

Astron. Astrophys. Suppl. Ser. **81**, 151-186 (1989)

Intensive photometry of southern Be variables. I. Winter objects (*)

J. Cuypers ⁽¹⁾, L. A. Balona ⁽²⁾ and F. Marang ⁽²⁾

⁽¹⁾ Koninklijke Sterrenwacht van België, Ringlaan 3, B-1180 Brussel, Belgium

⁽²⁾ South African Astronomical Observatory, P.O. Box 9, Observatory 7935, Cape, South Africa

Received May 29, accepted July 20, 1989

Summary. — We present results of an intensive photometric campaign on some bright southern Be stars to search for periodic light variations. In order to obtain good phase coverage, observations were conducted from two sites with different longitude : ESO and SAAO. Most of the stars observed are indeed variable with periods close to one day (the expected rotational period for these stars). We present our results for winter objects.

Key words : Be stars — variable stars — photometry — stellar pulsations.

1. Introduction.

Periodic variability in Be stars has attracted great interest in recent years. The discovery of unexpected short-period variations in these stars has renewed the search for a viable mechanism of mass ejection. The current consensus of opinion is that non-radial oscillations, driven by an unknown mechanism, is the likely cause of the mass loss leading to the Be phenomenon (see Percy, 1987, for a recent review). This conclusion is mostly based on line-profile variations seen in some Be stars.

The small-scale “moving bumps” seen in the line profiles of some Be stars is attributed to high-order non-radial pulsations (NRP). But moving bumps are also seen in B stars without emission lines and even in δ Scuti stars (Walker and Yang, 1987). It seems that NRP is not solely responsible for the mass ejection. It is also known that one of the distinguishing features of line-profile variations in Be stars is the presence of low-order line-profile variations which some attribute to low-order NRP (Penrod, 1987). But again, low-order NRP is present in non-Be stars as well (β Cep and δ Sct). Many β Cep variables are rotating more rapidly than some Be stars. High rotation coupled with NRP does not seem to be a sufficient condition for the Be phenomenon. The information from photometric observations has not been adequately considered and it is possible that the conclusion that NRP is responsible for the mass loss could be premature.

A viable hypothesis should be able to explain the distribution of observed periods, the fraction of Be stars which are periodic, their photometric amplitudes and evolution of the shapes of their light curves. Such information can only be obtained by intensive photometric monitoring of a large sample of Be stars. Owing to the fact that periods are close to one day, it is very important to obtain observations from different longitudes. These observations should be obtained over a short time as the light curve can evolve quite rapidly. These demands can only be met by an international campaign.

In this paper we present results of such a campaign conducted from ESO and SAAO as well as some observations obtained at SAAO only. We observed 17 bright southern Be stars ; most of them turned out to be short-period variables. These stars were chosen from the *Bright Star* catalogue on the sole criterion that they should be well placed for observing. A list of candidate and comparison stars is given in table I. Results for variable program stars are given in tables IIa and IIb.

2. Observations and reductions.

The observations at SAAO were obtained with the Volks photometer attached to the 0.5-m reflector at Sutherland. These were made mostly through a single filter, Strömgren b , which was chosen as a compromise between the expected larger amplitude towards shorter wavelength and the

(*) Observations collected at the European Southern Observatory and the South African Astronomical Observatory.
Send offprint requests to : L.A. Balona.

decreased sensitivity to atmospheric extinction at longer wavelengths. Some four-colour observations were also made. For the brightest stars, a neutral density filter had to be used to prevent over-illumination of the photocathode. In these circumstances we calibrated the ND filter by observing the comparison stars with and without the filter. The rms deviation suggests that this calibration is good to within 3 millimag (the expected observational error of one measurement). Systematic errors due to the use of the neutral density filters are probably negligible owing to the narrow bandpass of Strömgren *b*.

In the reduction procedure, mean extinction coefficients were obtained for each night using the comparison stars. Corrections for transparency changes during the night were made by calculating the deviation for each comparison star. Stars in close proximity were treated as a group and mean transparency corrections were applied to the group, which included the Be stars. Comparison stars which were variable were detected at this stage and could be eliminated. We searched for micro-variability in the comparison stars by calculating their periodograms; again such stars could be identified and eliminated. Finally, the mean magnitude of all comparison stars was adjusted to be in exact agreement with the mean *b*-magnitude for the same stars observed at ESO. This zero-point correction ensured that the Be stars observed from the two sites were on the same system.

The observations at ESO were made between 13 June and 6 July 1987. They were obtained with the six-channel *uvbyH β* photometer on the Danish 0.5-m telescope. This instrument uses a grating and slots to define the passbands; in consequence some systematic differences could be expected between the *b* magnitudes of the two observatories. In practice, these differences were very small (a few millimagnitudes) and posed no problem when combining the data sets. Because a new telescope control system with autocentering and automated observations had just been installed and still had to be checked during the observing run, the quality and quantity of the observations was not as high as could be expected at the start of the run. Later on, most of the problems were solved and nowadays the new, automated system has proved to be very reliable.

For the brightest stars, the same procedure was used as at SAAO. The reduction procedure was similar except that a polynomial model for transparency changes and zero-point drift was calculated for each night. Furthermore, a general least-squares solution which included all observations for all comparison stars provided best estimates for their mean magnitudes. Variable stars among them were easily identified by the large standard deviations. They were not included in obtaining the final magnitudes for the comparison stars. A small zero-point difference between these results and the standard *uvby* system is still possible, but this is unimportant for the study of their light variations. The rms scatter for the comparison stars is in all cases less than 5 millimag.

3. Period finding.

We employed two methods of searching for periodicities. One is the standard Fourier periodogram analysis technique for unequally spaced data; the other is a modification of Stellingwerf's (1978) phase-dispersion minimization (PDM) technique. The first method is suited to the case where the underlying variation is approximately sinusoidal, but it can give misleading results when the variations are highly non-sinusoidal or when a multiple-wave light curve is present. It is known that many Be stars have a double-wave light curve which will not be detected by this method. Clearly it will be impossible to discriminate between a single-wave light curve and a double-wave light curve with nearly equal amplitudes.

The PDM technique will identify the period for all kinds of light curves, but since the general noise level is rather high in this method (Swingler, 1989), one has to be careful of spurious periodicities. Our final choice was based on a careful examination of the phase diagram resulting from the periods produced by the Fourier and PDM techniques. Our criterion was by nature rather subjective, but we attempted to choose the period which resulted in the curve with the least scatter. Except for one or two cases, there is little doubt that our choice is the correct one.

Besides the complication discussed above, a serious difficulty arises when the irregular light variations which are sometimes present in Be stars (attributed to effects of the circumstellar material) has a time scale comparable to the underlying periodic variation. In these cases it is not possible to obtain the correct period with any certainty. As some long-term activity is nearly always present, we normally removed the long-term trend by fitting a low-order polynomial to the data. This of course has the effect of reducing the scatter at the expense of some uncertainty in the shape and amplitude of the light curve. The derived period will in general be unaffected by this procedure.

It is difficult to derive reliable standard errors for the periods, but an estimate can be obtained from the half-width of the Fourier periodogram peaks. Typically, we observed for about four weeks in each season which leads to a standard deviation of about 0.02 d^{-1} in frequency which can be taken as a reliable guide to the error in the quoted frequencies. Sometimes we have combined many seasons in order to obtain a more accurate period. This is always possible to do, but invariably there will be many possible choices of frequency (each separated by approximately one cycle per year) which will produce as good a fit. In this case the real error is not reduced by combining the data.

4. The results.

v Cen

This star is a single-line spectroscopic binary (Palmer, 1906). Wilson (1914) determined the orbital elements and obtained a circular orbit with $P = 2.62516 \text{ d}$. Hendry and Bahng (1981) found H α to have double emission

surrounding an absorption core. This observation indicates that ν Cen could be classified as a Be star, but it is probable that the emission originates as a result of binary interaction rather than an intrinsic property of the star itself. Spectroscopic observations suggest that this star may be a β Cep variable with a period of about 0.17 d and an amplitude of 5 km s⁻¹ (Rajamohan, 1977 ; Kubiak and Seggewiss, 1982 ; Ashoka *et al.*, 1985). Waelkens and Rufener (1983) do not find any evidence for short period light variability, except for a variation with the same period as the binary period.

Our photometric observations (SAAO only) extend over two seasons, 1987 and 1988. The 1988 observations are particularly numerous and clearly show a variation with the same period as the orbital period (Fig. 1). Minimum light corresponds to maximum positive radial velocity. As discussed by Waelkens and Rufener (1983), the resulting light curve is most probably a reflection effect. A careful investigation did not reveal any signs of a variation at any other frequency with an amplitude exceeding 2 millimags.

μ Cen

This star is known as a "pole-on" Be star (v in $i = 155$ km s⁻¹). Ghosh *et al.* (1987) have summarized the history of H α emission ; they reported an outburst in March/April 1987, about three months before the first set of our observations. A full discussion of an earlier outburst (February 1987) is given in Baade *et al.* (1988). Thimm (private communication) reported from spectra taken on May 29, 30 and on June 1, 1987 that H α changed from pure absorption to emission in 3 days. This was only a few days before the start of our observations. All indications are that this Be Star was in a very active phase during this time.

μ Cen was one of the first Be stars in which periodic line-profile variability was found (Baade, 1984). Baade observed three or four nearly equally-spaced moving bumps. There were also low-order changes in the line profile in which the symmetry of the whole line varied, but with a longer time scale. The period of the low-order variations was estimated to be 0.505 d, five times the period of the moving bumps. Baade interpreted these results in terms of NRP with $\ell = m = 2$ and $\ell = m = 10$. In our opinion the poor time coverage places considerable doubt on the accuracy of this period which is practically equal to half a day, since the aliasing problem must be quite considerable. More recently, Baade (1987a) obtained 9 nights of line profile observations in which this period is apparently confirmed.

Since two kinds of line-profile variations are seen in this star, Baade (1987b) claims that it is multiperiodic and hence proof that the NRP interpretation is the only feasible one. Harmanec (1987a) has suggested that the short period structure in Be stars can be represented by integer ratios, in which case we are dealing not with incompatible periods but harmonics of a longer period (which he identifies as the period of rotation). Baade (1987b) has criticized this suggestion. Since the period ratio in μ Cen is 5 : 1 according

to Baade, this star is perhaps not a convincing case of multi-periodicity.

We observed μ Cen for 29 nights in June/July 1987, 25 nights in March/April 1988 and 7 nights in June 1988. The latter two data sets are SAAO observations only. The 1987 data show that the star was photometrically very active at this time. There are two major excursions in brightness, each with a light range of 0.07 mag, separated by five nights.

A Fourier periodogram analysis shows no obvious trace of coherent periodicity : the three largest peaks being at frequencies of 0.05, 0.95 and 0.16 d⁻¹. These first frequencies appear to be artifacts caused by the large brightness excursions. The PDM periodogram shows essentially the same structure except for a large peak at 0.48 d⁻¹. This does not seem to be associated with the random brightness fluctuation and is the only coherent period we can extract from our 1987 data. The resulting light curve is a double-wave with a period of 2.10 d (Fig. 2).

Since our results from 1987 did not lead to a convincing period determination, we re-observed the star from SAAO in 1988. This time the star was less active, but surprisingly the periodograms look very similar to those of 1987. In particular the frequency at 0.48 d⁻¹ was very strongly present. This seems to indicate that the 2.10 d period is physically real. As a final check we combined the data from both seasons which allowed us to refine the period to 2.1017 d. A plot of the light curve is shown in figure 2. The double-wave nature is very clear, but the large scatter indicates that there is an additional source of variation. A large increase in amplitude followed by a swift decrease can explain some of the apparent scatter in 1987. The peak-to-peak amplitude decreased from 0.08 mag in 1987 to 0.04 mag in 1988, though the large amplitude in 1987 could be mostly a result of a non-periodic brightness excursion during a few nights. But this cannot be invoked as the cause of the large scatter in 1988 ; this phenomenon is very common in the light curves of Be stars and can be termed "flickering".

In the light of these results, we cannot confirm Baade's 0.5 d period (frequency 2.0 d⁻¹) in μ Cen. We can possibly reconcile our results with this period if we note that the single-wave frequency (0.96 d⁻¹) is the one-day alias of Baade's period. We certainly did not find any indication of a 0.5 d period in our data. As a test, we analyzed Baade's (1984) radial velocity observations for the HeI line and constructed the phase diagram using a period of 2.1017 d. The resulting velocity curve has considerably greater scatter than the one constructed with 0.505 d, but is still quite reasonable and gives a double-wave curve. However, there are only 19 observations – far too few for a reliable period estimation.

η Cen

A description of recent emission-line activity in the shell star η Cen is given by Dachs *et al.* (1986). They conclude that at times the internal self-absorption by the shell is so

severe that it may obscure a certain fraction of the light from the photosphere. Mennickent and Vogt (1988) observed the star during 1987. The lower order Balmer lines showed weak variable emission wings and variable shell absorption cores.

Baade (1983) discovered significant line-profile variations with a time scale of hours. During a search for rapid spectroscopic variations, Ghosh *et al.*, (1988) discovered a continuum flux variability with a time scale of hours.

Our results clearly show large short-period variability. A nightly range in brightness of nearly 0.1 mag is not uncommon. The Fourier periodogram shows a very strong peak at 1.56 d^{-1} , but the resulting light curve has a large scatter. The PDM periodogram shows, in addition, large peaks at one-half and one-third this frequency. An examination of the phase plots show that a better fit is obtained by assuming a double-wave light curve with $f = 0.78 \text{ d}^{-1}$, but even then the scatter is large. By far the best phase plot is obtained with $f = 0.52 \text{ d}^{-1}$ which results in a triple-wave light curve. While double-wave light curves are common, only one case of a triple-wave light curve (α Eri, Balona *et al.*, 1987) is known.

This surprising result prompted us to re-observe η Cen for a further season in 1988 (SAAO observations only). Our analysis of these data showed quite convincingly that $f = 0.52 \text{ d}^{-1}$ is indeed present once again and gives a light curve with the least scatter. On this basis we feel that a triple-wave light curve with period of 1.927 d is likely to be the correct interpretation (Fig. 3). The peculiar light curve and the very large amplitude (0.14 mag in 1987, 0.10 mag in 1988) makes this star unique among the periodic Be variables. Line profile observations of η Cen are potentially of great importance in view of the large amplitude. If NRP is involved, the geometrical distortions and/or temperature variation must be exceedingly severe to give rise to such a large amplitude for an $\ell = |m| = 3$ mode.

HD 137518

This star was included in our program at the suggestion of C. Waelkens who found it to be a large-amplitude variable. Very little is known about this star. It has been classified as a luminous blue supergiant (B5Ia) by McConnell and Bidelman (1976). On the other hand, Garrison *et al.*, (1977) obtain a classification B1IIInep with a comment that the HeI lines are strongly veiled and the suspicion that it is a double-line spectroscopic binary. In the *Michigan Spectral Catalogue* (Houk, 1978), it is given the classification B1/2(I/IIIn), while Mermilliod (1987) tabulates it as O9Vn.

The photometric variability was discovered by Strohmeier *et al.* (1964) who found a range of 0.5 mag from sky patrol plates of the Bamberg Southern Station. The variability was later confirmed by Kozok (1985).

Our 1987 observations show a light range of 0.5 mag, but we could not find a definite periodicity. The Fourier periodogram shows a strong peak at $f = 0.11 \text{ d}^{-1}$ or its

one-day alias at 0.89 d^{-1} . Smaller peaks are present at $f = 0.36$ and 1.29 d^{-1} . The PDM periodogram is similar except that $f = 0.12 \text{ d}^{-1}$ is by far the strongest. We examined the light curve at many different frequencies given by the Fourier and PDM periodograms, but no single choice produced a reasonable light curve. We could assume multiperiodicity of course. If we prewhiten the data by removing a sinusoid with $f_1 = 0.115 \text{ d}^{-1}$, we obtain a noisy Fourier periodogram in which the highest peak is at $f_2 = 0.55 \text{ d}^{-1}$. A multiperiodic Fourier fit with f_1 and f_2 leads to a rms scatter of 0.08 mag which is clearly unacceptable. Removing these two frequencies leads to a strong peak at $f_3 = 0.35 \text{ d}^{-1}$, which is the third harmonic of f_1 . A Fourier fit with f_1 , f_2 and f_3 gives an rms scatter of 0.044 which is still very much higher than the expected observational error. Further prewhitening just gives a periodogram in which the noise level steadily increases towards low frequencies.

It is clear that no direct evidence for coherent multiperiodicity exists as even a solution with three frequencies is quite insufficient to describe the data. A solution can always be found by including more and more frequencies until the resulting rms scatter is sufficiently small, but there is no guarantee that the resulting solution will bear any resemblance to the true physics of the problem. We prefer to believe that the only likely periodicity in HD137518 is $P = 8.70 \text{ d}$ ($f = 0.115 \text{ d}^{-1}$) and its harmonics (Fig. 4). Prewhitening by a Fourier curve leads to a further possible periodicity of 0.685 d, but the reality of this period is far from convincing. The residual variability can be attributed to random fluctuations.

This star is a very puzzling object. Without detailed spectroscopic observations it seems impossible to determine its nature. The large light amplitude suggests some type of eclipsing phenomenon.

κ^1 Aps

Slettebak (1982) describes how two spectra, taken one day apart, show striking differences in the widths of the absorption lines (from $v \sin i = 250 \text{ km s}^{-1}$ to 350 km s^{-1}), suggesting that this star may be a double-line spectroscopic binary. Mennickent and Vogt (1988) observed this star during 1988 and found the spectrum characterized by narrow absorption cores flanked by weak emission at H β .

The Fourier periodogram of this star shows it to be another periodic variable with a frequency of 1.61 d^{-1} . The PDM periodogram also gives a strong signal at half this frequency, indicating a double-wave light curve. However, the phase plot shows that this effect arises from only five data points which deepens the second minimum; apart from this there is no strong reason to believe that the light curve is double-wave.

Observations at SAAO during 1988 give much the same results, but this time the double-wave nature is more convincing. This time the amplitude difference in the two waves is based on more observations and there are also distinct differences in their shapes. By combining the data

from the two seasons we find the best double-wave period to be 1.238 d (Fig. 5). There has been relatively little change in overall amplitude between the two seasons.

η^1 TrA

This star turned out to be one of the few constant Be stars in our sample. No sign of periodicity is evident in either the Fourier or PDM periodograms. The rms scatter of all observations is 3 millimag – the value expected for a constant star. It was not observed in 1988.

48 Lib

This star has a long history of documented spectroscopic observations. It is well known for the long-lasting presence of numerous sharp shell absorption lines and for strong regular variations of Balmer emission line profiles and radial velocities of shell absorption lines with a quasi-period of about ten years (cf., e.g. Aydin and Faraggiana, 1978). Dachs *et al.* (1986) has described the emission-line history in recent years. Mennickent and Vogt (1988) found the spectrum of 48 Lib in 1987 to be characterized by many faint FeII shell lines, some of them surrounded by emission. H β and H γ had deep central absorption cores with emission flanks ($V < R$). The V/R ratio was variable.

The 1987 data show a very strong peak at 2.49 d $^{-1}$ in the Fourier periodogram. The PDM periodogram shows, in addition, smaller peaks at one-half and one-third of this frequency. Examination of the phase plots gives us no reason to suspect that any of the lower frequencies are physically real. There is no appreciable reduction in scatter or differences in the waves. The phase plot for the single-wave period of 0.40 d shows a rather asymmetrical light curve with a gradual decline and sharp rise (Fig. 6).

To confirm the period, we re-observed the star during 1988. We obtained the same results. The best period from the combined data is 0.4017 d. This is the shortest period we have found among the periodic Be stars.

Ringuelet-Kaswalder (1963) reported periodic radial velocity variations with $P = 0.1154295$ d suggesting a pulsation similar to that of the β Cep variables. Such a variation has never been confirmed, though rapid changes in spectral appearance during one night is not unknown (Aydin and Faraggiana, 1978). We phased Ringuelet-Kaswalder's data on our photometric period, but were unable to produce a satisfactory velocity curve.

χ Oph

Recent variations in the emission lines in this well-observed Be star are discussed by Dachs *et al.* (1986). The considerable number of radial velocities in the literature were analyzed by Harmanec (1987b) who tentatively proposed that χ Oph is a spectroscopic binary with a period of 34.12 d in a highly eccentric orbit ($e = 0.699$).

Balona and Engelbrecht (1987) observed this star photometrically during 1985 and suggested a possible period of 0.935 d. The available data were insufficient to discriminate between this period and its one-day alias at 14.3 d.

The shorter period was considered the more likely owing to the considerable brightness variations during the course of a night. In neither case is the resulting light curve particularly convincing.

Our 1987 data show much the same behaviour. Again, the strongest frequencies in both the Fourier and PDM periodograms occur at either 0.935 d or 14.3 d, but this time the longer period has a much higher peak. The resulting phase plot shows much less scatter for 14.3 d than for 0.935 d. Thus, by combining data from two separate sites we eliminate the alias problem and conclude that the true period is 14.3 d. Combining the 1985 and 1987 data yields a best period of 13.774 d (Fig. 7). A small number of observations were obtained in 1988.

This is the longest period so far detected for a periodic Be star. The question arises as to whether it represents an orbital variation. We looked for, but could not find, evidence for a radial velocity variation with this period from the data collected by Harmanec (1987b).

In spite of the fact that the 13.774 d period fits the photometric data, it is clear that there are variations on a much shorter time scale. However, these cannot be periodic as they would leave a signature in the periodogram. We are forced to conclude that this is one more example of the "flickering" phenomenon so common amongst these stars.

ζ Oph

ζ Oph (O9.5Ve) is one of the most rapidly rotating stars known. Emission lines at H α and HeI were observed between July 1973 and April 1974 (Irvine, 1974; Niemela and Mendez, 1974; Barker and Brown, 1974). During 1979 there was no hint of emission at H α , but by March 1980 H α appeared as a well-developed double-emission feature quite similar to that observed during 1974 (Ebbets 1981).

Walker *et al.* (1979) were the first to discover the by now well-known phenomenon of "moving bumps" in the line profile of B stars. It was in this star, ζ Oph, that they were first detected. At that time they were interpreted as nonuniformities in the stellar photosphere carried across the line of sight by rotation. In this way a rotational period of 21.7 hours was deduced, leading to an estimated $v \sin i = 560$ km s $^{-1}$. This value is considerably higher than observed ($v \sin i = 370$ km s $^{-1}$), but it is sensitive to the adopted stellar radius. Harmanec (1989) has shown that a somewhat smaller radius, which is still consistent with the physical parameters for this star, will bring the expected projected rotational velocity into good agreement with the observed value.

Walker *et al.* (1981) presented further observations of this new phenomenon and proposed an alternative model in which the moving bumps arise from obscurations in the circumstellar material. They also suggested NRP as a possible explanation.

Vogt and Penrod (1983) obtained extensive high-resolution spectra of ζ Oph and confirmed the moving bumps seen

by the earlier workers. They rejected the interpretation in terms of rotational modulation or obscuration by circumstellar material. While these models give rise to moving bumps in agreement with the observations, the expected light variations are not seen. They favour the NRP model. Harmanec (1989) has criticized this conclusion as it is based on only one night of photometry which does indeed show some variability. Nevertheless, the NRP model has been accepted by most groups as the most probable explanation in this and other Be stars.

It is clear that intensive photometric observations should shed considerable light on the nature of moving bumps. According to Vogt and Penrod (1983) the NRP interpretation should give rise to periodic light variations of amplitude 0.02 mag with a period of a few hours. Such periodic variations are easily detectable.

We started our photometric campaign on this star in 1985. The results of this first season were briefly reported in Balona and Engelbrecht (1987). During 1985 ζ Oph showed a clear indication of pulsation with a period of 0.193 d and amplitude of 0.02 mag. There was also a longer period of 1.075 d with amplitude 0.03 mag (Fig. 8a). The short-period pulsations were interpreted as NRP, but not as a new phenomenon. Since ζ Oph lies on the hot end of the β Cep instability strip, it seemed reasonable to classify the star as a β Cep variable. However, the presence of the longer period in a β Cep star is most unusual ; rather it is indicative of a periodic Be star. It is also of particular interest to note that if the light curve is phased on the long period of 1.075 d the short-period variation is easily visible as a sinusoidal modulation of the light curve (Fig. 8b). It looks very much as if the short period is practically one-sixth of the long period to within observational errors. This seems to confirm Harmanec's (1987a) suspicion that whenever a short-period is present it turns out to be a sub-multiple of the rotation period.

We observed ζ Oph during 1987 expecting to confirm the provisional results announced by Balona and Engelbrecht (1987). The Fourier and PDM periodograms both show some power at frequencies of 1.00 or 2.00 d^{-1} with a peak-to-peak amplitude of 0.01 mag. There is no sign of the short period variation seen in 1985. While the long period is close to the value seen in 1985, it is practically the same as one cycle per day and, as such, must give rise to a grave suspicion that this is an artifact. Indeed, since we had to observe ζ Oph at SAAO using a neutral density filter it would not be entirely surprising if some slight systematic error was present in the data. Under these conditions we can understand the results of the periodogram and we do not consider the one day period to be real. Indeed, the rms error of one observation from both sites is only 4 millimags., i.e. close to the expected rms error of a constant star. We conclude that ζ Oph was constant in light during June/July 1987.

Some observations at SAAO during early 1988 again

indicated a constant light. But in June of that year a dramatic change took place : our results of four photometric nights show marked photometric activity with night-to-night variations exceeding 0.01 mag. It would be very interesting to discover if this is associated with the development of emission at $H\alpha$. Further photometric monitoring is planned for 1989.

ι Ara

Our photometric observations of this star now cover four seasons (1985 - 1988). The 1985 light variations were particularly difficult to interpret in spite of the fact that there were no long-term trends or other difficulties. The problem is the very large flickering associated with this star which tends to mask the underlying variation almost completely. Balona and Engelbrecht (1987) obtained a double-wave light curve with period 0.515 d as the most likely solution for this season. During 1986 the amplitude appeared to increase steadily during the observing period of one month, but a period of 0.56 d could be extracted. This is the same as the period obtained in 1985 within the observational error and results in a single-wave light curve.

Our 1987 data showed that once more the star was active with a short time scale. The strongest peak in the Fourier periodogram occurs at 0.265 d, but the PDM periodogram indicates that a double-wave solution is better as the strongest peak occurs at 0.53 d. The 1988 observations are too few to derive a period, but a solution with $P = 0.56$ d fits the data well.

The conclusion is that in spite of the low signal-to-noise ratio, a period close to 0.55 d is derived independently from each observing season. On this account we feel confident that this is indeed the correct period. Combining all the data gives a best solution of $P = 0.5565$ d (Fig. 9). The flickering in this star is one of the largest we have encountered. As a consequence the periodicity in this star is somewhat blurred, but periodicity implies coherence over many cycles and this period was the only one where such a coherence was observed. The rms deviation from the light curve is as large as 0.02 mag though the amplitude is as large as 0.10 mag.

α Ara

Mennickent and Vogt (1988) found the spectrum in 1986-7 to show double emission in $H\beta$ and $H\gamma$ ($V \simeq R$) with small variations in the V/R ratio on a short time scale.

This star was immediately recognized as a large-amplitude short-period variable during the first observing run in 1985. The Fourier periodogram shows strong peaks at frequencies of 2.04 or 3.04 d^{-1} and its one day aliases. These two frequencies are also the strongest in the PDM periodogram, but the double-wave solution at 1.52 d^{-1} is nearly as strong. An examination of the light curve does indicate that there is a significant difference in amplitude between the two waves in the double-wave solution. Thus we adopted a period of 0.658 d ($f = 1.52$ d^{-1} , Balona and Engelbrecht, 1987).

