SURVEY OF ASTEROIDS*

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ABSTRACT

A systematic survey of asteroids down to photographic magnitude 16.5 is described. The ecliptic belt was photographed nearly twice around in 1950–1952, to a width of 40° . The 10-inch f/7 Ross-Fecker telescope on loan from the Cook Observatory was used, and 1094 pairs of plates were taken, each 8×10 inches in size and covering $6^\circ 5\times 8^\circ.1$. In addition, 149 plates were taken on Selected Areas for magnitude calibration, as were special sequences for the determination of field corrections, etc. The plate pairs were blinked independently of previous knowledge and only afterward were re-examined for known objects missed. The asteroids found were measured for position, daily motion, and magnitude; and the subsequent identification work with the Ephemeris asteroids and objects having provisional designations was done with great care.

The statistics of the Survey are summarized in Table 1. Previously announced objects, not found in the Survey and either below the plate limit or, in some cases, probably spurious, are listed in Table 2. Asteroids missed because they were outside the 40° belt are given in Table 3. Ephemeris asteroids not found, presumably because they were too faint, are listed in Table 4; in addition, 182 objects were not observed because they were definitely too faint. Six new objects are probably Trojans. For 2 of them and for 2 other new asteroids, circular orbits are given in Table 5. For 33 additional new objects our data

suffice to compute circular orbits; they are listed in Table 6A.

The measures resulting from the Survey are contained in Table A. The positions have a probable error of about $\pm 3''$. The Survey magnitudes of Table A are combined with other magnitude data in Table 7. This table is on the International Photographic System and represents the final compilation of this paper; both the mean photographic opposition magnitude, p_0 , and the absolute magnitude, p_0 , are given. The resulting magnitude system was calibrated photoelectrically afterward, and the scale was found to be precise over the entire range, 7–16 mag. Table 7 is recommended for future use, with one reservation: for some three hundred fainter asteroids, present data are still inadequate; for these objects new measures will be published as Paper VIII.

It was found that magnitudes derived during a single opposition are not representative, no matter how accurate, because of fluctuations due to the aspect of the asteroid amounting to about ± 0.11 mag. (p.e.). This comparatively large effect indicates that a good fraction of the asteroids have large obliquities. The importance of good magnitudes in future identification work is stressed. Numerous controls and re-

visions were made which are described in Section IX.

The results of the Survey are not limited to an inventory for the years 1950–1952 of asteroid positions, identifications, and magnitudes on the photometric system. Since the blinking was carried out independently of previous knowledge, the *completeness* of the Survey could be determined in two independent ways: from overlapping Survey regions and from comparison with the Ephemeris asteroids. The degree of completeness of different Survey fields is found in Table 11; the asteroid numbers corrected for incompleteness are given in Table 12; and a quadratic interpolation formula representing these numbers as a function of apparent photographic magnitude is given in equation (5). The representation of the data by equation (5) is shown in Table 13. The counted numbers in the 1957 Ephemeris, arranged by mean opposition magnitude, p_0 , are found in Table 14. The figures are essentially complete for $p_0 < 14.5$. The representation by two interpolation formulae, equations (6) and (7), is also given in Table 14. These formulae are estimated to give approximate minimum and maximum numbers of asteroids for 14 < $p_0 < 18$, and lead to estimates of the completeness factors of the Ephemeris asteroids for this interval (Tables 16 and 17 and Fig. 4). These factors, in turn, are used in Tables 15 and 19 to derive the distributions in absolute magnitude, g, for six distance groups of asteroids 1.85–2.00–2.60–3.00–3.50–4.30 astronomical units and the Trojans.

The results for the three main zones, between 2.0 and 3.5 a.u., are plotted in Figure 5. Remarkable population differences are found, and the frequency-curves appear to consist of two parts, separated by a flat portion near g = 11, which corresponds to asteroid sizes near d = 30 km. One could surmise that this flat portion separates two modes of asteroid formation (condensation by accretion and collisional breakup), but it is considered premature to conclude this. Because of the population differences between the zones (Fig. 5), the center of gravity of the asteroid zone shifts toward the larger a-values for increasing

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g (smaller sizes). The ring 3.0 < a < 3.5 contributes 23 per cent of the 2.0 < a < 3.5 ring for 4.0 < g < 8.0; 39 per cent for 8.0 < g < 10.0; 70 per cent for 10.0 < g < 11.0; 89 per cent for 11.0 < g < 12.0; and 95 per cent for 12.0 < g < 13.0; the geometric-mean diameters of these five subgroups are about 300, 80, 40, 25, and 15 km. This result has important implications for the collisional production of meteorites.

80, 40, 25, and 15 km. This result has important implications for the collisional production of meteorites. The results for the fringe zones are as follows. The 3.5 < a < 4.3 group, of which 27 members are known, allows a fair analysis, which shows this group to have the same composition with g or diameter as the main asteroid zone (range $8\frac{1}{2} < g < 12$), with an abundance of 3 per cent of the main zone. The 1.85 < a < 2.00 group, with 11 known members, is inadequate for statistical treatment. Around g = 14 the abundance appears to be about $\frac{1}{2}$ -1 per cent of either the 2.0 < a < 2.6 or the 2.6 < a < 3.0 zone, but at g = 15 the fraction seems smaller. The Trojans ($a \cong 5.2$) are represented by 13 members, but their degree of completeness is uncertain because of special searches that have been made for them.

Because of the rapid increase of faint asteroids, it is not possible at this time to estimate the total mass of the asteroid ring.

I. PURPOSE AND PLAN OF SURVEY

A very large amount of observational and computational work has been done on asteroids, particularly during the last half-century; but this great effort has not included a systematic photographic survey, coupled with a determination of the asteroid magnitudes on a photometric system, as is required for statistical studies of the asteroids. The present Survey was organized by Mr. Kuiper with the aim of supplying such information.

The photographic survey became feasible through the co-operation of Dr. Charles Olivier and the University of Pennsylvania, who agreed to make the 10-inch (f/7) Ross-Fecker telescope of the Cook Observatory available on a loan basis. A 5- and a 4-inch telescope of the same design are placed on the same mounting. The instrument was transported by our staff to the McDonald Observatory in the summer of 1949. After complete overhaul, largely by Mr. A. Shatzel, and the installation of a new crystal-controlled drive, the telescope was put into operation in June, 1950. The cost of shipment and repairs, as well as that of a new fireproof housing equipped with a sliding roof, was covered by a grant made by the Research Corporation of America.

Two brief preliminary reports on the Survey have been published. The first is incorporated in the 1952 Report of Commission 20 of the International Astronomical Union (1954), which had appointed a subcommittee under Professor A. Kopff concerned with the improvement of the magnitude system of the asteroids. Professor Kopff recommended that no changes in the Ephemeris magnitudes be introduced prior to the completion of the present Survey (op. cit., p. 291). The other report was issued by Groeneveld and Kuiper (1954, pp. 200–201).

II. OBSERVATIONS

Tests of the remounted telescope in June, 1950, showed that the four components of the objective were not aligned. Later it was found that this had been caused by shrinkage of the hard-rubber separators between the lenses. Mr. Kuiper spent about a week at the McDonald Observatory in efforts to realign the components, after which the performance was considered adequate for a program using plates no larger than 8×10 inches. Since the full 20×24 -inch plates for which the telescope was designed could not be used in the blink comparators at the two observatories in any case, it was decided to proceed with the program, using 8×10 -inch plates. This plate size, covering 6.5×8.1 of the sky, plus the requirements of overlap between consecutive monthly oppositions, led to the following observing program.

During each dark of the moon, a field $40^{\circ} \times 40^{\circ}$, centered on the opposition point computed for new moon, was to be taken, covered by approximately 48 plate pairs. The number was not quite constant because of the eccentricity of the moon's orbit and the resulting non-linear motion of the above-mentioned opposition point. The long dimension of the plates was always placed north-south, and 8×6 adjacent plate fields sufficed to cover the $40^{\circ} \times 40^{\circ}$ field. This provided an overlap of about 10° in R.A. between consecutive months and some overlap between adjacent plates. Twelve $40^{\circ} \times 40^{\circ}$ fields

covered the entire ecliptic belt to a width of 40°; thereafter, the opposition fields would essentially repeat themselves. The oppositions were called A, B, \ldots, X , and the plate fields were designated by two digits, the first for right ascension and the second for declination. The plate field B26 meant the southernmost field in the second strip of R.A. of opposition field B. It was found that 10-minute exposures caused very little trailing of the asteroid images and that an interval of 1 hour was adequate for the discovery of asteroids in the blink comparator. Accordingly, 10-minute exposure times on 103-aO plates were used, which led to a limiting magnitude of 16.5-17.5 pg for the faintest visible stars and 16.0-17.0 for discoverable asteroids. This limit did not include the faintest known asteroids, but this was regarded as inevitable. The plates were taken in the order, e.g., B11, B12, B13; B11, B12, B13; B14, etc. This allowed 10 minutes for changing between fields. Later, as the observer became more experienced, four or even five plates would be taken consecutively before the second set was started. The plate centers were selected on suitable guiding stars, preferably not fainter than the eighth magnitude. This made the plate centers depart slightly from the desired geometric pattern, but this caused no serious problem. The roughly 430 plate centers were selected by Mr. Kuiper, and the observing charts were prepared under his supervision. The plates of each pair were designated by the field number, followed by a and b. The displacement of consecutive monthly opposition fields being about 30°, two of the eight plate strips in R.A. were repeated as part of the next month's program. Plate fields may thus have two designations: e.g., E83 = F23. After the ecliptic had been covered once, the plate fields were repeated with their new designations; thus field B became field N during the second cycle. In all, 23 consecutive monthly oppositions were covered, called B-X, the surveys on field A were too incomplete to be useful, and even the B and C fields were only about two-thirds complete. Since the synodic period of an asteroid at 2.8 a.u. is 15 months, nearly half the asteroids in the 40° zone were covered twice by the present Survey. The selected belt width of 40° was, of course, a compromise. Some objects of high inclination were thus not included; but increasing the width for complete coverage would have meant a disproportionate increase in the amount of work.

The observations were made almost entirely by two night assistants—first, Mr. W. C. Braun and, after February, 1951, Mr. H. Rubingh. The observing program for each opposition field had to be carried out within 2 or 3 hours from midnight, during an interval of about 2 weeks (the dark of the moon). Since about 96 asteroid plates and at least 10 calibration plates on Selected Areas had to be taken, this meant that, depending on the declination, 7–9 perfect nights were needed for each monthly quota. Periods of bad weather, of course, caused difficulties, and it was only just possible to maintain a record of unbroken observations for the entire period of 23 months. With a climate less satisfactory than that at the McDonald Observatory, this assignment would have been impossible. However, the record is by no means perfect. The observers had often to compromise and use nights which were not of standard quality. This has resulted in some loss of uniformity in the Survey. This trouble would have been reduced if larger plates could have been used.

Another source of lack of uniformity is the variability of the photographic emulsion. Five opposition fields were taken on plates that showed chemical fog (M-Q). While the image quality and the guiding were, on the whole, satisfactory, the images near the corners of the plates were not good, and this will have led to some loss of limiting magnitude.

The total number of plates taken between August 10, 1950, and May 26, 1952, is 2404. This includes 1094 pairs of plates taken with time intervals varying between 50 and 90 minutes and 149 plates centered on Selected Areas. The latter were taken for magnitude calibrations of the asteroids; at least one was taken each night. The remainder consists of rejected plates and incompleted pairs, usually caused by intervening clouds. The statistics of the entire material are found in Table 1.

It will be clear that the unbroken series of observations extending over about 2 years required unusual energy on the part of the observers. We are much indebted to Messrs. Braun and Rubingh for their devotion to this project.

III. THE BLINKING OF THE SURVEY PLATES

The asteroids were found by blinking the plates in the blink-comparator. The blink observers are listed in Table 1. The number of asteroids found on one plate varied between 0 and 18, and the average was 3. In these numbers the results of reblinking are included. The efficiency of the blinking depended on the scanning speed; high speed resulted in the missing of faint asteroids. Plates taken in regions with a dense star background took 3–5 hours to blink; far from the Milky Way the time approached 2 hours. As expected, the probability of finding the asteroids showed a magnitude dependence, but

TABLE 1
STATISTICS ON THE ASTEROID PROGRAM

			Овл Fou			Asteron Found†			FAINT-		
FELD	Opposition	PLATE PAIRS	Total	By Re- exam.	Ident.	Un- ident.	Per Cent Un- ident.	Av. PER PLATE‡	EST MAG. MEAS- URED	BLINKED BY§	Re- marks
B C D E	1950 Aug. Sept. Oct. Nov.	32 30 52 50 45	83 88 143 128 119	10 13 29 21 15	53 55 95 74 75	8 17 25 21 12	15 24 21 22 14	2.6 2.9 2.7 2.5 2.6	16.1 16.3 16.4 16.5 16.3	Fu Fu Fu Fu Fu	
F G	Dec. Dec.\ 1951 Jan.∫	51	153	31	82	29	26	3.0	16.5	Ke	
H	Feb. Mar. MarApr. AprMay May-June June-July July-Aug. AugSept. SeptOct. OctNov.	59 37 62 46 49 50 48 48 45 50	131 123 191 159 130 108 95 105 65 90	21 07 18 16 14 31 25 15 22 #	78 67 114 90 74 78 69 65 47 61	22 31 31 18 30 09 03 10 06 04	22 32 21 17 29 10 06 13 12 06	2.2 3.3 3.1 3.5 2.7 2.2 2.0 2.2 1.4 1.8	16.6 16.9 16.8 16.6 16.2 16.4 15.8 15.9	Th, Ge Gr Gr, Th Gr VH Th VH Th VH Th VH	1 1 1 1
R S T	NovDec. Dec. 1952 Jan Feb.	32 41 73	160 238 207	$egin{array}{c} 60 \\ 20 \\ 64 \\ \end{array}$	59 103 112	69 81 46	54 44 29	5.0 5.8 2.8	17.0 17.4 17.6	Th VH Th	1 2 2 2
$egin{array}{c} U \ldots \ldots \ V \ldots \ldots \ W \ldots \ldots \ X \ldots \ldots \end{array}$	Feb. Feb. Mar. Apr. May	55 36 - 53 50	254 128 223 126	54 27 26 #	138 72 110 71	69 21 65 26	34 23 37 27	4.6 3.5 4.2 2.5	16.7 16.6 17.2 16.6	Th Th Th, VH, Gr Gr, VH	3 3 3
Total program		1094	3247				26	3.0			

^{*} All objects numbered on the plates.

[†] Asteroids found, excluding recurrences within an opposition field.

[‡] Refers to number of Objects.

[§] Fu = Fujita; Ge = Gehrels; Gr = Groeneveld; Ke = Kent; Th = Thorson; VH = Van Houten.

^{| 1 =} poor plates; 2 = excellent plates; 3 = some poor nights.

[#] See text.

personal effects played a role also. A dense sky background was a disturbing element, and especially in the southernmost fields (near 18^h R.A.) this posed a problem. Also the quality of the images (fuzziness, trail), the proximity of stars, and local plate fogging had their influence. Under favorable conditions and with proper care it was possible to get an almost complete inventory down to about 1 mag. above the plate limit. This was accomplished only rarely, however. Further, there was some difference in appearance between the left and the right fields of the blink microscope, which could not be readily corrected. Besides, for plates taken at low altitude, differential refraction caused some scale variation.

In a program of this size it is not possible to avoid all spurious objects. The working lists of suspected asteroids were made to include, obviously, only such image pairs as seemed real by careful inspection at the blink microscope and as showed retrograde, asteroidal-type motions. About two-thirds of the image pairs recorded could be checked afterward because of overlapping plate fields, either within the same opposition field or between consecutive monthly oppositions. These checks were always made when possible and resulted in the rejection as spurious of about 2 per cent of the image pairs. In some cases the check was inconclusive because of differences in the limiting magnitudes of the plates; these objects were retained if they looked real, in spite of their absence on the companion plates. Any remaining spurious image pairs will appear as unidentified objects in our tables; on the basis of the overlapping fields, their total number may be of the order of 20.

The counterpart of inclusion of spurious image pairs is the omission of real asteroids. The completeness of the Survey may be checked from the overlapping areas and by comparison with the lists of previously known asteroids. Table 1 shows that 74 per cent of the retained image pairs can be identified with asteroids that are either numbered or had previously assigned provisional designations. However, a number of objects with provisional designations could not be detected on our plates in spite of the fact that the published magnitudes indicated that they should have been present. This matter is discussed in Section V.

The total task of blinking the nearly 1100 pairs of plates was very large and arduous, and more than one-third of it was carried out by Mrs. Helen E. Thorson. We are much indebted to her for her great contribution to this program.

IV. THE TEN-DAILY MOTIONS

The ten-daily motions were determined with an eyepiece micrometer attached to the blink comparator, from measurements of the displacements in polar co-ordinates. The screw value was calibrated three times independently, with accordant results. However, the measures show small systematic plate errors in R.A. The angles were measured with respect to one edge of the plate; this assumes that the non-parallelism of the meridians may be neglected up to $\pm 40^{\circ}$, which is a rough approximation, although not too serious if the motions are used for identification only. The B, C, D, E, and F fields were measured by Mr. Kent; and the remaining 18 fields by Mr. Van Houten and Miss Groeneveld. The measures are recorded in Table A, below, the main table of observations. Survey numbers followed by X were found by reblinking; the column "VAR" gives O-C in declination computed from O-C in R.A. by using the variation. In a few fields (e.g., G, H, I, R, and S) some Survey numbers were inadvertently used twice, in separate observing runs; e.g., G76.3 = 1136; G76.3 = 667.

V. THE IDENTIFICATION PROBLEM

Approximate positions (to about $\pm 1'$) were determined for nearly all asteroids found, by comparing the plates with the BD or CD charts or by measuring the plates with a scale, starting from a known BD star. These positions, together with the measured daily

motions and rough magnitudes, were compared with the Ephemeris positions of the "numbered" asteroids (i.e., those in the Ephemeris); for 1950 and 1951 the Cincinnati edition was used, and for 1952 the Russian edition. Uncertain identifications are given in Table 6B. In some cases improved ephemerides were available and used; these are contained in Table 6C. The positions for the Survey dates were computed from the Ephemeris graphically or by using second differences in the interpolation. Account was taken of the variation line. Additional criteria for identification were the daily motion and the magnitude, though the latter was not very reliable. The largest O - C found was 29 minutes, and several were in excess of 20 minutes. The ten-daily motions for the fields B-F are of low accuracy; they were in part remeasured, when special identification problems arose. In addition, comparisons were made with published positions of "unnumbered" asteroids (i.e., asteroids having provisional designations). When these objects had not been found during the blinking of the plates and should have been above the plate limit according to the published magnitudes, the plates were re-examined, provided that the time interval was less than 5 days, unless these objects had published motions as well, in which case the allowed interval was increased. These positions were taken from the Minor Planet Circulars (MPC) issued by the Minor Planet Center at Cincinnati. Furthermore, ephemerides for unnumbered asteroids were occasionally available, also taken from the MPC's. Finally, Dr. Herget generously made available his "Index" of asteroid positions, arranged in order of asteroid number. This Index has proved most helpful in both the identification work and the study of the magnitudes.

