mert who established the coordinate systems for all of the target objects. The maps we have of Mercury, Venus, Mars and the satellites of the outer planets are all based on his work in establishing the point of zero longitude or the prime meridian for each object. As Bruce Murray has commented, to do so for even one such object would be a "major career achievement by any scientist," but to be credited for having done so for essentially every large solid object in the solar system except Earth and Pluto provides "an instructive lens through which to view Davies accomplishment." (EOS, 82, 46 (13 November 2001):551-552.)

Recognizing that there was no scientific society overseeing this activity, Mert became the founding chairman of the IAU/IAG Working Group on Cartographic Coordinates and Rotational Elements of the Planets and Satellites in 1976. At about this same time, he became a member of the newly