In 1986 the highest peak in the periodograms occurs at

2.04 d^{-1} ; the solution for 1985 thus appears to be a one-day alias of the true period. That a wrong choice was made is not surprising in view of the closeness of the frequency to a half a day and the severe aliasing problem.

Our 1987 data shows that the ambiguity is completely resolved thanks to the different longitudes of the observing sites. The true period is either a single-wave with $P = 0.49 \text{ d}$ or a double-wave with $P = 0.98 \text{ d}$. It is very difficult to be sure of the double-wave solution since the two waves are of nearly equal amplitude, but it is a distinct possibility. Unfortunately, the period is so close to one day that we cannot use the other observations (all made at SAAO) to settle this problem. By combining all the data we find a best period of 0.9807 d for the double-wave solution (Fig. 10). Results from 1988 indicate that this period fits the data well.

66 Oph

During 1987 this star showed double emission at $\text{H}\beta$ and $\text{H}\gamma$ (Mennickent and Vogt, 1988).

We observed this star for one season only (1987). Apart from a gradual increase in brightness, there is no indication of periodicity. Nevertheless, the star is variable as the scatter (0.01 mag) is much larger than expected. It appears that we are seeing only the flickering component.

V986 Oph

This star is not known as a Be star, but we have included it here owing to its great interest in connection with the role of NRP in the Be phenomenon. V986 Oph has a long history of photometric variability. Lynds (1959) determined a period of 0.289 d and classified it as a β Cep variable. Jerzykiewicz (1975) confirmed this period from seven nights of photometry and deduced a peak-to-peak amplitude of 0.014 mag. However, not all the data could be described by a simple sinusoidal variation with this period. Pike and Lloyd (1979) failed to detect any radial velocity variations with the photometric period, though a systematic decrease in nightly mean velocities was observed.

To clarify the nature of this star, C. D. Pike organized an international campaign for July 1980. Fullerton *et al.* (1985) analyzed both the photometry and radial velocities obtained during the campaign and found a period of 0.325 d with a light amplitude of 0.01 mag in B . No significant radial velocity variations were found, but the systematic variation in mean nightly velocity was confirmed and orbital elements derived ($P = 25.56 \text{ d}$, $2K = 35 \text{ km s}^{-1}$, $e = 0.23$). They also obtained an extensive set of high dispersion line-profile observations and discovered high-order profile variations progressing from blue to red with a period of about 0.3 d . They deduced $\ell = 6$, but at times $\ell = 4$ and $\ell = 8$ are excited.

This confusing picture of V986 Oph prompted us to include it in our observing project. Our photometry for 1985 was obtained over 10 nights and showed light variations at a low level. Fourier periodogram analysis showed peaks at 1.74 and 1.16 d^{-1} and their one-day aliases (Fig. 11a - top panel). An unique period or periods could not be

determined from these data alone. Combining these data with all available photometry showed the most likely period to be 0.30 d or 0.23 d , its alias.

Our 1987 data again shows micro-variability, but owing to the spacing in longitude the aliasing problem largely disappears. The Fourier periodogram shows three strong peaks at $f_1 = 3.30$, $f_2 = 1.41$ and $f_3 = 0.77 \text{ d}^{-1}$ with semi-amplitudes of 4.7, 4.6 and 3.8 millimags respectively (Fig. 11a - middle panel). Prewhitenning by f_1 leaves f_2 and f_3 with f_2 slightly stronger (Fig. 11a - bottom panel). Further prewhitening by either f_2 or f_3 leads to a pure noise spectrum. Our conclusion is that there are two periods present : the short period is well determined to be 0.303 d (Fig. 11b), but the long period is uncertain, it is either 1.299 d or 0.709 d .

The short period is the same as the one originally suggested by Lynds and Jerzykiewicz and also found in our 1985 data. We cannot confirm the period found by Fullerton *et al.* (1985); indeed there is no indication of a 0.325 d period in our data at all. We note, however, that the 0.303 d period is the same as their estimate of the period of the high-order line-profile variations. The low photometric amplitude is consistent with a high-order NRP mode. The nature of the long period is uncertain. It could be another NRP mode (either rotationally split p -mode or a g -mode), but the period is close to the expected period of rotation so that some kind of rotational modulation is not ruled out.

There appears to be no compelling reason to classify this star as anything other than a β Cep variable as originally suggested by Lynds (1959). Admittedly, it has the longest known period of a star in this class, but this is not surprising in view of its early spectral type and luminosity. It does apparently oscillate with somewhat higher order modes than the majority of other β Cep variables, but this is not a criterion for reclassification. The only reasonable criterion for classification is a definite indication of a different pulsation mechanism. Since there is no reason to believe that the pulsation mechanism in this star is any different from other β Cep variables, we feel that reclassification serves merely to confuse the issue.

ϵ Cap

Mennickent and Vogt (1988) observed this star in 1986 and found deep central absorption with no traces of emission at $\text{H}\beta$. Dachs *et al.* (1986) found $\text{H}\alpha$ to have a central absorption and emission wings in 1982.

Our 1985 photometry showed that ϵ Cap underwent a sharp drop in brightness of nearly 0.2 mag in just three days followed by a rapid recovery. Excluding this period, the periodogram shows a set of peaks with $f = 2.60 \text{ d}^{-1}$ and its one-day aliases, but a double-wave light curve with half this frequency is possible as the waves have distinctly different amplitudes. This solution ($P = 0.769 \text{ d}$) was proposed by Balona and Engelbrecht (1987).

The 1986 photometry again showed indications of peri-

odicity, but we could not confirm the solution derived for 1985. Instead a frequency close to two cycles per day was indicated. The severe aliasing problem prevented a meaningful solution.

The combined ESO and SAAO data of 1987 showed a slow brightening over the observing run. Removing this trend showed a possible period very close to one day. By combining all available data we determined the best value to be 1.030 d (Fig. 12). Because it is so close to one day and because of the discrepancy obtained in 1985, this period must be regarded as very uncertain.

η PsA

Mennickent and Vogt (1988) find H β to be very diffuse with weak double emission ($V \approx R$) around a broad absorption core.

We were unable to detect any clear periodicity in η PsA during 1985, but the star was clearly a low-amplitude variable. A Fourier periodogram suggested a possible period of 0.774 d, which is the value quoted by Balona and Engelbrecht (1987). During 1986, the low-level variability (flickering) was again evident. No reliable period could be extracted. The same is true for the combined ESO and SAAO data of 1987.

HR 8408

Slettebak (1982) discovered HR 8408 to be a Be star showing faint double emission ($V \approx R$) at H β surrounding an absorption core. Mennickent and Vogt (1988) did not see emission during 1986, but suspect incipient emission at H β . Buscombe and Morris (1961) found a radial velocity range of 40 km s $^{-1}$.

This star was originally used as a photometric standard before we discovered its variability. The 1985 SAAO photometry shows clear indications of periodicity with a frequency $f = 2.53$ d $^{-1}$, though a double-wave solution with half this frequency shows the two waves to have significantly different amplitudes.

The 1986 data produced the highest peak in the Fourier periodogram at $f = 1.53$ d $^{-1}$, i.e. the one-day alias of the solution adopted for 1985. Again, a double-wave solution appeared to be significant.

The combined ESO and SAAO data of 1987 removed the aliasing problem. The correct frequency was found to be $f = 1.53$ d $^{-1}$ or the double-wave solution at $f = 0.765$ d $^{-1}$. By combining all available data we refined the period to 1.3106 if we adopt the double-wave solution (Fig. 13). This value (or possibly the single-wave period of 0.6553 d) seems to be well established.

\circ Aqr

The emission features in this star have been relatively static since 1975. The spectrum shows double emission in the Balmer lines ($V \approx R$) and a central shell absorption feature (Mennickent and Vogt 1988).

Photometry during 1985 showed a steady brightening of the star over a period of three months with a large excursion

in magnitude on one occasion. These complications made period extraction very difficult. A subset of the data indicated a possible double-wave light curve with $P = 1.449$ d (Balona and Engelbrecht 1987). During 1986 a much larger amount of data was obtained and the star was more quiescent. Both the Fourier and PDM periodograms show a strong signal at $f = 1.40$ d $^{-1}$ ($P = 0.71$ d) - half the period found in 1985. However, a double-wave light curve with the period found in 1985 also described the 1986 data well.

The combined ESO and SAAO data for 1987 again shows a strong signal with $f = 1.39$ d $^{-1}$ or the double-wave equivalent at $f = 0.70$ d $^{-1}$. This time the amplitude difference in the waves did not appear to be significant. Considering the good agreement of the period determinations of three seasons, we can state with little doubt that the period of this star is $P = 1.4325$ d from all available data (Fig. 14). This gives a double-wave light curve with distinctly different amplitudes for the two waves in 1985 and 1986. An increase in amplitude from 0.01 mag in 1986 to 0.02 mag in 1987 seems to have occurred.

One of the aims of this project was to determine the temperature range at which periodic Be stars could be found. \circ Aqr is one of the coolest stars (B7IVe) which has been found to be periodic. Our impression is that periodicity is most likely to be found amongst earlier spectral types.

ψ^2 Aqr

This star is not known as an emission-line star. We decided to include it in our observing program owing to its very rapid rotation ($v \sin i = 332$ km s $^{-1}$) and to test the hypothesis that rapidly-rotating non-Be stars do not show low-order profile variations (which are expected to give rise to periodic light variations). Abt and Levy (1978) regard the star as a probable spectroscopic binary.

The SAAO data for 1986 showed a strong peak at $f = 1.87$ d $^{-1}$ in the Fourier periodogram. The PDM periodogram suggested that a double-wave solution with half this frequency is more appropriate. The ESO and SAAO data of 1987 confirmed these conclusions. We have adopted a double-wave solution with $P = 1.073$ d obtained from an analysis of all existing photometry. A large decrease in amplitude seems to have occurred between the two seasons (Fig. 15).

The light variation in this star is in every respect similar to those in Be stars. It would be of great+ value to look for possible emission at H α in this star. If it can be shown that emission is definitely not present but that short-period light variations are present, it could offer an important clue to the relationship between the short-period variations and the Be phenomenon.

5. Colour variations.

There are very few observations of periodic Be stars observed in more than one colour. Van Vuuren *et al.* (1988) found that the colour variations in periodic Be stars in the

open cluster NGC 3766 were very small. The evidence indicates that the star is bluest (hotest) when brightest. The ESO observations produced simultaneous *uvby* colours. In figures 16 - 20 we show light and colour curves for some of the periodic Be stars observed at ESO. In general, the conclusions of van Vuuren *et al.*, (1988) are supported. The best evidence comes from observations of 48 Lib (Fig. 17). There is a small phase shift between *b* and *u-b*, but the star is brightest when bluest.

6. Conclusions.

A detailed interpretation of the results found in this paper and others in the series will be presented elsewhere. Here we summarize the main findings.

We found periodic light variations in almost all Be stars. Of the 17 candidates, only three stars were constant in light, three others had uncertain periods. The Be star HD137518 has a very unusual light curve which may not be strictly periodic or else could be another example of a triple-wave light curve as found in η Cen and α Eri.

In most cases the light curves are double waves with unequal maxima and minima. Sometimes the waves are of nearly equal amplitude, so it is quite possible that the periods of some apparently single-wave variables may be in error by a factor of two. Such a situation was found for two Be stars in the cluster NGC3766 (Ahmed 1 and Ahmed 15) where an apparently single-wave light curve became a double-wave light curve with twice the period in the course of a year (van Vuuren *et al.* 1988). The triple-wave light curve of η Cen is remarkable for its large amplitude, but has a counterpart in the bright star α Eri (Balona *et al.*, 1987).

It has become increasingly clear, both in this work and in previous studies of the light variability of periodic Be stars, that the light variations are never adequately described

by purely periodic variations. There is always a residual scatter which is often many times larger than the expected observational error. We have called these random, short-term variations *flickering*. It seems that flickering is closely associated with periodic variability. The constant Be stars are really constant and do not show this flickering (or at least not to the same extent).

In all cases we found the period of light variation to be consistent with the expected period of rotation of the star. There is no clear evidence for multiperiodicity in any of the stars. In the case of ζ Oph we found a β Cep - like variation with a period of 0.193 d superimposed on a longer period of 1.075 d which is practically six times the length of the shorter period and could be the rotation period. This behaviour was only seen for one season. This star is clearly an exception to the rule and deserves closer study. For χ Oph (and possibly HD137518) we found evidence for a period longer than the expected period of rotation. It is not very clear whether this variation is intrinsic to the star (as is the case in general) or a result of a close companion or circumstellar material.

The results of this work clearly show the importance of multi-site observations for these short-period variables. The aliasing problem that still existed in some stars was completely resolved by combining results from the two sites.

Acknowledgements.

At the time that these results were obtained and for part of the duration of the reductions, J.C. was at the *Astronomisch Instituut van de Katholieke Universiteit Leuven* under grant OT/86/28 of the *Onderzoeksfonds van de Katholieke Universiteit Leuven*. This support is gratefully acknowledged. We are also grateful to Dr. C. Waelkens for his role in organizing the observing campaign and for his stimulating comments.

References

- ABT H. A. and LEVY S. G. : 1978, *Astrophys. J. Suppl. Ser.* **36**, 241.
- ASHOKA B. N., SURENDIRANATH R. and KAMESWARA RAO N. : 1985, *Acta Astron.* **35**, 395.
- AYDIN C. and FARAGGIANA R. : 1978, *Astron. Astrophys. Suppl. Ser.* **34**, 51.
- BAADE D. : 1983, *Astron. Astrophys.* **124**, 283.
- BAADE D. : 1984, *Astron. Astrophys.* **135**, 101.
- BAADE D. : 1987a, in *IAU Symp. No. 132*, The Impact of very high S/N Spectroscopy on Stellar Physics, G. Cayrell and M. Spite Eds. (Kluwer Academic Publ., Dordrecht) p. 193.
- BAADE D. : 1987b, *Inf. Bull. var. Stars* **3124**.
- BAADE D., DACHS J., VAN DE WEYGAERT R. and STEEMAN F. : 1988, *Astron. Astrophys.* **198**, 211.
- BALONA L. A. and ENGELBRECHT C. A. : 1986, *Mon. Not. R. Astron. Soc.* **219**, 131.
- BALONA L. A., ENGELBRECHT C. A. and MARANG F. : 1987, *Mon. Not. R. Astron. Soc.* **227**, 123.
- BALONA L. A. and ENGELBRECHT C. A. : 1987, in *IAU Coll. 92*, The Physics of Be Stars A. Slettebak and T. P. Snow Eds. (Cambridge

- University Press, Cambridge) p. 87.
- BARKER P. K. and BROWN T. : 1974, *Astrophys. J. Lett.* **192**, L11.
- BUSCOMBE W. and MORRIS P. M. : 1961, *Mon. Not. R. Astron. Soc.* **123**, 233.
- DACHS J., HANUSCHIK R., KAISER D., BALLEREAU D., BOUCHET P., KIEHLING R., KOZOK J., RUDOLPH R. and SCHLOSSER W. : 1986, *Astron. Astrophys. Suppl. Ser.* **63**, 87.
- EBBETS D. : 1981, *Publ. Astron. Soc. Pac.* **93**, 119.
- FULLERTON A. W., BOLTON C. T. and PENROD G. D. : 1985, *J. R. Astron. Soc. Canada* **79**, 236.
- GARRISON R. F., HILTNER W. A. and SCHILD R. E. : 1977, *Astrophys. J. Suppl. Ser.* **35**, 111.
- GHOSH K. K., VELU C., KUPPUSWAMY K., JAYKUMAR K. and ROSARIO M. J. : 1987, *Inf. Bull. Var. Stars* **3056**.
- GHOSH K. K., PUCALENTHI S. and JAYKUMAR K. : 1988, *Be star Newslett.* **18**, 7.
- HARMANEC P. : 1987a, *Inf. Bull. Var. Stars* **3097**.
- HARMANEC P. : 1987b, *Bull. Astron. Inst. Czech.* **38**, 283.
- HARMANEC P. : 1989, *Bull. Astron. Inst. Czech.*, in press.
- HENDRY E. M. and BAHNG J.D.R. : 1981, *J. Astron. Astrophys.* **2**, 141.
- HOUK N. : 1978, Michigan catalogue of two-dimensional spectral types for HD stars, **2**.
- IRVINE N. J. : 1974, *Astrophys. J. Lett.* **188**, L19.
- JERZYKIEWICZ M. : 1975, *Acta Astron.* **25**, 81.
- KOZOK J. R. : 1985, *Astron. Astrophys. Suppl. Ser.* **62**, 7.
- KUBIAK M. and SEGGEWISS W. : 1982, *Acta Astron.* **32**, 371.
- LYNDS C. R. : 1959, *Astrophys. J.* **130**, 577.
- MCCONNELL D. J. and BIDELMAN W. P. : 1976, *Astron. J.* **81**, 225.
- MENNICKENT R. I. and VOGT N. : 1988, *Astron. Astrophys. Suppl. Ser.* **74**, 497.
- MERMILLIOD J. C. : 1987, *Astron. Astrophys. Suppl. Ser.* **71**, 119.
- NIEMELA V. S. and MENDEZ R. H. : 1974, *Astrophys. J. Lett.* **187**, L23.
- PALMER H. K. : 1906, *Lick Obs. Bull.* **4**, 97.
- PENROD G. D. : 1987, in *IAU Coll. 92*, The Physics of Be Stars A. Slettebak and T.P. Snow Eds. (Cambridge University Press, Cambridge) p. 463.
- PERCY J. R. : 1987, in *IAU Coll. 92*, The Physics of Be stars A. Slettebak and T. P. Snow Eds. (Cambridge University Press, Cambridge) p. 49.
- PIKE C. D. and LLOYD C. : 1979, *Inf. Bull. Var. Stars* **1716**.
- RAJAMOHAN R. : 1977, *Kodaikanal Obs. Bull. Ser. A* **2**, 6.
- RINGUELET-KASWALDER A. : 1963, *Astrophys. J.* **137**, 1310.
- SLETTEBAK A. : 1982, *Astrophys. J. Suppl. Ser.* **50**, 55.
- STELLINGWERF R. F. : 1978, *Astrophys. J.* **224**, 953.
- STROHMEIER W., KNIGGE R. and OTT H. : 1964, *Inf. Bull. Var. Stars* **62**.
- SWINGLER D. N. : 1989, *Astron. J.* **97**, 280.
- VAN VUUREN G. W., BALONA L. A. and MARANG F. : 1988, *Mon. Not. R. Astron. Soc.* **234**, 373.
- VOGT S. S. and PENROD G. D. : 1983, *Astrophys. J.* **275**, 661.
- WAELKENS C. and RUFENER F. : 1983, *Astron. Astrophys.* **121**, 45.
- WALKER G. A. H., YANG S. and FAHLMAN G. G. : 1979, *Astrophys. J.* **233**, 199.
- WALKER G. A. H., YANG S. and FAHLMAN G. G. : 1981, in Proc. Workshop on Pulsating B Stars G.E.V.O.N. and C. Sterken Eds. (Nice Obs.) p. 261.
- WALKER G. A. H. and YANG S. : 1987, Observational astrophysics with very high precision data (Univ. Liege) p. 337.
- WILSON R. E. : 1914, *Lick Obs. Bull.* **8**, 130.

TABLE I.—*A summary of stars observed. Notes are given for program stars ; the others are comparison stars. The abbreviation 2W and 3W stands for double-wave and triple -wave light curve respectively. The grouping includes stars which are in close proximity and for which the same transparency corrections were applied. The projected rotational velocity, the mean Strömgren b magnitude and the total number of observations is shown.*

HR	HD	Name	MK type	vsini	$\langle b \rangle$	N	Notes
HR 5190	HD 120307	ν Cen	B2IV	91	3.33	279	P = 2.621 d.
HR 5193	HD 120324	μ Cen	B2IV-Ve	175	3.39	360	P = 2.1017 d, 2W.
HR 5439	HD 127971		B7V		5.872	202	
HR 5440	HD 127972	η Cen	B1.5Vne	333	2.32	352	P = 1.927 d, 3W.
HR 5668	HD 135348		B3IV		6.026	178	
	HD 137518		B1IIIInep		7.94	183	P = 8.696 d?
HR 5730	HD 137387	κ^1 Aps	B1pne		5.39	308	P = 1.238 d, 2W.
HR 5786	HD 138867		B9V		5.950	246	
HR 6172	HD 149671	η^1 TrA	B7IVe		5.89	202	Constant.
HR 6233	HD 151441		B8II-III		6.150	215	
HR 5902	HD 142096		B2.5V	197	5.065	230	
HR 5941	HD 142983	48 Lib	B5IIIpe	393	4.81	232	P = 0.4017 d.
HR 5993	HD 144470		B1V	142	3.986	401	
HR 6112	HD 147933		B2V	303	4.781	134	
HR 6118	HD 148184	χ Oph	B2IV:pe	134	4.61	346	P = 13.774 d.
HR 6175	HD 149757	ζ Oph	O9.5Vne	379	2.67	313	P ₁ = 1.075, P ₂ = 0.193 d.
HR 6224	HD 151133		B9.5III	100	6.050	131	
HR 6451	HD 157042	ι Ara	B2IIIIne	369	5.25	542	P = 0.556 d, 2W.
HR 6460	HD 157243		B7III	150	5.120	684	
HR 6475	HD 157599		B8-9V		6.190	405	
HR 6510	HD 158427	α Ara	B2IIIIne	298	2.81	545	P = 0.9807 d, 2W.
HR 6712	HD 164284	66 Oph	B2Ve	221	4.77	160	Constant.
HR 6719	HD 164432		B2IV		6.352	287	
HR 6732	HD 164716		B9V		6.878	66	
HR 6747	HD 165174	V986 Oph	B0IIIIn	434	6.20	274	P ₁ =0.303, P ₂ =1.3/0.7 d.
HR 8260	HD 205637	ϵ Cap	B2.5Vpe	293	4.36	567	P = 1.03 d?
HR 8293	HD 206546		A3m	47	6.336	581	
HR 8386	HD 209014	η PsA	B8Ve		5.39	598	Constant.
HR 8408	HD 209522		B4IVne	300	5.90	597	P = 1.3106 d, 2W.
HR 8446	HD 210300		A5V		6.500	482	
HR 8402	HD 209409	\circ Aqr	B7IVe	227	4.67	526	P = 1.433 d, 2W.
HR 8451	HD 210419		A1Vnn	254	6.270	455	
HR 8533	HD 212404		A0V	65	5.750	406	
HR 8840	HD 219402		A3V		5.561	389	
HR 8858	HD 219688	ψ^2 Aqr	B5V	332	4.32	388	P = 1.073 d, 2W.

TABLE IIa. — *Strömgren-b observations for periodic Be stars observed at SAAO. The heliocentric Julian day is with respect to JD2440000.000.*

HR 5190	HJD	b	HJD	b	HJD	b	HR 5190
HJD	b	HJD	b	HJD	b	HJD	b
6959.2227 3.326	6981.2317 3.329	7241.5998 3.301	7250.6334 3.302	7251.3468 3.287	7256.6034 3.302	7266.6034 3.302	
6959.2679 3.314	6981.3039 3.331	7241.6288 3.295	7251.3967 3.301	7251.4350 3.304	7256.6284 3.305	7267.2784 3.305	
6959.3058 3.315	6982.2613 3.322	7242.2955 3.301	7251.4794 3.305	7252.3293 3.295	7256.4691 3.309	7267.3253 3.299	
6959.3447 3.322	6982.3033 3.324	7242.3293 3.305	7251.5175 3.306	7252.3639 3.303	7256.5561 3.305	7268.3740 3.298	
6959.3817 3.329	6982.3944 3.317	7242.4022 3.303	7251.5553 3.298	7252.6244 3.289	7256.3204 3.301	7268.3712 3.297	
6959.4169 3.331	6983.2248 3.328	7242.4351 3.301	7251.6244 3.299	7252.6468 3.299	7256.4254 3.295	7268.4744 3.302	
6959.4671 3.328	6983.4623 3.333	7242.4691 3.302	7252.3409 3.297	7252.7367 3.300	7256.4354 3.295	7268.4744 3.302	
6960.4225 3.341	6984.2046 3.320	7242.5014 3.301	7252.4116 3.296	7252.7417 3.302	7256.4354 3.295	7268.4744 3.302	
6961.2187 3.314	6983.3515 3.320	7242.5813 3.304	7252.4416 3.296	7252.7450 3.303	7256.4354 3.295	7268.4744 3.302	
6961.2457 3.322	7236.3839 3.297	7242.6059 3.301	7252.4501 3.293	7252.7533 3.302	7256.4354 3.295	7268.4744 3.302	
6961.2717 3.320	7236.4252 3.292	7242.4817 3.302	7252.4933 3.293	7252.7533 3.302	7256.4354 3.295	7268.4744 3.302	
6961.3051 3.321	7236.4468 3.306	7243.3035 3.313	7252.5343 3.290	7252.7700 3.299	7256.4354 3.295	7268.4744 3.302	
6961.3389 3.318	7236.5013 3.296	7243.3364 3.308	7252.5716 3.288	7252.7700 3.299	7256.4354 3.295	7268.4744 3.302	
6961.3701 3.321	7236.5363 3.300	7243.3684 3.308	7252.6084 3.293	7252.7700 3.299	7256.4354 3.295	7268.4744 3.302	
6961.4039 3.323	7236.5711 3.297	7243.4018 3.306	7252.6448 3.288	7252.7700 3.299	7256.4354 3.295	7268.4744 3.302	
6961.4493 3.320	7236.6005 3.300	7243.4337 3.305	7252.6801 3.290	7252.7700 3.299	7256.4354 3.295	7268.4744 3.302	
6962.2118 3.307	7236.6294 3.294	7243.4663 3.303	7252.7022 3.287	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6962.2566 3.318	7237.2946 3.304	7243.5014 3.306	7252.7474 3.311	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6962.2845 3.324	7237.3298 3.305	7243.5338 3.306	7252.7613 3.310	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6962.3188 3.324	7237.3642 3.306	7243.5665 3.307	7252.7813 3.309	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6965.2383 3.363	7237.3980 3.301	7243.5939 3.306	7252.8108 3.305	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6965.3849 3.343	7237.4322 3.306	7243.6132 3.311	7252.8452 3.308	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6965.4322 3.337	7237.4647 3.300	7243.6293 3.311	7252.8644 3.311	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6966.2226 3.319	7238.2962 3.310	7245.5135 3.302	7252.8964 3.298	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6966.2535 3.318	7238.3271 3.311	7246.3011 3.312	7252.9359 3.297	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6966.2820 3.323	7238.3598 3.305	7246.3329 3.305	7252.9741 3.311	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6966.3101 3.323	7238.3944 3.306	7246.3650 3.304	7252.9942 3.310	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6966.3409 3.324	7238.4249 3.303	7246.3962 3.307	7252.9943 3.309	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6966.3923 3.328	7238.4630 3.307	7246.4297 3.308	7252.9943 3.297	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6966.4238 3.321	7238.4967 3.303	7246.4662 3.303	7252.9943 3.299	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6967.2232 3.336	7238.5291 3.300	7246.4984 3.305	7252.9871 3.295	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6967.2512 3.333	7238.5564 3.304	7246.5304 3.304	7252.9943 3.295	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6967.2836 3.335	7238.5966 3.305	7246.5628 3.308	7252.9943 3.295	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6967.3203 3.337	7238.6240 3.304	7246.6060 3.300	7252.9943 3.295	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6967.3495 3.346	7238.6599 3.305	7246.6331 3.301	7252.9943 3.295	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6967.3813 3.341	7239.2931 3.311	7246.6495 3.304	7252.9943 3.295	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6967.4218 3.334	7240.2931 3.311	7247.3082 3.309	7252.9943 3.295	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6967.4628 3.335	7240.3257 3.309	7247.3408 3.309	7252.9943 3.295	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6967.5032 3.341	7240.3599 3.309	7247.3791 3.296	7252.9943 3.295	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6968.2458 3.332	7240.4288 3.307	7247.4125 3.299	7252.9943 3.295	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6968.2735 3.335	7240.4636 3.306	7247.4498 3.300	7252.9943 3.295	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6968.3018 3.337	7240.4964 3.309	7247.4839 3.298	7252.9943 3.295	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6968.3318 3.357	7240.5293 3.310	7247.5213 3.299	7252.9943 3.295	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6972.2623 3.327	7240.5626 3.306	7247.5558 3.299	7252.9943 3.295	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6972.3076 3.330	7240.5956 3.307	7247.5891 3.302	7252.9943 3.295	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6972.3585 3.346	7240.6216 3.309	7247.6141 3.309	7252.9943 3.295	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6978.2279 3.330	7240.3599 3.309	7247.3791 3.296	7252.9943 3.295	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6978.2617 3.326	7240.4200 3.302	7247.4125 3.299	7252.9943 3.295	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6978.3011 3.341	7241.3289 3.303	7248.4083 3.304	7252.9943 3.295	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6978.3564 3.344	7241.3624 3.298	7248.5753 3.315	7252.9943 3.295	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6978.4156 3.334	7241.3961 3.303	7250.3382 3.284	7252.9943 3.295	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6980.2514 3.334	7241.4290 3.300	7250.3889 3.290	7252.9943 3.295	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6980.2988 3.324	7241.4622 3.304	7250.4330 3.288	7252.9943 3.295	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6980.3467 3.342	7241.4971 3.299	7250.4904 3.298	7252.9943 3.295	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6980.3858 3.334	7241.5310 3.300	7250.5435 3.307	7252.9943 3.295	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6980.4386 3.340	7241.5663 3.300	7250.5967 3.318	7252.9943 3.295	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6981.2331 3.411	7241.5672 3.379	7250.6093 3.387	7252.9943 3.295	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6981.2666 3.415	7241.6013 3.376	7241.6234 3.378	7251.3480 3.371	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6981.3052 3.416	7241.6234 3.376	7241.6504 3.373	7251.3890 3.366	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6981.3479 3.407	7241.6504 3.421	7241.6874 3.377	7251.4099 3.384	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6981.3822 3.393	7242.3493 3.382	7242.3651 3.380	7251.5184 3.384	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6981.3822 3.394	7242.3495 3.386	7242.4032 3.382	7251.5545 3.380	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6981.4181 3.399	7243.2261 3.408	7242.4161 3.382	7251.5814 3.376	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6981.4483 3.395	7243.2637 3.409	7242.4700 3.389	7251.6274 3.376	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6981.4842 3.408	7243.3082 3.406	7242.5025 3.387	7252.3777 3.389	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6981.5217 3.393	7243.3528 3.393	7242.5071 3.373	7252.4126 3.384	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6981.5557 3.393	7243.3984 3.393	7242.5404 3.378	7252.4515 3.380	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6981.6003 3.394	7243.4321 3.393	7242.5731 3.377	7252.4950 3.382	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6981.6360 3.396	7243.4640 3.394	7242.6014 3.378	7252.5354 3.380	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6981.6711 3.396	7243.5076 3.395	7242.6366 3.378	7252.5723 3.384	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6981.7111 3.396	7243.5376 3.395	7242.6693 3.378	7252.6093 3.386	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6981.7522 3.392	7243.5659 3.364	7242.7021 3.360	7252.6462 3.383	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6981.7827 3.357	7243.5914 3.361	7242.7347 3.361	7252.6824 3.383	7256.4354 3.295	7256.4354 3.295	7268.4744 3.302	
6981.8227 3.357	7243.6103 3.361	7242.7670 3.361	7252.7170 3.381	725			