Within the list of Survey asteroids itself, there were many cases of the same object, known or unknown, having been observed more than once. Identities of objects found on different plates of the same opposition field were established by superposing the overlapping regions and correcting for motion where necessary. For identities of new objects in successive opposition fields, the ten-daily motion was extrapolated.

Miss Groeneveld and Mr. Van Houten shared equally in the identification work and are responsible for the identifications assigned; preliminary work on the first five opposition fields had been done by Messrs. Fujita and Kent. Table A contains the results.

After the plates had been blinked and all identifications made, the plates were reexamined for the following groups of missing objects: (a) "numbered" asteroids that were not found on the plates and (b) "unnumbered" asteroids, found at other observatories but missed on our plates, although taken within 5 days from the published positions (cf. three paragraphs above). For group a the plates were reblinked around the Ephemeris position and, if necessary, along the variation line up to a distance of roughly 3° in each direction. If the asteroid was still not found, it may have been too faint or covered by star images or have an ephemeris that was grossly in error (say, more than 8^m in R.A.). For the good fields, objects looked for in the re-examination included all asteroids, numbered and unnumbered, which were expected to be 16.5 photographic or brighter on the International scale. For the Q field, which had strong chemical fog, the limit 15.5 was used, and for the M-P fields, which were less fogged, 16.0 mag. A similar limit was used on some inferior plates of other fields. For objects in group b the searches were confined to about 1° from the expected position. Several of these objects could not be found, in spite of the fact that the time difference was small enough and the magnitudes probably bright enough; a list of such objects (not claimed to be complete) is given in Table 2. Some of these published objects may not have been real.

The measurements are collected in Table A, which was printed at the Cincinnati Observatory under the direction of Dr. Herget from the data sheets prepared at Yerkes. The statistics of the Survey are shown in Table 1. The number of "objects found" includes recurrences within the same opposition field, while the number of "asteroids found" counts each object only once. The column "Ident." gives the number of asteroids that could be identified within a given opposition field (either "numbered" or "unnumbered"); recurrences between different opposition fields may occur, so that the total of

the column "Ident." is larger than the total number of identified asteroids. Also, the numbers listed under "Unident." refer to different objects of the same opposition field, because adjacent plates were always sufficiently close in time to identify new objects common to both plates. The percentages of unidentified objects are therefore correct, except for the occasional inclusion of a spurious image pair, already referred to.

The Q and X fields were only partly blinked; most of the Q plates were badly fogged, while 12 pairs of the X field appeared to have been tilted during the exposures, causing them to be partly out of focus. These defective plates were not blinked independently but were used only to locate the known asteroids. For these reasons, only the total numbers of asteroids found are given for these fields; and the percentage of unidentified asteroids is not representative, particularly for the Q field.

TABLE 2
UNNUMBERED ASTEROIDS ANNOUNCED BY OTHER OBSERVATORIES
BUT NOT FOUND IN SURVEY

Preliminary Designation	Mag.*	Institution	Preliminary Designation	Mag.*	Institution
1950 OE	15.5 pg 14.2 astr. 15.5 pg 15.2 pg 14.9 pg 14.8 pg 14.2 pg 15.1 pg	Indiana Indiana Indiana Indiana Heidelberg Indiana Indiana Indiana Indiana Indiana Indiana Uccle Uccle Nice Nice Johannesburg	1951 WF	14.0 astr. 12.5 astr. 15.6 pg 14.3 astr. 14.5 astr. 14.4 astr. 14.0 astr. 14.4 astr. 14.0 astr. 14.6 astr. 14.6 astr. 14.1 astr. 14.5 astr. 14.5 astr.	Nice Nice Indiana Heidelberg Nice Nice Nice Nice Heidelberg Heidelberg Heidelberg Heidelberg Heidelberg Heidelberg Heidelberg La Plata

^{*} pg is approximately on the International scale; astr. on the Ephemeris scale (needs correction of about +2 mag.).

If the recurrences in Table 1 are allowed for, a total number of asteroids, numbered, unnumbered, and new, is estimated to be 1550. Of this total, 1167 are numbered asteroids. The 1952 Ephemeris contains 1568 numbered asteroids; the 1956 Ephemeris, 1605. By February, 1956, the count had reached 1615. These additions to the numbered objects could be included as such, because of the current information supplied by the MPC's, including observations and ephemerides of unnumbered objects.

It should be pointed out that not all 1615 numbered asteroids have published ephemerides. In the 1956 edition there are 15 objects without ephemerides (155, 330, 452, 473, 525, 531, 612, 682, 719, 724, 831, 843, 864, 879, and 903), most of which have not been observed for decades and may in fact be identical with unnumbered or even numbered asteroids. As stated before, 23 monthly oppositions were taken, although the synodic period of an asteroid at 2.8 a.u. is 15 months. As a result, some asteroids were covered twice by the Survey, while others were covered once; further, they must have been either inside or outside the Survey fields and, if inside, either above the plate limit or below it (outside, bright, or faint). For asteroids at longitudes covered twice, all six combinations between o, b, and f occur; the three oo, bb, ff are classified as o, b, f, occurring singly; while ob and bf were counted in Table 1 as b and of as f.

Of the 1615-15 = 1600 asteroids, 111, or 7 per cent, were outside the regions photo-

graphed in the Survey, according to the ephemerides; they are listed in Table 3. Further, 182, or 12 per cent of the 1489 objects within the boundaries of the Survey, are found to have been too faint for our plates at the time of observation by the magnitude criteria stated before. It is noted, in this connection, that asteroids marked 15.0 mag. in the Ephemeris are, in reality, about 17.0 pg. Two asteroids (561, 920) appear to have been lost between adjacent plate fields. There remain 137 numbered asteroids within the boundaries of the Survey (9 per cent of the 1489) that were at least once above the supposed plate limit but were not found. They are listed in Table 4. It appears that most of these objects are close to the plate limits; none of them were expected to be brighter than 14.0 and only 11 were expected to be brighter than 15.0 mag. Intrinsic variability plus the large accidental error of the magnitudes in the Ephemeris, coupled with the rapid increase of asteroid numbers with increasing magnitude, will have caused some objects actually below the plate limit to have come out above it according to the reduced magnitudes. The five fields M-Q account for 38 per cent of the cases; the limiting magnitude

TABLE 3
ASTEROIDS OUTSIDE REGIONS PHOTOGRAPHED

					1	Ī
2	323	605	768	978	1166	1362
25	329	617	771	998	1170	1373
36	372	626	779	1019	1191	1437
85	386	634	785	1025	1208	1474
89	391	663	787	1031	1215	1477
99	413	679	806	1035	1222	1508
130	433	692	849	1036	1241	1509
132	434	697	860	1049	1252	1521
157	471	704	862	1050	1263	1547
164	475	706	880	1093	1264	1548
170	483	714	881	1101	1276	1566
176	493	729	911	1103	1303	1568
181	536	733	926	1108	1310	1584
225	564	747	930	1139	1317	1585
290	594	751	950	1140	1318	1612
292	596	754	977	1146	1341	
- 0						

on these fogged plates may have been slightly overestimated. Other objects may have large O-C compared to the Ephemeris and may instead be listed among the unnumbered asteroids. A few objects may have been missed even in the reblinking.

From the daily motion, 6 new objects were suspected to be Trojans: N73.1, R36.5, R46.3, R54.15, S13.11, S26.2. Mr. Van Houten has derived circular orbits for two of them, Nos. 3 and 4, based on two observations each, separated by 18 days, made in December, 1951. The circular elements are given in Table 5. He has also computed circular orbits for two non-Trojans given in Table 5. Finally, he computed ephemerides for 1606, identical with B41.2; 1614, identical with G36.1; and 1605, probably identical with H64.1. Miss Groeneveld computed an ephemeris for 1586, identical with L13.4. Table 6A lists additional new objects for which circular orbits could be computed on the basis of our observations. Table 6B lists asteroids in Table A with uncertain identifications; and Table 6C lists asteroids identified from special ephemerides.

VI. THE MEASUREMENTS OF POSITION

The original Survey plan did not include the determination of accurate positions for the more than 3000 asteroids on the plates. The task of such a determination was considered beyond the available resources in time and manpower. This situation changed when Mr. Van Biesbroeck (1955) developed the simple method of measurement with a

TABLE 4
NUMBERED ASTEROIDS NOT FOUND

No.	Field, Year	Mag.	No.	Field, Year	Mag.	No.	Field, Year	Mag.
343 603 616 630 641 698 699 725 730 745 765 802 810 812 821 822 833 837 840 842 855 857 869 870 890 897 942 959 963 968 970 983 988 993 1000 1017 1020 1022 1026 1037 1052 1064 1068 1069	M 51 J 51 Q 51 N 51 DC 51, V 52 PO 51 W 52 K 51 T 52 P 51 G 51 KJ 51 E 50, U 52 S 51 K 51 F 50, T 52 QR 51 F 50, T 52 QR 51 F 50, T 52 QR 51 F 50, T 52 G 50 O 51 J 51, X 52 F 50, T 52 G 50 O 51 P 51 F 50 D 50 N 51 P 51 F 50, V 52 H 51 DC 50 G 51 M 51 F 50, V 52 H 51 DC 50 G 51 M 51 F 50, V 52 H 51 DC 50 G 51 M 51 F 50, V 52 H 51 DC 50 G 51 M 51 F 50, V 52 H 51 DC 50 G 51 M 51 F 50, V 52	16.1 14.7 14.3 15.7 15.9, 16.5 16.0 16.5 16.4 16.2 15.1 15.9 15.5 15.8 14.9, 16.4 14.9 15.2, 16.1 14.8 16.3 15.9 15.8 16.4, 16.3 15.9 15.8 16.4, 16.1 15.9 15.8 16.4, 16.1 15.9 15.8 16.3, 16.5 16.4, 16.3 16.5 16.4, 16.5 16.5 16.4, 16.3 16.5 16.4, 16.3 16.5 16.4, 16.3 16.5 16.4, 16.5 16.5 16.4, 16.5 16.5 16.6, 16.5 16.6, 16.5 16.6, 16.5 16.6, 16.5 16.7 16.0, 15.7 16.0, 15.7 16.0, 15.7 16.0, 15.7 16.2 15.6 16.2 15.6	1081 1082 1095 1127 1138 1151 1152 1156 1168 1181 1190 1192 1209 1224 1225 1226 1228 1231 1233 1234 1260 1265 1285 1290 1311 1313 1314 1328 1330 1332 1335 1337 1342 1355 1357 1358 1359 1360 1361 1371 1371 1392 1401 1392 1401 1392 1401 1392	H 51 G 51, U 52 M 51 F 51 M 51 O 51 D 50, VU 52 S 51 N 51 IJ 51 O 51 M 51 O 51 KL 51 F 50, U 52 I 51 O 51 K 51 H 51, W 52 T 52 T 52 T 52 P 51 B 50 U 52 T 52 T 52 F 50, U 52 T 51 D 50, V 52 T 51 D 50, V 52 T 51 D 50, V 52 T 51 D 50, U 52 T 52 T 52 T 52 T 52 F 50 U 52 T 52 T 52 T 52 T 51 B 50 U 52 T 52 T 52 T 52 T 52 T 52 T 52 T 52 T	14.9 16.4, 16.5 14.2 16.3 15.5 16.3 15.3 16.0, 15.8 16.2 15.7 16.4 16.0 15.8 15.5 15.9 16.0 16.3, 15.9 15.8 15.1 16.0 15.3 14.7, 15.0 16.3 16.3 15.5 16.2 15.8, 15.1 16.1 16.5 15.6 15.7 15.3 15.7 15.3 15.7 15.3 15.7 15.3	1406 1408 1408 1414 1420 1425 1430 1431 1433 1442 1444 1446 1449 1451 1456 1462 1463 1466 1471 1475 1481 1488 1495 1497 1514 1517 1526 1529 1531 1533 1535 1540 1559 1561 1573 1576 1579 1588 1595 1598 1610 1611	G 51 N 51 I 51 G 51, U 52 O 51 EF 50 T 52 Q 51 M 51 W 52 H 51 I 51 I 51 F 50 H 51, X 52 O 51 M 51 C 50 N 51 D 51 M 51 F 50 H 51, X 52 I 51 M 51 F 50 H 51, X 52 I 51 M 51 F 50 F 51 M 51 F 50 F 51 M 51 F 51 I 51 I 51 I 51 I 51 I 51 I 51 I 51 I	16.0 16.0 16.2 16.4, 16.4 16.0 16.4 15.2 15.7 15.5 15.4 15.7 15.3 16.5 15.9 16.0 16.4 15.3 16.2 15.7 16.0 15.9 16.2 16.4 15.3 15.2 15.6 15.5 15.7 16.0 16.3 16.3 16.2 17.6 18.9 18.9 18.9 18.9 18.9 18.9 18.9 18.9

TABLE 5

NEW ELEMENTS, CIRCULAR ORBITS (VAN HOUTEN)

Object	a	i	Ω	u_0	t_0
D75.4 = E16.3 = 1950 TH ₂	3.066	14°21	175°56	209°61	1950 Oct. 13,2299
B63.2 = C14.6 = 1950 OC1	4.220	4.09	346.22	345.63	1950 Aug. 18.3389
R54.15 = S54.3 = 1951 XK	5.076	18.55	85.6	346.8	1951 Dec. 5.2535
R46.3 = S46.2 = 1951 XJ	4.912	21.38	106.6	317.9	1951 Dec. 4.2104

precision theodolite which gave equatorial co-ordinates with an accuracy of about 2''-3'', satisfactory to the computers of asteroid orbits. It was then decided that systematic position measurement, which would greatly enhance the value of the Survey, was feasible, and Mr. Van Biesbroeck was interested to undertake this task himself. Much of the uncertainty in the positions was found to be due to the lack of reliable proper motions for the comparison stars. For declinations from -30° to $+30^{\circ}$ these were usually taken from the Yale Zone catalogues; for declinations north of $+30^{\circ}$ from the AGK2 catalogue or Prager's catalogue; and for zones south of -30° from the Cordoba AG catalogue. Occasionally, near plate edges or corners, these brighter stars were not suitably located, and fainter stars had to be used, taken from the Astrographic Catalogue. The asteroid positions, for the equinox of 1950.0, are given in Table A to $0^{\circ}1$ in R.A.

TABLE 6A
ASTEROIDS FOR WHICH CIRCULAR ELEMENTS CAN BE COMPUTED
(Found in Overlaps of Opposition Fields)

1950 QB 1 QD 1 QE 1 QF 1 QG 1 SS	1950 ST TV 3 TZ 3 TB 4 WB WC	1950 WD WE WF XS 1951 EY EX 2	1951 WL XR XW XX XX XY XB 1	1951 XD 1 XE 1 XF 1 YR 2 1952 CB HU 3	1952 HW¶3 HY 3 HZ 3
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TABLE 6BAsteroids with Uncertain Identifications

285, <i>U</i> 24.2	1053, T32.2	1505, W66.1
428, <i>X</i> 34.2	1113, R41.1	1505, W76.1
428, <i>X</i> 35.2	1183, U73.4	1532, X16.1
580, <i>C</i> 34.1	1218, R33.11	1605, H64.1
603, <i>I</i> 44.10	1218, R34.5	1605, H65.1
1053, T31.2	1321, S12.3	

TABLE 6CASTEROIDS IDENTIFIED FROM SPECIAL EPHEMERIDES

350, M75.2 Russ. Eph. 1951	1590, T45.5 MPC 660
396, J54.3 Russ. Eph. 1951	1599, E34.4 MPC 704
919, D12.4 MPC 460	T82.3
1099, S41.3 MPC 602	1601, X14.6 MPC 736
1349, B13.5 MPC 595	X23.2
B14.1a	1603, X42.2 MPC 736
1523, R32.1 MPC 727 1560, S53.8 MPC 528	X52.1 1604, U95.5 MPC 659
1569, W21.1 La Plata Circ. 9	V16.3
1570, K34.3 MPC 659	V25.2
1572, U73.8 Astr. Circ. U.S.S.R., 120,	1605, X21.4 MPC 736
1578, U63.9 MPC 692	1607, T93.5 MPC 697
1589, T92.1 MPC 731	U13.12 MPC 697
U12.2	1951QA, P43.1 MPC 671
U13.7	P53.1

and 1" in declination; they have probable errors of 2"-3". The measurements and reductions were made by Mr. Van Biesbroeck; the preparatory selection of most of the comparison stars was made by Miss Groeneveld and Mr. Van Houten.

VII. THE MEASUREMENTS OF MAGNITUDE

During the Survey a calibration plate on a Selected Area was usually taken each night. In addition, of course, all opposition fields north of -19° declination contained Mount Wilson Selected Areas among the Survey plates themselves; and on nights on which fields south of -19° were taken, plates north of this limit were usually taken also. The Selected Areas on the Survey plates themselves were, on the whole, more satisfactory than the special plates (always taken at $+15^{\circ}$ declination) because they were closer to the asteroids and the sequences occurred at random positions on the plates, so that field corrections were reduced.

For the first three opposition fields, August 10-November 7, 1950, the asteroid magnitudes were derived as follows. Fifteen stars were selected on each of the overlapping edges, five each near the fifteenth, thirteenth, and twelfth magnitudes, and the system was transferred outward, starting with the Selected Area field, until all plate fields were covered. The overlapping edges are about 2 cm wide. Obviously, the differences in the quality between the different plates, as well as changes in the over-all absorption due to haze, are thus eliminated; differential extinction corrections were applied separately.

A less laborious method was followed for the remaining opposition fields, in which the asteroids were compared directly with the nearest available Selected Area sequence. Actually, two sequences could be used for most nights, and they were both measured. This transfer method relies, of course, on the approximate constancy of the atmospheric transparency, developing, etc., for the plates to be compared. In case no Selected Area sequence was available, a special sequence was set up afterward. Such sequences were established photographically with the Yerkes 24-inch reflector by Mr. Gehrels and with the 10-inch Survey telescope at McDonald by Mr. Van Houten. The latter were taken at -28° during the summer of 1955, by matching 10-minute exposures in this zone with similar plates of Selected Areas at -15° , but taken at the same zenith distance. From these plates seven special sequences near -28° were set up by Mr. Gehrels.