TABLE IIa (*continued*)

HR 5440	HJD	b	HJD	b	HJD	b	HR 5440	HJD	b	
	HJD	b	6980.3898	2.284	7241.604	2.191	7250.6384	2.194		
6959.	2277	2.298	6980.4417	2.296	7241.6265	2.193	7251.4028	2.180	7268.2777	2.173
6959.	2714	2.287	6981.2354	2.311	7241.6340	2.193	7251.4390	2.188	7268.3242	2.182
6959.	3088	2.290	6981.3074	2.311	7242.3327	2.187	7251.5215	2.178	7268.3755	2.195
6959.	3427	2.292	6981.3793	2.313	7242.3685	2.181	7251.5571	2.186	7268.4290	2.210
6959.	3845	2.298	6982.2655	2.273	7242.4058	2.161	7251.6304	2.201	7268.4777	2.222
6959.	4197	2.306	6982.3064	2.287	7242.4386	2.150	7252.3649	2.206	7270.2797	2.205
6959.	4699	2.315	6982.3475	2.294	7242.4726	2.147	7252.3804	2.220	7270.3286	2.198
6960.	4272	2.286	6982.3973	2.298	7242.5050	2.149	7252.4153	2.227	7270.3810	2.200
6961.	2201	2.298	6982.4505	2.300	7242.5846	2.172	7252.4984	2.204	7270.4355	2.198
6961.	2744	2.306	6983.2286	2.312	7242.6094	2.181	7252.5385	2.193	7270.4890	2.203
6961.	3079	2.312	6983.2659	2.310	7242.6260	2.188	7252.5747	2.184	7270.5234	2.193
6961.	3407	2.316	6983.3080	2.316	7243.3069	2.192	7252.6119	2.178	7270.5590	2.186
6961.	3725	2.331	6983.3561	2.318	7243.3401	2.199	7252.6493	2.174	7270.5912	2.173
6961.	4065	2.333	7236.3883	2.213	7243.3720	2.211	7260.3150	2.200	7270.6222	2.176
6961.	4522	2.341	7236.4291	2.226	7243.4050	2.204	7260.3875	2.192	7274.2955	2.227
6962.	2143	2.319	7236.4642	2.225	7243.4376	2.193	7261.4797	2.235	7274.3519	2.210
6962.	2594	2.302	7236.5052	2.221	7243.4711	2.191	7261.5173	2.224	7274.4139	2.183
6962.	2873	2.290	7236.5400	2.216	7243.5047	2.191	7261.5517	2.217	7274.4716	2.160
6962.	3212	2.281	7236.5744	2.196	7243.5369	2.197	7261.5842	2.203	7274.5094	2.162
6962.	3561	2.282	7236.6124	2.175	7243.5698	2.202	7261.6192	2.202	7274.5469	2.160
6962.	3561	2.280	7236.6339	2.184	7243.5975	2.200	7261.6500	2.199	7274.5803	2.163
6965.	2412	2.342	7237.3331	2.166	7243.6181	2.204	7262.3017	2.200	7321.4189	2.197
6965.	2739	2.345	7237.3679	2.171	7243.6334	2.198	7262.3409	2.201	7322.2284	2.180
6965.	3874	2.372	7237.4019	2.178	7245.3170	2.211	7262.3767	2.194	7322.2885	2.198
6965.	4348	2.341	7237.4352	2.181	7246.3043	2.163	7262.4473	2.187	7322.3454	2.215
6965.	4784	2.296	7237.4680	2.192	7246.3362	2.150	7262.4844	2.188	7322.4024	2.219
6966.	2254	2.269	7238.3308	2.213	7246.3681	2.155	7262.5196	2.188	7322.4621	2.215
6966.	2563	2.272	7238.3634	2.218	7246.3993	2.157	7262.5562	2.194	7323.2346	2.215
6966.	2850	2.278	7238.3984	2.213	7246.4333	2.169	7262.5928	2.192	7323.2916	2.205
6966.	3328	2.298	7238.4322	2.202	7246.4696	2.187	7262.6250	2.201	7323.3463	2.187
6966.	3634	2.309	7238.4667	2.207	7246.5019	2.204	7263.3001	2.218	7323.4277	2.166
6966.	3948	2.317	7238.4999	2.195	7246.5336	2.217	7263.3377	2.234	7323.4827	2.195
6966.	4265	2.316	7238.5322	2.183	7246.5663	2.215	7263.3704	2.242	7328.2558	2.219
6966.	4665	2.318	7238.5672	2.178	7246.6095	2.217	7263.4057	2.243	7328.3399	2.226
6967.	2257	2.372	7238.5999	2.173	7246.6370	2.216	7263.4421	2.239	7328.4169	2.210
6967.	2540	2.384	7238.6289	2.171	7246.6534	2.215	7263.4772	2.228	7329.2549	2.166
6967.	2863	2.397	7238.6440	2.169	7247.3115	2.204	7263.5120	2.214	7329.3145	2.182
6967.	3227	2.393	7240.3299	2.211	7247.3485	2.195	7263.5489	2.204	7329.3651	2.201
6967.	3524	2.379	7240.3648	2.232	7247.3823	2.200	7263.5861	2.191	7329.4265	2.202
6967.	3838	2.350	7240.3998	2.219	7247.4157	2.193	7264.2809	2.180	7330.2599	2.206
6967.	4241	2.319	7240.4322	2.207	7247.4531	2.194	7265.3107	2.227	7330.3160	2.201
6967.	4651	2.287	7240.4673	2.189	7247.4879	2.184	7265.4204	2.203	7330.3629	2.194
6968.	2220	2.304	7240.5002	2.170	7247.5245	2.176	7265.4715	2.176	7330.4257	2.181
6968.	2485	2.312	7240.5323	2.155	7247.5590	2.173	7265.5225	2.173	7331.2561	2.181
6968.	2758	2.319	7240.5658	2.148	7247.5922	2.167	7265.5585	2.172	7331.2581	2.186
6968.	3042	2.328	7240.5980	2.151	7247.6180	2.169	7266.2798	2.179	7331.3168	2.209
6968.	3342	2.356	7240.6348	2.232	7247.6344	2.170	7266.3131	2.184	7331.3690	2.214
6969.	2683	2.337	7240.6448	2.164	7248.3164	2.169	7266.3458	2.186	7331.4383	2.217
6972.	3107	2.344	7241.3325	2.183	7248.4118	2.205	7266.3783	2.194		
6972.	3231	2.319	7241.3660	2.198	7248.5807	2.233	7266.4461	2.182		
6972.	3691	2.316	7241.3998	2.203	7250.3453	2.203	7266.5018	2.194		
6978.	3143	2.111	7241.4224	2.211	7250.3942	2.221	7266.5431	2.212		
6978.	3607	2.293	7241.4655	2.211	7250.4387	2.217	7266.5763	2.221		
6978.	4192	2.292	7241.5005	2.198	7250.4966	2.210	7266.6071	2.218		
6980.	3026	2.278	7241.5348	2.184	7250.5483	2.204	7267.2821	2.210		
6980.	3499	2.282	7241.5719	2.187	7250.6010	2.203	7267.3280	2.197		

TABLE IIa (continued)

HR 5941	HJD	b	HJD	b	HJD	b	HJD	b	HJD	b	HJD	b
6959.2407 4.840	6980.3161 4.791		6277.2342 4.575	6277.3850 4.603	6959.4358 4.593		6277.3710 4.603	6959.4923 4.589	6981.2547 4.629		6981.2881 4.628	
6959.2844 4.849	6980.3624 4.801		6271.2811 4.573	6277.4109 4.598	6960.4743 4.591		6271.2323 4.574	6960.5010 4.587	6981.3473 4.626		6981.3972 4.631	
6959.3223 4.834	6980.4019 4.810		6271.2551 4.574	6277.4233 4.598	6960.5010 4.587		6271.2848 4.576	6278.2331 4.611	6960.5460 4.595		6982.2852 4.620	
6959.3596 4.829	6980.4593 4.815		6271.3006 4.574	6278.2351 4.610	6961.2350 4.608		6271.2351 4.576	6278.2667 4.611	6961.2631 4.601		6982.3644 4.624	
6959.3966 4.807	6981.2490 4.827		6271.3302 4.577	6278.3642 4.609	6961.2900 4.593		6271.3302 4.576	6278.3774 4.607	6961.3233 4.588		6982.4174 4.623	
6959.4311 4.793	6981.2824 4.827		6271.3449 4.576	6278.3784 4.616	6961.3233 4.588		6271.3449 4.576	6278.3784 4.616	6962.4689 4.622		6982.4689 4.622	
6959.4869 4.804	6981.3259 4.830		6271.3627 4.581	6278.3905 4.604	6961.3556 4.590		6271.3627 4.581	6278.4005 4.602	6961.3877 4.591		6983.2866 4.629	
6959.5330 4.805	6981.3917 4.836		6271.3904 4.577	6278.4160 4.603	6961.4221 4.599		6271.3904 4.577	6278.4160 4.603	6961.4221 4.599		6983.3307 4.627	
6960.4382 4.846	6981.4454 4.815		6271.4040 4.576	6279.3228 4.619	6961.4675 4.600		6271.4040 4.576	6279.3228 4.619	6961.4675 4.600		6983.4155 4.626	
6960.4963 4.844	6982.2794 4.798		6271.4179 4.574	6279.3427 4.620	6962.2750 4.595		6271.4179 4.574	6279.3427 4.620	6962.2750 4.595		7321.4133 4.531	
6960.5418 4.834	6982.3199 4.793		6271.4309 4.576	6279.3563 4.618	6962.3033 4.591		6271.4309 4.576	6279.3563 4.618	6962.3033 4.591		7321.5075 4.522	
6961.2307 4.830	6982.3592 4.801		6273.2347 4.583	6279.3711 4.615	6962.3361 4.595		6273.2347 4.583	6279.3711 4.615	6962.3361 4.595		7322.2401 4.500	
6961.2589 4.827	6982.4117 4.809		6273.2557 4.582	6279.3843 4.616	6965.2584 4.612		6273.2557 4.582	6279.3843 4.616	6965.2584 4.612		7322.3007 4.500	
6961.2855 4.830	6982.4463 4.822		6273.2730 4.588	6279.4006 4.617	6965.2903 4.612		6273.2730 4.588	6279.4006 4.617	6965.2903 4.612		7322.3563 4.517	
6961.3190 4.834	6983.2423 4.809		6273.2895 4.587	6279.4141 4.611	6965.4033 4.619		6273.2895 4.587	6279.4141 4.611	6965.4033 4.619		7322.4133 4.528	
6961.3512 4.817	6983.2813 4.822		6273.3050 4.585	6281.2376 4.600	6965.4503 4.622		6273.3050 4.585	6281.2376 4.600	6965.4503 4.622		7322.4793 4.525	
6961.3833 4.806	6983.3222 4.829		6273.3230 4.586	6281.2531 4.603	6965.4942 4.616		6273.3230 4.586	6281.2531 4.603	6965.4942 4.616		7322.5658 4.516	
6961.4172 4.798	7321.5034 4.804		6273.3396 4.585	6281.2686 4.605	6965.5149 4.619		6273.3396 4.585	6281.2686 4.605	6965.5149 4.619		7323.2458 4.501	
6961.4630 4.801	7322.2370 4.804		6273.3459 4.584	6281.2822 4.611	6966.2426 4.626		6273.3459 4.584	6281.2822 4.611	6966.2426 4.626		7323.3031 4.507	
6962.2273 4.791	7322.2974 4.806		6273.3702 4.583	6281.2977 4.618	6966.3255 4.625		6273.3702 4.583	6281.2977 4.618	6966.3255 4.625		7323.3584 4.514	
6962.2707 4.804	7322.3532 4.819		6273.3850 4.586	6281.3132 4.621	6966.3036 4.622		6273.3850 4.586	6281.3132 4.621	6966.3036 4.622		7323.4389 4.516	
6962.2981 4.807	7322.4103 4.821		6273.3906 4.585	6281.3237 4.622	6966.3466 4.616		6273.3906 4.585	6281.3237 4.622	6966.3466 4.616		7323.4941 4.514	
6962.3316 4.813	7322.4704 4.827		6273.3974 4.585	6281.3343 4.623	6966.3795 4.617		6273.3974 4.585	6281.3343 4.623	6966.3795 4.617		7328.2743 4.525	
6965.2531 4.826	7322.5424 4.806		6274.2374 4.595	6281.3413 4.631	6966.4096 4.622		6274.2374 4.595	6281.3413 4.631	6966.4096 4.622		7328.3539 4.526	
6965.2853 4.817	7323.2424 4.827		6274.2396 4.595	6281.3551 4.630	6966.4407 4.621		6274.2396 4.595	6281.3551 4.630	6966.4407 4.621		7328.4618 4.520	
6965.3986 4.802	7323.2998 4.820		6274.2466 4.601	6281.3695 4.630	6966.4811 4.623		6274.2466 4.601	6281.3695 4.630	6966.4811 4.623		7329.2680 4.524	
6965.4457 4.806	7323.3548 4.804		6274.2546 4.601	6281.3843 4.630	6966.5191 4.623		6274.2546 4.601	6281.3843 4.630	6966.5191 4.623		7329.3255 4.525	
6965.4891 4.803	7323.4336 4.809		6274.2686 4.601	6281.3982 4.633	6966.5191 4.623		6274.2686 4.601	6281.3982 4.633	6966.5191 4.623		7329.3763 4.526	
6966.2376 4.806	7323.4910 4.810		6274.2990 4.603	6281.3984 4.638	6966.5191 4.623		6274.2990 4.603	6281.3984 4.638	6966.5191 4.623		7330.4369 4.526	
6966.2673 4.806	7328.3497 4.819		6274.3136 4.603	6281.4211 4.619	6967.2415 4.621		6274.3136 4.603	6281.4211 4.619	6967.2415 4.621		7329.3763 4.526	
6966.2980 4.799	7328.4582 4.825		6274.3560 4.608	6282.3268 4.614	6967.2703 4.621		6274.3560 4.608	6282.3268 4.614	6967.2703 4.621		7329.4389 4.526	
6966.3426 4.801	7329.2641 4.846		6274.3709 4.604	6282.3395 4.613	6967.3074 4.618		6274.3709 4.604	6282.3395 4.613	6967.3074 4.618		7329.4941 4.526	
6966.3742 4.806	7329.3224 4.827		6274.3843 4.604	6282.3518 4.612	6967.3370 4.617		6274.3843 4.604	6282.3518 4.612	6967.3370 4.617		7330.2726 4.514	
6966.4050 4.816	7329.3730 4.816		6274.3874 4.606	6282.3635 4.608	6967.3666 4.617		6274.3874 4.606	6282.3635 4.608	6967.3666 4.617		7330.3269 4.516	
6966.4365 4.813	7329.4342 4.803		6276.2803 4.592	6282.3901 4.597	6967.3994 4.616		6276.2803 4.592	6282.3901 4.597	6967.3994 4.616		7330.3746 4.524	
6966.4766 4.814	7329.4900 4.827		6276.2951 4.592	6282.3951 4.592	6967.4383 4.610		6276.2951 4.592	6282.3951 4.592	6967.4383 4.610		7330.4369 4.526	
6966.5146 4.810	7320.2693 4.814		6276.3096 4.595	6282.4148 4.588	6967.4827 4.609		6276.3096 4.595	6282.4148 4.588	6967.4827 4.609		7330.4923 4.526	
6967.2369 4.827	7320.3234 4.815		6276.3264 4.596	6282.2379 4.592	6967.5216 4.616		6276.3264 4.596	6282.2379 4.592	6967.5216 4.616		7331.2702 4.510	
6967.2657 4.823	7320.3714 4.824		6276.3435 4.605	6282.2518 4.594	6968.2374 4.610		6276.3435 4.605	6282.2518 4.594	6968.2374 4.610		7331.3287 4.523	
6967.2984 4.785	7320.4337 4.840		6276.3568 4.608	6282.2639 4.595	6968.2649 4.614		6276.3568 4.608	6282.2639 4.595	6968.2649 4.614		7331.3809 4.522	
6967.3328 4.817	7320.4891 4.834		6276.3697 4.607	6282.2801 4.595	6968.2924 4.609		6276.3697 4.607	6282.2801 4.595	6968.2924 4.609		7331.4504 4.519	
6967.3620 4.813	7321.2669 4.856		6276.3831 4.613	6282.2974 4.596	6968.3185 4.592		6276.3831 4.613	6282.2974 4.596	6968.3185 4.592		7331.5105 4.544	
6967.3939 4.815	7321.3252 4.832		6276.3972 4.604	6282.3110 4.593	6972.2905 4.586		6276.3972 4.604	6282.3110 4.593	6972.2905 4.586			
6967.4340 4.810	7331.4470 4.810		6276.4091 4.612	6282.3324 4.585	6972.3325 4.578		6276.4091 4.612	6282.3324 4.585	6972.3325 4.578			
6967.5175 4.805	7331.5073 4.814		6276.4226 4.616	6282.3375 4.587	6973.4027 4.580		6276.4226 4.616	6282.3375 4.587	6973.4027 4.580			
6968.2329 4.802			6277.2328 4.593	6282.3401 4.584	6978.2521 4.614		6277.2328 4.593	6282.3401 4.584	6978.2521 4.614			
6968.2600 4.804			6277.2473 4.598	6283.3633 4.580	6978.2902 4.610		6277.2473 4.598	6283.3633 4.580	6978.2902 4.610			
6968.2888 4.801			6277.2611 4.601	6283.3761 4.582	6978.3327 4.611		6277.2611 4.601	6283.3761 4.582	6978.3327 4.611			
6968.3146 4.813			6277.2757 4.601	6283.3897 4.578	6978.3822 4.616		6277.2757 4.601	6283.3897 4.578	6978.3822 4.616			
6972.2841 4.795			6277.2887 4.604	6284.4034 4.580	6978.4278 4.616		6277.2887 4.604	6284.4034 4.580	6978.4278 4.616			
6972.3259 4.803			6277.3168 4.604	6284.4255 4.585	6980.2932 5.253		6277.3168 4.604	6284.4255 4.585	6980.2932 5.253			
6972.3723 4.802			6277.3315 4.604	6284.4334 4.582	6980.3115 5.258		6277.3315 4.604	6284.4334 4.582	6980.3115 5.258			
6972.4056 4.770			6277.3500 4.605	6284.4520 4.582	6980.3295 5.260		6277.3500 4.605	6284.4520 4.582	6980.3295 5.260			
6972.4205 4.770			6277.3620 4.605	6284.4620 4.582	6980.3475 5.255		6277.3620 4.605	6284.4620 4.582	6980.3475 5.255			
6972.4625 4.770			6277.3700 4.605	6284.4720 4.582	6980.3475 5.255		6277.3700 4.605	6284.4720 4.582	6980.3475 5.255			
6972.4866 4.770			6277.3780 4.605	6284.4820 4.582	6980.3475 5.255		6277.3780 4.605	6284.4820 4.582	6980.3475 5.255			
6972.5066 4.770			6277.3867 4.605	6284.4920 4.582	6980.3475 5.255		6277.3867 4.605	6284.4920 4.582	6980.3475 5.255			
697												

PHOTOMETRY OF BE STARS

TABLE IIa (*continued*)

HR 6451	HJD	b	HJD	b								
6972.3387	5.280	7328.5936	5.168	6972.3737	2.833	6676.3866	3.124	6691.3507	3.089	6691.3742	3.093	
6972.5204	5.242	7329.2879	5.168	6977.4090	2.831	6676.4035	3.130	6691.3681	3.044	6692.2260	3.064	
6972.5434	5.247	7329.3419	5.203	6977.4214	2.826	6677.3404	3.110	6693.2259	3.049	6693.2492	3.042	
6972.5646	5.248	7329.3914	5.273	6977.4337	2.824	6677.3599	3.117	6693.2721	3.034	6693.2934	3.037	
6972.5868	5.252	7329.4514	5.271	6977.4458	2.825	6677.3830	3.124	6693.3185	3.046	6693.3423	3.039	
6973.4105	5.268	7329.5105	5.254	6977.4579	2.827	6677.4044	3.121	6693.3681	3.041	6693.3861	3.041	
6975.5615	5.249	7329.5476	5.236	6978.1624	2.832	6680.2181	3.104	6694.2254	3.063	6694.2543	3.053	
6978.2586	5.238	7329.6029	5.201	6978.2845	2.830	6680.2415	3.107	6694.3016	3.058	6694.3309	3.058	
6978.2973	5.238	7330.2909	5.219	6978.3185	2.824	6680.2580	3.113	6694.3586	3.062	6694.3861	3.044	
6978.3389	5.241	7330.3417	5.201	6978.4019	2.828	6680.2770	3.116	6695.2562	3.016	6695.2762	3.030	
6978.3886	5.249	7330.3905	5.199	6978.4145	2.822	6680.2951	3.118	6695.2993	3.005	6695.3254	3.043	
6978.4448	5.267	7330.4527	5.202	6978.4268	2.820	6680.3131	3.121	6695.3517	3.057	6695.3778	3.054	
6978.4971	5.268	7330.5089	5.170	6978.4357	2.818	6680.3312	3.118	6696.3016	3.053	6696.3393	3.122	
6978.5364	5.257	7330.5453	5.186	6979.3406	2.833	6680.3493	3.122	6696.3721	3.056	6696.3902	3.112	
6978.5564	5.249	7330.5817	5.190	6979.3546	2.829	6680.3712	3.122	6696.3956	3.062	6696.3982	3.062	
6978.5755	5.244	7331.2879	5.242	6979.3676	2.823	6680.3902	3.112	6696.4243	3.017	6696.4437	3.020	
6980.3451	5.250	7331.4013	5.229	6979.3882	2.824	6680.2220	3.120	6696.4552	3.016	6696.4747	3.016	
6980.3741	5.237	7331.4603	5.227	6979.3967	2.824	6680.2433	3.117	6696.4847	3.017	6696.5037	3.017	
6980.4123	5.242	7331.4665	5.226	6979.4122	2.822	6680.2612	3.106	6696.5125	3.012	6696.5312	3.012	
6980.4713	5.259	7331.4671	5.223	6979.4254	2.820	6680.2780	3.100	6696.5254	3.011	6696.5445	3.011	
6980.5113	5.263	7331.4677	5.221	6981.2509	2.844	6680.2946	3.095	6696.5534	3.010	6696.5725	3.010	
6980.5421	5.258	7331.4682	5.218	6981.2660	2.841	6680.3121	3.085	6696.5824	3.009	6696.5916	3.009	
6980.5607	5.251	7331.4687	5.221	6981.2799	2.836	6680.3288	3.082	6696.5986	3.008	6696.6175	3.008	
6980.5607	5.251	7331.4687	5.221	6981.2955	2.829	6680.3480	3.076	6696.6264	3.007	6696.6452	3.007	
6980.5783	5.247	7331.4692	5.221	6981.3107	2.821	6680.3667	3.076	6696.6372	3.006	6696.6561	3.006	
6981.2610	5.255	7331.5250	5.225	6981.3257	2.819	6680.3843	3.083	6696.6429	3.003	6696.6615	3.003	
6981.2937	5.271	7331.5665	5.218	6981.3394	2.818	6680.3939	3.094	6696.6701	3.073	6696.6891	3.073	
6981.3532	5.268	7332.5677	5.243	6982.1509	2.820	6680.4043	3.083	6696.6791	3.072	6696.6981	3.072	
6981.4029	5.239	7332.5683	5.240	6982.1556	2.815	6680.4148	3.081	6696.6881	3.071	6696.7081	3.071	
6981.4581	5.215	7332.5689	5.238	6982.1621	2.812	6680.4339	3.096	6696.7171	3.070	6696.7271	3.070	
6981.4607	5.251	7332.5694	5.240	6982.1806	2.813	6680.4529	3.083	6697.0371	3.076	6697.0567	3.076	
6981.4985	5.228	7332.6038	5.240	6982.2199	2.836	6680.4716	3.082	6697.3117	3.088	6697.3309	3.088	
6981.5313	5.239	7332.6038	5.240	6982.2955	2.829	6680.4897	3.087	6697.3555	3.089	6697.3747	3.089	
6981.5542	5.263	7332.6237	5.262	6982.4229	2.817	6680.5093	3.080	6698.2447	3.047	6698.2634	3.047	
6982.2194	5.222	7332.6237	5.262	6982.4324	2.819	6680.5278	3.066	6698.2861	3.066	6698.3052	3.066	
6982.3116	5.263	7332.6252	5.262	6982.4478	2.830	6680.5469	3.075	6698.2872	3.074	6698.3052	3.074	
6982.3702	5.282	7332.6254	5.224	6982.1501	2.820	6680.5651	3.067	6698.3111	3.071	6698.3302	3.071	
6982.4236	5.303	7332.6254	5.224	6982.1621	2.819	6680.5845	3.059	6698.3325	3.095	6698.3517	3.095	
6982.4745	5.279	7332.6254	5.224	6982.1806	2.813	6680.6035	3.049	6699.2411	3.065	6699.2604	3.065	
6983.2925	5.258	7332.6254	5.224	6982.4009	2.822	6680.6333	3.039	6699.2642	3.066	6699.2831	3.066	
6983.3381	5.249	7332.6254	5.224	6982.4130	2.824	6680.6352	3.024	6700.2415	3.072	6700.2640	3.064	
6983.4237	5.262	7332.6254	5.224	6982.4245	2.826	6680.6371	3.012	6700.2861	3.066	6700.3049	3.066	
7321.4397	5.240	7321.5138	5.243	6982.4323	2.827	6680.6390	3.005	6700.3231	3.084	6700.3412	3.084	
7321.5677	5.243	7322.5677	5.243	6982.4400	2.830	6680.6224	3.080	6700.3412	3.087	6700.3592	3.087	
7321.5852	5.246	7322.5683	5.246	6982.4478	2.830	6680.6246	3.096	6700.3592	3.096	6700.3777	3.096	
7321.6264	5.244	7322.5689	5.244	6982.4594	2.830	6680.6264	3.063	6703.2611	3.081	6703.2801	3.081	
7322.3054	5.244	7322.5694	5.244	6982.2632	2.829	6680.6284	3.073	6703.2831	3.096	6703.3059	3.098	
7322.3637	5.267	7322.5694	5.244	6982.2777	2.824	6680.6305	3.064	6703.3261	3.092	6703.3456	3.092	
7322.4184	5.265	7322.5694	5.244	6982.2952	2.819	6680.6319	3.066	6703.3481	3.092	6703.3677	3.092	
7322.4237	5.262	7322.5694	5.244	6982.3087	2.821	6680.6319	3.066	6703.3681	3.092	6703.3877	3.092	
7322.4397	5.240	7322.5694	5.244	6982.3225	2.814	6680.6330	3.099	6703.3891	3.092	6703.4081	3.092	
7322.4897	5.246	7322.5694	5.244	6982.3355	2.818	6680.6390	3.098	6703.3977	3.084	6703.4164	3.084	
7322.5138	5.243	7322.5694	5.244	6982.3482	2.818	6680.6414	3.095	6703.4164	3.080	6703.4354	3.080	
7322.5264	5.244	7322.5694	5.244	6982.3614	2.821	6680.6350	3.089	6703.4254	3.085	6703.4441	3.085	
7322.5747	5.256	7322.5694	5.244	6982.3748	2.824	6680.6378	3.068	6703.4524	3.086	6703.4712	3.086	
7322.6020	5.257	7322.5694	5.244	6982.3878	2.826	6680.6393	3.097	6703.4515	3.082	6703.4705	3.082	
7322.6232	5.256	7322.5694	5.244	6982.4013	2.829	6680.6411	3.094	6703.4598	3.082	6703.4788	3.082	
7322.6369	5.261	7322.5694	5.244	6982.4143	2.824	6680.6429	3.095	6703.4681	3.081	6703.4871	3.081	
7322.6741	5.261	7322.5694	5.244	6982.4277	2.825	6680.6447	3.096	6703.4761	3.081	6703.4951	3.081	
7322.7050	5.249	7322.5694	5.244	6982.4411	2.826	6680.6466	3.095	6703.4841	3.081	6703.5030	3.081	
7322.7367	5.261	7322.5694	5.244	6982.4546	2.827	6680.6484	3.095	6703.4931	3.081	6703.5121	3.081	
7322.7675	5.261	7322.5694	5.244	6982.4680	2.828	6680.6503	3.094	6703.5022	3.081	6703.5211	3.081	
7322.7984	5.261	7322.5694	5.244	6982.4814	2.829	6680.6521	3.094	6703.5111	3.081	6703.5301	3.081	
7322.8292	5.261	7322.5694	5.244	6982.4948	2.830	6680.6540	3.095	6703.5199	3.081	6703.5391	3.081	
7322.8601	5.201	7322.5694	5.244	6982.5084	2.824	6680.6556	3.094	6703.5281	3.081	6703.5471	3.081	
7322.8909	5.206	7322.5694	5.244	6982.5219	2.825	6680.6576	3.093	6703.5371	3.081	6703.5561	3.081	
7322.9101	5.201	7322.5694	5.244	6982.5353	2.826	6680.6595	3.092	6703.5461	3.081	6703.5651	3.081	
7322.9302	5.203	7322.5694	5.244	6982.5487	2.827	6680.6614	3.091	6703.5551	3.081	6703.5741	3.081	
7322.9504	5.203	7322.5694	5.244	6982.5621	2.828	6680.6633	3.090	6703.5641	3.081	6703.5831	3.081	
7322.9704	5.203	7322.5694	5.244	6982.5755	2.829	6680.6651	3.090	6703.5731	3.081	6703.5921	3.081	
7322.9904	5.203	7322.5694	5.244	6982.5889	2.830	6680.6669	3.090	6703.5821	3.081	6703.6011	3.081	
7323.0104	5.203	7322.5694	5.244	6982.6023	2.831	6680.6687	3.090	6703.5911	3.081	6703.6101	3.081	
7323.0304	5.203	7322.5694	5.244	6982.6157	2.832	6680.6705	3.090	6703.5901	3.081	6703.6091	3.081	
7323.0514	5.203	7322.5694	5.244	6982.6291	2.833	6680.6723	3.090	6703.5891	3.081	6703.6181	3.081	
7323.0714												

TABLE IIa (*continued*)

HR 8260	HJD	b										
	6278.6244	4.414	6363.3383	4.423	6682.4108	4.434			6696.3946	4.439	6962.5653	4.369
HJD	b		6363.3541	4.423	6682.4259	4.430	RJD	b	6966.4163	4.432	6962.5880	4.369
6271.4769	4.431		6363.3698	4.419	6682.4418	4.431	6691.2644	4.424	6966.4374	4.434	6962.6141	4.379
6271.4901	4.432		6281.4625	4.427	6682.4565	4.433	6691.2878	4.427	6967.2471	4.425	6965.4161	4.367
6271.5035	4.429		6281.4753	4.428	6363.4002	4.423	6691.3089	4.424	6967.2690	4.426	6965.4432	4.364
6271.5166	4.429		6281.4870	4.426	6364.2602	4.421	6691.3333	4.428	6967.2927	4.435	6965.5139	4.370
6272.3434	4.428		6281.4980	4.426	6364.2770	4.419	6691.3585	4.432	6967.3161	4.435	6965.5574	4.369
6272.3557	4.431		6281.5119	4.422	6367.2909	4.446	6691.3818	4.432	6967.3391	4.445	6965.6211	4.362
6272.3697	4.433		6281.5239	4.421	6367.3075	4.453	6691.4031	4.432	6967.3647	4.430	6965.6369	4.362
6272.5832	4.430		6281.5364	4.418	6367.3765	4.437	6691.4214	4.432	6967.3859	4.431	6965.6520	4.362
6274.5166	4.438		6281.5499	4.417	6367.3932	4.437	6691.4387	4.434	6967.4071	4.429	6966.4534	4.360
6274.5292	4.439		6281.5619	4.417	6367.4102	4.437	6691.4628	4.435	6968.2526	4.417	6966.4934	4.364
6274.5401	4.437		6281.5750	4.417	6367.4275	4.440	6691.4803	4.436	6968.2733	4.427	6966.5399	4.365
6274.5515	4.435		6281.5861	4.412	6367.4440	4.443	6691.4969	4.438	6968.2952	4.427	6966.5571	4.366
6276.4542	4.417		6281.5975	4.411	6367.4626	4.442	6691.5146	4.434	6968.3185	4.427	6966.5840	4.367
6276.4677	4.415		6281.6084	4.409	6367.4820	4.447	6692.4238	4.438	6968.3418	4.436	6966.5999	4.366
6276.4794	4.414		6281.6205	4.407	6367.5037	4.451	6692.4465	4.436	6968.3622	4.437	6966.6150	4.372
6276.4907	4.414		6281.6313	4.410	6367.5202	4.448	6692.4671	4.434	6969.2498	4.434	6966.6330	4.366
6276.5022	4.417		6281.4729	4.411	6367.5371	4.447	6692.4892	4.438	6969.2723	4.438	6966.6480	4.369
6276.5141	4.417		6281.4848	4.404	6367.5536	4.448	6692.5119	4.438	6970.2496	4.431	6967.4131	4.362
6276.5252	4.415		6281.4960	4.411	6367.5673	4.441	6693.2342	4.423	6970.2720	4.430	6967.4512	4.367
6276.5377	4.417		6281.5070	4.413	6367.3667	4.434	6693.2578	4.424	6970.2946	4.433	6967.4948	4.360
6276.5503	4.417		6281.5181	4.417	6368.2722	4.424	6693.2803	4.425	6970.3192	4.430	6967.5443	4.354
6276.5633	4.414		6281.5289	4.419	6367.4109	4.435	6693.3030	4.433	6970.3453	4.426	6967.5690	4.348
6276.5745	4.415		6282.5400	4.421	6367.4288	4.441	6693.3262	4.434	6970.4444	4.432	6967.5842	4.355
6276.5848	4.417		6282.5500	4.422	6367.4454	4.448	6693.3524	4.436	6970.4665	4.431	6967.6000	4.356
6276.5973	4.413		6282.5675	4.421	6367.4618	4.443	6693.3766	4.437	6970.4865	4.435	6967.6167	4.358
6276.6082	4.409		6282.5800	4.424	6367.4783	4.447	6693.3996	4.438	6970.5031	4.435	6967.6299	4.356
6276.6183	4.411		6282.5924	4.422	6367.4977	4.453	6693.4194	4.437	6970.5210	4.426	6967.6409	4.359
6276.6287	4.412		6282.6047	4.424	6367.5132	4.451	6693.4389	4.438	6970.5314	4.428	6967.6514	4.358
6276.6391	4.413		6282.6166	4.421	6367.5253	4.440	6693.4650	4.444	6970.5373	4.433	6970.5399	4.365
6276.6494	4.414		6282.6288	4.420	6367.5538	4.462	6693.4876	4.445	6970.5569	4.436	6970.5701	4.365
6277.4611	4.413		6282.6401	4.419	6368.2267	4.427	6693.5092	4.456	6970.5777	4.434	6970.5966	4.355
6277.4735	4.412		6334.3637	4.563	6368.2489	4.429	6694.2395	4.419	6970.5984	4.434	6970.6248	4.357
6277.4844	4.415		6334.3843	4.563	6368.2564	4.430	6694.2624	4.421	6970.6184	4.436	6970.6431	4.358
6277.4961	4.417		6334.4024	4.563	6368.2845	4.429	6694.2857	4.434	6970.6391	4.433	6972.5242	4.357
6277.5070	4.418		6334.4203	4.560	6368.3027	4.430	6695.2841	4.444	6970.6472	4.433	6972.5463	4.356
6277.5190	4.417		6334.4366	4.554	6368.3205	4.436	6695.3071	4.430	6970.6574	4.354	7322.5107	4.440
6277.5305	4.422		6334.4563	4.555	6368.3383	4.436	6695.3664	4.429	6970.6591	4.355	7322.5709	4.458
6277.5414	4.419		6334.4702	4.549	6368.3308	4.401	6694.3802	4.431	6970.6570	4.357	7322.6140	4.482
6277.5560	4.415		6334.4874	4.546	6368.3777	4.436	6694.3913	4.456	6970.6777	4.434	7322.6436	4.548
6277.6151	4.421		6338.3946	4.490	6368.3996	4.437	6694.4295	4.419	6970.7184	4.434	7328.5115	4.492
6277.6267	4.416		6338.4120	4.485	6368.4287	4.441	6695.2616	4.424	6970.7224	4.433	7328.5675	4.484
6277.6373	4.420		6338.4271	4.483	6368.4388	4.442	6695.2806	4.442	6970.7316	4.435	7328.6114	4.489
6277.6479	4.414		6338.4433	4.483	6368.4540	4.441	6695.3071	4.430	6970.7387	4.432	7328.6558	4.501
6277.6552	4.410		6338.4580	4.488	6368.4736	4.442	6695.3343	4.436	6970.7602	4.438	7329.4096	4.529
6278.4497	4.423		6338.4733	4.461	6368.4975	4.444	6695.3605	4.435	6970.7745	4.432	7329.4673	4.533
6278.4872	4.419		6339.3485	4.467	6368.5145	4.445	6695.3833	4.435	6970.7882	4.434	7329.5264	4.528
6278.5007	4.418		6339.3657	4.467	6368.5229	4.433	6695.4048	4.438	6970.8011	4.436	7329.5846	4.536
6278.5137	4.413		6339.3813	4.468	6368.5206	4.418	6695.4267	4.438	6970.8244	4.436	7329.6198	4.537
6278.5262	4.414		6339.4034	4.471	6368.5263	4.417	6695.4505	4.440	6970.8476	4.437	7329.6659	4.534
6278.5389	4.408		6339.4194	4.471	6368.5284	4.421	6695.4841	4.444	6970.8648	4.437	7328.6779	4.530
6278.5513	4.412		6339.4343	4.471	6368.5302	4.422	6695.5071	4.443	6970.8866	4.437	7330.4067	4.522
6278.5644	4.410		6339.4561	4.465	6368.5318	4.425	6696.2796	4.427	6970.9002	4.436	7330.4123	4.528
6278.5768	4.407		6363.2663	4.424	6368.3356	4.422	6696.3018	4.428	6970.9469	4.437	7330.4691	4.532
6278.5890	4.407		6363.2835	4.423	6368.3545	4.426	6696.5047	4.427	6970.9629	4.439	7330.5327	4.528
6278.6008	4.411		6363.3067	4.423	6368.3739	4.427	6696.5271	4.425	6970.9811	4.435	7330.5640	4.538
6278.6129	4.412		6363.3226	4.424	6368.3924	4.423	6696.3726	4.431	6978.5398	4.355	7330.5967	4.551

HR 8260	HJD	b										
	6278.6184	4.710	6363.3791	4.677	6682.3940	4.675			6682.5099	4.685		
HJD	b		6278.6138	4.709	6364.2374	4.673	RJD	b	6682.5250	4.684		
6271.4695	4.708		6278.6434	4.702	6364.2530	4.673	6271.4840	4.709	6364.2703	4.671	6271.5013	4.687
6271.4968	4.708		6281.4499	4.711	6367.2850	4.676	6271.5103	4.707	6367.3011	4.673	6272.4631	4.684
6271.5228	4.706		6281.4932	4.715	6367.3168	4.674	6272.5276	4.697	6367.3386	4.678	6272.5907	4.698
6272.5502	4.697		6281.5179	4.716	6367.3494	4.680	6272.5619	4.696	6367.3713	4.685	6272.6342	4.688
6272.5749	4.706		6281.5529	4.719	6367.3651	4.682	6272.5759	4.698	6367.3843	4.685	6272.6463	4.694
6272.5829	4.706		6281.5558	4.711	6367.3823	4.687	6272.5829	4.698	6367.3943	4.682	6272.6565	4.698
6272.5948	4.707		6281.5678	4.711	6367.3959	4.686	6272.5958	4.698	6367.4077	4.682	6272.6698	4.698
6272.6053	4.705		6281.5805	4.711	6367.4186	4.688	6272.6053	4.698	6367.4337	4.691	6272.6841	4.688
6272.6159	4.697		6281.5921	4.710	6367.4366	4.687	6272.6159	4.697	6367.4508	4.685	6272.6944	4.689
6272.6228	4.692		6281.6030	4.706	6367.4525	4.689	6272.6228	4.692	6367.4737	4.685	6272.6977	4.688
6272.6474	4.709		6281.61									

PHOTOMETRY OF BE STARS

TABLE IIa (*continued*)

HR 8402	HJD	b	HJD	b	HJD	b	HR 8408	HJD	b	HJD	b	HJD	b
HJD	b	HJD	b	HJD	b	HJD	b	HJD	b	HJD	b	HJD	b
6698.3691	4.692	6966.6191	4.659	6981.5856	4.665	6278.5217	5.895	6334.4154	5.896	6680.2855	5.884	6680.4007	5.895
6699.2565	4.693	6966.6404	4.660	6981.6074	4.670	6278.5341	5.898	6334.4320	5.897	6680.3037	5.887	6680.3215	5.890
6691.5027	4.679	6700.2559	4.678	6966.6558	4.659	6981.6224	4.668	6271.4724	5.896	6334.4482	5.899	6680.3369	5.897
6691.5201	4.683	6700.2782	4.677	6967.4561	4.669	6981.6369	4.664	6271.4856	5.897	6278.5592	5.898	6680.3394	5.894
6692.2414	4.687	6700.3009	4.680	6967.4997	4.668	6981.6522	4.664	6271.4993	5.896	6278.5722	5.900	6334.4823	5.901
6692.4314	4.686	6700.3255	4.679	6967.5492	4.677	6981.6646	4.661	6271.5121	5.900	6278.5845	5.901	6680.3786	5.897
6692.4528	4.681	6700.3517	4.680	6967.5729	4.677	6983.5971	4.673	6271.5246	5.900	6278.5945	5.903	6338.4064	5.899
6692.4738	4.684	6702.4504	4.676	6967.5882	4.675	6983.6142	4.677	6272.5392	5.893	6278.6083	5.904	6338.4227	5.896
6692.4963	4.685	6702.4734	4.672	6967.6051	4.679	6983.6336	4.671	6272.5518	5.895	6278.6203	5.908	6338.4387	5.896
6692.5195	4.689	6703.2767	4.688	6967.6230	4.676	6983.6486	4.669	6272.5634	5.897	6278.6335	5.907	6680.4551	5.881
6693.2409	4.681	6703.2985	4.689	6967.6353	4.677	6983.6601	4.669	6272.5791	5.896	6278.6448	5.910	6338.4682	5.892
6693.2641	4.686	6703.3209	4.689	6967.6460	4.678	7322.5149	4.669	6274.5119	5.885	6278.6569	5.914	6680.4988	5.877
6693.2867	4.671	6703.3435	4.690	6967.6562	4.679	7322.5751	4.664	6274.5245	5.887	6278.6635	5.910	6339.3613	5.906
6693.3090	4.688	6703.3630	4.692	6970.5504	4.682	7322.6182	4.667	6274.5364	5.886	6278.6691	5.904	6682.2314	5.880
6693.3321	4.691	6703.3819	4.691	6970.5769	4.681	7327.6153	4.670	6274.5477	5.885	6278.6752	5.909	6339.4144	5.887
6693.3582	4.689	6703.4046	4.693	6970.6098	4.682	7328.5772	4.674	6276.4492	5.884	6281.4572	5.896	6682.2693	5.876
6693.3829	4.690	6703.4247	4.691	6970.6338	4.680	7328.6177	4.673	6276.4637	5.887	6281.4712	5.897	6682.2857	5.879
6693.4058	4.692	6703.4592	4.692	6970.6501	4.687	7328.6623	4.668	6274.4757	5.883	6281.4829	5.892	6682.3021	5.879
6693.4258	4.696	6703.4783	4.689	6972.5302	4.674	7329.5888	4.667	6276.4870	5.886	6281.4947	5.890	6682.3195	5.879
6693.4447	4.689	6958.4507	4.654	6972.5521	4.676	7329.6242	4.672	6274.4983	5.885	6281.5080	5.893	6682.3366	5.881
6693.4711	4.700	6959.5155	4.672	6972.5749	4.674	7329.6704	4.659	6276.5103	5.889	6281.5194	5.893	6682.3555	5.883
6693.4939	4.704	6959.5362	4.675	6972.5941	4.674	7329.6818	4.668	6276.5216	5.889	6281.5326	5.891	6682.3747	5.880
6693.5153	4.705	6959.5394	4.665	6972.6075	4.675	7330.5685	4.670	6276.5340	5.893	6281.5458	5.893	6682.3933	5.879
6694.2466	4.680	6959.6164	4.671	6972.6206	4.674	7330.6022	4.667	6276.5463	5.893	6281.5575	5.892	6682.4116	5.886
6694.2686	4.684	6959.6358	4.668	6972.6340	4.677	7332.5907	4.671	6276.5576	5.900	6281.5693	5.893	6682.4267	5.880
6694.3164	4.682	6959.6642	4.666	6972.6590	4.678	6972.6722	4.675	6276.5706	5.901	6281.5820	5.893	6682.4427	5.882
6694.3731	4.679	6959.6770	4.669	6972.6845	4.677	6972.6845	4.678	6276.5813	5.900	6281.5935	5.893	6354.2552	5.890
6694.4352	4.689	6959.6915	4.671	6972.6963	4.671	6972.6974	4.675	6276.5937	5.897	6281.6045	5.894	6682.4575	5.877
6694.4964	4.691	6960.5676	4.676	6972.7080	4.677	6972.7080	4.676	6276.6040	5.902	6281.6155	5.896	6367.2865	5.892
6694.4977	4.701	6960.5930	4.674	6975.5713	4.680	6978.4733	4.684	6276.6147	5.905	6281.6273	5.894	6682.5057	5.871
6695.3043	4.685	6960.6100	4.679	6975.5888	4.679	6978.5453	4.679	6276.6251	5.904	6281.6374	5.898	6682.5208	5.871
6695.3670	4.686	6960.6282	4.684	6975.6034	4.677	6978.5888	4.675	6276.6355	5.905	6281.6433	5.895	6682.5365	5.871
6695.3894	4.684	6960.6410	4.677	6975.6186	4.676	6978.6186	4.676	6276.6456	5.902	6281.6495	5.895	6682.5544	5.873
6695.4112	4.686	6960.6523	4.674	6975.6332	4.676	6978.6472	4.675	6276.6565	5.906	6281.6635	5.894	6682.5730	5.870
6695.4332	4.688	6960.6648	4.676	6975.6472	4.675	6978.6673	4.676	6277.4562	5.903	6281.6773	5.896	6684.3424	5.893
6696.2369	4.692	6960.6766	4.663	6975.6613	4.678	6978.6877	4.677	6277.4696	5.904	6282.6482	5.887	6684.3639	5.889
6696.2860	4.691	6961.4431	4.687	6978.6733	4.684	6978.6959	4.676	6277.4805	5.903	6282.4807	5.890	6676.5047	5.902
6696.3081	4.690	6961.4860	4.680	6978.6924	4.681	6978.7143	4.680	6277.4925	5.902	6282.4922	5.892	6676.5211	5.901
6696.3310	4.694	6962.5431	4.667	6978.5453	4.677	6978.5531	4.680	6277.5031	5.901	6282.5035	5.895	6684.4270	5.886
6696.3573	4.692	6962.5709	4.669	6978.5647	4.677	6978.5711	4.678	6277.5145	5.900	6282.5154	5.897	6684.4425	5.875
6696.3793	4.689	6962.5946	4.675	6978.5844	4.677	6978.5914	4.678	6277.5268	5.902	6282.5257	5.899	6676.5693	5.896
6696.4008	4.664	6962.6235	4.671	6978.6002	4.674	6978.6102	4.672	6276.5379	5.906	6282.6365	5.900	6676.4466	5.883
6696.4223	4.693	6962.6440	4.678	6980.4999	4.676	6980.6415	4.679	6277.5512	5.902	6282.6470	5.905	6677.3482	5.866
6696.4436	4.685	6965.4234	4.680	6980.5312	4.681	6980.5499	4.682	6277.5636	5.904	6282.6482	5.887	6684.3639	5.889
6697.2534	4.682	6965.4699	4.676	6980.5499	4.682	6980.5765	4.694	6277.4805	5.903	6282.6576	5.904	6677.3907	5.867
6697.2758	4.685	6965.5199	4.682	6980.5689	4.681	6980.6024	4.681	6277.5337	5.890	6282.5890	5.905	6677.4119	5.869
6697.2944	4.685	6965.5617	4.672	6980.5880	4.680	6972.5236	5.906	6277.5444	5.895	6282.6132	5.906	6677.4463	5.873
6697.3224	4.687	6965.6022	4.683	6980.6099	4.680	6980.6138	4.681	6277.5565	5.906	6282.6252	5.905	6677.4627	5.872
6697.3456	4.676	6965.6262	4.677	6980.6138	4.681	6980.6248	4.682	6277.5682	5.907	6282.6366	5.908	6686.3415	5.872
6697.3712	4.677	6965.6423	4.678	6980.6284	4.684	6980.6415	4.679	6277.5741	5.907	6282.6471	5.879	6677.4985	5.883
6697.3921	4.678	6965.6570	4.679	6980.6570	4.679	6980.6645	4.682	6277.5826	5.908	6282.6545	5.907	6677.5141	5.873
6697.4133	4.675	6966.4584	4.672	6980.6527	4.677	6980.6745	4.677	6277.5938	5.909	6282.6620	5.904	6677.5334	5.872
6698.2592	4.686	6966.4987	4.669	6980.6653	4.675	6980.6877	4.687	6278.4446	5.887	6282.6670	5.904	6677.5334	5.872
6698.2806	4.690	6966.5345	4.667	6981.4841	4.675	6981.5181	4.674	6278.4582	5.889	6282.6690	5.903	6677.5547	5.880
6698.3017	4.691	6966.5712	4.661	6981.5181	4.674	6981.5591	5.900	6278.4829	5.895	6314.3564	5.894	6680.2283	5.878
6698.3246	4.690	6966.5883	4.662	6981.5414	4.674	6981.5893	5.904	6278.4957	5.893	6314.3792	5.894	6680.2500	5.879
6698.3481	4.689	6966.6037	4.661	6981.5674	4.670	6981.5974	5.903	6278.5089	5.893	6334.3977	5.895	6680.2665	5.883

HR 8408	HJD	b	HJD	b	HJD	b	HR 8408	HJD	b	HJD	b	HJD	b
6693.4661	5.884	6703.2714	5.885	6967.6174	5.886	6981.6219	5.886	6981.6606	5.894	6982.3878	5.906	6982.4000	5.887
HJD	b	6693.4987	5.880	6703.3159	5.883	6967.6431	5.886	6982.3878	5.906	6983.4482	5.899	6983.4682	5.884
6698.4901	5.879	6693.5292	5.892	6703.3385	5.891	6967.6651	5.880	6982.4040	5.903	6983.4738	5.897	6983.4959	5.886
6698.5076	5.879	6694.2409	5.884	67									

TABLE IIa (*continued*)