For purposes of identification, charts had to be prepared for 68 Selected Areas, each showing about four stars per magnitude interval between magnitudes 10 and 17. When bright stars were lacking in the Mount Wilson Catalogue of Photographic Magnitudes, such stars were taken from the Bergedorfer Spektral-Durchmusterung, with a magnitude correction of +0.04 mag. applied. The measures were all made by Mr. Gehrels with the Ross (1936) photometer of the Yerkes Observatory, after it was found that intercomparisons with a fixed scale of images (a "flyspanker") were unsatisfactory, apparently because such a fixed scale does not reproduce the effects of different image qualities. Since the thermopile of the Ross photometer gave only small deflections for the darker plates, it was replaced by a gas-filled cell; its output was fed through an M.I.T. amplifier to the galvanometer. The quotient, galvanometer reading of star-plus-sky to sky-only, was plotted for the Selected Area stars against their magnitudes; similar quotients for the asteroids could then be converted to observed magnitudes. The diaphragms employed in the Ross photometer were 0.15 and 0.25 or 0.28 mm (projected diameters). Stars fainter than photographic magnitude roughly 13 were measured with the 0.15-mm diaphragm, those brighter than 13 with one of the two larger diaphragms.

The observed magnitudes require four corrections to become apparent magnitudes: (1) differential extinction; (2) field effects; (3) asteroid trailing; and (4) a small correction to the system of the Selected Area magnitudes. For 1 a mean extinction coefficient of 0.30 mag. per unit air mass was used. The field correction of the 10-inch Survey telescope, 2, was determined from the existing calibration plates taken of Selected Area 68, sup-

plemented by a new series of 24 exposures of 20 minutes each, taken with the 5-inch Cook Observatory telescope at McDonald during the summer of 1953. The 10-inch plates had the Selected Area in the center; and sequences were selected elsewhere on the plates and calibrated by means of 5-inch plates, taken with different centers. In the reductions it was assumed that all field corrections are radially symmetrical and linear with distance from the center. On this basis the plot of Figure 1 was obtained, which gives the field correction for a standard distance interval of 10 cm and its dependence on the photographic magnitude. The interpretation of Figure 1 is probably approximately as follows: images near the edges of the plate will differ from those near the center in being slightly elongated; furthermore, there is a vignetting effect which becomes appreciable only near the corners of the plate. For faint objects the field effect will be small; it will increase, beginning about 2 mag. above the plate limit, as is true for the trail correction; but for bright objects, having large images, the effect will decrease, and for very bright objects vignetting in the extra-focal disk becomes effective.

The trail correction, 3, could not be determined from plates taken with the 10-inch Survey telescope because its optics were being overhauled after completion of the Survey;

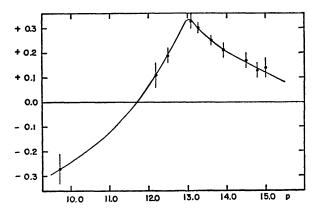


Fig. 1.—Field correction, 10-inch Survey telescope. Abscissae, photographic magnitudes; ordinates, field correction for difference of 10 cm distance to plate center. Vertical lines are probable errors.

instead, the similar 5-inch telescope was used; and it was assumed that the trail corrections were the same for a given length measured in millimeters on the plate. Fourteen plates were taken with the 5-inch, each bearing three exposures which were, respectively, trailed, non-trailed, and trailed. The separate exposure times were either 10 or 20 minutes. The trail lengths used were 0.22, 0.043, and 0.086 mm, made by using a Zeiss micrometer on the guiding telescope. In the reductions it was assumed that the corrections depended on trail length and photographic magnitude only and that they were linear with trail length. The average trail length of the asteroids on the Survey plates was 0.054 mm. For this value the trail effect was found to be a brightening of the image by 0.09 mag., except within 2 mag. from the plate limit, where the effect was negligible (less than 0.04 mag.). This correction was applied to the measured asteroid magnitudes except when the asteroid trail was appreciably above the 0.054-mm value, in which case it was increased proportionally. Correction 4 was applied to the system of the Mount Wilson catalogue of Selected Areas as found by Baum (1956), namely, +0.1(SA-14.0) for SA > 14.0 mag.

Both the special sequences and the determinations of field and trail corrections were made on 103a-O plates. Stars of exceptional color, such as No. 72 in Selected Area 68, showed on the Survey plates no systematic deviation from the characteristic curve of the sequence based on the system of Stebbins, Whitford, and Johnson (1950). It was concluded that the color correction to our magnitude system is negligible.

The apparent magnitudes, p, are listed under MAG in Table A. They are reduced to

absolute magnitudes, g, in column G, except where no orbital elements were available. The quantity $g = p - 5 \log r\rho - F(a)$, where r and ρ are the heliocentric and geocentric distances in astronomical units, respectively, and F(a) is the phase function. The method used for the reduction of p to g is described in Asteroid Paper VI (Gehrels 1957), and the phase correction applied was taken from Table 4 of that paper. The reductions to g were mostly made by Mr. W. D. Caldwell, to whom we are indebted for his careful work.

The best 2 per cent of the magnitudes in Table A are marked with A. These are objects which occurred on the same plate as, and within about 2° from, a Selected Area; further, the difference between the determinations on the two plates was less than 0.20 mag. and p was at least 0.5 mag. above the plate limit. The worse 21 per cent of the magnitudes are marked with C; for these the difference between the two determinations was greater than 0.5 mag., or the quality of the night was unsatisfactory, or else the object was close to the edge of the plates. If the reduction to "zero phase" by means of Table 4 of Paper VI was uncertain (phase angles ≤ 1 .0) the class of p was lowered from p to p to p to p to p magnitude determinations were made on plates taken on the poorest nights used in the Survey or for the brightest dozen or so asteroids; photoelectric observations are available for the latter.

The precision of the magnitudes was derived from internal evidence. In the first place, a determination could be made from the two Survey plates taken at each epoch. Other determinations could be made from asteroids in overlapping areas or those observed at consecutive oppositions. The latter, however, involved differences in aspect and also uncertainties due to the reduction to absolute magnitude; they were not used here but are considered in Section VIII. The determination from overlapping areas is somewhat uncertain and not entirely typical, since the images are of lesser quality; nevertheless, the result so obtained was consistent with the determination from the plate pairs. The plate pairs give an internal probable error of ± 0.09 mag. for the weighted mean; most of this amount is attributed to variable atmospheric effects. However, part of it must be due to rotational variation of the asteroids and to accidental errors in the zero points of the different Selected Areas, since often two different Areas were used during one night and an effort was made to make the two reductions for the image pair as nearly independent as possible. The probable error of the zero point of a Selected Area, after removal of the systematic effect mentioned before, is estimated to be ± 0.06 mag. (probable error) on the basis of Baum's measures for nine Selected Areas. Other sources of uncertainty, which must be added to the probable error of ± 0.09 mag. mentioned above, are differential extinction and the uncertainties in the field and trail corrections. It is estimated that these have contributed the following respective probable errors: ± 0.04 , 0.03, and 0.02 mag. Not all image pairs were reduced with separate Selected Areas; we therefore add the full uncertainty of the Selected Area zero points in computing the total uncertainty of the average Survey magnitudes; this value is found to be 0.12 ± 0.02 mag. In the comparison of the Survey magnitudes with other data it seems advisable to increase this value further for two reasons: (1) There have been some gradual changes in the quality of the Survey plates, and these changes somewhat affect the visibility of the asteroid trailing; the empirical trail corrections were found from a more homogeneous set of plates. (2) The pairs of Survey plates were taken at 1-hour intervals, and the rotational magnitude variations (cf. Papers I-V and VII) will have been less than between random epochs on the light-curves; it is the latter quantity which is needed in comparing Survey magnitudes with other values. The two effects are roughly allowed for by increasing the average probable error of the Survey magnitudes from ± 0.12 to ±0.15 mag. For the three categories in Table A the probable errors are then estimated to be ± 0.10 , ± 0.14 , and ± 0.20 mag. These values still do not include variations due to the asteroid aspect or uncertainties due to the reductions to absolute magnitude. The measurements and the discussion of the magnitudes were carried out by Mr. Gehrels.

VIII. THE COMBINATION OF MAGNITUDES

One of the main objectives of the Survey was to derive a table of accurate asteroid magnitudes on the International Photographic Scale. Since not all numbered asteroids were observed in the Survey, it was necessary to consider other available magnitude sources and reduce them to the International scale. Furthermore, since the asteroid magnitudes are variable, not merely because of rotation and phase variations, but because of aspect, data obtained during a single opposition are not necessarily representative. The reductions involved three steps:

- a) From intercomparisons of Survey magnitudes at the same opposition, a probable error of ± 0.15 mag. is derived for a Survey magnitude of unit weight, while intercomparisons at different oppositions give ± 0.22 mag. These values are obtained in the discussion accompanying Table 8. The increased probable error determined from different oppositions, amounting to an additional variation of ± 0.16 (p.e.), is found to be due to aspect variation.
- b) In Paper VI, comparisons were made between Survey and Ephemeris magnitudes, from which the probable error of the latter was found to be ± 0.29 mag. This value includes variations due to aspect. The systematic error is about 2.0 mag. at 15th mag.
- c) Also in Paper V1, the systematic and accidental errors of magnitude series made at several other observatories were derived with the aid of Survey magnitudes. Intercomparisons were made from observations made within one month, so that aspect variations were not appreciable.

For 22 of the brightest asteroids (Nos. 1, 2, 3, 4, 6, 7, 8, 9, 10, 11, 14, 15, 16, 17, 20, 25, 39, 44, 321, 354, 511, 532), photoelectric light-curves and colors had been derived previously, mostly at the McDonald Observatory, and accurate magnitudes, averaged over a rotational cycle, were available for them (Papers I–V and VII). Usually more than one opposition was covered. These values were reduced to photographic magnitudes with the relation

$$p = V - 0.176 + 1.090 (B - V), \tag{1}$$

and to zero phase angle by means of the phase coefficients listed in the papers referred to. The resulting p_0 - and g-values are included in Table 7, weight 2.0 having been assigned to the mean magnitude derived at one opposition (see below).

For all other asteroids, data from the various sources referred to were combined with the proper weights, including photoelectric data for 38 additional asteroids, given in Tables 9 and 10. Unit weight corresponds to a probable error of \pm 0.22 mag.; this definition is appropriate, as is seen from the discussion under a above. From the discussion referred to under b above, the weight of the reduced Ephemeris magnitudes is then found to be 0.6. This comparatively large weight results from the systematic effects due to aspect, present in even the best measures made during a single opposition. It was therefore decided to include the reduced Ephemeris magnitudes for all asteroids not having photoelectric light-curves, which is essentially those fainter than 10 mag. However, the aspect variation of ± 0.16 mag. found above was derived from magnitudes at consecutive oppositions. These are, on the average, 5/4 years apart and will, on the average, show the largest aspect effects. If the aspect variation is assumed to be sinusoidal, the average aspect variation will be $1/\sqrt{2}$ times ± 0.16 or ± 0.11 mag. (p.e.). Unit weight will now correspond to a probable error of ± 0.19 mag. instead of ± 0.22 mag., while the weight of the Ephemeris magnitudes will be 0.5 and of the italics magnitudes 0.4, slightly less than adopted in Table 7. Additional magnitude determinations for many asteroids will be needed before an accurate value of the aspect variation can be found. The systematic correction applied to the Ephemeris magnitudes is

$$p_0 = 1.136 \, m_0 - 0.16 \, . \tag{2}$$

TABLE 7

MEAN PHOTOGRAPHIC MAGNITUDES OF THE ASTEROIDS AND THEIR WEIGHTS

No.	Po	g	wt	No.	P _o	g	wt
1 2 3 4 5	7•45 8•45 9•60 6•74 10•98	4.00 4.99 6.36 4.20 7.94	4.0 4.0 4.0 4.0	56 57 58 59 60	12.58 12.42 13.03 11.97 12.57	9•49 8•25 9•72 8•63 9•95	3.4 2.4 3.1 2.6 2.1
6	9•29	6.60	3•5	61	12.45	8 • 59	3.2
7	9•35	6.76	6•0	62	13.77	9 • 66	4.0
8	9•56	7.45	4•0	63	10.96	8 • 34	2.3
9	9•77	7.17	5•5	64	12.03	8 • 75	1.8
10	10•61	6.45	4•0	65	12.44	7 • 87	2.5
11	10.43	7.67	4.3	66	13.64	10.44	3.9
12	11.17	8.71	2.2	67	12.40	9.72	1.2
13	10.94	7.90	1.0	68	11.67	8.19	2.2
14	10.37	7.31	2.0	69	12.04	8.18	1.8
15	9.25	6.08	4.0	70	12.18	9.05	1.8
16	10.48	6.74	4.1	71	11.80	8.38	3.5
17	11.39	8.59	4.0	72	12.52	10.23	2.5
18	10.11	7.74	4.1	73	13.43	10.19	1.7
19	11.23	8.50	1.9	74	13.49	10.02	3.6
20	10.03	7.38	2.0	75	13.15	2.90	2.2
21	11.30	8.58	2.3	76	13.43	8.88	1.9
22	11.08	7.35	2.1	77	12.75	9.51	2.6
23	11.40	8.24	3.9	78	12.14	9.00	2.6
24	12.20	8.08	2.5	79	11.92	9.18	2.1
2 5	11.60	8.98	4.0	80	11.56	9.19	2.6
26	11.80	8.58	2.6	81	13.26	9.65	2.2
27	10.96	8.46	2.2	82	12.79	9.35	1.9
28	11.56	8.09	2.2	83	12.43	9.72	2.1
29	10.15	7.16	4.4	84	12.73	10.19	2.3
30	11.23	8.68	4.8	85	12.23	9.02	1.6
31	11.91	7.73	1.8	86	13.84	9•75	2.5
32	11.72	8.65	2.4	87	13.03	8•34	2.8
33	13.35	9.70	1.6	88	11.57	8•12	2.6
34	12.83	9.55	3.4	89	11.35	8•37	1.3
35	13.52	9.64	1.9	90	13.40	9•25	3.1
36	13.33	9•92	1.0	91	12.71	9.64	2.1
37	11.55	8•36	5.1	92	12.19	7.95	2.2
38	12.83	9•1վ	2.1	93	12.12	8.70	2.9
39	10.75	7•30	8.6	94	12.88	8.72	1.8
40	10.64	8•35	2.4	95	12.80	8.79	1.8
41	11.76	8.33	2.1	96	12.97	8.98	2.2
42	11.43	8.70	2.2	97	11.81	8.57	2.1
43	11.20	9.08	2.1	98	13.81	10.53	1.6
44	10.59	7.91	3.9	99	14.81	11.58	1.0:
45	11.79	8.44	2.5	100	13.14	9.08	2.2
46	12.21	9.28	2.4	101	12.30	9.24	1.8
47	12.75	9.08	2.4	102	13.63	10.40	3.2
48	12.16	8.08	2.4	103	11.76	8.45	2.3
49	12.55	8.51	3.6	104	13.62	9.48	3.2
50	13.50	10.30	3.0	105	12.23	9.66	1.4
51	11.12	8•57	4.1	106	12.91	8.73	2.6
52	11.50	7•14	4.2	107	12.79	8.10	2.0
53	12.81	9•67	2.6	108	13.47	9.21	2.2
54	12.05	8•72	2.2	109	13.33	10.03	2.0
55	12.38	8•95	3.3	110	11.74	8.36	1.7

TABLE 7-Continued

No.	р _о	g	wt	No.	р _о	g	wt
111	12.04	8.96	3.6	171	13.78	9.66	1.7
112	13.33	10.62	2.6	172	12.11	9.53	2.6
113	12.15	9.58	2.6	173	12.29	8.89	1.9
114	12.67	9.41	3.6	174	13.10	9.47	3.6
115	11.60	9.02	1.5	175	13.78	9.50	2.8
116	12.13	8.68	2.6	176	13.61	9•42	1.3
117	12.94	9.07	2.4	177	13.96	10•50	2.2
118	12.57	9.85	2.6	178	13.33	10•55	2.6
119	12.35	9.30	2.3	179	13.79	9•95	2.6:
120	12.93	8.83	3.2	180	14.80	11•44	2.0
121	12.87	8.23	1.7	181	13.05	8.95	2.0
122	13.08	8.81	2.0	182	12.48	9.13	2.6:
123	13.40	10.10	2.7	183	14.37	10.87	2.5
124	12.18	9.02	2.6	184	13.76	9.56	2.1
125	12.91	9.51	2.1	185	11.82	8.43	3.1
126	13.04	10.31	3.3	186	12.82	10.28	2.4
127	12.90	9.48	4.0	187	12.80	9.42	1.8
128	12.06	8.65	4.6	188	13.80	10.36	2.0
129	11.35	7.70	2.2	189	13.09	10.34	2.3
130	11.96	7.85	1.8	190	13.81	8.48	2.6
131	13.65	10.94	2.6	191	13.59	9.90	3.2
132	13.33	10.21	1.8	192	10.94	8.30	1.9
133	12.97	8.97	2.6	193	14.26	11.16	2.1
134	12.39	9.37	2.2	194	11.85	8.72	4.0
135	11.68	8.98	2.3	195	13.83	10.16	2.3
136	12.94	10.60	1.9	196	11.67	7.58	2.1
137	13.00	8.89	1.1	197	14.24	10.85	2.0
138	13.35	10.60	2.0	198	12.33	9.56	1.8
139	12.59	9.11	1.5	199	14.03	9.86	2.6
140	12.74	9.36	3.8	200	12.73	9.34	1.6
142 143 144 145	12.79 14.02 13.89 12.10 12.62	9.55 11.34 10.45 8.88 9.37	2.0 2.2 2.4 2.6 1.8	201 202 203 201 205	12.64 12.87 13.50 13.47 13.83	9.38 8.85 10.11 10.23 10.36	2.6: 3.4 2.8 2.3 2.1
146	12.49	9.14	3.2	206	13.37	9.98	2.7
147	13.97	9.84	2.6	207	13.24	10.90	2.4
148	12.12	8.66	2.2:	208	14.05	10.36	2.1
149	14.22	12.18	2.0	209	12.98	8.83	2.2
150	13.01	9.15	2.0	210	13.82	10.46	2.2
151	13.48	10.40	2.4	211	12.87	8.91	2.1:
152	13.70	9.58	2.0	212	13.47	9.37	3.5
153	14.16	8.80	2.3	213	13.34	9.92	2.2
154	12.57	8.37	2.3	214	13.46	10.34	2.5
155	15.18	11.45	0.6	215	14.22	10.78	3.1
156	12.86	9.49	2.4	216	11.49	7.99	1.4
157	15.39	12.34	1.2	217	14.61	10.97	2.6
158	14.20	10.56	3.3	218	13.04	9.80	2.6
159	13.36	9.29	1.6	219	12.89	10.37	3.1
160	13.38	10.01	3.4	220	14.84	12.34	2.6
161	12.74	10.16	2.1	221	12.89	8.97	2.6
162	13.89	9.97	3.2	222	14.47	10.33	1.8
163	12.92	10.37	4.1	223	15.10	11.04	2.6
164	12.82	9.65	1.0	224	13.07	9.88	3.1
165	12.74	8.62	3.1	225	14.08	9.59	2.0
166	13.94	10.66	1.8	226	14.31	10.97	2.8
167	14.27	10.65	2.6	227	14.21	10.05	3.4
168	13.61	9.09	2.6	228	15.99	13.88	2.6
169	13.05	10.52	1.8	229	15.01	10.43	2.2
170	13.61	10.62	1.21	230	11.09	8.51	1.9:

TABLE 7-Continued

No.	p_{o}	g	wt	No.	р _о	g	wt
231	14.22	10.48	2.0	291	14.87	12.70	2.0
232	14.55	11.55	2.6	292	13.89	10.95	1.1
233	12.68	9.45	2.6	293	14.60	10.97	2.0
234	12.92	10.32	3.6	294	15.16	10.99	1.8
235	13.51	9.84	2.0	295	14.96	11.45	1.6
236	12.97	9.46	1.8	296	15.65	13.46	2.8
237	14.15	10.71	2.3	297	14.82	10.61	1.8
238	12.96	9.24	2.6	298	14.83	12.55	1.7
239	15.52	11.69	1.9	299	15.76	13.04	1.8
240	12.93	9.69	2.6:	300	14.69	10.44	1.8
24:1	12.66	8.68	2.5	301	14.48	11.12	2.2
24:2	14.12	10.48	3.6	302	14.97	12.32	2.3:
24:3	14.72	11.09	1.7	303	13.93	9.81	2.6
24:4	15.44	13.40	2.2	304	13.77	11.13	2.2
24:5	13.74	9.68	1.6	305	14.27	10.22	2.6
2146	13.14	9.84	2.1	306	12.66	10.13	2.4
2147	12.49	9.10	2.0	307	14.60	10.88	2.3
2148	14.22	11.42	3.2	308	12.25	8.84	3.0
2149	15.02	12.44	2.2	309	14.63	11.39	2.6
250	12.80	8.66	2.0	310	15.04	11.60	2.5
251	15.27	11.21	2.1	311	14.93	11.23	2.0
252	14.71	10.54	2.0	312	13.65	10.17	1.9
253	14.69	11.49	2.6	313	12.09	9.52	3.6
251 ₄	15.24	13.15	1.8	311,	15.27	11.10	2.0
255	14.76	11.36	2.1	315	15.97	13.75	1.3
256	14.85	10.96	2.6	316	14.89	10.69	2.3
257	14.23	10.13	2.1	317	13.64	11.30	2.1
258	12.56	9.43	2.1	318	14.53	10.28	2.5
259	13.12	8.99	2.6	319	15.80	11.25	1.2
260	15.07	10.43	2.7	320	15.55	11.63	2.4
261	12.97	10.51	2.6	321	14.93	11.26	5.1
262	15.72	12.73	1.0	322	13.81	10.33	2.3
263	15.19	11.51	2.6	323	13.89	11.30	1.3
264	13.43	9.92	2.2:	324	11.30	8.02	2.0
265	15.37	12.69	1.3	325	14.23	10.00	2.0
266	12.99	9.47	2.6	326	12.46	10.04	0.9
267	15.47	12.01	2.0	327	14.73	11.27	3.1
268	13.59	9.53	4.1	328	13.99	9.91	1.7
269	14.35	11.22	2.1	329	13.45	10.64	1.5
270	12.22	10.12	3.0	330	15.18	13.40	0.6
271	14.59	10.69	2.4	331	14.25	10.31	2.0
272	15.15	11.68	1.7	332	13.93	10.47	2.6
273	13.70	11.08	1.8:	333	14.48	10.37	2.3
274	15.14	11.18	2.1	334	13.58	8.33	2.1
275	13.40	9.95	2.6	335	12.82	10.01	3.0
276	13.41	9.31	2.1	336	13.14	10.89	2.6
277	14.72	11.04	2.1	337	12.53	9.95	1.8
278	13.89	10.47	3.1	338	13.28	9.55	2.0
279	15.42	9.68	4.6	339	14.29	10.37	3.6
280	15.58	11.79	1.6:	340	14.45	11.04	1.7
281 282 283 284 285	14.88 14.67 13.39 14.09 16.17	12.81 12.19 9.43 11.56 12.14	1.8 1.6 2.4 2.5 1.6	34:1 34:2 34:3 34:5	14.68 14.17 15.08 12.41 12.57	12.58 11.11 12.12 9.32 10.13	4.3 2.6 0.8 2.6 2.1
286	14.49	10.26	3•3	34.6	12.57	9.06	2.0
287	11.90	9.38	1•8	34.7	13.25	10.13	2.1
288	14.28	10.86	2•4	34.8	14.49	10.65	2.2
289	14.38	10.72	2•0	34.9	10.95	7.20	4.7
290	15.63	13.15	1•2	350	14.01	9.92	2.2

TABLE 7-Continued

No.	р _о	g	wt	No.	р _о	g	wt
351	13.67	10.23	2.3	411	13.55	9•78	2.6
352	13.57	11.48	2.6	412	13.65	10•21	2.6
353	15.66	12.28	2.2	413	13.70	10•64	0.6
354	10.98	7.47	3.8	414	15.17	10•145	1.8
355	14.60	11.64	2.0	415	13.77	10•28	3.6
356	12.47	9.04	2.0	416	12.74	9.26	2.5
357	13.56	9.42	1.7	417	14.14	10.63	2.5
358	13.86	10.20	2.6	418	13.78	10.70	2.6
359	13.82	10.45	2.6	419	12.44	9.36	3.0
360	13.38	9.49	2.3	420	13.87	9.28	2.5
361	14.81	9.50	1.5	421	15.89	12.93	1.0
362	12.77	9.72	1.8	422	14.18	12.00	2.6
363	13.28	9.87	2.6	423	12.43	8.42	2.6
364	13.19	11.02	3.4	424	14.15	10.69	2.3
365	13.83	10.32	2.2	425	14.58	10.90	2.6
366	13.87	9.73	2.2	426	13.18	9.49	1.8
367	14.13	11.96	2.6	427	14.33	10.49	2.6
368	15.04	11.03	3.4	428	15.30	12.91	1.4
369	12.72	9.52	1.8	429	13.93	10.82	2.2
370	14.09	11.65	3.1	430	15.26	11.66	2.4
371	13.23	9.86	3.1	431	14.17	10.05	2.3
372	12.20	8.03	1.2:	432	12.65	10.09	1.6
373	14.35	10.25	1.8	433	11.43	12.31	2.1
374	13.52	10.05	3.9	434	13.24	11.92	0.6
375	12.40	8.28	3.6	435	13.90	11.15	2.5
376	12.80	10.45	2.8	436	15.21	10.97	2.1
377	13.01	9.72	4.1	437	14.12	11.52	3.6
378	14.35	10.89	2.2	438	13.52	10.53	1.6
379	14.26	10.11	2.5	439	14.72	10.60	1.5
380	13.74	10.48	4.1	140	14.94	12.80	1.7
381 382 383 384 385	13.71 13.74 14.87 13.94 12.32	9.48 9.62 10.75 10.73 .8.72	2.3 1.2 1.3 2.0 2.1	1442 1443 1443 1444	13.15 13.56 13.63 12.76 14.28	9.62 11.06 11.48 9.31 10.08	2.4: 2.1 2.4 2.9 2.4
386 387 388 389 390	12.00 12.29 13.13 12.29 14.59	8.30 8.90 9.23 9.18 11.38	0.8 1.9 2.6 1.3 2.0	րիզ Իր 17 17 17 17 17 17	13.39 14.15 15.09 13.63 15.15	9.90 10.29 10.95 10.64 11.24	2.4 3.1 3.1 2.5: 1.6
391 392 393 394 395	14.67 14.20 12.59 14.42 14.73	12.24 10.52 9.12 10.98 11.25	1.7 1.6 2.1 2.6 2.1	451 453 454 455	12.16 16.99 13.92 12.89 13.18	8.16 13.35 11.86 9.73 9.96	1.9 0.6 3.1 1.4 2.6
396	14.56	11.16	2.4	456	14.36	10.87	3.5
397	13.34	10.17	2.2	457	16.99	12.94	0.6
398	15.25	11.86	0.8	458	14.36	10.49	2.6
399	14.45	10.47	2.0	459	14.95	11.81	2.5
400	15.34	11.22	1.9	460	15.26	11.91	3.6
40 1 403 404 405	14.64 12.89 13.72 12.88 12.56	10.18 9.89 10.19 9.81 9.50	2.1 2.1 2.5 1.1 2.4	461 462 463 465	15.62 14.48 15.47 13.80 14.92	11.52 10.83 12.84 10.28 10.87	2.0 2.3 1.4 2.1 1.8
406	15.06	11.32	2.6	466	13.60	9.09	2.3
407	13.35	10.20	2.6	467	15.80	12.02	3.6
408	14.90	10.74	2.2	468	14.66	10.53	2.2
409	11.59	8.55	2.3	469	14.05	9.89	2.6
410	12.98	9.62	1.8	470	13.94	11.30	3.0

TABLE 7-Continued

No.	р _о	g	wt	No.	P _o	g	wt
471	11.44	7.76	2.2	531	15.74	12.22	0.6
472	13.46	10.49	2.6	532	11.33	7.88	3.4:
473	14.95	11.10	0.6	533	14.92	11.06	1.1
474	14.62	11.86	2.0	534	14.54	10.87	2.6
475	15.40	12.32	1.2	535	13.46	10.43	2.3
476	12.86	9.66	2.l4	536	14.02	9•31	1.5:
477	14.04	11.37	3.6	537	14.08	10•07	2.1
478	12.59	8.67	2.0	538	14.66	10•48	3.1
479	14.26	10.91	2.l	539	14.41	11•02	2.6
480	13.05	9.86	2.l	540	14.07	11•91	4.1
481	13.13	9•74	3.6	541	14.63	11.09	2.3
482	13.83	9•94	2.3	542	13.91	10.20	2.6
483	14.17	9•57	1.7	543	14.51	10.51	3.5
484	14.61	11•37	2.3	544	14.30	11.22	2.6
485	12.90	9•49	1.9	545	13.72	9.53	2.2
486	14.69	12.18	3.6	546	13.87	10.78	2.3
487	12.77	9.52	3.4	547	14.33	10.88	2.6
488	13.10	8.94	2.6	548	14.80	12.47	2.3
489	13.79	9.63	2.5	549	15.18	11.90	0.6
490	13.64	9.44	4.4	550	13.36	10.29	2.9
49 1	14.13	9.89	1.5	551	14.27	10.14	2.6
492	14.94	10.85	3.5:	552	14.33	10.18	1.6:
493	16.00	11.90	1.5	553	15.56	13.37	2.0
494	13.85	9.99	3.2	554	12.06	9.50	2.5
495	14.29	11.45	1.8	555	15.81	11.63	2.0
496	15.03	12.92	3.8	556	13.24	10.46	2.6:
497	14.73	11.12	2.1	557	15.65	12.92	2.2
498	13.09	9.88	4.6	558	13.68	9.96	3.5
499	15.47	10.12	2.6	559	13.91	10.58	1.3
500	13.54	10.42	2.0	560	15.12	11.70	1.8
501	14.39	10.22	2.2	561	16.06	11.87	0.8
502	14.81	12.22	2.3:	562	14.74	10.82	2.6
503	13.46	10.10	2.6	563	12.79	9.45	2.5
504	14.36	11.01	2.1	564	15.36	11.95	0.8
505	13.35	10.08	1.3	565	14.75	12.01	1.4
506	13.84	9.87	2.0	566	13.59	9.05	3.0
507	14.54	10.39	3.0	567	14.54	10.41	2.7
508	13.59	9.42	2.4	568	14.00	10.33	2.1
509	13.35	9.35	2.4	569	14.04	10.82	3.3
510	14.06	10.94	6.4	570	14.54	9.94	2.6
511	11.26	7.02	6.5:	571	15.50	12.84	2.6
512	14.13	12.05	2.5	572	14.46	11.83	4.5
513	14.33	10.42	3.6	573	14.54	10.62	2.2
514	14.10	10.13	1.7	574	16.08	13.83	0.6
515	16.65	12.63	0.5	575	15.26	12.26	2.6
516	12.68	9.42	1.6	576	14.51	10.63	2.3
517	14.56	10.42	3.0	577	14.89	10.81	2.6
518	15.14	12.19	2.5:	578	13.85	10.43	1.7
519	13.68	10.18	2.1	579	13.17	9.25	3.1
520	15.85	11.94	2.6	580	15.22	10.95	2.6
521	13.26	9.86	2.1	581	15.14	10.88	2.5
522	14.81	9.90	3.6	582	13.49	10.38	3.1:
523	14.44	10.61	2.1	583	14.46	10.25	2.6
524	13.85	10.68	2.9	584	12.52	9.95	3.1
525	15.52	11.06	0.6	585	14.09	11.38	2.1
526	14.83	10.70	2.1	586	14.35	10.35	2.2
527	14.57	11.21	3.1	587	15.95	13.48	0.8
528	14.32	9.77	2.3	588	16.04	9.33	8.1
529	14.83	10.90	2.1	589	14.11	9.98	2.2
530	14.17	9.91	2.4	590	14.95	11.07	2.0

TABLE 7-Continued

No.	р _о	g	wt	No.	p _o	g	wt
591	15.09	11.83	2.2	651	15.09	11.15	2.6
592	14.44	10.52	2.3	652	15.29	12.29	2.0
593	13.71	10.40	3.2	653	14.46	10.55	2.3
594	16.63	13.47	1.0	654	12.15	9.78	2.0
595	13.41	9.16	2.0	655	14.37	10.50	1.8
596	13.43	9.67	1.5	656	15.37	11.19	3.0
5 97	13.94	10.69	1.6	657	15.04	11.93	1.8
598	13.25	9.80	2.6	658	15.26	11.64	4.6
599	13.44	9.98	2.0:	659	16.31	9.59	6.6
600	14.60	11.37	2.6	660	12.82	9.87	2.4:
601	14.48	10.35	2.lı	661	14.46	10.54	2.4
602	13.63	9.59	1.8	662	14.69	11.70	3.0
603	16.45	13.46	1.3	663	14.51	10.51	0.8
601	14.40	10.23	1.9	661,	15.69	11.51	2.0
605	14.41	10.52	2.3	665	13.85	9.70	2.0
606	14.50	11.43	2.0	666	15.03	11.95	2.6
607	14.38	10.77	4.1	667	14.80	10.57	2.2
608	15.72	11.78	2.6	668	16.78	13.28	1.6
609	14.95	10.90	3.1	669	15.24	11.33	3.6
610	17.24	13.21	0.8	670	14.68	11.16	2.0
611	14.35	10.50	2.1	671	15.10	11.04	2.0
612	16.43	12.29	0.6	672	15.19	12.19	1.4
613	14.64	10.90	2.3	673	14.76	11.22	2.6
614	15.28	11.99	2.6	674	12.14	8.39	1.3
615	14.18	11.02	2.8	675	12.53	9.08	3.6
616	14.21	11.22	0.8	676	14.47	10.48	2.6
617	15.79	9.09	8.2:	677	14.62	10.81	2.0
618	13.78	9.55	2.8	678	13.56	10.52	2.2:
619	14.04	11.12	2.6	679	12.20	9.13	1.0
620	15.11	12.40	1.8	680	14.93	10.81	2.6
621	15.66	11.54	2.4	681	15.94	11.86	1.6
622	14.36	11.69	1.8	682	16.65	13.47	0.6
623	14.21	11.44	1.5	683	13.80	9.71	2.3
624	15.17	8.55	8.8	684	14.76	12.05	3.3
625	14.18	10.98	2.3	685	15.07	12.86	2.2
626	13.03	9.99	1.2	68 6	14.50	11.43	2.1:
627	14.70	11.00	2.6	687	16.30	12.95	1.6
628	13.42	10.36	2.1	688	14.94	11.63	2.6
629	14.91	10.80	2.6	689	15.70	13.28	2.6
630	15.57	12.42	1.5	690	12.96	8.79	1.6
631	13.55	10.06	2.1	691	14.34	10.43	2.6
632	16.41	13.18	1.8	692	14.83	10.34	0.8
633	14.93	11.01	2.6	693	14.16	10.37	3.2
634	15.01	11.04	1.3	694	13.32	10.07	3.6
635	14.22	10.08	2.5	695	12.68	9.72	2.3
636	14.33	10.61	2.6	696	14.52	10.28	4.1
637	16.04	11.87	2.0	697	13.96	10.29	1.0
638	14.54	11.17	2.6:	698	15.43	11.77	1.0
639	13.21	9.30	3.1	699	16.16	13.03	0.8
640	14.48	10.30	1.7	700	14.62	12.43	1.7
९७२ ९५३ ९५३ ९५२	16.28 14.82 14.86 14.78 15.22	14.12 10.64 10.40 11.69 11.01	0.8 1.8 2.6: 2.6 2.0	701 702 703 7014 705	14.41 12.84 15.63 11.50 13.46	10.50 8.62 13.60 7.50 9.72	3.2 2.2: 0.6 1.1 2.0
646 647 648 649 650	16.64 15.31 14.89 17.08 16.18	14.20 12.58 10.71 14.10 13.41	1.0 1.8 2.6 1.0	706 70 7 708 709 710	15.41 15.41 14.99 13.51 16.28	12.00 13.36 11.74 9.77 12.15	1.3 1.7 2.0 2.0 3.2