HR 8858	HJD	b	HJD	b	HJD	b	HR 8858	HJD	b
	6682.4665 4.346		6692.5017 4.339		6700.2841 4.344			6983.6001 4.328	
HJD	b		6682.4823 4.363		6692.5257 4.346		6983.6180 4.320		
6676.3843 4.324			6682.4971 4.341		6693.2458 4.346		6983.6388 4.324		
6676.4019 4.329			6682.5135 4.343		6693.2693 4.345		6983.6510 4.322		
6676.4192 4.330			6682.5289 4.344		6693.2913 4.354		6983.6623 4.326		
6676.4355 4.333			6682.5453 4.346		6693.3143 4.352		6983.6777 4.327		
6676.4359 4.337			6682.5615 4.347		6693.3366 4.356		6983.6878 4.328		
6676.4716 4.339			6682.5776 4.347		6693.3627 4.349		6983.6977 4.328		
6676.4939 4.343			6682.5908 4.344		6693.3874 4.353		6983.7022 4.328		
6676.5127 4.342			6682.6036 4.344		6693.4107 4.346		6983.7139 4.327		
6676.5294 4.349			6684.3513 4.330		6693.4305 4.348		6983.7228 4.327		
6676.5464 4.343			6684.3726 4.329		6693.4502 4.339		6983.7328 4.327		
6676.5623 4.342			6684.3962 4.322		6693.4755 4.335		6983.7428 4.327		
6676.5747 4.347			6684.4185 4.331		6693.4987 4.335		6983.7529 4.327		
6676.5898 4.341			6684.4332 4.329		6693.5202 4.325		6983.7629 4.327		
6676.4024 4.341			6684.4512 4.338		6693.5372 4.316		6983.7729 4.327		
6677.3567 4.344			6684.4962 4.339		6693.5524 4.312		6983.7830 4.327		
6677.3754 4.353			6686.2812 4.343		6693.5647 4.309		6983.7930 4.327		
6677.3980 4.328			6686.2987 4.343		6694.2514 4.320		6983.8030 4.327		
6677.4197 4.327			6686.3137 4.342		6694.2981 4.333		6983.8130 4.327		
6677.4380 4.319			6686.3317 4.345		6694.3221 4.338		6983.8230 4.327		
6677.4543 4.324			6686.3492 4.347		6694.3513 4.341		6983.8330 4.327		
6677.4705 4.310			6686.3680 4.347		6694.3779 4.324		6983.8430 4.327		
6677.4879 4.322			6686.3874 4.351		6694.3993 4.334		6983.8530 4.327		
6677.3059 4.319			6686.4054 4.349		6694.4226 4.336		6983.8630 4.327		
6677.5223 4.317			6686.4219 4.352		6694.4471 4.327		6983.8730 4.327		
6677.5412 4.333			6686.4365 4.353		6695.2731 4.336		6983.8830 4.327		
6677.5563 4.332			6686.4519 4.354		6695.3186 4.330		6983.8930 4.327		
6677.5792 4.326			6686.4666 4.352		6695.3453 4.330		6983.9030 4.327		
6680.2589 4.324			6686.4823 4.348		6695.3725 4.345		6983.9130 4.327		
6680.2748 4.322			6686.4994 4.348		6695.3940 4.347		6983.9230 4.327		
6680.2937 4.334			6686.5167 4.339		6695.4162 4.349		6983.9330 4.327		
6680.3115 4.334			6686.5325 4.337		6695.4384 4.349		6983.9430 4.327		
6680.3296 4.329			6686.5485 4.332		6695.5187 4.339		6983.9530 4.327		
6680.3475 4.330			6686.5592 4.333		6695.5354 4.370		6983.9630 4.327		
6680.3650 4.338			6686.5702 4.328		6696.2699 4.332		6983.9730 4.327		
6680.3865 4.343			6686.5829 4.330		6696.2919 4.333		6983.9830 4.327		
6680.4088 4.348			6686.5947 4.329		6696.3132 4.342		6983.9930 4.327		
6680.4312 4.352			6688.2442 4.306		6696.3362 4.344		6980.3454 4.327		
6680.4471 4.362			6688.2636 4.314		6696.3629 4.341		6980.3556 4.327		
6680.4639 4.359			6688.2995 4.336		6696.3845 4.341		6980.3713 4.327		
6680.4906 4.353			6688.3175 4.331		6696.4058 4.336		6980.3907 4.327		
6680.5070 4.354			6688.3360 4.335		6696.4280 4.345		6980.6034 4.327		
6680.5412 4.353			6690.3101 4.330		6696.4491 4.334		6980.6165 4.327		
6680.5580 4.350			6690.3343 4.336		6697.2588 4.334		6980.6293 4.327		
6680.5804 4.366			6690.3702 4.323		6697.2817 4.331		6980.6441 4.327		
6682.2413 4.346			6690.3996 4.335		6697.3046 4.332		6980.6592 4.327		
6682.2603 4.342			6690.4272 4.327		6697.3287 4.335		6980.6676 4.327		
6682.2768 4.337			6690.4497 4.327		6697.3509 4.334		6981.4877 4.324		
6682.2937 4.339			6690.4753 4.340		6697.3769 4.334		6981.5214 4.333		
6682.3106 4.335			6690.4962 4.335		6697.3983 4.342		6981.5465 4.335		
6682.3271 4.335			6690.5167 4.342		6697.4195 4.339		6981.5722 4.334		
6682.3439 4.338			6690.5384 4.340		6698.2647 4.349		6981.5970 4.331		
6682.3614 4.336			6690.5549 4.349		6698.2863 4.349		6981.6115 4.339		
6682.4011 4.328			6690.5787 4.353		6698.3081 4.344		6981.6253 4.336		
6682.4189 4.352			6692.4370 4.346		6698.3303 4.340		6981.6408 4.334		
6682.4346 4.338			6692.4575 4.344		6698.3537 4.340		6981.6549 4.331		
6682.4501 4.339			6692.4789 4.342		6700.2636 4.353		6981.6669 4.325		

TABLE IIb. — Four-colour Strömgren photometry for periodic Be stars observed at ESO. Observations between JD2446991 and 2447000 were obtained at SAAO. The heliocentric Julian day is with respect to JD 2440000.000.

HR 5193	HJD	u	v	b	y	HR 5440	HJD	u	v	b	y			
6961.6036	3.954	3.687	3.418	3.444	6973.5925	3.874	3.641	3.365	3.383	6973.5925	2.839	2.599	2.326	2.348
6961.6320	3.953	3.685	3.416	3.441	6973.6791	3.849	3.634	3.358	3.378	6965.5139	2.768	2.542	2.275	2.300
6961.6522	3.952	3.685	3.417	3.446	6975.5990	3.942	3.674	3.399	3.430	6965.5241	2.751	2.531	2.264	2.288
6961.6780	3.950	3.684	3.416	3.444	6975.6156	3.928	3.659	3.394	3.428	6965.5919	2.759	2.535	2.267	2.292
6961.7008	3.948	3.679	3.410	3.438	6975.6326	3.946	3.671	3.399	3.427	6965.6300	2.776	2.545	2.278	2.307
6961.7379	3.947	3.680	3.412	3.442	6975.6553	3.923	3.661	3.386	3.417	6965.6919	2.803	2.564	2.297	2.330
6963.5516	3.926	3.682	3.400	3.434	6979.4812	3.959	3.688	3.413	3.443	6965.7219	2.819	2.572	2.307	2.338
6963.5475	3.929	3.665	3.394	3.420	6979.5282	3.955	3.685	3.408	3.439	6979.5722	2.822	2.592	2.318	2.342
6963.5856	3.929	3.668	3.395	3.421	6979.5701	3.957	3.688	3.411	3.441	6979.6116	2.800	2.569	2.294	2.321
6963.6102	3.929	3.670	3.397	3.423	6979.6095	3.957	3.685	3.406	3.441	6965.6300	2.776	2.545	2.278	2.307
6963.6385	3.933	3.668	3.395	3.422	6979.6422	3.957	3.683	3.406	3.438	6979.6442	2.790	2.561	2.287	2.314
6963.6633	3.933	3.670	3.395	3.422	6980.4793	3.953	3.678	3.409	3.443	6965.6919	2.803	2.564	2.297	2.330
6963.6956	3.935	3.673	3.401	3.442	6980.4823	3.958	3.680	3.413	3.446	6980.5227	2.832	2.582	2.316	2.346
6963.7329	3.935	3.672	3.398	3.425	6980.5364	3.962	3.682	3.414	3.444	6965.7645	2.832	2.583	2.317	2.350
6965.5115	3.918	3.650	3.384	3.410	6980.6022	3.960	3.682	3.412	3.443	6980.6065	2.835	2.582	2.315	3.43
6965.5525	3.910	3.650	3.378	3.405	6980.6627	3.955	3.679	3.409	3.444	6967.5373	2.787	2.557	2.289	2.313
6965.5905	3.907	3.649	3.378	3.404	6981.4869	3.978	3.694	3.424	3.453	6980.6690	2.839	2.584	2.317	3.45
6965.6299	3.907	3.647	3.378	3.402	6981.5449	3.969	3.691	3.419	3.450	6987.5962	2.791	2.562	2.292	2.318
6965.6895	3.899	3.644	3.376	3.400	6981.5840	3.968	3.691	3.420	3.451	6981.5468	2.783	2.558	2.289	2.315
6965.7197	3.908	3.646	3.375	3.401	6981.6247	3.969	3.691	3.417	3.450	6967.6731	2.826	2.583	2.313	2.340
6967.4974	3.882	3.647	3.370	3.391	6981.6698	3.968	3.690	3.416	3.455	6967.7011	2.829	2.585	2.315	2.344
6967.5349	3.887	3.673	3.385	3.412	6982.6634	3.977	3.688	3.416	3.451	6987.7403	2.832	2.580	2.308	2.337
6967.5629	3.883	3.642	3.369	3.388	6986.5214	2.931	2.657	2.380	2.406	6980.6656	2.848	2.593	2.327	2.362
6967.5947	3.872	3.642	3.365	3.386	6986.5560	2.944	2.675	2.382	2.422	6981.4893	2.816	2.584	2.316	2.341
6967.6322	3.873	3.642	3.363	3.382	6986.5938	2.948	2.682	2.388	2.429	6986.5938	2.948	2.682	2.388	2.429
6967.6716	3.872	3.642	3.364	3.380	6986.6294	2.976	2.689	2.427	2.442	6986.6294	2.976	2.689	2.427	2.442
6967.6989	3.876	3.642	3.365	3.383	6986.6322	2.946	2.683	2.411	2.432	6986.6666	2.920	2.658	2.388	2.412
6968.4851	3.934	3.677	3.402	3.433	6986.7123	2.857	2.601	2.333	2.360	6986.7123	2.857	2.601	2.333	2.360
6968.5178	3.928	3.673	3.399	3.428	6986.7513	2.887	2.627	2.371	2.397	6986.7513	2.887	2.627	2.371	2.397
6968.5533	3.922	3.666	3.395	3.425	6987.0007	2.917	2.626	2.358	2.382	6970.6722	2.845	2.597	2.324	2.346
6968.5920	3.918	3.666	3.392	3.422	6987.0442	2.917	2.626	2.352	2.375	6970.6642	2.863	2.595	2.325	2.341
6968.6279	3.921	3.663	3.391	3.421	6987.0980	2.917	2.626	2.352	2.375	6970.6980	2.865	2.597	2.307	2.334
6968.6646	3.919	3.660	3.388	3.419	6987.1483	2.894	2.612	2.343	2.368	6971.5183	2.881	2.613	2.344	2.368
6968.7102	3.917	3.653	3.386	3.418	6987.1507	2.876	2.605	2.335	2.357	6971.5231	2.917	2.626	2.358	2.382
6969.5799	3.973	3.671	3.412	3.434	6987.1607	2.881	2.613	2.344	2.365	6969.6507	2.876	2.609	2.341	2.361
6969.6168	3.947	3.676	3.404	3.433	6987.1679	2.870	2.605	2.334	2.357	6969.6799	2.870	2.605	2.334	2.357
6969.6491	3.947	3.673	3.403	3.430	6987.1711	2.872	2.604	2.334	2.354	6970.6722	2.845	2.597	2.324	2.346
6969.6773	3.948	3.675	3.406	3.433	6987.1755	2.853	2.607	2.330	2.350	6970.6778	2.833	2.570	2.302	2.330
6970.6216	3.936	3.665	3.388	3.418	6987.2055	2.853	2.584	2.317	2.345	6972.7255	2.853	2.584	2.317	2.345
6970.6626	3.929	3.657	3.381	3.413	6987.2487	2.857	2.581	2.317	2.346	6971.7175	2.936	2.637	2.367	2.386
6970.6933	3.919	3.652	3.375	3.404	6987.2488	2.900	2.623	2.337	2.378	6972.4878	2.900	2.623	2.337	2.378
6971.4822	3.893	3.638	3.366	3.389	6987.2546	2.879	2.615	2.346	2.368	6972.4878	2.900	2.623	2.337	2.378
6971.5219	3.889	3.635	3.361	3.383	6987.2552	2.852	2.597	2.326	2.351	6972.5517	2.836	2.581	2.312	2.334
6971.5517	3.884	3.629	3.354	3.376	6987.2559	2.836	2.581	2.312	2.334	6972.5570	2.836	2.581	2.312	2.334
6971.5934	3.875	3.620	3.346	3.364	6987.2625	2.825	2.567	2.299	2.325	6972.6020	2.825	2.562	2.302	2.325
6971.6301	3.876	3.625	3.367	3.363	6987.2678	2.833	2.570	2.302	2.330	6972.6778	2.833	2.570	2.302	2.330
6971.6729	3.878	3.627	3.352	3.369	6987.2755	2.853	2.584	2.317	2.345	6973.4827	2.878	2.613	2.343	2.369
6971.7157	3.892	3.639	3.365	3.379	6987.3147	2.873	2.613	2.343	2.369	6973.5124	2.873	2.613	2.343	2.369
6972.4852	3.847	3.605	3.333	3.371	6987.3150	2.874	2.614	2.343	2.369	6973.5124	2.873	2.613	2.343	2.369
6972.5230	3.858	3.614	3.337	3.355	6987.3170	2.874	2.614	2.343	2.369	6973.5137	2.874	2.613	2.343	2.369
6972.5521	3.857	3.617	3.340	3.356	6987.3175	2.874	2.614	2.343	2.369	6973.5137	2.874	2.613	2.343	2.369
6972.5948	3.864	3.627	3.348	3.362	6987.3180	2.874	2.614	2.343	2.369	6973.5137	2.874	2.613	2.343	2.369
6972.6371	3.871	3.635	3.356	3.373	6987.3185	2.874	2.614	2.343	2.369	6973.5137	2.874	2.613	2.343	2.369
6972.6762	3.878	3.640	3.363	3.380	6987.3190	2.874	2.614	2.343	2.369	6973.5137	2.874	2.613	2.343	2.369
6972.7199	3.878	3.643	3.364	3.387	6987.3195	2.874	2.614	2.343	2.369	6973.5137	2.874	2.613	2.343	2.369
6972.7568	3.878	3.646	3.365	3.388	6987.3200	2.874	2.614	2.343	2.369	6973.5137	2.874	2.613	2.343	2.369
6972.7918	3.878	3.647	3.367	3.389	6987.3205	2.874	2.614	2.343	2.369	6973.5137	2.874	2.613	2.343	2.369
6972.8348	3.878	3.648	3.368	3.390	6987.3210	2.874	2.614	2.343	2.369	6973.5137	2.874	2.613	2.343	2.369
6972.8718	3.878	3.649	3.369	3.391	6987.3215	2.874	2.614	2.343	2.369	6973.5137	2.874	2.613	2.343	2.369
6972.9148	3.879	3.650	3.370	3.392	6987.3220	2.874	2.614	2.343	2.369	6973.5137	2.874	2.613	2.343	2.369
6972.9579	3.879	3.651	3.371	3.393	6987.3225	2.874	2.614	2.343	2.369	6973.5137	2.874	2.613	2.343	2.369
6972.9918	3.880	3.652	3.372	3.394	6987.3230	2.874	2.614	2.343	2.369	6973.5137	2.874	2.613	2.343	2.369
6972.5275	3.875	3.652	3.373	3.395	6987.3235	2.874	2.614	2.343	2.369	6973.5137	2.874	2.613	2.343	2.369
6972.5626	3.876	3.653	3.374	3.396	6987.3240	2.874	2.614	2.343	2.369	6973.5137	2.874	2.613	2.343	2.369
6972.6093	3.876	3.654	3.375	3.397	6987.3245	2.874	2.614	2.343	2.369	6973.5137	2.874	2.613	2.343	2.369
6972.6407	3.876	3.654	3.374	3.398	6987.3250	2.874	2.614	2.343	2.369	6973				

TABLE IIb (continued)

HR 5941	HJD	u	v	b	y	HR 5941	HJD	u	v	b	y			
6960.5071	5.862	5.185	4.836	4.793	6970.7108	5.806	5.147	4.796	4.766	6994.2634	5.423	4.861	4.808	4.820
6960.5455	5.839	5.171	4.824	4.787	6971.5018	5.817	5.149	4.797	4.770	6998.3200	5.412	4.860	4.799	4.810
6960.5681	5.826	5.162	4.814	4.780	6971.5357	5.853	5.183	4.830	4.802	6998.3693	5.422	4.853	4.799	4.815
6960.5956	5.818	5.150	4.805	4.772	6971.5649	5.841	5.162	4.812	4.782	6998.4158	5.416	4.862	4.804	4.816
6960.6264	5.818	5.140	4.792	4.763	6971.6061	5.845	5.164	4.811	4.781	6999.2235	5.415	4.855	4.813	4.815
6960.6538	5.824	5.145	4.800	4.768	6971.6437	5.857	5.172	4.819	4.789	6999.2713	5.437	4.869	4.813	4.828
6960.6830	5.834	5.158	4.810	4.780	6971.7301	5.913	5.213	4.835	4.819	6999.3188	5.454	4.881	4.821	4.823
6960.7196	5.839	5.160	4.813	4.783	6971.7792	5.880	5.185	4.834	4.794	6999.3625	5.444	4.872	4.821	4.827
6960.7407	5.847	5.159	4.812	4.783	6972.3022	5.901	5.203	4.850	4.818	6999.4065	5.454	4.876	4.823	4.832
6960.7848	5.856	5.166	4.817	4.784	6972.5372	5.909	5.210	4.856	4.820	7000.2236	5.450	4.881	4.819	4.828
6960.8184	5.856	5.166	4.817	4.784	6972.5649	5.958	5.202	4.851	4.815	7000.2676	5.449	4.885	4.823	4.828
6962.5198	5.876	5.192	4.842	4.805	6972.6129	5.869	5.178	4.827	4.794	7000.3115	5.427	4.875	4.816	4.819
6962.5211	5.868	5.176	4.829	4.791	6972.6496	5.829	5.150	4.800	4.770	7000.3552	5.414	4.865	4.807	4.807
6962.5231	5.858	5.178	4.826	4.791	6972.6878	5.816	5.140	4.790	4.761	7000.4134	5.392	4.845	4.793	4.811
6962.5424	5.850	5.167	4.818	4.785	6972.7350	5.829	5.146	4.797	4.768					
6962.5562	5.841	5.157	4.808	4.776	6972.7682	5.842	5.158	4.805	4.776					
6962.5799	5.843	5.158	4.809	4.779	6973.4952	5.817	5.158	4.794	4.764					
6962.5851	5.836	5.149	4.802	4.772	6973.5276	5.827	5.151	4.793	4.771					
6962.6103	5.836	5.150	4.803	4.774	6973.5660	5.842	5.162	4.812	4.782					
6962.6279	5.832	5.149	4.801	4.771	6973.6106	5.838	5.164	4.813	4.783					
6962.6484	5.836	5.150	4.803	4.773	6973.6925	5.896	5.202	4.850	4.816					
6962.7015	5.847	5.162	4.814	4.784	6973.7482	5.837	5.160	4.812	4.778					
6962.7228	5.853	5.166	4.818	4.787	6975.6435	5.839	5.163	4.809	4.779					
6962.7444	5.858	5.156	4.821	4.790	6975.6670	5.846	5.168	4.815	4.784					
6962.7699	5.867	5.172	4.825	4.790	6979.3052	5.801	5.133	4.780	4.753					
6962.8030	5.868	5.178	4.827	4.788	6979.5453	5.814	5.137	4.787	4.760					
6965.5312	5.836	5.158	4.805	4.774	6979.5886	5.832	5.153	4.799	4.773					
6965.5610	5.848	5.166	4.813	4.784	6979.6228	5.847	5.159	4.808	4.770					
6965.6061	5.855	5.172	4.820	4.788	6979.6544	5.849	5.160	4.808	4.779					
6965.6448	5.851	5.170	4.819	4.785	6979.7150	5.867	5.148	4.818	4.765					
6965.7327	5.846	5.160	4.810	4.777	6980.5021	5.884	5.173	4.824	4.796					
6965.7808	5.837	5.158	4.810	4.775	6980.5439	5.880	5.196	4.835	4.810					
6967.5179	5.843	5.158	4.809	4.782	6980.5803	5.877	5.190	4.837	4.805					
6967.5483	5.835	5.159	4.804	4.776	6980.6282	5.851	5.181	4.829	4.793					
6967.5773	5.845	5.159	4.806	4.779	6980.6831	5.873	5.171	4.824	4.792					
6967.6113	5.852	5.167	4.816	4.786	6980.7362	5.851	5.154	4.804	4.773					
6967.6528	5.846	5.171	4.818	4.790	6981.5244	5.817	5.143	4.793	4.764					
6967.6846	5.866	5.178	4.825	4.794	6981.5604	5.826	5.146	4.796	4.766					
6967.7116	5.876	5.182	4.830	4.797	6981.6003	5.839	5.161	4.808	4.779					
6967.7518	5.849	5.173	4.822	4.790	6981.6393	5.858	5.169	4.817	4.788					
6967.7912	5.829	5.156	4.804	4.769	6981.6853	5.879	5.183	4.828	4.801					
6968.5027	5.889	5.204	4.847	4.814	6982.6826	5.826	5.145	4.795	4.767					
6968.5344	5.889	5.201	4.846	4.814	6991.2590	5.454	4.867	4.817	4.827					
6968.5673	5.864	5.184	4.829	4.799	6991.3026	5.468	4.884	4.831	4.832					
6968.6057	5.841	5.160	4.807	4.778	6991.3503	5.473	4.896	4.837	4.841					
6968.6480	5.843	5.153	4.802	4.773	6991.3955	5.465	4.888	4.830	4.834					
6968.6767	5.834	5.151	4.800	4.772	6991.4399	5.450	4.880	4.823	4.831					
6968.7244	5.851	5.152	4.811	4.781	6991.4931	5.445	4.871	4.804	4.829					
6968.7636	5.857	5.156	4.814	4.783	6994.2336	5.510	4.919	4.853	4.865					
6969.5975	5.866	5.164	4.812	4.785	6994.2788	5.432	4.887	4.827	4.833					
6969.6331	5.861	5.164	4.810	4.784	6994.3213	5.394	4.844	4.798	4.809					
6969.6622	5.880	5.182	4.829	4.799	6994.3660	5.402	4.860	4.792	4.799					
6969.6898	5.888	5.196	4.842	4.809	6994.4104	5.430	4.875	4.834	4.817					
6969.7277	5.892	5.215	4.859	4.818	6995.2274	5.392	4.828	4.776	4.789					
6970.6364	5.817	5.153	4.800	4.771	6995.2720	5.398	4.848	4.792	4.802					
6970.6766	5.793	5.135	4.782	4.754	6995.3222	5.410	4.853	4.795	4.807					
6961.5051	5.266	5.138	4.612	4.258	6971.5061	5.261	5.155	4.625	4.289	6995.3052	4.913	4.857	4.614	4.294
6961.6140	5.265	5.142	4.608	4.270	6971.5684	5.262	5.130	4.601	4.264	6995.3548	4.910	4.836	4.592	4.274
6961.6423	5.261	5.141	4.608	4.267	6971.6099	5.259	5.126	4.597	4.257	6998.2979	4.924	4.854	4.627	4.317
6961.6634	5.260	5.141	4.606	4.264	6971.6475	5.258	5.126	4.594	4.254	6998.3502	4.918	4.857	4.624	4.315
6961.6870	5.261	5.141	4.608	4.263	6971.6903	5.257	5.125	4.594	4.250	6998.3990	4.915	4.840	4.614	4.305
6961.7488	5.258	5.137	4.604	4.258	6971.7357	5.254	5.125	4.596	4.250	6998.4445	4.892	4.839	4.601	4.300
6961.7676	5.255	5.138	4.603	4.258	6972.5061	5.253	5.120	4.590	4.249	6999.2534	4.897	4.821	4.598	4.296
6961.7994	5.253	5.133	4.602	4.257	6972.5404	5.251	5.118	4.586	4.244	6999.3014	4.904	4.821	4.598	4.289
6961.8193	5.256	5.139	4.604	4.258	6972.5688	5.249	5.114	4.583	4.240	6999.3458	4.904	4.828	4.599	4.297
6963.5617	5.272	5.156	4.621	4.283	6972.6163	5.249	5.118	4.584	4.241	6999.3886	4.887	4.820	4.595	4.282
6963.5794	5.274	5.155	4.623	4.282	6972.6357	5.246	5.114	4.581	4.238	6999.4315	4.894	4.828	4.591	4.289
6963.5965	5.277	5.155	4.623	4.283	6972.6908	5.243	5.115	4.584	4.239	7000.2507	4.897	4.800	4.569	4.271
6963.6250	5.274	5.156	4.622	4.281	6972.7380	5.244	5.116	4.582	4.236	7000.2951	4.880	4.793	4.561	4.255
6963.6493	5.274	5.157	4.623	4.282	6972.7718	5.242	5.117	4.587	4.244	7000.3382	4.886	4.795	4.558	4.254
6963.6727	5.273	5.158	4.623	4.283	6973.4987	5.255	5.125	4.591	4.246	7000.3939	4.886	4.800	4.572	4.263
6963.7125	5.273	5.155	4.623	4.281	6973.5310	5.249	5.121	4.581	4.245					
6963.7426	5.268	5.152	4.619	4.277	6973.5699	5.246	5.120	4.586	4.243					
6963.7635	5.271	5.154	4.621	4.278	6973.6151	5.245	5.119	4.589	4.246					
6964.6479	5.271	5.159	4.624	4.284	6973.6960	5								

PHOTOMETRY OF BE STARS

HR 6175	HJD	u	v	b	y	HR 6451	HJD	u	v	b	y			
6965.5464	3.179	3.057	2.668	2.540	6979.5250	3.185	3.059	2.669	2.543	6968.8680	5.786	5.588	5.266	5.236
6965.5803	3.175	3.051	2.662	2.535	6979.5625	3.183	3.061	2.670	2.544	6968.8079	5.775	5.581	5.257	5.231
6965.6241	3.174	3.051	2.662	2.536	6979.6048	3.180	3.061	2.670	2.546	6969.6081	5.761	5.568	5.243	5.225
6965.6656	3.173	3.053	2.662	2.535	6979.6701	3.188	3.065	2.674	2.549	6969.6413	5.753	5.568	5.243	5.226
6965.7590	3.184	3.061	2.672	2.543	6980.5166	3.191	3.065	2.677	2.552	6969.6705	5.755	5.566	5.241	5.223
6965.8076	3.181	3.057	2.669	2.536	6980.5623	3.182	3.060	2.671	2.546	6969.6978	5.741	5.551	5.225	5.207
6967.5305	3.186	3.065	2.672	2.550	6980.5986	3.179	3.057	2.667	2.541	6969.7373	5.733	5.542	5.217	5.195
6967.5578	3.179	3.059	2.665	2.542	6980.6479	3.186	3.059	2.671	2.543	6970.7063	5.789	5.590	5.265	5.239
6967.5911	3.181	3.059	2.667	2.543	6980.7042	3.184	3.059	2.670	2.543	6970.7254	5.788	5.585	5.262	5.236
6967.6281	3.181	3.058	2.666	2.543	6980.7538	3.179	3.056	2.665	2.539	6970.7279	5.782	5.583	5.259	5.235
6967.6663	3.177	3.052	2.662	2.537	6981.5406	3.184	3.082	2.679	2.560	6970.7482	5.777	5.573	5.249	5.227
6967.6969	3.179	3.054	2.665	2.539	6981.5805	3.181	3.057	2.666	2.541	6970.7702	5.796	5.593	5.270	5.248
6967.7238	3.181	3.057	2.667	2.541	6981.6178	3.189	3.063	2.672	2.546	6970.7702	5.796	5.593	5.270	5.248
6967.7659	3.181	3.057	2.666	2.538	6981.6534	3.191	3.065	2.675	2.548	6970.7802	5.784	5.549	5.224	5.205
6967.8058	3.187	3.062	2.672	2.544	6982.6475	3.181	3.054	2.665	2.541	6971.5072	5.805	5.599	5.256	5.236
6968.5135	3.188	3.066	2.672	2.548	6982.7029	3.190	3.063	2.674	2.547	6971.5427	5.815	5.610	5.263	5.265
6968.5663	3.184	3.066	2.672	2.544	6991.2089	2.789	2.745	2.652	2.568	6971.5802	5.800	5.598	5.246	5.243
6968.5815	3.179	3.057	2.667	2.543	6991.3332	2.778	2.752	2.654	2.567	6971.6191	5.790	5.580	5.235	5.228
6968.6240	3.182	3.061	2.669	2.546	6991.3824	2.783	2.764	2.665	2.574	6971.6322	5.764	5.560	5.236	5.210
6968.6612	3.179	3.058	2.668	2.542	6991.4262	2.775	2.767	2.664	2.574	6971.6531	5.766	5.567	5.243	5.214
6968.6889	3.181	3.059	2.666	2.542	6991.4747	2.788	2.763	2.666	2.580	6971.7068	5.773	5.576	5.252	5.229
6968.7373	3.194	3.072	2.679	2.548	6991.5284	2.804	2.756	2.680	2.567	6971.7282	5.774	5.549	5.224	5.205
6968.7768	3.184	3.054	2.665	2.536	6994.2640	2.798	2.766	2.662	2.570	6971.7499	5.770	5.577	5.227	5.203
6969.5128	3.178	3.053	2.662	2.538	6994.3079	2.805	2.774	2.671	2.587	6972.0006	5.785	5.581	5.224	5.202
6969.5663	3.178	3.055	2.663	2.540	6994.3521	2.794	2.783	2.684	2.592	6972.0434	2.777	2.767	2.665	2.579
6969.6753	3.184	3.058	2.668	2.543	6994.3969	2.804	2.775	2.669	2.572	6972.1022	5.775	5.555	5.232	5.212
6969.7020	3.182	3.059	2.667	2.541	6994.4398	2.781	2.764	2.678	2.591	6972.1502	5.776	5.559	5.232	5.210
6969.7423	3.185	3.063	2.671	2.544	6995.2572	2.811	2.791	2.692	2.600	6972.1747	5.767	5.559	5.232	5.215
6969.7454	3.201	3.080	2.688	2.564	6999.2550	2.772	2.779	2.675	2.588	6972.2068	5.773	5.565	5.232	5.214
6970.5177	3.182	3.061	2.669	2.547	6999.3568	2.807	2.787	2.682	2.594	6972.2582	5.777	5.577	5.233	5.213
6970.5518	3.181	3.060	2.666	2.546	6999.3904	2.777	2.767	2.665	2.579	6972.2950	5.777	5.577	5.230	5.212
6970.6111	3.181	3.058	2.668	2.541	6999.3924	2.777	2.767	2.665	2.576	6972.3324	5.778	5.577	5.231	5.211
6970.7242	3.181	3.059	2.669	2.543	7000.2327	2.775	2.760	2.668	2.583	6972.4001	2.788	2.764	2.672	2.586
6970.7654	3.185	3.062	2.670	2.544	7000.2970	2.775	2.761	2.665	2.579	6972.4470	2.784	2.762	2.676	2.580
6971.5132	3.191	3.070	2.676	2.553	7000.3403	2.778	2.761	2.663	2.575	6972.5272	2.781	2.762	2.671	2.584
6971.5474	3.201	3.080	2.688	2.564	7000.3967	2.790	2.762	2.671	2.577	6972.5832	2.782	2.763	2.670	2.582
6972.5881	3.189	3.066	2.675	2.549	6999.3035	2.794	2.782	2.690	2.596	6972.6249	2.781	2.763	2.671	2.581
6972.6243	3.183	3.062	2.670	2.546	6999.3479	2.785	2.772	2.675	2.589	6972.6592	2.782	2.763	2.670	2.588
6972.6624	3.192	3.066	2.674	2.547	6999.3909	2.776	2.763	2.669	2.580	6972.6980	2.783	2.765	2.672	2.586
6973.7004	3.189	3.061	2.671	2.541	6999.4336	2.772	2.763	2.668	2.577	6972.7422	2.780	2.765	2.670	2.585
6973.7540	3.183	3.057	2.666	2.536	7000.2327	2.775	2.760	2.668	2.583	6972.7744	2.782	2.767	2.671	2.584
6973.8015	3.180	3.056	2.669	2.532	7000.2970	2.775	2.761	2.665	2.579	6972.8178	2.782	2.768	2.673	2.587
6972.5122	3.188	3.059	2.670	2.541	7000.3403	2.778	2.761	2.663	2.575	6972.8474	2.785	2.769	2.674	2.586
6972.5479	3.185	3.056	2.667	2.540	7000.3967	2.790	2.762	2.671	2.577	6972.8843	2.789	2.771	2.672	2.585
6972.5881	3.181	3.060	2.666	2.539	7000.4470	2.794	2.764	2.667	2.580	6972.9202	2.788	2.770	2.675	2.584
6972.6294	3.179	3.057	2.665	2.539	6999.3035	2.794	2.782	2.665	2.584	6972.9599	2.789	2.771	2.674	2.583
6972.6651	3.179	3.057	2.666	2.541	6999.3524	2.776	2.763	2.667	2.580	6972.9938	2.787	2.770	2.673	2.582
6972.7020	3.176	3.056	2.664	2.538	6999.4064	2.771	2.761	2.662	2.579	6972.9949	2.786	2.769	2.672	2.581
6972.7476	3.184	3.058	2.670	2.540	6999.4555	2.777	2.767	2.664	2.583	6972.9950	2.787	2.770	2.674	2.583
6972.7911	3.181	3.057	2.666	2.538	6999.5082	2.772	2.763	2.661	2.571	6972.9951	2.788	2.771	2.673	2.582
6972.8328	3.189	3.059	2.667	2.541	6999.5538	2.774	2.765	2.663	2.582	6972.9952	2.789	2.772	2.674	2.581
6973.5398	3.188	3.059	2.668	2.541	6999.6059	2.775	2.766	2.664	2.583	6972.9953	2.790	2.773	2.675	2.582
6973.5749	3.179	3.055	2.668	2.542	6999.6554	2.776	2.767	2.665	2.584	6972.9954	2.791	2.774	2.676	2.583
6973.6260	3.183	3.062	2.670	2.546	6999.7056	2.777	2.768	2.666	2.585	6972.9955	2.792	2.775	2.677	2.584
6973.6612	3.186	3.066	2.673	2.542	6999.7547	2.778	2.769	2.667	2.586	6972.9956	2.793	2.776	2.678	2.585
6973.7134	3.187	3.067	2.674	2.543	6999.8048	2.779	2.770	2.668	2.587	6972.9957	2.794	2.777	2.679	2.586
6973.7620	3.183	3.061	2.671	2.541	6999.8548	2.780	2.771	2.669	2.588	6972.9958	2.795	2.778	2.680	2.587
6973.8054	3.179	3.059	2.666	2.538	6999.9072	2.781	2.772	2.667	2.587	6972.9959	2.796	2.779	2.681	2.586
6973.8527	3.177	3.057	2.667	2.538	6999.9599	2.782	2.773	2.668	2.586	6972.9960	2.797	2.780	2.682	2.585
6973.9022	3.179	3.057	2.668	2.538	6999.9938	2.783	2.774	2.669	2.587	6972.9961	2.798	2.781	2.683	2.586
6974.0522	3.179	3.057	2.668	2.538	6999.9938	2.784	2.775	2.669	2.587	6972.9962	2.799	2.782	2.684	2.585
6974.1021	3.179	3.057	2.668	2.538	6999.9938	2.785	2.776	2.669	2.587	6972.9963	2.800	2.783	2.685	2.586
6974.1521	3.179	3.057	2.668	2.538	6999.9938	2.786	2.777	2.669	2.587	6972.9964	2.801	2.784	2.686	2.585
6974.2021	3.179	3.057	2.668	2.538	6999.9938	2.787	2.778	2.669	2.587	6972.9965	2.802	2.785	2.687	2.586
6974.2521	3.179	3.057	2.668	2.538	6999.9938	2.788	2.779	2.669	2.587	6972.9966	2.803	2.786	2.688	2.587
6974.3021	3.179	3.057	2.668	2.538	6999.9938	2.789	2.780	2.669	2.587	6972.9967	2.804	2.787	2.689	2.588
6974.3521	3.179	3.057	2.668	2.538										

TABLE IIb (continued)

HR 6510	HJD	u	v	b	y	HR 6747	HJD	u	v	b	y
6979. 5612 3.379 3.137 2.808 2.789	7000.2380 2.978 2.839 2.810 2.829	6979. 2826 2.983 2.839 2.806 2.837	6979. 6384 6.641 6.578 6.207 6.112	6971.7881 6.622 6.565 6.192 6.096	6971.8227 6.617 6.563 6.189 6.092						
6979. 6038 3.376 3.132 2.803 2.785	7000.3259 2.984 2.834 2.809 2.833	6979. 6734 6.640 6.582 6.207 6.115	6972.5718 6.640 6.582 6.206 6.110	6972. 5718 6.640 6.582 6.206 6.110	6972. 5718 6.640 6.582 6.206 6.110						
6979. 6389 3.368 3.126 2.797 2.778	7000.3787 2.986 2.843 2.815 2.841	6979. 7090 6.642 6.580 6.207 6.114	6972.6567 6.629 6.574 6.200 6.107	6972. 6567 6.629 6.574 6.200 6.107	6972. 6567 6.629 6.574 6.200 6.107						
6979. 6691 3.372 3.129 2.800 2.780	7000.4304 2.997 2.850 2.820 2.843	6979. 7307 6.646 6.583 6.212 6.121	6972.6952 6.625 6.573 6.198 6.107	6972. 6952 6.625 6.573 6.198 6.107	6972. 6952 6.625 6.573 6.198 6.107						
6979. 7630 3.382 3.142 2.814 2.792	7000.4828 2.998 2.861 2.826 2.863	6979. 7540 6.643 6.581 6.213 6.122	6972.7409 6.628 6.575 6.202 6.108	6972. 7409 6.628 6.575 6.202 6.108	6972. 7409 6.628 6.575 6.202 6.108						
6979. 7849 3.380 3.139 2.810 2.791	6979. 8055 3.381 3.145 2.816 2.797	6979. 8014 6.641 6.581 6.209 6.115	6972.7829 6.628 6.574 6.203 6.108	6972. 7829 6.628 6.574 6.203 6.108	6972. 7829 6.628 6.574 6.203 6.108						
6979. 8224 3.381 3.138 2.807 2.789	6980. 5152 3.390 3.146 2.816 2.803	6979. 8388 6.644 6.579 6.207 6.111	6972.8149 6.633 6.580 6.206 6.111	6972. 8149 6.633 6.580 6.206 6.111	6972. 8149 6.633 6.580 6.206 6.111						
6980. 8786 3.386 3.144 2.812 2.793	6980. 8208 3.387 3.143 2.811 2.796	6979. 8646 6.643 6.582 6.210 6.117	6973.5720 6.633 6.591 6.217 6.121	6973. 5720 6.633 6.591 6.217 6.121	6973. 5720 6.633 6.591 6.217 6.121						
6980. 8422 3.392 3.148 2.814 2.797	6980. 8589 3.395 3.151 2.816 2.798	6979. 8659 6.647 6.580 6.207 6.113	6973.6193 6.645 6.584 6.209 6.116	6973. 6193 6.645 6.584 6.209 6.116	6973. 6193 6.645 6.584 6.209 6.116						
6980. 8455 3.363 3.119 2.790 2.771	6980. 7029 3.374 3.134 2.804 2.782	6979. 6031 6.642 6.582 6.209 6.110	6973.6990 6.660 6.598 6.222 6.122	6973. 6990 6.660 6.598 6.222 6.122	6973. 6990 6.660 6.598 6.222 6.122						
6980. 7524 3.385 3.142 2.814 2.793	6980. 8204 3.387 3.143 2.811 2.796	6979. 6208 6.634 6.574 6.201 6.106	6973.7557 6.621 6.570 6.196 6.105	6973. 7557 6.621 6.570 6.196 6.105	6973. 7557 6.621 6.570 6.196 6.105						
6980. 7841 3.386 3.144 2.812 2.793	6980. 8422 3.392 3.148 2.814 2.797	6979. 6352 6.633 6.575 6.202 6.104	6973.7997 6.668 6.609 6.234 6.132	6973. 7997 6.668 6.609 6.234 6.132	6973. 7997 6.668 6.609 6.234 6.132						
6980. 8581 3.378 3.132 2.801 2.785	6981. 5793 3.361 3.119 2.789 2.775	6979. 6677 6.637 6.568 6.205 6.107	6975.6747 6.648 6.584 6.212 6.114	6975. 6747 6.648 6.584 6.212 6.114	6975. 6747 6.648 6.584 6.212 6.114						
6981. 6158 3.365 3.121 2.792 2.775	6981. 6555 3.371 3.127 2.797 2.778	6979. 7103 6.626 6.564 6.195 6.102	6977.7287 6.633 6.580 6.207 6.111	6977. 7287 6.633 6.580 6.207 6.111	6977. 7287 6.633 6.580 6.207 6.111						
6981. 7019 3.381 3.141 2.809 2.791	6982. 6454 3.385 3.137 2.807 2.788	6979. 7310 6.624 6.565 6.194 6.102	6977.7601 6.637 6.585 6.211 6.115	6977. 7601 6.637 6.585 6.211 6.115	6977. 7601 6.637 6.585 6.211 6.115						
6982. 7014 3.390 3.145 2.816 2.798	6981. 2734 2.999 2.847 2.818 2.838	6979. 7523 6.630 6.567 6.198 6.103	6977.7903 6.631 6.583 6.209 6.112	6977. 7903 6.631 6.583 6.209 6.112	6977. 7903 6.631 6.583 6.209 6.112						
6981. 3172 2.984 2.835 2.810 2.826	6981. 3172 2.984 2.835 2.810 2.826	6979. 7771 6.631 6.572 6.202 6.106	6978.7537 6.621 6.570 6.196 6.105	6978. 7537 6.621 6.570 6.196 6.105	6978. 7537 6.621 6.570 6.196 6.105						
6981. 3672 2.972 2.826 2.795 2.819	6981. 4108 2.967 2.827 2.802 2.823	6979. 8125 6.637 6.575 6.206 6.110	6978.7831 6.622 6.569 6.195 6.104	6978. 7831 6.622 6.569 6.195 6.104	6978. 7831 6.622 6.569 6.195 6.104						
6981. 4586 2.980 2.840 2.811 2.835	6981. 5105 2.994 2.844 2.813 2.836	6979. 8323 6.637 6.576 6.207 6.110	6979.5535 6.642 6.580 6.207 6.108	6979. 5535 6.642 6.580 6.207 6.108	6979. 5535 6.642 6.580 6.207 6.108						
6981. 5463 3.002 2.836 2.810 2.837	6984. 2482 2.996 2.838 2.821 2.833	6979. 8496 6.633 6.580 6.205 6.116	6979.5972 6.636 6.580 6.207 6.113	6979. 5972 6.636 6.580 6.207 6.113	6979. 5972 6.636 6.580 6.207 6.113						
6984. 2492 2.982 2.839 2.801 2.826	6984. 3365 3.004 2.857 2.805 2.829	6979. 8573 6.634 6.581 6.208 6.110	6979.6322 6.636 6.578 6.203 6.113	6979. 6322 6.636 6.578 6.203 6.113	6979. 6322 6.636 6.578 6.203 6.113						
6984. 3805 2.973 2.836 2.808 2.835	6984. 4248 2.993 2.833 2.810 2.829	6979. 8654 6.635 6.582 6.212 6.116	6979.6323 6.637 6.580 6.205 6.112	6979. 6323 6.637 6.580 6.205 6.112	6979. 6323 6.637 6.580 6.205 6.112						
6984. 5173 3.003 2.846 2.806 2.838	6985. 2419 2.997 2.836 2.805 2.833	6979. 8706 6.635 6.583 6.209 6.115	6979.6623 6.637 6.580 6.205 6.112	6979. 6623 6.637 6.580 6.205 6.112	6979. 6623 6.637 6.580 6.205 6.112						
6985. 2867 3.003 2.845 2.807 2.845	6985. 2914 2.981 2.835 2.810 2.829	6979. 8727 6.634 6.583 6.211 6.116	6979.7562 6.629 6.570 6.195 6.104	6979. 7562 6.629 6.570 6.195 6.104	6979. 7562 6.629 6.570 6.195 6.104						
6985. 3362 2.991 2.846 2.814 2.848	6985. 3404 3.002 2.844 2.809 2.835	6979. 8733 6.634 6.584 6.210 6.109	6980.5533 6.657 6.580 6.210 6.116	6980. 5533 6.657 6.580 6.210 6.116	6980. 5533 6.657 6.580 6.210 6.116						
6985. 3764 3.002 2.847 2.812 2.843	6986. 5166 2.984 2.835 2.805 2.829	6979. 8747 6.635 6.585 6.212 6.117	6980.5869 6.654 6.576 6.212 6.116	6980. 5869 6.654 6.576 6.212 6.116	6980. 5869 6.654 6.576 6.212 6.116						
6986. 4189 2.998 2.858 2.826 2.850	6986. 4661 3.003 2.854 2.821 2.843	6979. 8752 6.636 6.587 6.207 6.110	6980.6385 6.636 6.568 6.191 6.098	6980. 6385 6.636 6.568 6.191 6.098	6980. 6385 6.636 6.568 6.191 6.098						
6986. 4672 3.000 2.863 2.835 2.852	6989. 5172 3.021 2.865 2.833 2.844	6979. 8760 6.637 6.588 6.206 6.112	6980.6970 6.647 6.575 6.201 6.109	6980. 6970 6.647 6.575 6.201 6.109	6980. 6970 6.647 6.575 6.201 6.109						
6986. 4894 3.004 2.857 2.805 2.829	6987. 8016 4.950 4.662 4.342 4.407	6979. 8770 6.638 6.592 6.211 6.119	6980.7454 6.650 6.578 6.204 6.111	6980. 7454 6.650 6.578 6.204 6.111	6980. 7454 6.650 6.578 6.204 6.111						
6987. 3360 2.996 2.848 2.823 2.847	6987. 8016 4.955 4.667 4.347 4.407	6979. 8780 6.639 6.594 6.207 6.111	6980.8132 6.649 6.580 6.207 6.108	6980. 8132 6.649 6.580 6.207 6.108	6980. 8132 6.649 6.580 6.207 6.108						
6987. 3853 2.994 2.846 2.813 2.836	6987. 8016 4.957 4.668 4.348 4.407	6979. 8782 6.643 6.583 6.211 6.116	6981.5319 6.643 6.589 6.211 6.114	6981. 5319 6.643 6.589 6.211 6.114	6981. 5319 6.643 6.589 6.211 6.114						
6988. 4308 3.002 2.843 2.816 2.837	6988. 8114 4.954 4.669 4.349 4.407	6979. 8792 6.642 6.583 6.212 6.117	6981.5733 6.643 6.579 6.204 6.108	6981. 5733 6.643 6.579 6.204 6.108	6981. 5733 6.643 6.579 6.204 6.108						
6988. 4661 3.003 2.854 2.821 2.843	6988. 8114 4.955 4.670 4.349 4.407	6979. 8803 6.644 6.585 6.210 6.109	6981.6093 6.641 6.581 6.206 6.110	6981. 6093 6.641 6.581 6.206 6.110	6981. 6093 6.641 6.581 6.206 6.110						
6988. 5166 3.027 2.873 2.843 2.862	6988. 8114 4.956 4.671 4.349 4.407	6979. 8814 6.645 6.586 6.211 6.117	6981.6474 6.647 6.562 6.214 6.116	6981. 6474 6.647 6.562 6.214 6.116	6981. 6474 6.647 6.562 6.214 6.116						
6988. 5365 3.027 2.873 2.843 2.862	6988. 8114 4.957 4.672 4.349 4.407	6979. 8820 6.646 6.587 6.212 6.117	6981.6474 6.648 6.563 6.214 6.117	6981. 6474 6.648 6.563 6.214 6.117	6981. 6474 6.648 6.563 6.214 6.117						
6988. 5806 3.027 2.873 2.843 2.862	6988. 8114 4.958 4.673 4.349 4.407	6979. 8827 6.647 6.588 6.213 6.118	6981.6474 6.649 6.564 6.215 6.117	6981. 6474 6.649 6.564 6.215 6.117	6981. 6474 6.649 6.564 6.215 6.117						
6988. 6365 3.027 2.873 2.843 2.862	6988. 8114 4.959 4.674 4.349 4.407	6979. 8837 6.648 6.589 6.214 6.118	6981.6474 6.650 6.565 6.216 6.117	6981. 6474 6.650 6.565 6.216 6.117	6981. 6474 6.650 6.565 6.216 6.117						
6988. 6862 3.027 2.873 2.843 2.862	6988. 8114 4.960 4.675 4.349 4.407	6979. 8847 6.649 6.590 6.215 6.119	6981.6474 6.652 6.566 6.218 6.117	6981. 6474 6.652 6.566 6.218 6.117	6981. 6474 6.652 6.566 6.218 6.117						
6988. 7329 3.027 2.873 2.843 2.862	6988. 8114 4.961 4.676 4.349 4.407	6979. 8857 6.650 6.591 6.216 6.120	6981.6474 6.654 6.567 6.220 6.118	6981. 6474 6.654 6.567 6.220 6.118	6981. 6474 6.654 6.567 6.220 6.118						
6988. 7849 3.027 2.873 2.843 2.862	6988. 8114 4.962 4.677 4.349 4.407	6979. 8867 6.651 6.592 6.217 6.121	6981.6474 6.655 6.568 6.221 6.119	6981. 6474 6.655 6.568 6.221 6.119	6981. 6474 6.655 6.568 6.221 6.119						
6988. 8364 3.027 2.873 2.843 2.862	6988. 8114 4.963 4.678 4.349 4.407	6979. 8876 6.652 6.593 6.218 6.122	6981.6474 6.658 6.570 6.222 6.120	6981. 6474 6.658 6.570 6.222 6.120	6981. 6474 6.658 6.570 6.222 6.120						
6988. 8874 3.027 2.873 2.843 2.862	6988. 8114 4.964 4.679 4.349 4.407	6979. 8886 6.653 6.594 6.219 6.123	6981.6474 6.662 6.569 6.223 6.121	6981. 6474 6.662 6.569 6.223 6.121	6981. 6474 6.662 6.569 6.223 6.121						
6988. 9374 3.027 2.873 2.843 2.862	6988. 8114 4.965 4.680 4.349 4.407	6979. 8895 6.654 6.595 6.220 6.124	6981.6474 6.666 6.571 6.224 6.122	6981. 6474 6.666 6.571 6.224 6.122	6981. 6474 6.666 6.571 6.224 6.122						
6989. 4302 3.027 2.873 2.843 2.862	6989. 8114 4.966 4.681 4.349 4.407	6979. 8905 6.655 6.596 6.221 6.125	6981.6474 6.667 6.572 6.225 6.123	6981. 6474 6.667 6.572 6.225 6.123	6981. 6474 6.667 6.572 6.225 6.123						
6989. 4842 3.027 2.873 2.843 2.862	6989. 8114 4.967 4.682 4.349 4.407	6979. 8914 6.656 6.597 6.222 6.126	6981.6474 6.668 6.573 6.226 6.124	6981. 6474 6.668 6.573 6.226 6.124	6981. 6474 6.668 6.573 6.226 6.124						
6989. 5342 3.027 2.873 2.843 2.862	6989. 8114 4.968 4.683 4.349 4.407	6979. 8923 6.657 6.598 6.223 6.127	6981.6474 6.671 6.574 6.227 6.125	6981. 6474 6.671 6.574 6.227 6.125	6981. 6474 6.671 6.574 6.227 6.125						
6989. 5852 3.027 2.873 2.843 2.862	6989. 8114 4.969 4.684 4.349 4.407										

TABLE IIb (*continued*)