TABLE 7-Continued

No.	$\mathbf{p}_{\mathbf{o}}$	g	wt	No.	р _о	g	wt
711	14.75	12.54	1.7	771	15.06	11.85	0.6
712	12.51	9.47	2.3	772	13.54	9.65	1.0
713	14.44	9.85	2.2	773	14.16	10.53	2.3
714	12.85	9.90	1.6	774	13.86	9.89	2.1
715	14.53	11.08	2.6	775	15.19	11.28	2.6
716	15.29	11.76	2.6	776	12.63	8.86	1.7
717	16.0h	11.91	3.6	777	15.53	11.28	2.2
718	14.71	10.73	2.6	778	15.68	11.50	1.8
719	19.83	16.77	0.6	779	12.89	9.65	0.8
720	14.5h	10.86	2.1	780	14.23	10.14	2.6
721	15.23	10.42	2.2	781	14.75	10.45	2.0
722	15.08	13.05	2.0	782	14.59	12.54	2.3
723	15.32	11.44	1.3	783	14.64	12.16	2.3
724	17.45	14.74	0.6	784	14.36	10.30	3.4
725	15.30	12.27	1.5	785	13.50	10.46	1.4:
726	15.14	12.12	2.1	786	14.21	10.02	3.1
727	14.20	11.18	3.0	787	14.28	11.32	1.3
728	16.08	13.78	0.6	788	13.56	9.46	2.1:
729	14.38	10.95	0.8	789	15.55	12.27	2.4
730	16.89	14.66	1.4	790	13.81	9.27	1.8
731	14.37	10.51	2.6	791	14.73	10.62	2.1
732	14.65	11.88	1.1	792	14.21	11.07	1.3
733	14.64	10.09	0.8	793	14.26	10.76	1.7
734	15.07	10.91	3.1	794	16.42	12.29	2.4
735	14.23	10.86	2.0	795	14.10	10.69	2.7
736	14.19	12.08	2.0	796	13.46	10.29	2.2
737	12.75	9.67	2.3	797	14.50	11.55	2.6
738	14.95	11.00	2.2	798	14.50	10.58	2.6
739	13.38	10.00	1.9	799	14.39	11.42	3.6
740	14.28	10.30	2.2	800	14.70	12.61	2.4
741	14.60	11.25	2.0	801	15.51	12.40	2.0
742	14.38	10.46	2.1;	802	15.42	13.32	0.8
743	14.70	11.20	3.1;	803	14.79	10.54	1.8
744	15.35	11.14	2.1;	801	12.76	9.17	2.2
745	15.35	11.05	0.8	805	14.78	10.50	2.4
746	14.60	10.51	2.6	806	15.32	11.09	1.0
747	12.66	8.78	1.8	807	15.65	11.72	3.1
748	15.13	9.82	3.4	808	14.36	10.96	1.8
749	15.08	12.85	1.0	809	15.41	13.08	2.6
750	15.71	12.98	2.0	810	16.20	14.15	0.6
751	12.95	9.96	1.6	811	15.35	11.65	2.6
752	14.19	11.41	2.0	812	15.63	12.41	0.6
753	14.26	11.81	2.6	813	15.28	13.11	2.0:
754	14.30	10.43	1.4	814	14.05	9.85	2.4
755	14.77	10.59	1.9	815	15.15	11.93	3.0
756	15.46	11.16	0.8	816	15.20	11.30	2.6
757	14.05	11.48	1.8	817	15.00	11.93	2.0
758	13.48	9.23	2.4	818	14.52	10.32	2.3
759	15.10	11.96	2.0	819	15.18	13.08	2.6
760	13.62	9.47	2.9	820	15.26	11.14	1.8
761	15.49	11.85	2.0	821	15.69	12.22	1.0
762	13.34	9.18	2.0	822	15.06	12.80	0.6
763	15.95	13.73	1.6	823	14.85	12.68	2.6
764	14.86	10.66	2.0	824	14.87	11.37	3.1
765	16.99	14.01	0.6	825	15.21	13.03	4.6
766	14.72	10.79	2.6	826	15.72	12.38	1.8:
767	15.35	11.26	1.7	827	16.20	13.89	1.6
768	15.38	11.26	1.0	828	15.36	11.14	2.6
769	14.24	10.02	2.0	829	14.94	11.89	2.6
770	14.26	12.09	2.0	8 30	14.74	10.50	2.0

TABLE 7-Continued

No.	P _o	g	wt	No.	p _o	g	wt
831	15.63	13.48	0.6	891	114.95	11.32	2.1
832	15.67	12.04	2.6	892	114.81	10.51	2.1
833	16.14	12.23	1.2	893	114.72	10.74	3.4
834	14.56	10.39	3.1	894	114.99	10.89	2.5
835	16.13	11.89	2.3	895	114.08	9.83	2.0
836	16.35	14.27	0.8	896	15.32	12.99	2.0
837	15.18	12.81	0.6	897	15.13	12.16	1.0
838	14.90	11.20	3.0	898	16.76	13.39	0.8
839	14.79	11.66	2.0	899	15.10	11.38	1.7
840	14.61	10.50	0.6	900	15.96	13.15	1.2
841 843 844 845	15.60 15.43 16.514 14.86 14.89	13.34 11.17 11.22 10.64 11.11	1.6: 0.8 0.6 3.1 1.8	901 902 903 904 905	15.34 16.31 15.06 15.22 14.39	12,59 13,56 10,75 11,34 12,24	2.2 0.6 0.6 1.3 1.8
846	15.54	11.42	2.4	906	14.32	10.62	2.7:
847	14.81	11.33	2.2	90 7	14.15	10.64	2.9
848	15.92	11.84	2.0	908	14.83	12.02	3.0
849	13.05	8.90	1.5	909	14.31	9.54	2.0
850	14.51	10.62	2.6	910	15.02	11.25	3.0
851	15.04	12.86	2.6	911	15.44	8.81	2.7
852	13.85	11.31	1.8	912	13.40	9.27	2.4
85 3	14.93	12.53	2.6	913	15.79	13.69	1.0
854	15.90	13.35	4.3	914	13.14	10.39	1.6
855	15.39	12.68	2.4	915	15.30	13.12	2.3
856	14.73	12.02	2.0	916	15.11	12.57	1.3
857	14.68	12.60	1.2	917	15.16	12.57	3.0
858	15.01	11.48	2.2	918	15.53	11.89	2.1
859	15.11	10.86	3.6:	919	15.76	12.30	2.4
860	14.34	10.83	1.0	920	15.07	11.93	1.7
861.	14.95	10.81	2.4	921	15.35	11.15	2.3
862	14.74	11.23	1.4	922	16.30	13.01	2.6
863	14.45	10.22	1.8	923	15.77	12.65	1.0
864	15.86	13.09	0.6	921	14.15	10.37	3.6
865	15.72	13.05	2.9	925	11.95	8.64	2.4
866 867 868 869 870	14.39 15.70 14.37 16.34 15.15	10.28 11.71 11.05 13.05 12.72	2.2 0.8 3.6 1.0 0.8	926 927 928 929 930	15.60 12.15 14.90 15.54 15.08	11.75 7.90 10.75 13.32 12.37	1.0 1.6 1.8 1.6
871	15.95	13.78	0.8	931	14.62	10.45	2.2
872	14.55	11.18	2.1	932	13.36	10.68	2.3
873	15.44	12.28	1.8	933	16.02	13.47	2.2
874	15.03	10.87	1.6	934	15.51	12.10	0.8:
875	15.72	12.73	1.6	935	16.43	14.27	2.0
876	15.77	11.86	2.0	936	15.23	11.09	4.1
877	14.59	11.75	2.3	937	15.27	13.07	2.6
878	19.04	16.50	0.6	938	16.47	12.29	0.8
879	15.63	12.67	0.6	939	15.52	13.29	2.0
880	16.85	12.96	0.8	940	14.96	10.40	2.0
881	16.65	13.53	0.6	941	16.20	12.72	0.6
882	15.66	11.55	2.1	942	15.63	11.44	0.6
883	15.78	13.57	1.5	943	14.97	10.88	2.5
884	16.48	9.77	4.3:	944	19.16	11.95	0.9
885	15.80	11.75	2.2	945	14.50	11.32	2.5
886	14.32	10.17	1.8:	946	15.40	11.30	2.0
887	19.27	16.35	0.6	947	14.27	10.85	1.8
888	14.18	10.85	2.2	948	16.22	12.27	2.6
889	14.93	12.19	2.2	949	14.76	10.88	2.6
890	15.21	11.27	1.5	950	14.96	12.41	1.0

TABLE 7-Continued

No.	р _о	g	wt	No.	P_{o}	g	wt
951 952 953 954 955	15.14 14.18 14.98 15.82 15.71	13.01 10.30 11.49 11.70 12.63	2.3 2.3 2.6 1.0	1011 1012 1013 1014 1015	17.70 15.95 13.70 16.65 14.47	15.09 13.13 10.43 13.12 10.23	0.8 2.0 1.4 1.1 2.6
956	15.90	13.53	1.8	1016	15.34	13.18	1.8
957	14.71	10.97	2.5	1017	15.34	12.24	0.8
958	16.10	10.79	1.1	1018	14.55	11.59	3.0
959	15.99	11.80	0.8	1019	15.06	13.85	0.6
960	16.28	14.04	2.8	1020	15.55	12.07	0.8
961	15.59	12.30	1.3	1021	13.27	9.89	3.1
962	16.30	12.59	4.0	1022	14.77	11.24	0.8
963	15.96	13.73	0.9	1023	15.24	11.05	2.8
964	15.92	11.94	1.4	1024	15.51	11.87	1.4
965	16.04	11.87	1.7:	1025	15.48	14.04	0.8
966	14.43	11.08	1.4	1026	16.75	14.52	0.8
967	15.42	13.24	2.0	1027	16.65	12.47	0.6
968	14.96	11.32	0.8	1028	14.90	10.33	2.6
969	16.11	13.33	1.3	1029	15.56	11.88	3.1
970	16.50	13.49	1.1	1030	15.64	11.54	1.8
971	14.23	11.05	2.1	1031	14.66	10.69	1.2
972	14.65	10.64	2.0	1032	15.13	11.01	1.5:
973	15.13	10.84	3.6	1033	16.02	12.13	3.6
974	14.61	11.66	2.2	1034	16.21	13.85	1.0
975	14.57	10.99	2.0	1035	15.78	11.64	1.0
976	14.66	10.46	4.1	1036	13.98	10.76	1.8
977	14.77	10.67	1.5	1037	17.12	15.05	0.8
978	15.15	10.90	1.4	1038	16.88	11.59	0.6
979	15.20	11.05	1.3	1039	15.74	10.45	3.3
980	12.59	9.19	1.6	1040	15.62	11.53	1.5
981	16.03	11.96	1.8	1041	14.94	10.93	1.8
982	15.28	11.27	2.6	1042	15.20	10.95	2.3
983	14.71	10.52	0.8	1043	15.03	10.98	4.6
984	14.15	10.63	3.6	1044	15.16	12.12	2.6
985	16.53	14.15	1.9	1045	17.12	14.60	1.3
986	14.82	10.70	3.4	1046	15.40	11.54	2.6
98 7	14.83	10.70	3.4	1047	15.61	13.39	1.8
988	16.52	12.39	0.8	1048	13.98	10.60	2.6
989	16.54	13.32	0.6	1049	15.81	11.75	0.8
990	16.10	12.85	1.6	1050	16.99	13.84	0.6
99 1	15.95	11.83	1.0	1051	15.45	11.19	2.2
992	16.09	12.15	1.6	1052	14.49	12.28	0.6
993	16.65	13.02	0.6	1053	16.56	13.43	1.8
994	14.44	11.50	1.4	1054	15.38	11.63	3.6
995	14.53	11.40	3.1	1055	14.73	12.63	2.4
996	15.68	11.63	1.8	1056	14.98	12.79	1.8
997	16.60	13.35	0.8	1057	15.57	11.88	2.6
998	16.20	12.11	0.6	1058	15.22	13.12	2.8
999	15.48	12.36	2.0	1059	15.45	12.27	2.4:
1000	15.52	11.29	0.6	1060	16.46	14.25	0.8
1001	14.73	10.51	2.6	1061	16.14	12.03	1.6
1002	15.53	12.04	1.2	1062	15.18	11.27	2.6
1003	15.31	11.17	1.1	1063	14.70	12.28	1.6
1004	15.23	10.70	2.6	1064	15.23	12.25	2.0
1005	15.03	10.86	2.2	1065	16.96	14.42	0.8
1006	16.91	12.79	2.6	1066	16.74	14.10	1.2
1007	15.91	12.58	1.6	1067	15.69	12.03	2.4
1008	15.91	11.79	1.6	1068	14.83	11.10	1.0
1009	19.95	16.62	0.6	1069	14.84	10.73	0.6
1010	15.42	11.66	1.8	1070	16.34	12.08	1.6

TABLE 7-Continued

No.	p _o	g	wt	No.	р _о	g	, wt
1071 1072 1073 1074 1075	14.78 15.89 16.66 15.33 15.40	11.68 12.47 11.17 11.48	2.6 1.3 2.0 1.1 1.8	1131 1132 1133 1134 1135	17.50 15.10 15.17 18.56 14.89	15.32 11.82 13.10 15.29 11.65	1.0 1.7 2.5 1.6: 2.0
1076	15.97	13.16	2.8	11.36	15.18	12.16	2.6
1077	16.54	13.93	0.6	11.37	14.69	12.00	2.1
1078	15.09	12.79	3.3	11.38	16.41	12.26	1.2
1079	15.69	12.03	3.0	11.39	15.58	14.25	1.1
1080	16.22	13.55	1.6	11140	14.82	11.37	0.8
1081 1082 1083 1084 1085	16.43 15.67 16.36 15.01 15.03	12.37 11.55 13.91 11.72 10.83	0.6 1.0 1.0 2.6 3.1	11/12 11/13 11/15 11/11	16.81 15.59 16.00 15.97 14.91	14.51 11.39 9.32 10.90 12.22	0.8 2.0 4.5 2.1 2.6
1086	14.81	10.62	2.6	1146	14.92	10.93	0.8
1087	14.81	10.89	2.3	1147	15.75	13.45	2.0
1088	14.79	12.68	2.6	1148	15.25	11.34	1.2
1089	14.96	12.82	2.6	1149	15.12	18.42	3.6
1090	16.43	13.90	0.6	1150	16.99	14.91	0.6
1091	16.65	12.07	0.6	1151	17.45	14.81	0.6
1092	15.34	11.63	2.2	1152	15.06	12.37	1.0
1093	13.77	9.63	0.8	1153	15.42	13.32	2.6
1094	15.91	12.93	1.8	1154	15.91	11.35	2.4
1095	14.53	12.56	1.1	1155	15.78	13.00	2.5
1096	14.39	11.30	2.1	1156	16.08	13.81	0.6
1097	16.26	13.08	1.8	1157	15.38	11.13	2.3
1098	15.16	11.88	1.6	1158	15.18	12.17	1.8
1099	15.83	11.67	1.3	1159	15.54	12.96	2.4
1100	16.06	12.36	2.0	1160	15.78	12.77	1.5
1101	16.44	12.10	0.8	1161	16.88	12.70	0.6
1102	14.86	10.85	3.2	1162	15.53	10.16	1.5
1103	14.79	13.51	1.2	1163	15.83	11.58	3.1
1104	16.53	13.37	1.0:	1164	16.42	14.03	1.3
1105	15.07	11.16	2.0	1165	15.62	11.49	1.8
1106	15.96	12.87	1.0	1166	15.53	12.58	1.5
1107	14.27	10.04	2.6	1167	15.53	10.93	2.6
1108	15.04	12.34	1.6:	1168	16.31	13.32	0.6
1109	15.17	10.94	3.6	1169	16.88	14.45	0.6
1110	15.42	13.27	2.3	1170	15.66	13.22	1.0
1111	15.35	11.47	2.6	1171	14.91	10.75	2.2
1112	14.86	10.93	2.1	1172	16.01	9.32	2.7
1113	14.75	10.65	3.6	1173	16.61	10.01	3.6
11114	14.72	10.67	2.6	1174	16.77	12.84	0.6
1115	14.61	10.54	2.6	1175	15.86	11.59	1.6
1116	14.56	10.81	2.0	1176	15.40	12.10	1.0
1117	15.35	13.11	2.5	1177	14.92	10.44	2.4
1118	15.22	10.98	2.0	1178	16.15	12.89	0.7
1119	15.50	12.38	2.0	1179	18.13	15.00	0.6
1120	15.48	13.33	1.2	1180	15.53	10.15	2.1
1121	15.50	12.52	0.8	1181	15.86	12.63	0.6
1122	15.69	12.58	1.0:	1182	14.85	12.58	2.3
1123	15.05	12.87	2.0	1183	15.68	13.09	2.1
1124	15.81	12.05	1.6	1184	15.64	12.40	2.8
1125	18.50	14.20	0.8	1185	15.40	13.19	2.4
1126	16.15	13.84	1.0	1186	14.72	10.80	1.2
1127	14.74	11.65	0.8	1187	15.98	12.80	2.1:
1128	15.22	11.74	3.1	1188	15.10	13.02	2.4
1129	14.96	11.03	2.6	1189	14.87	11.11	3.1
1130	15.79	13.60	1.5	1190	15.88	13.17	1.1