HR 8408	HJD	u	v	b	y	HR 8858	HJD	u	v	b	y				
	HJD	u	v	b	y		HJD	u	v	b	y				
6960.7646	6.688	6.223	5.491	5.942	6977.7457	6.687	6.220	5.486	5.943	6978.7713	5.229	4.673	4.316	4.379	
6960.7657	6.689	6.220	5.488	5.939	6977.7809	6.697	6.232	5.486	5.952	6978.8033	5.260	4.689	4.329	4.390	
6960.8037	6.699	6.232	5.500	5.951	6977.8109	6.701	6.234	5.501	5.952	6978.8292	5.258	4.691	4.331	4.394	
6960.8410	6.699	6.231	5.500	5.952	6977.8336	6.696	6.233	5.489	5.953	6979.7956	5.276	4.695	4.341	4.398	
6960.8676	6.695	6.230	5.500	5.951	6977.8498	6.697	6.231	5.485	5.953	6979.8131	5.263	4.689	4.330	4.394	
6960.8785	6.697	6.231	5.501	5.953	6977.8818	6.693	6.230	5.485	5.950	6979.8163	5.263	4.689	4.330	4.394	
6960.8869	6.692	6.229	5.499	5.951	6977.8927	6.696	6.226	5.484	5.947	6979.8177	5.264	4.685	4.331	4.391	
6960.8946	6.695	6.226	5.499	5.949	6978.7668	6.687	6.220	5.488	5.939	6979.8446	5.266	4.687	4.330	4.393	
6960.9011	6.693	6.229	5.498	5.949	6978.7797	6.680	6.220	5.489	5.940	6979.8648	5.264	4.687	4.328	4.389	
6960.9084	6.696	6.229	5.501	5.950	6978.8208	6.681	6.218	5.486	5.939	6979.8775	5.257	4.687	4.325	4.389	
6960.9141	6.697	6.230	5.498	5.951	6979.7685	6.701	6.232	5.903	5.951	6979.8936	5.256	4.680	4.324	4.387	
6960.9203	6.697	6.230	5.501	5.951	6979.7918	6.696	6.231	5.900	5.947	6979.9086	5.250	4.674	4.318	4.380	
6965.8142	6.673	6.213	5.482	5.936	6979.8120	6.695	6.231	5.902	5.951	6980.7984	5.263	4.683	4.326	4.387	
6965.8165	6.679	6.217	5.486	5.938	6979.8272	6.696	6.228	5.899	5.948	6980.8307	5.263	4.685	4.329	4.391	
6965.8397	6.676	6.216	5.486	5.939	6979.8398	6.695	6.233	5.904	5.953	6980.8520	5.258	4.683	4.327	4.388	
6965.8571	6.675	6.215	5.484	5.936	6979.8591	6.695	6.232	5.901	5.950	6980.8701	5.255	4.678	4.324	4.385	
6965.8799	6.671	6.216	5.484	5.936	6979.8727	6.696	6.233	5.901	5.952	6980.8936	5.253	4.677	4.322	4.382	
6965.8939	6.674	6.217	5.483	5.938	6979.8891	6.694	6.232	5.901	5.950	6985.8447	5.268	4.691	4.334	4.385	
6967.7317	6.680	6.216	5.481	5.937	6979.9039	6.697	6.233	5.905	5.956	6985.8626	5.264	4.687	4.331	4.386	
6967.7711	6.682	6.222	5.487	5.942	6980.6571	6.694	6.227	5.892	5.946	6991.6139	4.877	4.391	4.361	4.424	
6967.8108	6.684	6.228	5.492	5.946	6980.7127	6.682	6.213	5.880	5.936	6995.8841	5.263	4.686	4.330	4.381	
6967.8441	6.679	6.223	5.488	5.941	6980.7941	6.679	6.212	5.880	5.932	6999.5703	4.857	4.369	4.329	4.397	
6967.8674	6.676	6.219	5.486	5.935	6980.8254	6.680	6.214	5.882	5.935	6999.6287	4.854	4.372	4.338	4.393	
6968.7034	6.679	6.215	5.482	5.938	6980.8467	6.682	6.217	5.886	5.938	6999.5803	4.844	4.357	4.322	4.380	
6968.7425	6.697	6.232	5.497	5.953	6980.8670	6.683	6.219	5.886	5.939	6999.6321	4.842	4.365	4.331	4.391	
6968.7800	6.692	6.223	5.489	5.946	6980.8896	6.679	6.219	5.886	5.940	7000.5960	4.832	4.371	4.324	4.380	
6968.8194	6.687	6.222	5.487	5.942	6980.9005	6.680	6.218	5.883	5.943	6968.7853	5.239	4.673	4.316	4.378	
6968.8485	6.697	6.230	5.487	5.949	6981.6654	6.692	6.222	5.887	5.936	6968.8226	5.245	4.676	4.320	4.382	
6968.8753	6.685	6.217	5.484	5.937	6982.6586	6.721	6.225	5.894	5.943	6968.8511	5.257	4.689	4.332	4.393	
6968.8973	6.686	6.220	5.487	5.940	6982.7129	6.728	6.234	5.906	5.959	6968.8782	5.256	4.685	4.329	4.382	
6969.7085	6.685	6.223	5.491	5.943	6982.7584	6.701	6.208	5.881	5.932	6970.7738	5.253	4.682	4.327	4.386	
6969.7479	6.690	6.221	5.491	5.942	6991.5826	6.300	5.950	5.939	6.007	6970.8097	5.252	4.682	4.324	4.384	
6970.7308	6.689	6.226	5.492	5.944	6991.6348	6.297	5.956	5.934	6.003	6970.8133	5.253	4.681	4.326	4.386	
6970.7703	6.696	6.218	5.492	5.944	6998.4855	6.267	5.921	5.908	5.983	6970.8593	5.248	4.676	4.319	4.380	
6970.8057	6.686	6.221	5.488	5.941	6998.5373	6.268	5.909	5.899	5.963	6970.8698	5.248	4.673	4.319	4.379	
6970.8280	6.685	6.222	5.492	5.942	6998.5599	6.280	5.919	5.905	5.972	6970.8863	5.250	4.675	4.321	4.380	
6970.8571	6.686	6.221	5.490	5.944	6999.6524	6.286	5.922	5.905	5.981	6970.8956	5.252	4.677	4.322	4.382	
6970.8669	6.690	6.227	5.489	5.950	6999.6879	6.285	5.925	5.904	5.975	6971.7635	5.276	4.693	4.341	4.400	
6970.8780	6.690	6.227	5.485	5.949	6999.7515	6.291	5.930	5.920	5.984	6971.8075	5.255	4.682	4.326	4.386	
6971.7058	6.689	6.226	5.482	5.944	6999.8545	6.281	5.935	5.907	5.984	6971.8378	5.250	4.677	4.321	4.384	
6971.7605	6.695	6.224	5.488	5.944	7000.5010	6.270	5.905	5.911	5.962	6971.8616	5.246	4.674	4.316	4.382	
6971.8049	6.689	6.219	5.485	5.942	7000.5293	6.252	5.903	5.887	5.958	6971.8842	5.248	4.672	4.317	4.379	
6971.8347	6.696	6.220	5.489	5.945	7000.5648	6.262	5.915	5.898	5.963	6971.8957	5.245	4.671	4.315	4.378	
6971.8582	6.694	6.222	5.490	5.945	7000.6218	6.265	5.917	5.894	5.965	6971.9073	5.246	4.671	4.317	4.377	
6971.8813	6.698	6.226	5.493	5.947						6972.7551	5.269	4.690	4.334	4.392	
6971.8926	6.700	6.227	5.493	5.947						6972.8013	5.263	4.684	4.330	4.391	
6971.9044	6.698	6.223	5.492	5.949						6972.8357	5.255	4.680	4.324	4.384	
6972.6706	6.701	6.224	5.494	5.942						6972.8763	5.246	4.676	4.322	4.383	
6972.7073	6.699	6.224	5.496	5.942						6972.8985	5.249	4.677	4.323	4.382	
6972.7523	6.689	6.216	5.488	5.940						6973.7722	5.279	4.693	4.344	4.402	
6972.7981	6.686	6.214	5.486	5.937						6977.7843	5.238	4.679	4.324	4.384	
6972.8313	6.694	6.220	5.490	5.943						6977.8139	5.235	4.679	4.323	4.385	
6972.8698	6.695	6.224	5.495	5.943						6977.8367	5.239	4.677	4.322	4.385	
6972.8952	6.699	6.230	5.489	5.950						6977.8534	5.235	4.673	4.320	4.382	
6973.6677	6.673	6.209	5.486	5.934						6977.8660	5.228	4.673	4.318	4.381	
											6977.8859	5.233	4.674	4.320	4.381

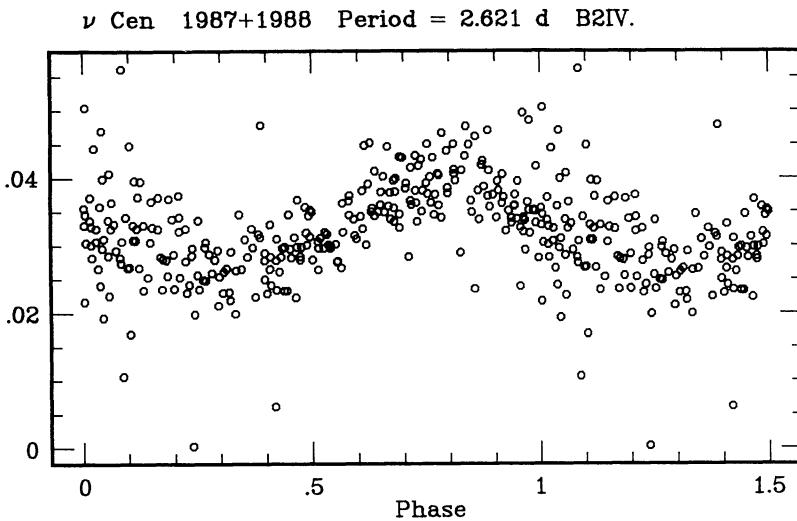


FIGURE 1. — Light curve of ν Cen ($P = 2.621$ d). In this and subsequent figures, the scale is in magnitudes and the epoch of phase zero is JD2446000.000.

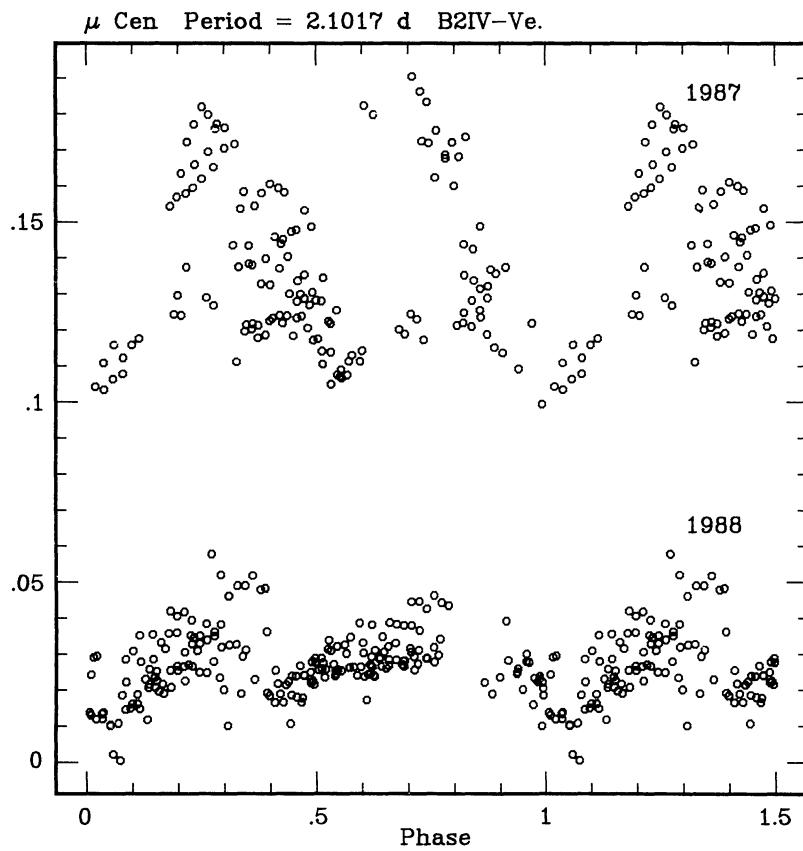
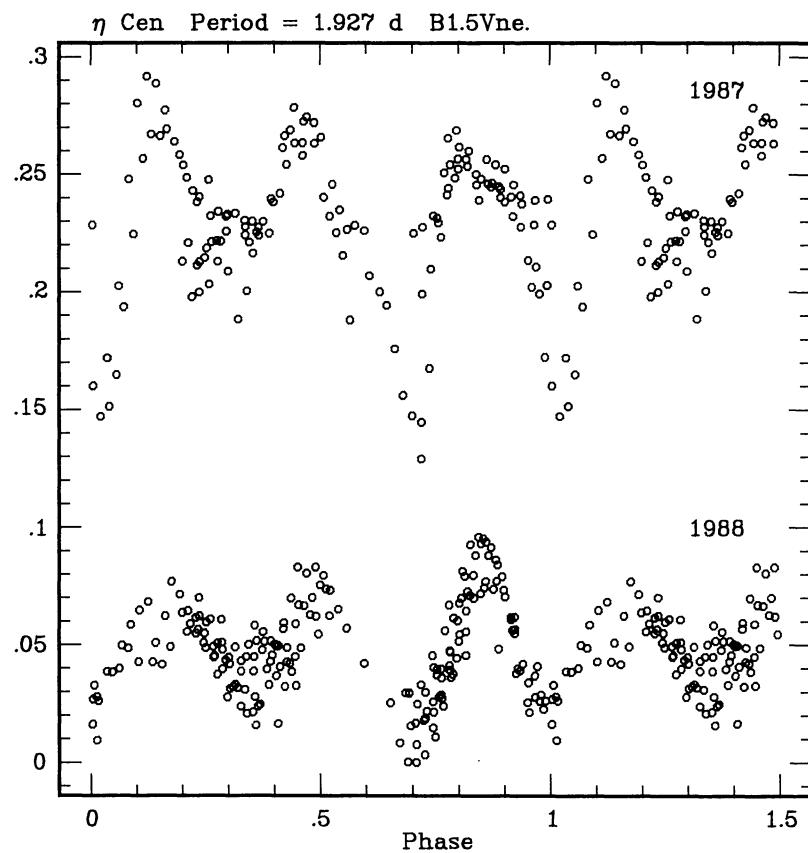
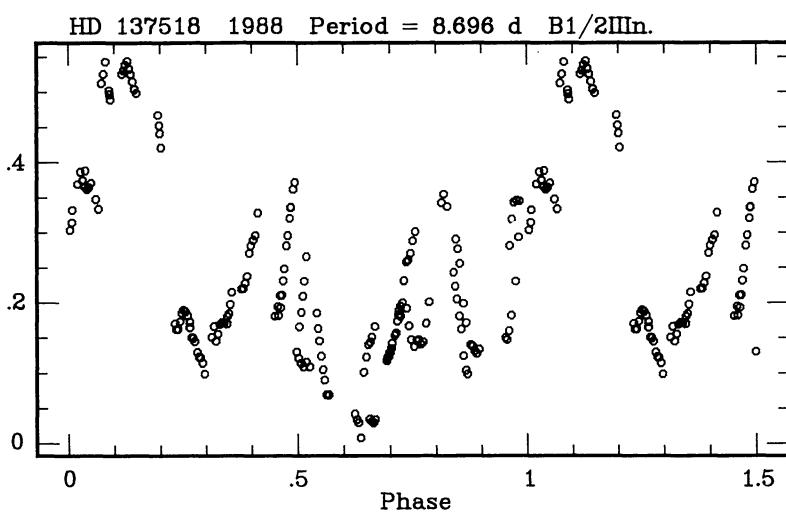
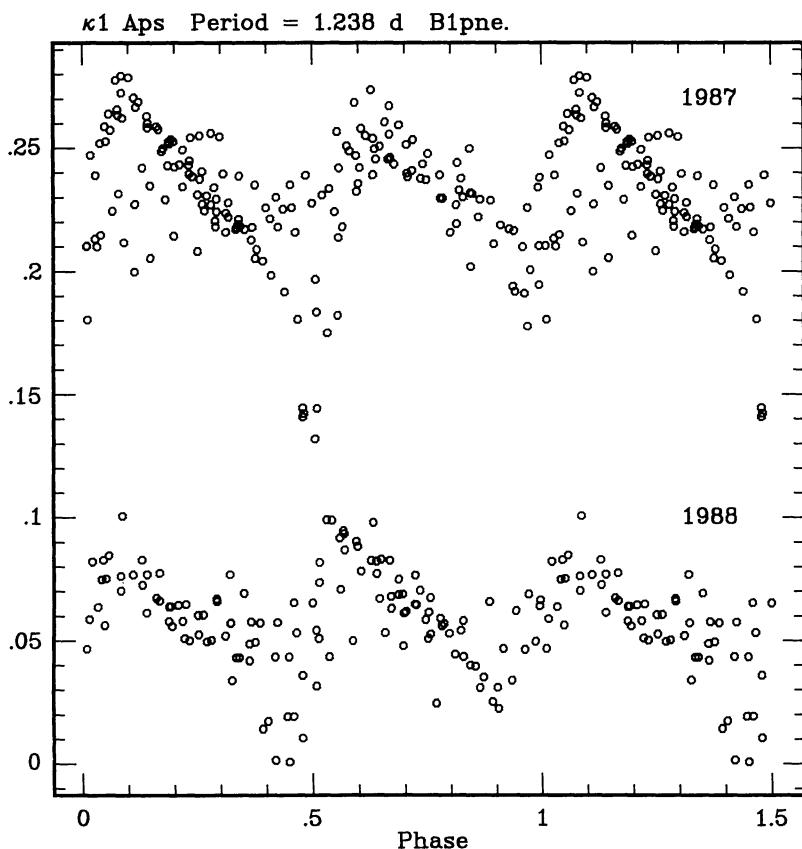
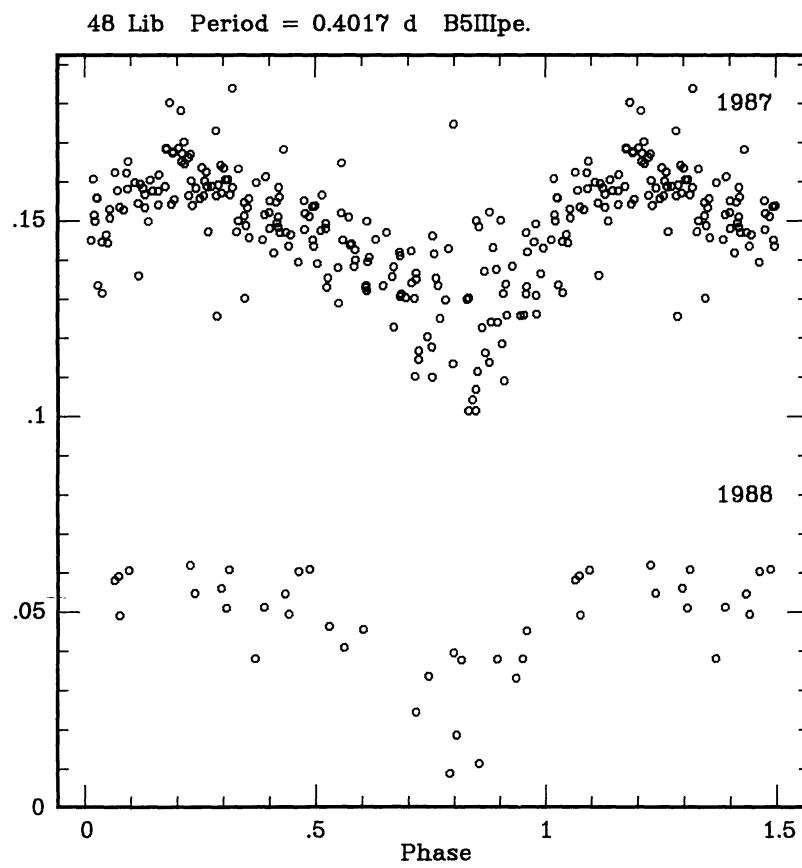
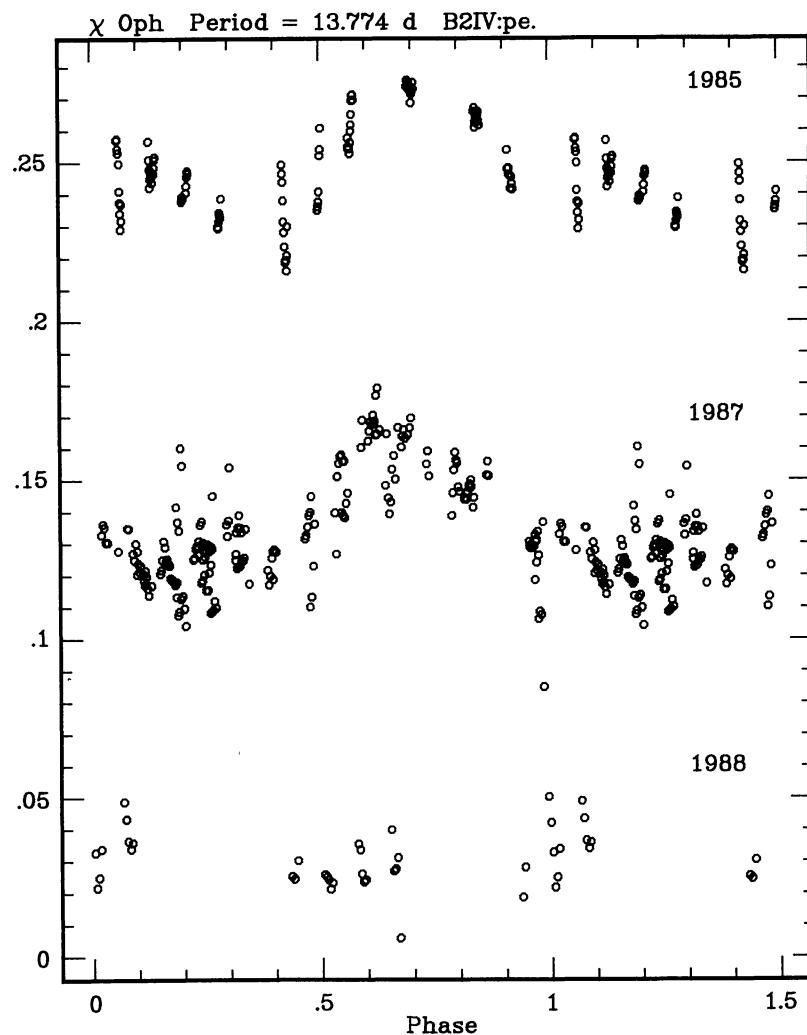
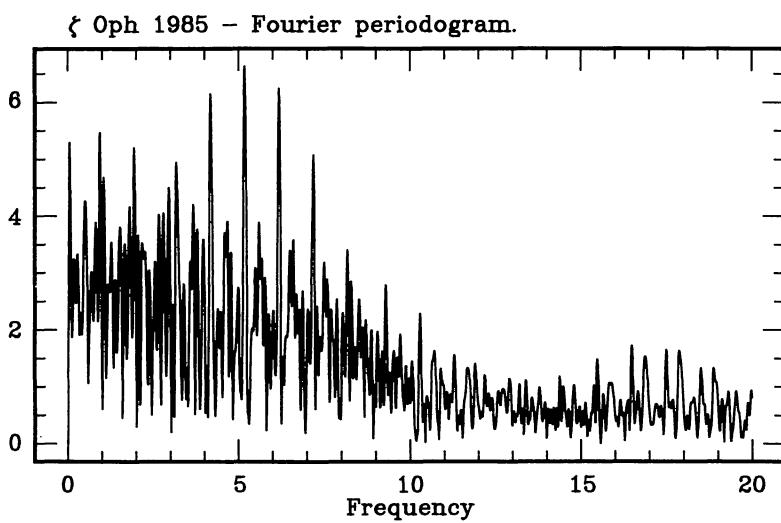
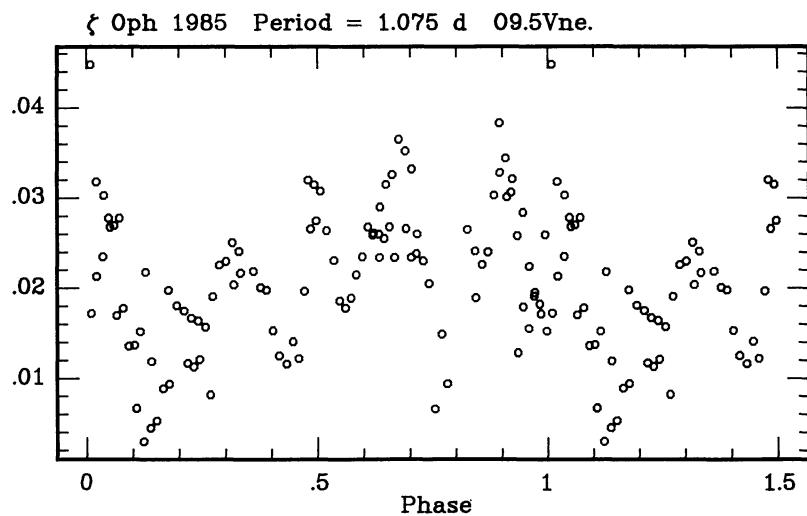
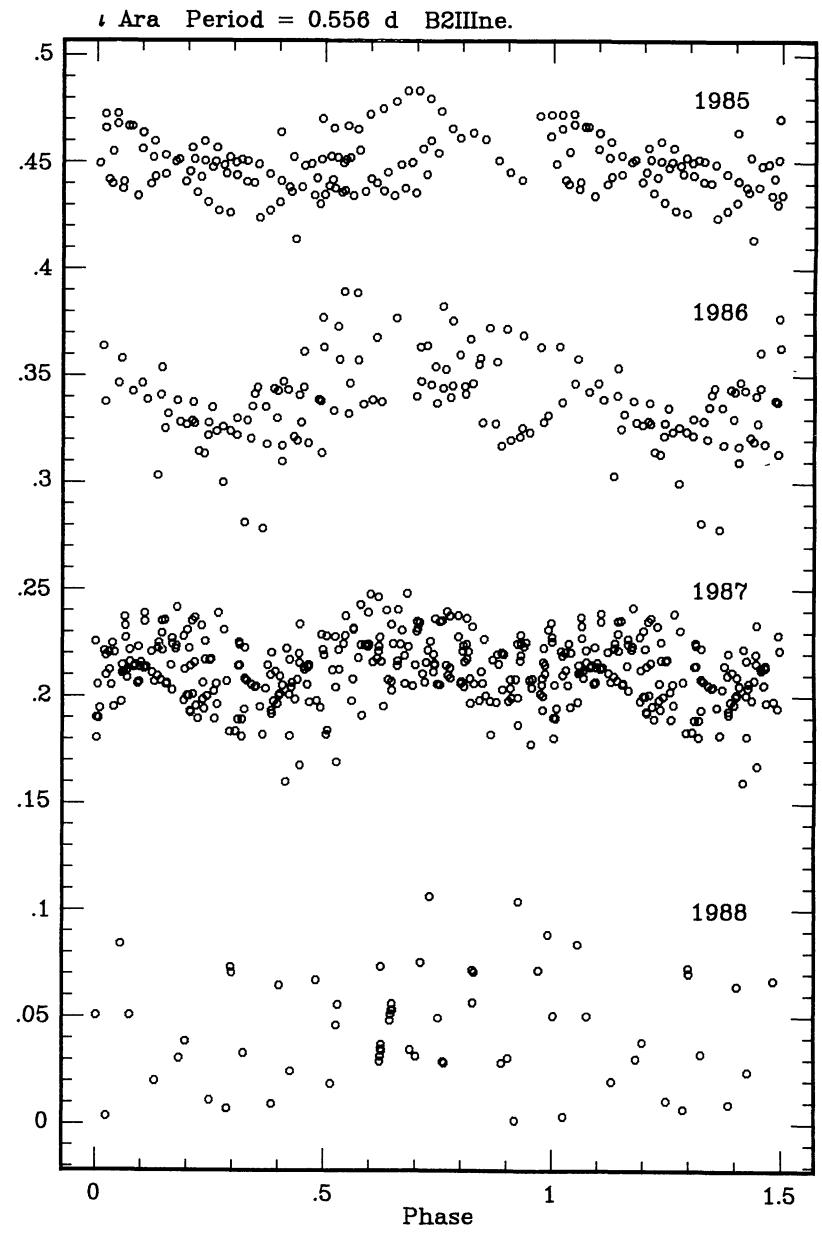


FIGURE 2. — Light curve of μ Cen ($P = 2.1017$ d).

FIGURE 3. — Light curve of η Cen ($P = 1.927$ d).FIGURE 4. — Light curve of HD137518 ($P = 8.696$ d).

FIGURE 5. — Light curve of κ^1 Aps ($P = 1.238$ d).FIGURE 6. — Light curve of 48 Lib ($P = 0.4017$ d).

FIGURE 7. — Light curve of χ Oph ($P = 13.774$ d).FIGURE 8a. — Fourier periodogram for 1985 data of ζ Oph. The frequency is in cycles d^{-1} and the semi-amplitude in millimagnitudes.

FIGURE 8b. — Light curve of ζ Oph for 1985 ($P = 1.075$ d).FIGURE 9. — Light curve of ι Ara ($P = 0.556$ d).