TABLE 7-Continued

No.	р _о	g	wt	No.	p _o	g	wt
1191	15.40	11.71	0.6	1251	15.12	11.77	2.6
1192	16.12	13.58	1.0	1252	14.91	11.61	1.3
1193	16.08	12.88	0.6	1253	17.40	13.22	0.8
1194	15.16	11.43	1.8	1254	15.68	11.55	1.5
1195	16.77	14.51	0.6	1255	15.76	11.62	2.6:
1196	14.68	11.47	4.0	1256	16.07	10.80	1.8
1197	14.79	11.12	1.8	1257	15.69	12.85	2.2
1198	18.92	16.69	0.6	1258	15.94	11.73	4.1
1199	15.31	11.39	2.0	1259	15.81	11.74	2.3
1200	15.70	11.71	3.4	1260	16.02	12.90	1.0
1201	15.89	12.58	2.6	1261	15.96	11.83	1.0
1202	16.77	11.46	0.6	1262	14.96	11.07	1.0
1203	16.80	13.12	3.6	1263	13.61	10.42	0.8:
1204	16.10	13.82	2.0	1264	14.47	10.84	1.3
1205	18.13	15.19	0.6	1265	14.95	11.03	0.6
1206	15.03	11.38	1.6:	1266	14.83	10.32	2.3
1207	16.77	12.05	0.6	1267	16.18	13.39	2.0
1208	16.34	9.69	2.4	1268	15.24	9.94	2.1
1209	16.02	11.81	0.8	1269	14.92	9.59	2.3
1210	15.17	11.26	1.8	1270	16.25	14.05	2.0
1211	15.91	12.15	2.8	1271	15.85	11.71	1.8
1212	16.23	10.88	1.4	1272	16.99	13.51	0.6
1213	16.34	12.21	1.2	1273	16.76	14.14	2.1
1214	15.33	12.00	2.3:	1274	15.24	13.05	1.8
1215	14.86	11.81	1.9	1275	15.11	11.84	2.6
1216	15.49	13.29	1.6	1276	16.01	11.83	1.0
1217	16.99	14.48	0.6	1277	15.82	12.52	2.2
1218	16.62	14.34	1.6	1278	15.19	12.54	2.0:
1219	15.40	13.26	2.6	1279	16.31	13.75	2.2
1220	16.37	12.47	0.8	1280	15.75	11.18	3.0
1221	20.3	19.06	0.6	1281	15.46	12.46	2.3
1222	16.65	13.16	0.6	1282	15.50	11.39	1.5
1223	15.28	11.63	2.1	1283	16.20	11.92	0.6
1224	15.22	12.83	0.8	1284	14.59	11.40	3.6
1225	15.46	13.27	0.8	1285	15.10	11.23	1.7
1226	16.08	13.03	0.6	1286	15.46	11.53	2.1
1227	15.67	11.45	2.8	1287	15.94	12.03	3.6
1228	16.16	12.71	1.0	1288	16.65	12.98	0.6
1229	17.13	12.91	0.8	1289	15.12	11.49	1.7
1230	17.68	14.64	0.6	1290	16.20	13.65	0.8
1231	15.88	12.64	1.2	1291	15.28	11.37	4.4
1232	15.47	11.28	2.6	1292	15.37	12.40	2.0
1233	15.32	12.32	1.2	1293	17.07	14.89	0.8
1234	16.22	12.31	0.8	1294	15.20	11.92	2.8
1235	16.67	15.47	0.8	1295	16.54	12.04	1.2
1236	15.52	12.81	2.6	1296	15.50	12.82	1.2
1237	15.08	11.96	2.14	1297	16.30	12.38	1.6
1238	16.25	13.01	1.0	1298	15.68	11.56	2.3
1239	16.43	13.20	0.6	1299	16.41	12.89	2.4
1240	14.65	11.01	3.6	1300	15.73	12.25	2.6
1241	14.63	10.41	1.3	1301	15.22	11.78	1.4
1242	14.45	11.06	2.2:	1302	16.08	11.98	0.6
1243	15.29	11.23	3.6:	1303	14.68	10.43	1.5
1244	14.88	12.39	1.3	1304	14.64	10.41	2.9
1245	14.67	10.98	2.3	1305	15.38	11.47	2.1
1246	16.54	13.40	0.6	1306	14.90	10.75	1.6
1247	15.89	11.77	1.8	1307	15.41	13.17	1.1
1248	14.31	10.96	2.4	1308	15.65	11.93	2.1
1249	15.06	12.88	1.6	1309	15.49	11.23	2.1
1250	17.04	14.05	0.8	1310	15.29	12.68	0.6

TABLE 7-Continued

No.	P _o	g	wt	No.	P _o	g	wt
1311	16.45	13.75	1.1	1371	16.77	12.52	0.6
1312	16.79	12.74	1.6:	1372	16.07	12.63	1.2
1313	16.20	12.99	0.6	1373	18.81	14.23	0.6
1314	16.54	14.18	0.6	1374	16.95	14.70	0.8
1315	15.26	10.99	2.3	1375	15.59	12.84	2.1
1316 1317 1318 1319 1320	17.45 13.85 15.52 15.57 15.57	14.79 9.65 13.13 11.71 11.71	0.6 0.8 0.6 1.1	1376 1377 1378 1379 1380	15.92 16.45 15.72 15.04 17.22	13.73 14.18 13.16 12.10 13.08	2.2 2.0 2.0 2.6 0.6
1321	14.89	11.10	2.5	1381	15.77	12.93	2.6
1322	16.76	14.08	1.0	1382	15.62	13.46	2.0
1323	15.45	11.23	3.0	1383	16.86	12.84	1.6
1324	15.63	13.56	0.6	1384	16.08	12.82	0.6
1325	16.17	13.21	1.2	1385	15.29	11.90	2.1
1326	15.14	11.90	0.7	1386	17.22	14.68	0.6
1327	16.37	12.89	2.0	1387	16.88	14.60	0.6
1328	15.99	11.26	1.0	1388	15.75	11.82	1.4
1329	14.51	11.37	2.2	1389	16.16	12.52	2.6
1330	15.74	11.55	0.6	1390	14.64	10.03	2.1
1331	15.80	11.73	1.6	1391	17.79	14.81	0.6
1332	15.01	11.00	1.11	1392	16.01	12.89	0.9
1333	16.00	12.83	2.8	1393	15.82	13.10	3.0:
1334	15.09	11.36	2.6	1394	15.57	12.84	2.3
1335	17.11	14.89	0.6	1395	16.91	12.67	0.8
1336	15.63	12.02	2.0	1396	15.24	13.00	3.6
1337	15.9h	12.22	1.8	1397	16.00	12.72	2.6
1338	16.30	14.02	2.0	1398	15.55	11.37	1.7
1339	15.43	11.50	2.h	1399	17.33	15.17	0.6
1340	16.66	12.47	1.6	1400	16.99	12.88	0.6
1341 1342 1343 1344 1345	15.40 15.70 15.52 16.26 16.21	12.00 13.35 12.49 14.03 10.86	0.6 0.8 2.lı 2.0 0.8	1401 1402 1403 1404 1405	14.72 17.79 16.88 16.83 16.55	12.41 14.52 13.56 10.17 14.31	0.6 0.6 0.9 0.5
1346	15.52	12.36	0.6	1406	16.05	12.75	0.5
1347	15.20	12.17	1.6	1407	15.72	12.28	1.9
1348	15.58	12.09	2.6	1408	16.18	12.09	1.7
1349	15.47	11.51,	1.8	1409	14.98	11.72	1.8
1350	15.83	12.20	2.0:	1410	16.28	12.36	3.1
1351	15.19	10.98	1.8	1411	15.82	11.93	2.6
1352	15.74	12.27	1.8	1412	15.91	13.76	1.6
1353	14.99	11.08	2.4	1413	16.37	12.44	3.2
1354	16.23	12.11	1.3	1414	17.22	13.74	0.6
1355	14.80	13.30	1.2	1415	15.63	13.46	2.4
1356	15•30	11.26	2.0	1416	15.65	11.73	1.9
1357	14•84	10.61	0.6	1417	16.18	12.34	2.0
1358	15•64	12.83	1.5	1418	15.22	13.04	3.6
1359	14•96	10.86	1.0	1419	15.10	12.74	1.6
1360	15•63	12.46	0.8	1420	16.08	12.67	0.6
1361	15.98	11.93	0.8	1421	15.49	11.44	1.9
1362	16.20	11.82	0.6	1422	17.34	15.10	0.8
1363	16.26	12.55	2.6	1423	16.17	12.54	2.1
1364	15.89	11.98	2.2	1424	14.93	10.72	1.9
1365	15.46	13.22	2.8	1425	15.75	12.64	0.5
1366 1367 1368 1369 1370	14.84 16.69 14.80 15.48 17.11	11.18 14.20 11.88 11.38 14.86	2.5 1.0 1.5 1.7 0.6	1426 1427 1428 1429 1430	15.18 15.29 15.05 16.63 16.16	12.13 11.88 11.52 13.64 13.16	1.0 1.8 2.2 1.0

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No.	P _o	g	wt	No.	P _o	g	wt
1431	15.64	12.51	1.0	1491	16.74	12.51	1.0
1432	15.87	13.28	1.5	1492	16.36	14.33	2.7:
1433	16.26	12.76	1.2	1493	15.29	12.58	2.3:
1434	15.34	11.42	2.4	1494	16.06	13.98	1.4
1435	17.90	14.70	0.6	1495	16.60	13.42	0.9
1436	15.52	11.37	2.0	1496	15.94	13.82	1.8
1437	15.76	9.12	5.6	1497	16.42	12.72	0.9
1438	16.88	12.70	1.0	1498	17.06	13.00	0.7
1439	16.31	10.93	0.6	1499	16.52	13.27	1.5
1440	17.22	13.04	0.8	1500	16.85	14.63	0.5
1441	17.33	14.16	0.6	1501	16.61	13.64	2.5
1442	16.21	12.55	1.5	1502	16.54	13.16	0.6
1443	16.06	12.28	1.0	1503	14.99	11.84	2.5
1444	16.31	12.14	0.6	1504	15.75	13.12	0.9
1445	16.03	11.93	1.3	1505	15.58	12.36	1.8
1446	16.35	14.12	0.5	1506	16.26	13.24	1.8
1447	15.78	12.83	1.6	1507	17.25	14.79	0.5
1448	16.85	14.29	0.5	1508	16.85	13.40	0.5
1449	15.89	13.72	0.9	1509	15.04	14.00	1.8
1450	15.67	12.55	1.5	1510	15.77	12.52	2.4
1451	15.85	13.72	0.5	1511	16.60	14.07	1.0
1452	17.09	13.00	1.0	1512	15.84	10.50	2.0
1453	14.85	13.70	1.0	1514	16.69	14.60	0.8
1454	17.03	14.49	0.8	1514	15.76	13.54	1.0
1455	16.65	14.41	0.6	1515	16.88	13.85	0.6
1456	15.98	11.78	1.0	1516	16.06	12.91	2.0
1457	15.66	12.36	0.8	1517	15.52	12.17	0.6
1458	15.72	12.57	2.0	1518	15.84	13.66	1.9
1459	16.14	11.97	2.6	1519	16.95	12.84	0.5
1460	16.79	13.83	0.7	1520	15.84	11.76	1.9
11461	14.66	10.55	1.2	1521	16.75	13.14	0.5
11462	16.25	12.09	0.5	1522	16.23	13.69	0.9
11463	16.13	12.00	0.9:	1523	15.56	13.34	1.6
11464	16.12	12.23	2.5	1524	15.97	11.89	2.5
11465	16.08	12.14	0.6	1525	17.57	14.27	0.9
14,66	16.60	14.02	0.8	1526	17.15	14.73	0.5
14,67	14.33	9.79	1.8	1527	15.79	13.61	1.1
14,68	16.75	14.54	3.0	1528	16.19	13.52	1.6
14,69	14.98	10.87	4.6	1529	16.58	11.19	1.1
14,70	17.33	13.16	0.6	1530	16.95	14.71	0.5
1471	17.33	13.99	0.6	1531	16.18	13.02	0.8
1472	16.08	13.88	1.7	1532	15.78	11.88	2.5
1473	16.55	13.51	0.5	1533	15.80	11.89	0.7
1474	16.30	12.92	0.8	1534	16.32	12.95	1.5
1475	16.59	14.09	0.8	1535	16.96	12.80	0.9
1476	17.12	14.79	0.8	1536	16.61	14.49	0.7
1477	16.47	12.27	0.7	1537	17.79	13.82	0.6
1478	16.35	13.56	0.5	1538	18.02	15.48	0.6
1479	15.68	12.42	2.0	1539	16.11	12.01	2.7
1480	16.40	14.29	1.2	1540	15.55	11.94	0.5
14,81	15.90	11.98	0.8	1541	15.91	12.46	2.0
14,82	15.71	12.06	3.1	1542	15.63	11.57	1.8
14,83	15.89	12.55	2.5	1543	16.99	13.83	0.6
14,84	15.61	12.22	0.8	1544	15.43	12.86	1.8
14,85	16.31	12.37	0.6	1545	16.26	12.80	2.1
1486	16.76	14.66	1.6	1546	15.83	11.65	1.2
1487	15.86	11.73	2.9	1547	15.92	12.73	1.0:
1488	15.95	11.99	0.5	1548	15.15	11.67	0.5
1489	17.22	13.02	0.6	1549	15.55	13.36	0.5
1490	15.17	12.66	1.7	1550	16.74	13.76	0.8

TABLE 7-Continu

No.	р _о	g	wt	No.	р _о	g	wt
3 449	_		0.0				
1551	16.20	13.58	2.9	1586	15.74	13.04	0.6
1 552	16.53 16.44	12.63 12.72	1.1 1.0	1587 1588	15.78	12.81 12.10	1.4
1553 1554	15.85	12.71	2.0	1589	16.0Ц 15.90	13.22	o•fi
1555	15.82	12.53	1.3	1590	15.23	13.04	3.6 2.6
±333	15,02	12.55	ر•1	1990	19.23	13.04	2.0
1556	15.95	11.35	1.2	1591	15.79	13.18	0.9
1557	16.13	12.22	2.1	1592	16.32	12.86	1.4
1558	15.78	11.52	2.2	1593	16.76	14.58	1.3
1559	15.88	13.27	0.9	1594	15.68	13.38	1.2
1560	16.07	12.79	2.1	1595	15.76	12.57	0.9
	- 4			- 4-4			
1561	16.17	11.98	0.8	1 596	15.86	12.17	1.4
1562	15.69	13.51	1.0	1597	16.89	13.27	1.6:
1563	15.84	13.76	1.0	1598	16.78	32 بلا	2.0
1564	16.21	12.05	0.8	1599	16.33	12.19	2.6
1565	16.29	13.68	1.3	1600	16.99	16.01	0.6
1566	12.35	17.7և	0.5	1601	16.01	13.81	1.9
1567	15.17	10.90	1.4	1602	15.91	13.68	ī.ú
1568	15.63	13.12	1.7	1603	15.46	12.04	2.4
1569	16.59	12.43	2.2:	160 <u>L</u>	15.64	11.70	3.6
1570	16.13	12.53	2.2	1605	15.27	11.36	1.6
	-						
1571	17.26	13.13	1.2	1606	15.95	12,66	1.5
157 2	15.40	11.31	2.2	1607	15.97	13.00	4.4
157 3	16.43	13.87	0.6	1608	1 5 .7 4	13.60	0.6
1574	16.30	11.54	1.6	1609	14 . 99	11.93	2.6
157 5	16.47	13.90	0.8	16 1 0	16.88	14.77	0.6
1576	. 15.78	11.65	1.5	1611	15.97	11.73	0.6
1577	17.43	15.24	1.8	1612	16.08	12.01	0.6
1578	17.05	11.71	2.0	1613	16.31	12.92	2.6
1579	15.63	11.02	1.1	1614	15.52	11.64	2.6
1580	17.65	15.56	0.5	1615	16.43	12.33	0.6
	-1100	-,-,-		202)	2002)	20000	0.0
1581	15.34	11.14	0.2	1616	16.65	12.92	0.6
1582	17.20	13.02	0.8				
1 583	16.48	9.71	3.4				
1584	14.82	12.25	1.3				
158 5	15.51	11.75	1.0				

The weight of magnitudes printed in italics in the Ephemeris was determined from the plot of differences, italics minus combined magnitudes, against combined magnitudes, shown in Figure 2. The average of the differences is found to be -0.35 mag. The average deviation of the dots in Figure 2 from -0.35 is ± 0.43 mag. The combined values that were used have an average weight of 1.2, or ± 0.20 mag. probable error. The probable error of the italic values is then ± 0.30 mag., and their weight is 0.5. Italic magnitudes from the Ephemeris were therefore used in the combination, with weight 0.5 after correction of +0.35 mag.

The accidental errors of magnitude determinations at other observatories were derived in Paper VI by avoiding aspect variations; thus unit weight in Paper VI corresponds to ± 0.14 mag. not ± 0.22 mag. Partly for convenience, the relative weights so established were retained in the present compilation, although it includes the scatter from aspect variations. The relative weights of the more uncertain series will thus have been slightly reduced.

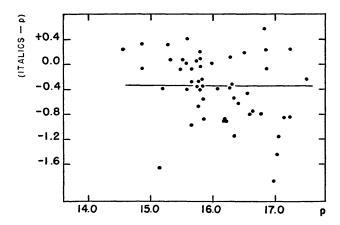


Fig. 2.—Plot of differences (italics minus p) versus p

For about 300 asteroids, magnitudes were obtained in the Survey at two oppositions. Some 800 asteroids were observed in the Survey during one opposition only; the precision of the final magnitude for these objects was increased by inclusion of two magnitude observations made at other observatories in addition to the reduced Ephemeris magnitudes mentioned before. For some 500 asteroids no magnitudes could be determined in the Survey; for each of these, four magnitude determinations made elsewhere were added to the reduced Ephemeris magnitudes, if such additional values were available. The magnitude values determined at other observatories were taken as much as possible from oppositions different from the Survey and from each other. They were collected from the Cincinnati Index, which was kindly loaned to us by Dr. Herget. The Index magnitudes were reduced to absolute magnitude by the same method and the same phase corrections as those of the Survey magnitudes. Usually only observations of the years 1948–1954 were included, but in a few cases 1947 observations were used also. The series used are discussed in Paper VI, with an example of the combination procedure given in Table 5 of that paper.

The result of the compilation is found in Table 7, with the weighted mean absolute magnitude, g, given in column 3. The total weights of the combinations are shown in column 4; they are the sum of the weights of the separate determinations if only one Survey magnitude was available. If more than one Survey magnitude was derived within a period of 10 days, the average rather than the sum of the weights of the Survey magnitudes is given. This was done to allow for common uncertainties, such as in aspect variation and in the reductions to absolute magnitude; furthermore, these asteroids oc-

curred on the overlapping edges of the plates where the images were somewhat inferior. Because the aspect variations correspond to a p.e. roughly of ± 0.16 mag., it is obvious that an absolute magnitude derived during a single opposition cannot have a weight in excess of roughly 2.0 (unit weight corresponding to ± 0.22 mag.).

For some 38 asteroids, photoelectric data other than complete light-curves had been obtained at the time of this discussion; some of these are based on observations made on two nights. They are collected in Tables 9 and 10 and have been added to the data in Table 7 with weight 2.0, after reduction to the International Photographic Scale by means of equation (1). In about 150 cases neither Survey magnitudes nor Index magnitudes were available; they can be recognized from the assigned weight, 0.6, in column 4 of Table 7. For the Trojans the magnitudes were taken from Paper VI, after inclusion of the corrected Ephemeris magnitudes.

Column 2 of Table 7, giving the mean opposition magnitudes, was derived from column 3 by means of the relation

$$p_0 = g + 5 \log a (a - 1).$$
(3)

The differences between columns 2 and 3 were computed twice independently, based on the 1955 and 1957 Ephemeris, respectively. To preserve the computational result, they are given in two decimals even in cases of low weight.