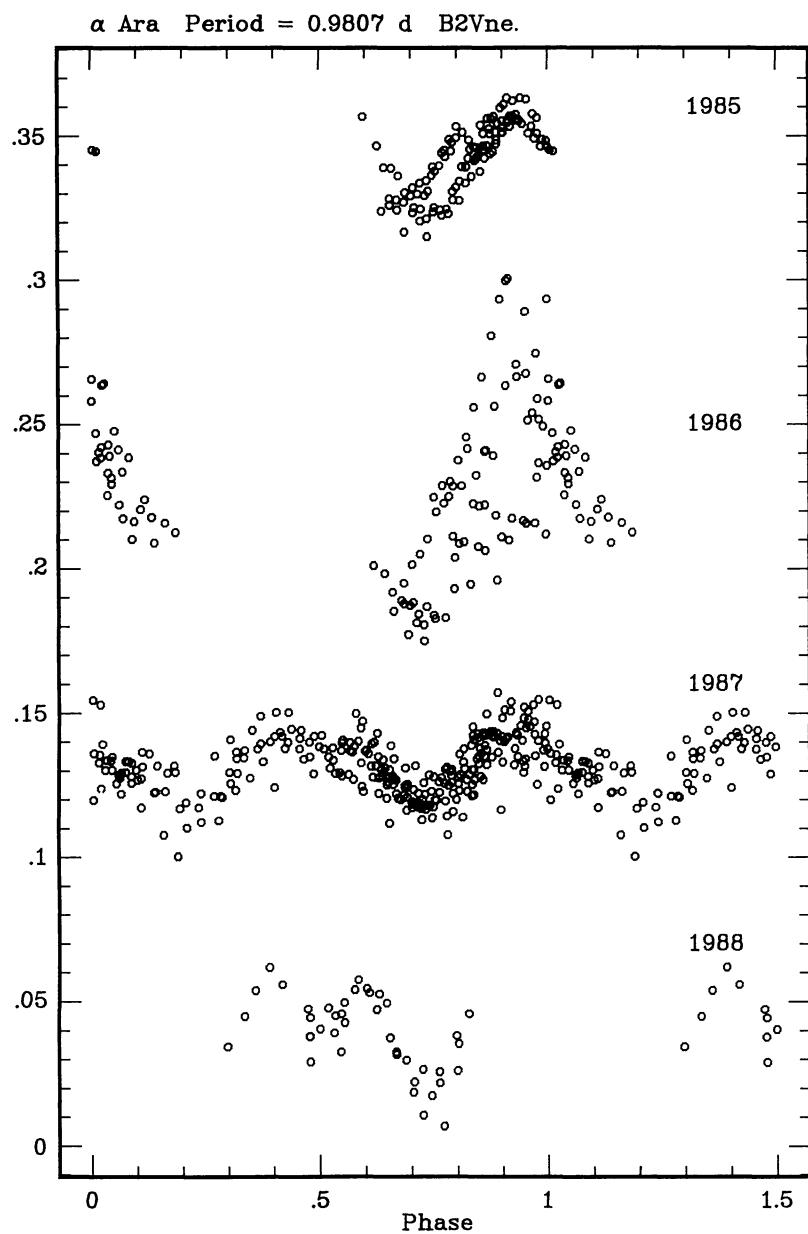


FIGURE 10. — Light curve of α Ara ($P = 0.9807$ d).

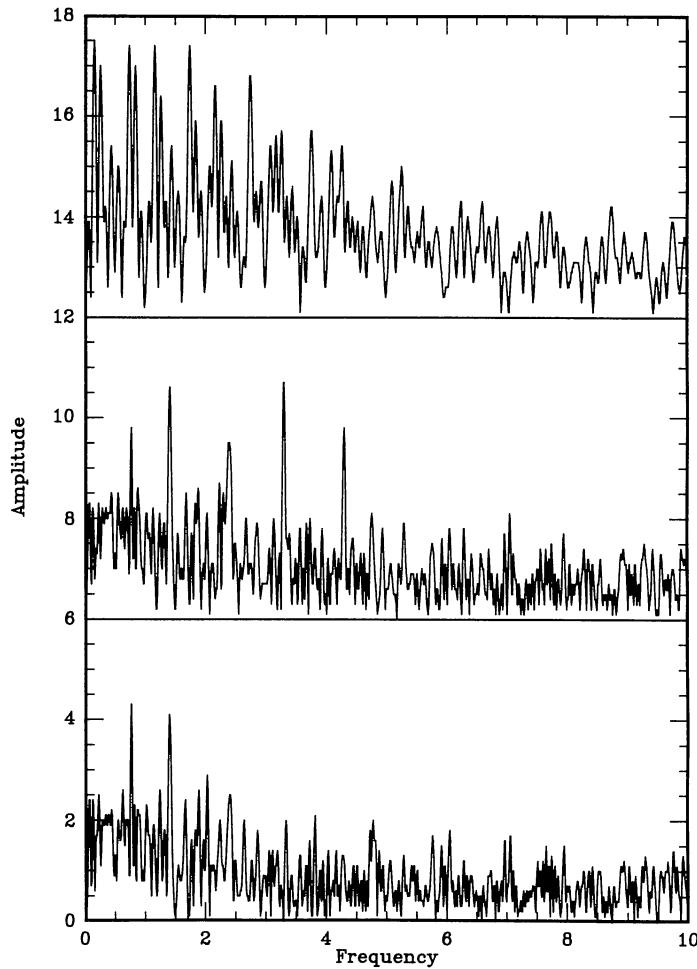


FIGURE 11a. — Fourier periodograms for V986 Oph. Top panel - 1985 data ; middle panel - 1987 data ; bottom panel - 1987 data prewhitened by 3.30 d^{-1} . The frequency is in cycles d^{-1} and the semi-amplitude in millimagnitudes.

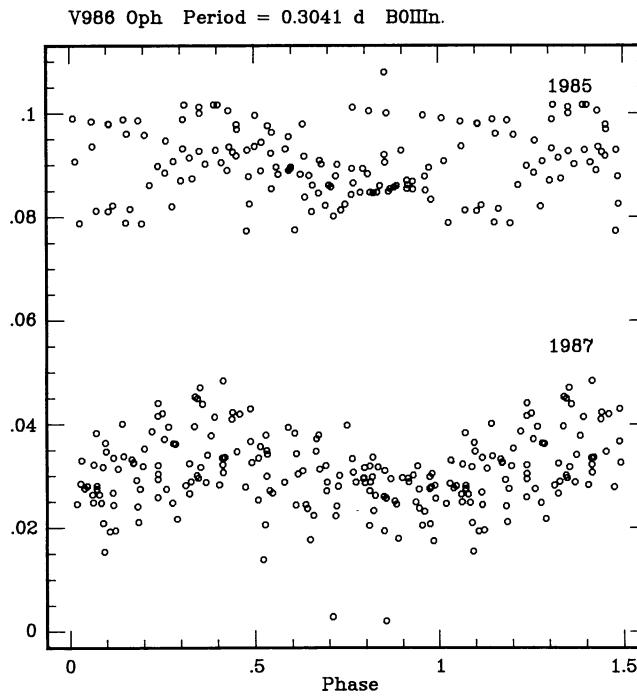


FIGURE 11b. — Light curve of V986 Oph ($P = 0.3041 \text{ d}$).

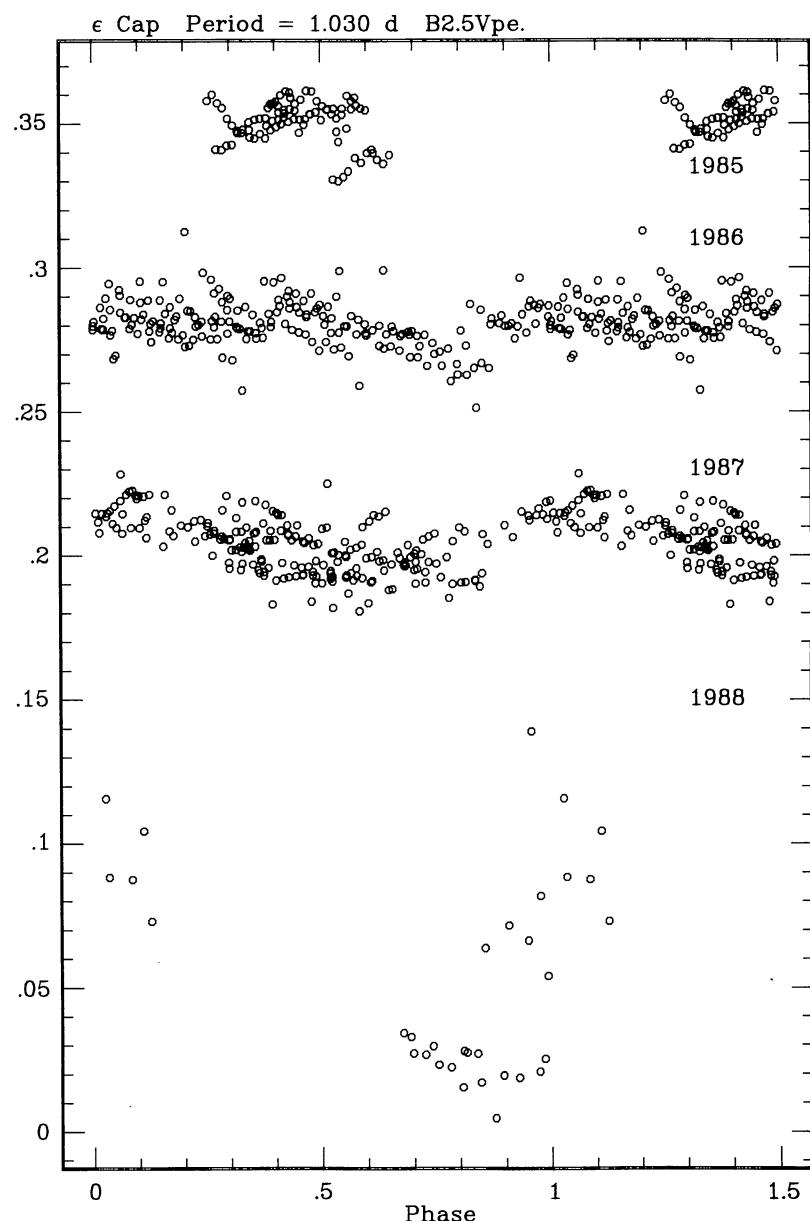
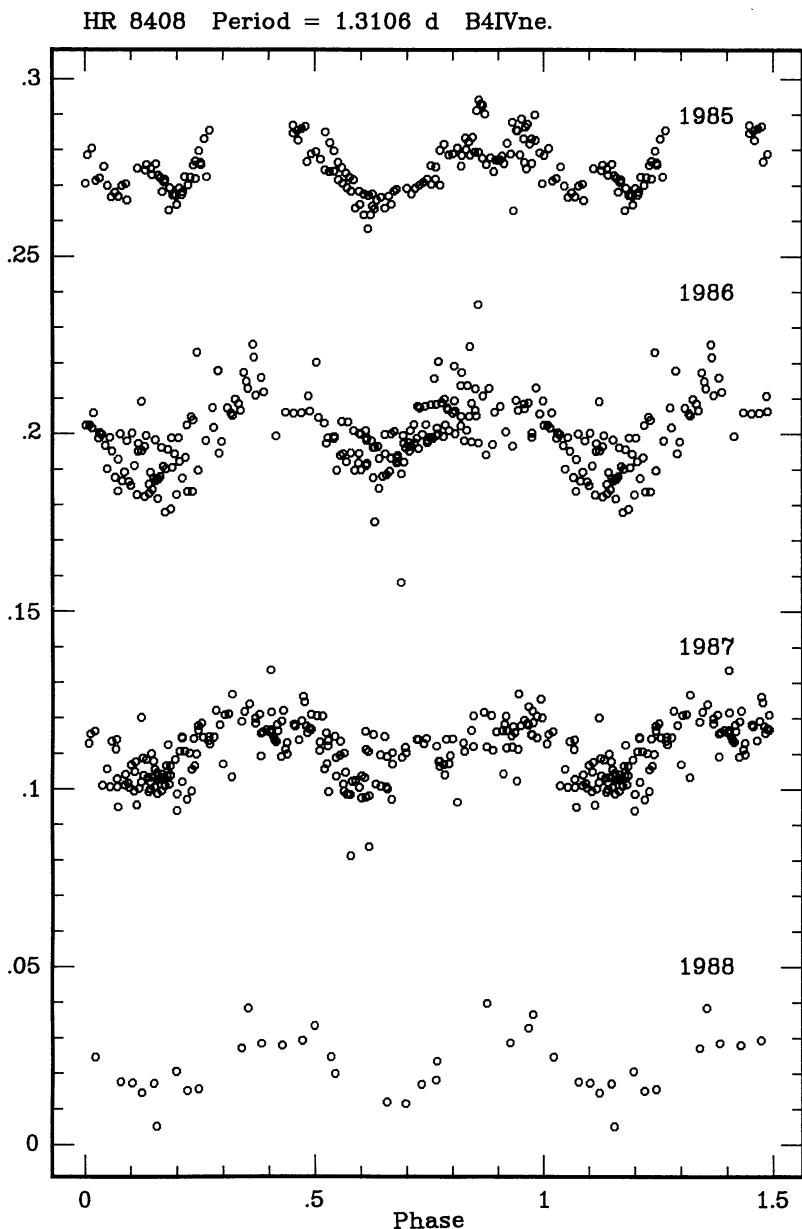


FIGURE 12. — Light curve of ϵ Cap ($P = 1.030$ d).

FIGURE 13. — Light curve of HR 8408 ($P = 1.3106$ d).

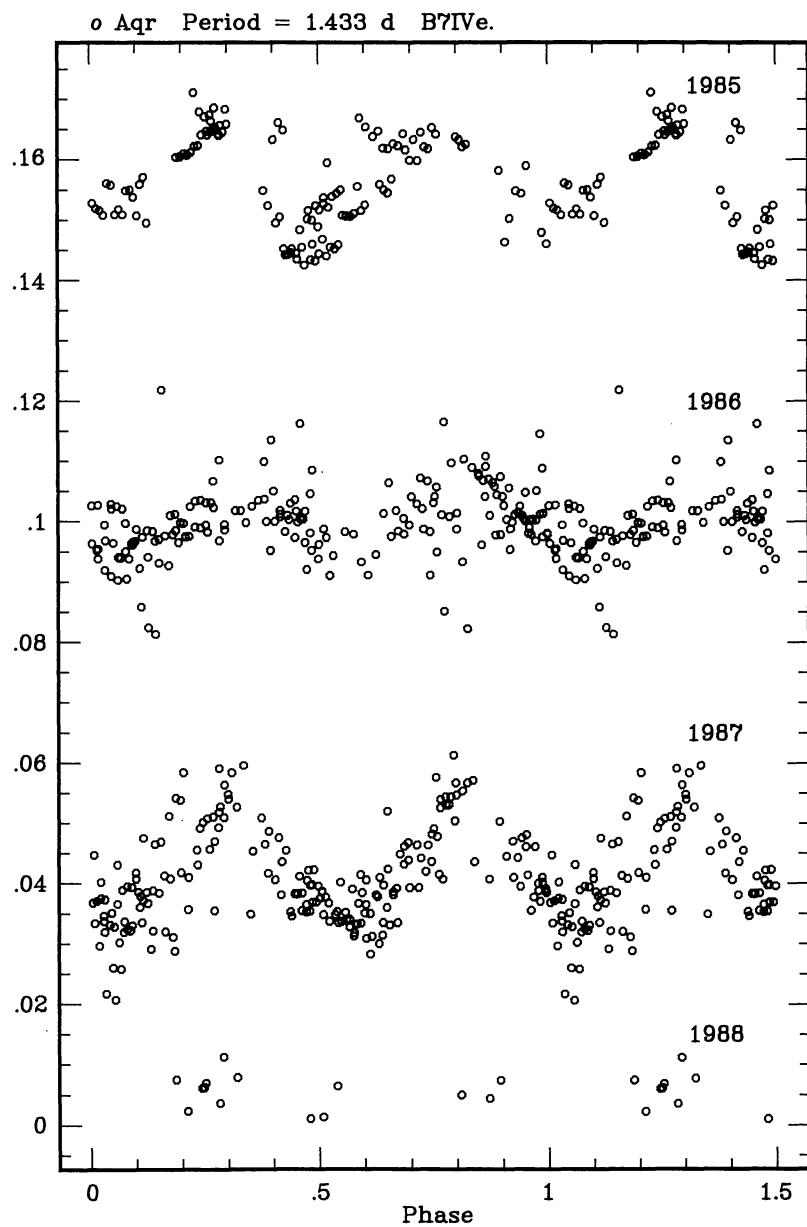
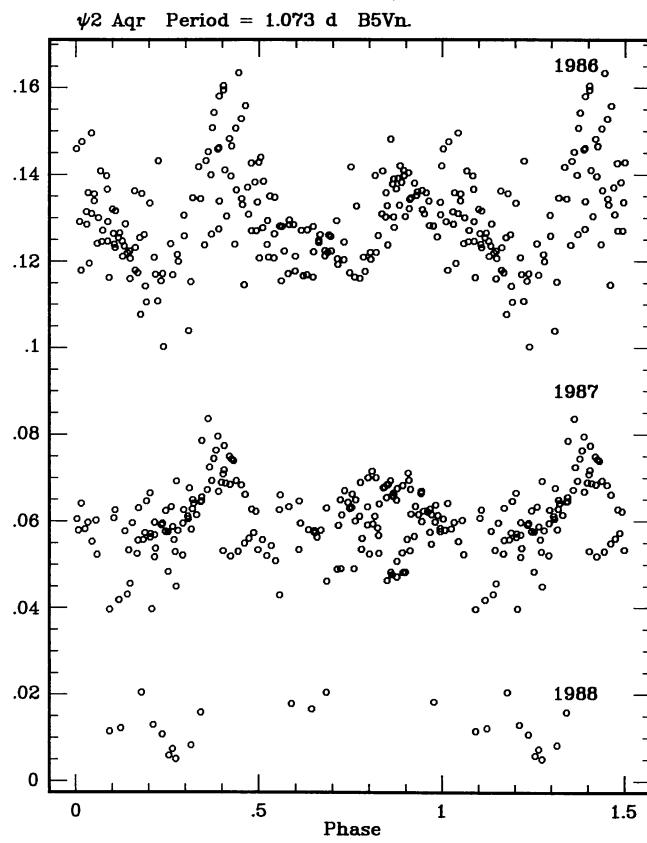
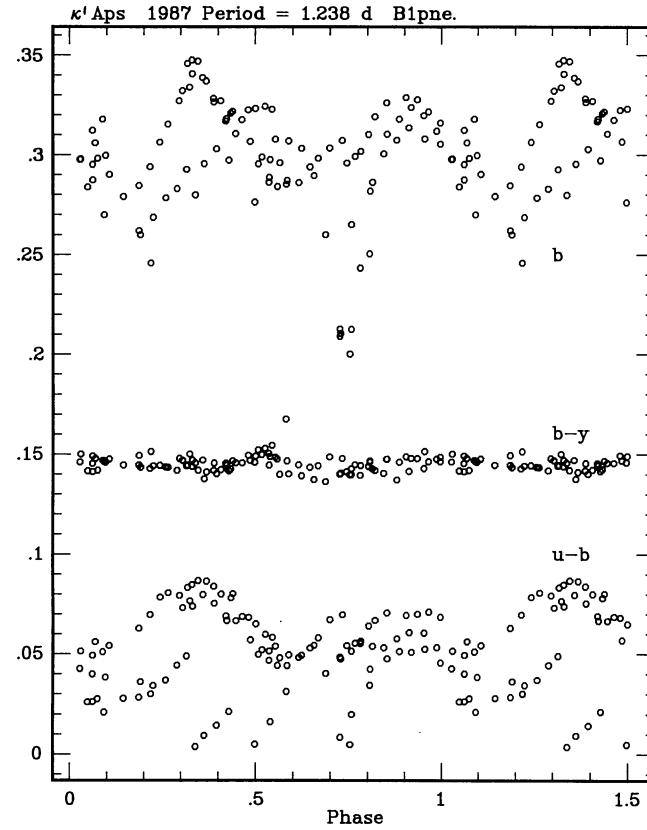
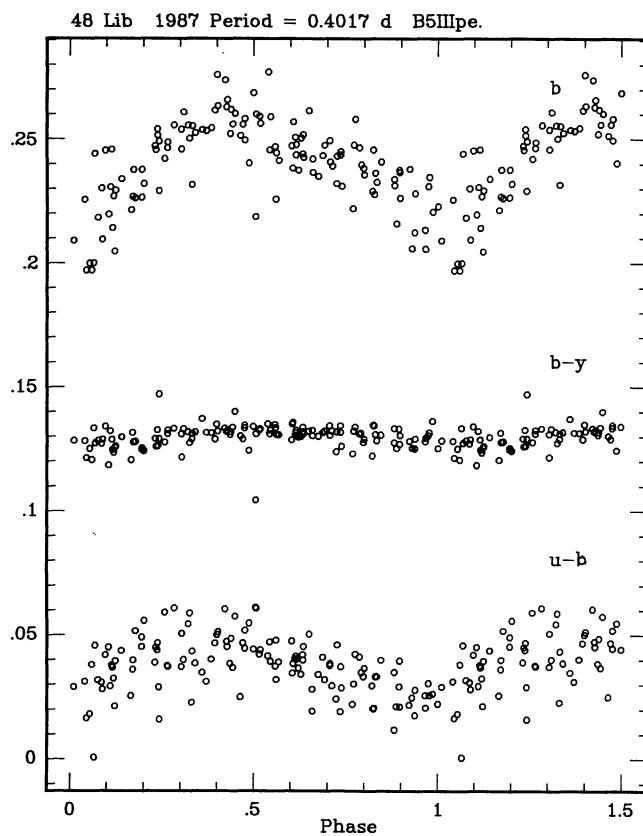
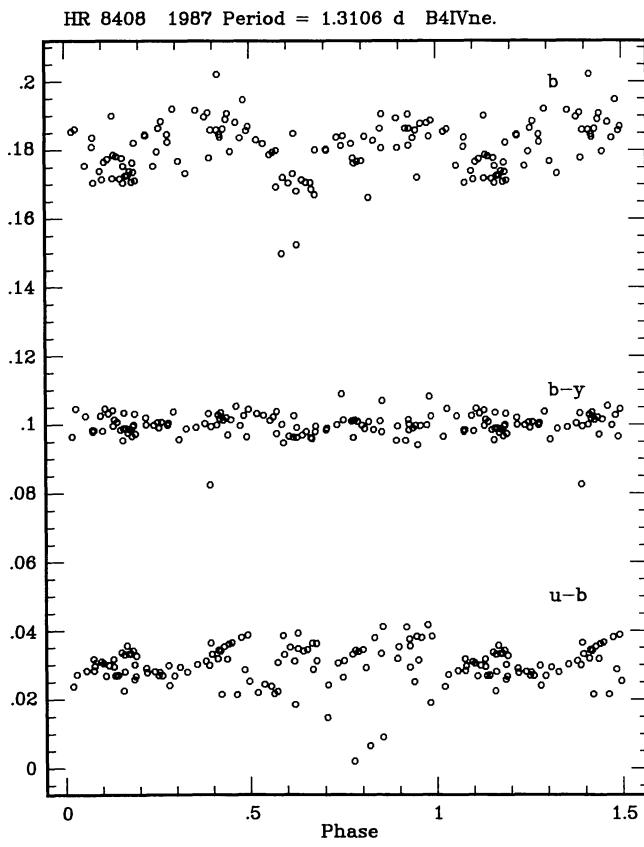
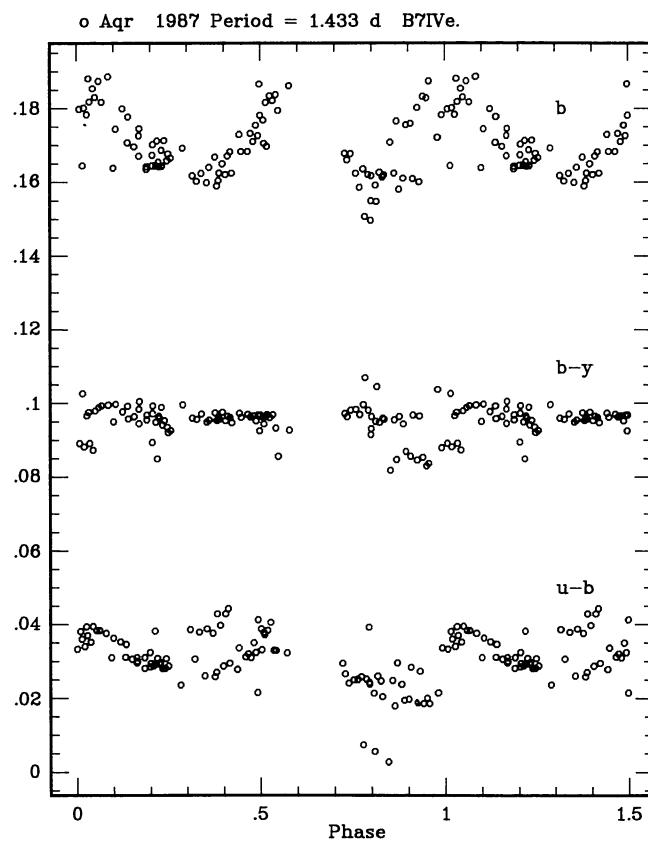
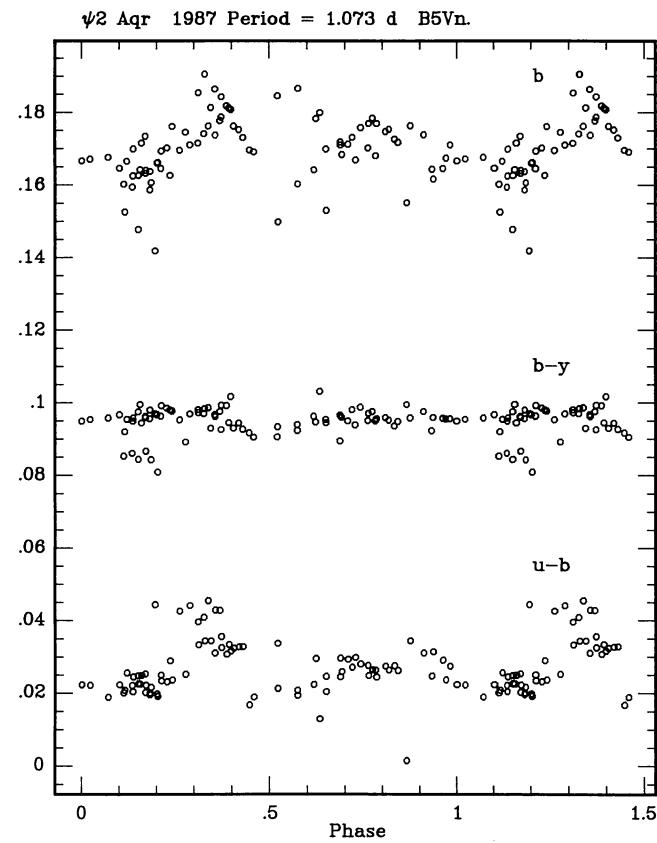


FIGURE 14. — Light curve of *o* Aqr ($P = 1.433$ d).

FIGURE 15. — Light curve of ψ^2 Aqr ($P = 1.073$ d).FIGURE 16. — Light and colour variations in κ^1 Aps ($P = 1.238$ d). In this and subsequent figures the colour and light variations are plotted according to the astronomical convention, i.e. with more negative values (bluer colour) upwards.

FIGURE 17. — Light and colour variations in 48 Lib ($P = 0.4017$ d).FIGURE 18. — Light and colour variations in HR 8408 ($P = 1.3106$ d).

FIGURE 19. — Light and colour variations in o Aqr ($P = 1.433$ d).FIGURE 20. — Light and colour variations in ψ^2 Aqr ($P = 1.073$ d).