TABLE 8

COMPARISON OF INDIVIDUAL g-VALUES OF UNIT WEIGHT

Interval of Comparison	Average Difference of Two Values	p.e. of One Value	No. of Compari- sons	Interval of Comparison	Average Difference of Two Values	p.e. of One Value	No. of Compari- sons
0 ^d	0.232	0.14	90	14 ^d -25 ^d	0.254	0.15	62
1 ^d –10 ^d	0.276	0.17	58	1 -2 years	0.362	0.22	295

The precision of the Survey magnitudes referred to above will now be determined. All asteroids were used for which two determinations were made. They were subdivided into groups according to the interval elapsed between the two measures, as given in Table 8. It is noted that the probable error of one value is essentially constant, at ± 0.15 mag., for comparisons between values obtained during the same opposition, while the probable error is ± 0.22 mag. for comparisons between different oppositions. The former result is consistent with the discussion in Section VII in which a probable error of ± 0.14 mag. was found for one determination of unit weight (which consists of two measures made on pairs of plates taken approximately 1 hour apart). This value depends on internal evidence between the measures on one pair of plates, while ± 0.15 mag. was found from intercomparisons between different plate pairs. The latter value includes certain additional effects, such as differential corrections for phase. The extra dispersion between magnitudes obtained at different oppositions and derived above on the basis of Survey magnitudes alone, was qualitatively confirmed by comparisons of Survey magnitudes with observations made at Heidelberg and Simeis. In both instances the dispersion of the differences, Observatory minus Survey, was substantially larger if the magnitudes had been obtained at different oppositions.

As remarked before, the difference in the two probable errors, ± 0.22 and ± 0.15 mag., corresponding to a probable error of ± 0.16 mag., may be attributed to aspect variations. This may be shown as follows: Intercomparisons between different oppositions involve a comparison between the zero point of the 1950–1951 opposition with that of 1951–1952. A variety of checks made in connection with the preparation of Paper VI as well as

Figures 1 and 2 of that paper show that the difference in the zero points of the two oppositions is not in excess of 0.05 mag.; actually, since the same Selected Areas were used in the two oppositions, no systematic differences would be expected. The next possibility was that the larger difference between consecutive oppositions might be due to occasional errors in identification. The identifications were examined with care whenever the g-values of the two oppositions differed more than 0.35 mag. The criteria used were O-C in both position and motion, with allowance for the variation, as well as the similarity of the O-C values between the two oppositions. No identifications were found to be erroneous. Parenthetically, for future identifications the importance is stressed of accurate magnitude determinations, especially for fainter objects, for which the numbers increase so rapidly that the probability for similar orbits is greatly increased. A third possibility is that phase effects are responsible for the increased scatter between oppositions. Examination showed, however, that this was not the case. The mean phase angle of asteroid magnitudes for the Survey is 5°; the mean difference between pairs of observations was 40, for pairs made both during the same opposition and at different oppositions. The only remaining possibility appears that variations in aspect

The rather surprisingly large scatter due to aspect can be understood if certain conditions are satisfied. One expects theoretically that the asteroids will rotate around the shortest figure axis, i.e., around that axis about which the moment of inertia is largest. If, for simplicity, the asteroid is assumed to have the shape of an ellipsoid, with semiaxes a > b > c, one expects it to rotate around the c-axis. As the asteroid moves in its orbit, the aspect can change at most from pole-on (area πab) to equatorial (area varying between πac and πbc). The rotational variation will be zero pole-on and by the ratio b/aequatorially. Maximum rotational variation will be observed in the plane of the asteroid equator, being 2.5 log b/a magnitudes; while maximum aspect variation, 2.5 log $\frac{1}{2}(c/b +$ c/a), will occur only if the obliquity is so high that the earth can pass near the asteroid pole. Now for the brightest of the asteroids the average rotational semiamplitude may be found from Papers I-V and VII, ± 0.09 mag.; while the average aspect variation, based on Table 8 (which gives p.e.) is about ± 0.19 mag. The latter value may actually approach the maximum time-average, since the interval between consecutive oppositions is nearly 5/4 years, making the two aspects nearly 90° apart. In any case, the comparatively large variation attributed to aspect appears to require that the asteroids as a group have large obliquities unless the c values should be much smaller than the a and b values. Since flat disks are improbable, it is concluded that the asteroids have large (possibly random) obliquities. This subject, including such corollaries as the relation between the phase of the aspect variation and the amplitude of the rotational variation, will not be pursued here; when the maximum aspect is presented, the rotational variation should be at minimum, as seems to be true for asteroid 511, described in Paper I.

Asteroids showing large differences between the magnitudes determined at different oppositions are of special interest; they are designated by a colon following the weight in column 4 of Table 7. Included are the objects whose average deviation per unit weight of the individual g-values from their mean was greater than about 0.35 mag. This value may be compared with that expected on the basis of the probable error of ± 0.22 mag. per individual observation at different oppositions. If the average is based on four observations, typical for Table 7, the computed mean deviation is 0.32 mag., and the fraction for which the mean deviation is larger than 0.35 mag. is about 23 per cent, which is roughly the percentage of such deviations actually found. The aspect deviations may therefore roughly follow a Gaussian distribution. The reason for marking with a colon the large differences between oppositions is that these objects as a class will have appreciable obliquities and departures from spherical shape and are therefore of special interest. It is noted that the objects in question are by no means a complete inventory of asteroids with large aspect variations, since their selection was based on two adjacent

oppositions for the Survey, to which a few observations made at other observatories and at different oppositions were added. In addition, a few cases are uncertain for other reasons, as is shown by internal inconsistencies between the measures. If the asteroids in Table 7 marked with a colon in column 4 are omitted, a rediscussion of the material used in Table 8 gives the probable errors ± 0.13 , 0.16, 0.14, and 0.17 mag., respectively.

Some cases were found where the magnitudes in the Ephemeris should perhaps have been printed in italics, while they were not; then the Ephemeris magnitudes were not used, or else a colon was added in column 4 of Table 7. Furthermore, there are cases where the Ephemeris magnitudes were exceptionally discrepant, so that they were not used for combination; examples are 127, 369, 830, 1178, 1186, 1192, 1206, 1220, 1235, 1243, 1263, 1312, 1345, 1552, 1555, 1581, 1588, 1591, 1592, 1593, 1602, and 1606. It is therefore likely that, for some cases for which we did not have recent observations and for which, therefore, the Ephemeris value transformed to the p system was taken, the magnitudes are actually too faint by about $1\frac{1}{2}$ mag. because the Ephemeris value should have been printed in italics but was not. The opposite may have occurred also; an example may be 1463, where two observations are available, although of low weight and only in one opposition, that give an absolute magnitude of 12.9, while the Ephemeris q is 11.0.

The asteroids marked with colons in column 4 of Table 7 were investigated for possible orbital peculiarities. No appreciable difference was found between the average orbital eccentricities of these bodies and those of asteroids in general, for the semimajor axis a, or in the average absolute magnitude. The average orbital inclination, however, was 8.3 versus 10.5 for the other asteroids, a difference that may just be significant. If this effect is real, it may mean that these bodies have experienced more collisions. The scatter of the g-values for each asteroid, within one opposition or in different oppositions, appears not correlated with p, unless the scatter be slightly less for fainter asteroids.

The magnitude scale of Table 7 was checked by a special photoelectric program which supplemented the photoelectric checks already available from asteroids whose lightcurves had been observed. Since the latter group consisted mostly of bright asteroids, the supplementary observations were extended down to about 16.0 mag. The useful comparisons between the photoelectric and photographic scales are listed in Table 9. In this table are included all asteroids observed photoelectrically for which photographic observations existed for more than one opposition and whose total weight exceeded 1.8. The table is arranged by increasing P_0 . Column 2 gives the year, month, date, and decimal in such cases where the results have not been published before. If the decimal of the day is omitted, it signifies that a mean magnitude has been used, taken from light-curves published in Papers I-V and VII. The observers and telescopes are listed in column 3. Column 4, giving V, gives the mean magnitude determined from lightcurves or the mean derived from different observations during the same night. Column 7 has been derived from columns 4 and 5 by means of equation (1). Column 8 gives the quantity p_0 taken from Table 7 after the photoelectric data used in the compilation had been removed.

The identification of moving objects near the limit of visibility, as required here, poses a special problem in photoelectric photometry. Finding charts were prepared from Survey plates, and the Ephemeris positions were corrected with the residuals derived from 10-inch McDonald or 24-inch Yerkes plates taken for the purpose. The 24-inch objects were identified on search plates taken just prior to the photoelectric observations. The 24-inch photoelectric records were made with the aid of an integrator designed and built by Mr. R. H. Weitbrecht. The estimated probable errors of the photoelectric magnitudes are small compared to those of the photographic magnitudes in Table 9. The photoelectric data listed in Table 10 could not be matched with accurate photographic observations; but they were, of course, used in Table 7.

TABLE 9
PHOTOELECTRIC CHECK OF THE MAGNITUDES OF TABLE 7

		1			_ 		1		
Asteroid (1)	Date U.T. (2)	Observ- er* (3)	\overline{V} (4)	B-V (5)	<i>U-B</i> (6)	P_0 Photoel. (7)	#0 Table 7 (8)	$p_0 - P_0$ (9)	
29 18 9	56.3.4.3 56.8.22.1 {54.1.3 54.1.16	K K Gr Gr	9.83 11.39 8.67 9.08	+0.88 + .83 + .84 + .85	+0.30 + .50 + .48	10.01 10.10 10.19 10.22	10.27 10.12 9.91	$ \begin{array}{c} +0.26 \\ + .02 \\30 \end{array} $	
16 11	\$55.12.26 \$56.1.2 \$6.1.3 \$52.1.29	Gr Gr Gr K	9.78 9.92 10.59 10.28	+ .71 + .81 + .89	+ .25 + .39 + .52	10.32 10.44 10.60	10.65	$\begin{array}{ c c c c c c c c c c c c c c c c c c c$	+0.05
39	53.4.10 55.12.18 55.12.19 55.12.28	Gr Gr Gr Gr	10.65 10.65 10.82	+ .88	+ .49 $+ .50$	10.67	10.76	09	
349	\$\\ \{ 56.8.23.2 \\ 56.8.24.2 \end{array}	K K	9.74: 9.80	+ .91: + .98	+ .50 + .59	10.78: 10.92	11.01	+ .14	
30 51 23 511 52 37 128 194	\$\begin{cases} 56.8.23.3 \\ 56.8.24.3 \\ 56.3.4.3 \\ 56.3.4.3 \\ 53.4.8 \\ 56.3.4.3 \\ 56.3.4.3 \\ 56.3.4.3 \\ 56.3.2.2 \end{cases}\$	K K K K Gr K Ge K	9.73: 9.83 10.62 10.27 9.98 11.44 10.67 11.32 11.92 11.02	+ .84: + .88 + .83 + .91 + .71 + .72 + .70 + .89 + .84 + .70	+ .45 + .46 	11.04: 11.18 11.15 11.29 11.32 11.54 11.59 11.68 12.01 12.04	11.29 11.11 11.53 11.01 11.43 11.47 12.11 11.66	+ .16 04 + .24 42 16 21 + .10 38	09
122 498 268 490 380 62 510	$\begin{array}{c} 56.2.16.2 \\ 56.3.13.3 \\ 56.1.4.2 \\ 56.3.13.2 \\ 56.3.3.2 \\ 52.2.16.2 \\ 56.1.4.2 \\ 56.1.6.1 \end{array}$	Ge Ge Gr Ge K Ge Gr Gr	12.61 13.77 13.32 13.29 13.62 13.08 14.50 14.40	+ .68 + .77 + .70 + .78 + .72 + .76 + .74 + .73:	+ .41 + .36 + .30 + .69 	13.11 13.14 13.32 13.61 13.70 13.76 14.01 13.91	13.08 13.06 13.85 13.68 13.78 13.77 14.12	$ \begin{array}{cccc} & - & .03 \\ & - & .08 \\ & + & .53 \\ & + & .07 \\ & + & .08 \\ & + & .01 \\ & + & .17 \end{array} $	+ .11
540 976 321 1043 658 1291	56.3.13.3 56.3.13.2 \$55.12.19 \$55.12.20 56.3.4.2 \$56.1.4.2 \$56.1.6.2 \$56.1.6.2 \$56.1.6.2 \$56.1.6.3	Ge Ge Gr Gr Gr Gr Gr Gr	12.96 13.41 14.17 15.27 14.52 14.57 14.31 14.74 15.17 15.26:	+ .95 + .74 + .82 + .92 + .87 + .87 + .86 + .81 +0.85	+ .06 + .25 + .45 	14.26 14.46 14.73 14.77 15.02 15.05 15.25 15.06 15.46 16.15:	13.89 14.86 15.07 14.61 15.03 15.26 15.30 15.87	$ \begin{array}{rrr}37 \\ + .40 \\ + .34 \\16 \\01 \\ + .01 \\ + .04 \\ - 0.28 \end{array} $	0.00

^{*} Ge = Gehrels, 24-inch; Gr = Groeneveld, 82-inch; K = Kuiper, 82-inch.

Group averages per $1\frac{1}{2}$ -mag. interval of the magnitude differences in column 9 of Table 9 are found in column 10. It is noted that there is no dependence on apparent magnitude. The over-all average of $(p_0 - P_0)$ is $+0.01 \pm 0.03$ mag. (p.e.). Both the zero point and the scale of Table 7 are thus found to be satisfactory. The average deviation of single values from the mean difference is ± 0.18 mag. This quantity is estimated to have a probable error itself of ± 0.03 mag. The probable error thus found cannot be compared

TABLE 10
Additional Photoelectric Data

Aster- oid (1)	Date U.T. (2)	Observer	<i>V</i> (4)	B-V (5)	<i>U-B</i> (6)	P ₀ Photoelectric (7)
7	55.12.28 55.12.29 56.1.2	Gr \ Gr \ Gr \	9.81	0.86	0.47	9.29
	56.1.5 56.3.8 (55.12.20 55.12.27	Gr∫ Gr Gr Gr}	8.99			
8	55.12.28 55.12.29 56.1.1	Gr\ Gr∫ Gr\ Gr∫	8.82	.88	.48	9.53
15	\begin{cases} 55.12.21 \\ 55.12.24 \\ 55.12.27 \end{cases}	Gr Gr Gr	8.65 8.74 8.83	.84	.44	9.15
17	$ \begin{cases} 55.12.23 \\ 55.12.24 \\ 55.12.25 \end{cases} $	Gr Gr Gr	11.47	.84	.40	11.22
25	(56.1.3)56.1.4)56.1.5 (56.1.6	Gr Gr Gr Gr	12.54	.93	.51	11.59
3	\$56.8.22.2 \$56.8.23.2	K K	9.95 10.07:	.77 .83:	.35 .38	9.45 9.64:
10	\$56.8.23.2 \$56.8.24.2	K K	10.39: 10.39	.78: .71	. 47 . 40	10.90: 10.82
2	\$56.8.23.2 \$56.8.24.2	K K	9.36: 9.32	.62: 0.65	0.27 0.27	8.59: 8.58

directly with the values listed in Table 8 because, in comparisons with observations taken at different oppositions, the photoelectric data have a scatter due to aspect, while the asteroids to be observed photoelectrically were somewhat selected to favor objects having consistent photographic determinations. Nevertheless, the quantity is of the order of magnitude to be expected on the basis of previous discussions.

Mr. Gehrels is largely responsible for this section. Part of the reductions were carried out by him at Indiana University, after his appointment there in August, 1956.

IX. CONTROLS AND REVISIONS

The measured positions, motions, and magnitudes were entered on sheets similar to Table A, except that the latter is arranged in order of asteroid number, while the original

Survey records were kept separate for each opposition field and were arranged by the plate-field number. For each plate pair, the asteroids were assigned numbers added to the plate-field number, e.g., B24.3. The Survey records were checked carefully against the original measures and computations. The positions were measured once, but the reductions were all checked independently; furthermore, the previously derived approximate positions were used as a check on the theodolite positions, and the latter were remeasured in a number of doubtful cases. The identifications of the first eight opposition fields were checked by Dr. E. K. Rabe, as consultant to the project, during a month's stay at the Yerkes Observatory in August, 1954; later he made additional checks at Cincinnati and found that none of the identifications checked needed revision. In the early stages of the project, advice was given by Dr. W. Strobel, of the Rechen Institute at Heidelberg. The O—C's for the positions were computed in duplicate. The daily motions were measured and computed only once, and only in cases where the O—C's with the Ephemeris motions were large were control measures and reductions made.

The magnitude measures and reductions for the a and b plates of each pair were made independently; in case of serious discrepancy they were repeated. Similarly, the g-values were derived independently for each plate pair and controlled if necessary. This latter check led to the discovery of a few misidentifications or clerical errors. As has been stated, the differences between columns 2 and 3 of Table 7 were computed in duplicate.

X. COMPLETENESS OF SURVEY; FREQUENCY-CURVE OF ASTEROIDS

One of the principal aims of the Survey was the determination of the frequency-curve of asteroid magnitudes to the limit of the plates. Such a determination requires a fairly homogeneous collection of plates. All factors that influence the apparent brightness of an asteroid, other than its size, such as the distances from the sun and the earth and the phase correction, will, to a first approximation, be the same for faint and bright asteroids, so that the rate of increase with apparent magnitude is a significant quantity. Because the discovery of asteroids by blinking is never entirely complete, particularly for fainter asteroids, the completeness factors must be determined. In the present Survey this has been done in two ways: (a) from overlapping Survey regions and (b) from comparison with the Ephemeris asteroids.

Method a can be used because the blinking was done without previous knowledge of asteroids present. A fraction of the objects will have been recorded twice, another fraction once, and a third fraction not at all. The ratio of the first two fractions is readily found to be

$$\frac{\text{Found twice}}{\text{Found once}} = \frac{k^2}{2k(1-k)},$$
(4)

in which k is the completeness factor for single blinking. The method depends, however, on the outer areas of the plates, which, because of the somewhat inferior image quality, will tend to make k slightly too small for the plate average.

The completeness of the entire Survey will be larger than k because of the areas blinked twice. If μ is the fraction of the 40° belt blinked twice, the average probability of finding an asteroid is clearly

$$K = (1 - \mu) k + \mu [k^2 + 2k (1 - k)].$$

Method b also depends on the fact that the blinking was done independently of previous knowledge, so that the fraction of Ephemeris asteroids recorded gives a true measure of the degree of completeness of each magnitude group. Objects found by reblinking must, of course, be excluded. On the other hand, if an object is not found by reblinking, its position might be grossly in error; then the object should not be counted at all. If it were missed for any other reason, it should be counted. Half the objects not found by reblinking were regarded here as "lost" and were not counted.

Method a was used as follows: All objects were counted that were found in the Survey, irrespective of recurrences; the completeness, K, as defined will apply to these counts. This approach has the advantage that the largest possible numbers are used. The quantity μ was found to be 0.25. The material was divided into groups of fields blinked by the same person and of about the same quality. For these groups the degree of completeness was computed; the results are found in Table 11. The figures in parentheses are found by method b. It is seen that the two methods give generally accordant results.

The numbers of asteroids, counted per half-magnitude intervals and corrected for incompleteness, are given in Table 12. The last column of that table shows the number of declination strips, each one plate (6.5) wide, which contributed to the statistics. The totals are found at the bottom, with the values in parentheses adjusted to the full num-

TABLE 11

DEGREE OF COMPLETENESS OF SURVEY

Fields	Blinked by	14.0 14	5 19	5.0 15	5.5 10	6.0 16.5	Effective Plate Limit
$B, C, D, E, F \dots$ $G, \dots, H, J, \dots, I, K, \dots, I$ $L, \dots, M, N, M, N, M, M,$	K T Gr VH T VH T	0.85(0.94) 0.77 1.00	0.80(1.00) 0.90(0.89) 1.00 0.50 0.50 0.86	0.80(0.79) 0.91(0.87) 0.67	0.50(0.41) 0.73(0.69) 0.80(0.79)	0.86	16.5 16.7

ber, 154, of strips. The logarithms of the adjusted numbers are found in Table 13, together with the residuals with respect to the equation

$$\log n = 1.983 + 0.344 (p - p_1) - 0.014 (p - p_1)^{2},$$

$$\pm 0.023 \pm 0.007 \qquad \pm 0.005 \text{ (mean errors)}$$
(5)

which was obtained by least squares; $p_1 = 12.75$. The numbers and equation (5) are shown in Figure 3. The deviation from linearity is small and hardly significant.

The constant term in equation (5) must now be reduced to the proper unit. Subtraction by $\log 154 = 2.188$ leads to values applicable per one strip; subtraction of $\log 1.10 = 0.046$ allows for overlaps between strips. Addition of $\log 360^{\circ}/6.5 = 1.744$ reduces the count to the entire ecliptic belt. A correction by the factor 1.10 (logarithm + 0.046) is required, to allow for objects missed because they were outside the 40° belt (the correction is somewhat larger than the 7 per cent noted above for the entire Survey because recurrences are now excluded and the Survey covered the asteroids, on the average, nearly 1.5 times). The four reductions combine to make the constant term in equation (5) equal to +1.539. Reduction to the interval of 1 mag. requires multiplication by a factor slightly in excess of 2. More precisely, the factor is $(10^{s/2} - 10^{-s/2})/10^{s/4} - 10^{-s/4}) = 2.04$ (logarithm 0.309), where s is the coefficient 0.344 in equation (5). The constant term thus becomes 1.848.

Equation (5), thus modified, still does not allow for the fact that one year (360°) does not bring all asteroids into opposition; the fraction will be the mean of the reciprocal synodic periods. The average mean motion of asteroids 1–250 and 1001–1250 is 798″.4/

TABLE 12

Numbers of Asteroids in Magnitude Intervals (Corrected for Incompleteness)

Field	9.0 9.	.5 10	.0 10.	5 11	.0 11	.5 12	.0 12	.5 13	.0 13	3.5 14	.0 14	5 1	5.0 15	5.5 10	6.0	No. Vert Str
B-F G H-J I-K L M-N O R-V X	0 1 0 0 0	1 1 0 0 1 0 1	3 1 2 1 1 0 2 3 0	3 0 2 0 0 1 4 5 1	2 1 3 4 1 2 0 6 4 2	5 4 3 6 5 5 2 10 6 2	14 6 8 6 2 6 4 15 4 2	23 5 7 13 3 8 9 18 8	27 10 16 20 11 7 5 30 16 5	41 15 14 28 8 22 9 45 22 8	53 12 11 28 15 31 10 50 19	66 16 31 45 21 60 36 90 29 13	147 42 39 60 30 102 41 22	45 70 147 51	82	36 8 12 17 8 16 8 31
Total	4	6	13	16	25	48	67	100	147	212	235	407	483 (571)	313 (690)	82 (1204)	15

TABLE 13
REPRESENTATION OF EQUATION (5)

Þ	log n Obs.	log n Comp.	0-C	Þ	log n Obs.	log n Comp.	0-C
9.25	0.60 0.78 1.11 1.20 1.40 1.68 1.83 2.00	0.60 0.82 1.04 1.24 1.43 1.63 1.81 1.98	0.00 04 + .07 04 03 + .05 + .02 +0.02	13.25	2.17 2.33 2.37 2.61 2.76 2.84 3.08	2.15 2.31 2.47 2.61 2.76 2.89 3.01	+0.02 + .02 10 .00 .00 05 +0.07

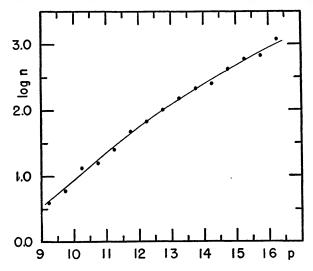


Fig. 3.—Observed numbers of asteroids, corrected for incompleteness, and representation by equation (5).

day compared to 3548''.2/day for the earth; the average differential motion is therefore 2749''.8/day, and the annual fraction 2749.8/3548.2 = 1/1.290 (add log 1.29 or 0.111); Further, the observed magnitudes are not quite so bright as the opposition magnitudes; if the mean correction is $\frac{1}{4}$ mag., the correction to the constant in equation (5) is 0.344/4 = 0.086. The two effects together will make the constant term in equation (5) about +2.045.

Yet another effect results from the finite orbital eccentricities. They cause asteroids of a given p_0 to be observed at opposition magnitudes that scatter by amounts Δp of roughly ± 1 mag., because, for a=2.8 and e=0.2, the perihelion opposition magnitude is 1.29 mag. brighter and the aphelion magnitude 0.98 mag. fainter than p_0 . Now the law of areas will make the number of asteroids near perihelion less than that near aphelion. It is readily shown that the ratio of the probabilities of occurrence of r < a and of r > a is $(1 - 2e/\pi)/(1 + 2e/\pi)$, which, for e=0.2, equals 0.773. Since the ratio 0.98/1.29 = 0.760, almost the same value, the mean of Δp will actually be nearly zero;

TABLE 14
FREQUENCY OF ASTEROIDS IN THE 1957 EPHEMERIS, BY MEAN
PHOTOGRAPHIC OPPOSITION MAGNITUDE

p 0	$N(p_0)$	E q. (6)	Eq. (7)	O-C, Eq. (7)
7	2	$\frac{1}{2}$	1	$+ 1 \pm 1$
8	1	$1\frac{1}{2}$	3	-2 ± 2
9	3	4	6	-3 ± 2
10	9	11	13	-4 ± 4
11	30	28	30	0 ± 5
12	83	67	66	$+17 \pm 8$
13	185	150	148	$+37 \pm 12$
14	269	315	332	-57 ± 18
15	478	620	740	
16	401	1150	1660	
17	133	1990	3700	1
18	12	3240	8300	1
19	7	4940	18600	1
20	3	7050	42000	1

but, because the asteroid numbers increase exponentially with p_0 , the scatter in Δp will indirectly cause a small increase in the observed numbers in p. A 10 per cent increase will make the coefficient 2.09.

The statistics of the Survey thus lead to the following equation, giving the total numbers of asteroids N_p between limits $p_0 - \frac{1}{2}$ and $p_0 + \frac{1}{2}$ for the mean photographic opposition magnitude, p_0 :

$$\log N(p_0) = 2.09 + 0.344(p_0 - p_1) - 0.014(p_0 - p_1)^2, \tag{6}$$

in which $p_1 = 12.75$, as in equation (5). Equation (6) may now be checked by direct comparison with the known numbers of the brighter asteroids where completeness is expected. These numbers are found from counts in Table 7 and are shown in Table 14, second column. The completeness limit may be found as follows: The brightest *new* asteroid found in the Survey was 13.7 mag., and only a few objects were as bright as 14.0. Between 14.0 and 15.0 there were 42 new asteroids in the Survey among a total of 642, or $6\frac{1}{2}$ per cent. There will be a strong statistical preference for these brighter new objects to be asteroids with eccentric orbits observed near perihelion. Accordingly, the Ephemeris and Table 7 are probably complete down to $p_0 = 14.0$, while between 14.0 and 15.0 the incompleteness will probably not exceed half of the $6\frac{1}{2}$ per cent, and 2 per

cent for $13.5 < p_0 < 14.5$. The entry in Table 14 for $p_0 = 14$ may therefore be corrected to about 275, while the entry for $p_0 = 15$ will be the first to be distinctly incomplete.

The third column of Table 14 shows the numbers of asteroids computed according to equation (6). The sums of the bright, the middle, and the faint magnitude groups, 7+8+9, 10+11, 12+13+14, are almost perfectly represented, while the fluctuations within each group are probably merely random. Equation (6), derived statistically, is thus in full accord with the individually observed asteroids brighter than $p_0 = 14$, while on the basis of Tables 12 and 13 it is established to $p_0 = 16$.

The question remains whether the limited information on asteroids fainter than 16 is compatible with equation (6). Baade (1934) derived the number of 44000 asteroids brighter than 19.0 photographic for 360° of longitude. Since the synodic period of the asteroids is about 15 months, this number must be multiplied by $\frac{5}{4}$, to 55000, to represent the *total* number of asteroids. The uncertainty in this figure is probably between 15 and 20 per cent. Recent work on the photographic magnitude scale leads to a corrected limiting magnitude of 19.5 (Baade 1957). Now it is seen from Table 14 that the number of asteroids computed by equation (6) would give 12500 objects brighter than 19.5, much smaller than observed. Also, the quadratic term in equation (6) will cause the computed numbers to pass through a maximum, at $p_0 = 25$ mag., after which the numbers will decrease. At the fortieth magnitude, which is the apparent magnitude of moderate-sized meteorites moving in the asteroid ring, the computed number would be a mere 10 over the entire sky, while the actual number must be enormous. These considerations suggest that the quadratic term (whose reality was in doubt in any case) be dropped and an appropriate linear relation be used instead:

$$\log N(p_0) = -2.38 + 0.35 p_0. \tag{7}$$

The numbers thus computed are shown in the fourth column of Table 14 and the residuals in the fifth; the assigned uncertainties are mean errors (square roots of the computed numbers). While the representation of equation (7) is not quite so good as that of equation (6) (which had an extra parameter), it is fair. The computed number brighter than 19.5 is now 33600, which is of the right order of magnitude. Equation (7) therefore appears to represent the entire range 7 .

If the space density of asteroids increases everywhere with absolute magnitude g as

$$\log N_q = c + b g , (8)$$

then, regardless of the dependence of c on the semimajor axis, a, the mean-opposition magnitudes, p, will increase as

$$\log N(p_0) = c' + bp_0. \tag{9}$$

The proof is readily found. It is therefore not possible to derive from equation (7) anything on the space distribution in the asteroid ring. If in equation (8) a term fluctuating in g is added, this will be smoothed out in equation (9) because, for different a, the fluctuations will appear at different p_0 -values. The possible presence of such deviations from equation (8) must therefore be studied directly from the g distributions for different groups of a.

The material of Table 7 was subdivided into three distance groups:

$$2.0 < a_1 < 2.6 < a_2 < 3.0 < a_3 < 3.5$$
 a.u.

The comparatively few objects with a < 2.0 and a > 3.5 are discussed separately. The counted numbers for half-magnitude intervals in the absolute magnitude g are shown in Table 15. The mean values of a for the three groups were found to be approximately 2.43, 2.75, and 3.17 a.u., corresponding to absolute-magnitude corrections, $p_0 - g$, of 2.71,

3.41, and 4.19 mag. Now the counts in Table 15 are essentially complete to g-values corresponding to $p_0 \leq 14.0$, as shown above. For fainter objects we rely on the discussion accompanying Table 14, from which it is inferred that equations (6) and (7) yield rough lower and upper limits, respectively, to the numbers of asteroids between $14 < p_0 < 18$ mag. The second, third, and fourth columns of Table 14 then give the completeness factors; they are listed in Table 16. These quantities are plotted in Figure 4.

Because of the steep decline in f with magnitude, there is some danger that the completeness-curve so derived may be distorted. Accordingly, the counts and computations were repeated for half-magnitude intervals. Table 17 gives the results. The counts in

 ${\it TABLE~15}$ Observed and Rectified Numbers of Asteroids in Three Zones of a

		2.0 < a < 2	2.6		2.6 < a < 3	3.0		3.0 < a <	3.5	2	2.0 < a < 3	3.5
g	N (Obs.)	N* (Min.)	N* (Max.)	N (Obs.)	N (Min.)	N (Max.)	N (Obs.)	N (Min.)	N (Max.)	N (Obs.)	N (Min.)	N (Max.)
4.25. 4.75. 5.25. 5.75. 6.25. 6.75. 7.25. 7.75. 8.25. 8.75. 9.25. 9.75. 10.25. 10.75. 11.25. 11.75. 12.25. 11.75. 12.25. 13.25. 13.25. 14.25. 14.75. 15.25. 15.75.	0 0 0 0 2 5 5 5 13 15 24 24 29 28 27 54			1 1 0 0 2 1 4 4 15 20 39 51 62 68 78 64 51 55 28 7 2		69 91 104 134 300 360 (360) (550) (3800)	0 0 0 0 1 0 2 5 11 24 30 39 64 93 78 77 45 17 10 0 0 0	30 47 107 246 420 1100 2700 4400	32 54 133 350 680 2050 5700 11000	2 1 0 0 3 3 3 11 14 31 57 84 114 150 180 184 150 138 120 62 42 29 10	84 122 193 333 530 1210 2870 4700	86 129 219 438 800 2180 5900 11400

^{*} Observed numbers are not repeated in adjacent columns if complete.

Table 7 for the intervals 14.00-14.49, 14.50-14.99, etc., are found in the second column. The computed numbers per half-magnitude interval based on equations (6) and (7) are found in the third and fifth columns; the constant terms in equations (6) and (7) were diminished by 0.31 in accordance with the discussion following equation (5). The curves in Figure 4 are based primarily on the half-magnitude intervals. They may now be used to correct the statistics of the three distance groups in Table 15. This was done by dividing the column N(Obs.) by f(max.) and f(min.), leading to the columns N(Min.) and N(Max.) of Table 15, respectively. The f's were found graphically from Figure 4 for values larger than 0.2 and by logarithmic interpolation in Table 17 for smaller values. If N(obs.) was less than 10 and therefore statistically very uncertain, the rectified number was put in parentheses.

The results are plotted in Figure 5. The maximum and minimum values are shown

separately, connected with vertical lines; the statistically uncertain values are connected with broken lines. The numbers below g=7 are very small, and their statistical uncertainties cause difficulties on a logarithmic plot. For this reason, two minor changes were made over Table 15 for g=6.25 and 6.75: the numbers 2 and 1 in the second distance group were interchanged, and the numbers 1 and 0 in the next group were replaced by $\frac{1}{2}$ and $\frac{1}{2}$. This, of course, is statistically quite legitimate.

So far it has been assumed that the completeness of the numbered asteroids (i.e., Table 7) depends on apparent photographic magnitude only. Actually, there is probably

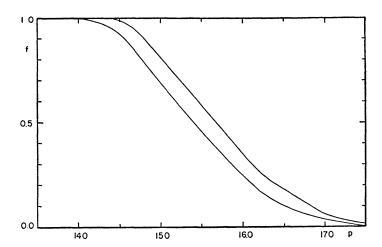


Fig. 4.—Estimated upper and lower limits of completeness factors for Table 7, the Ephemeris asteroids

TABLE 16
Completeness Factors for Table 7 (the Numbered Asteroids)

Þυ	f (Max.)	f (Min.)	Þо	f (Max.)	f (Max.)
14 15 16		0.98 .65 0.24	17 18	0.067 0.004	0.036 0.0016

TABLE 17

DERIVATION OF MINIMUM AND MAXIMUM VALUES OF COMPLETENESS
FACTORS OF TABLE 7 (NUMBERED ASTEROIDS) FOR
HALF-MAGNITUDE INTERVALS

p_0	$N(p_0)$	Eq. (6)*	f (Max.)	Eq. (7)*	f (Min.)
14.25	156	183	(0.853)	198	(0.79)
14.75	234	258	`.907 [′]	297	`.79
15.25	244	357	. 683	444	. 550
15.75	242	520	. 465	665	.364
16.25	159	650	. 245	995	. 160
16.75	100	855	. 117	1490	. 067
17.25	33	1110	. 030	2230	.015
17.75	9	1410	. 0064	3330	. 0027
18.25	3	1770	0.0017	4980	0.0006

^{*} For half-magnitude intervals; hence 0.31 was subtracted from the constant term (see text).

a slight dependence on a as well. If the asteroids were all discovered on plates guided at the sidereal rate (which is not the case), then the ratio of the trail lengths for different a could be readily computed. The trail length for an asteroid in a circular orbit observed at opposition is proportional to $(1 - a^{-1/2})/(a - 1)$. For a = 2.43 and a = 3.17, the mean values of the inner and outer groups, respectively, the ratio is then 1.24, so that the limiting magnitude of the outer group might be about 0.2 mag. fainter than for the inner group. This must be an upper limit, since for short trails the correction is found to be less; the guiding was usually done at the average rate of asteroid motion. From the graphs in Figure 4 and Table 17 it is then found that the differential distortion in the right-hand portions of the graphs for the inner and outer asteroid zones will be less than about 0.2 in $\log N_g$, in the sense that $\log N_g$ in the outer zone will be slightly less, and in the inner zone slightly more, than is shown in Figure 5. The middle zone should be very nearly correct.

It is concluded that there are marked differences in the absolute-magnitude distributions

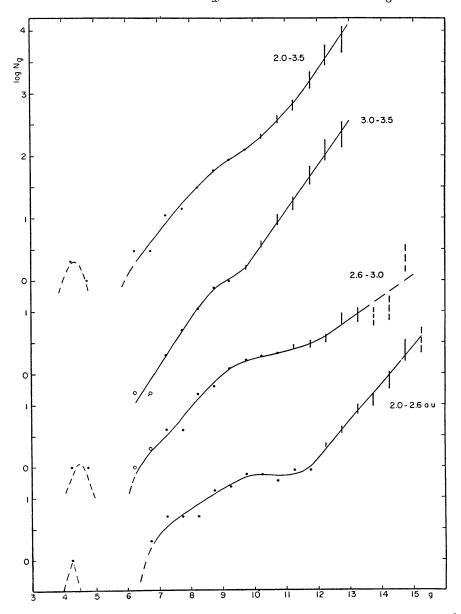


Fig. 5.—Frequency distribution of absolute magnitudes, g, for three distance zones and their sum

in the three zones and, consequently, in the distribution of asteroid dimensions. The halt in the 2.0–2.6 group between 10 < g < 12 is of special interest. In the 2.6–3.0 group there is no halt but a greatly diminished rate of increase, while in the outer group there is a discontinuity near g = 9. The asteroid dimensions corresponding to these g-values may be found on the plausible assumption that the mean albedo of the fainter asteroids is the same as that of the brightest asteroids (except for Vesta, whose albedo is known to be higher). From the values for Ceres (g = 4.00; d = 770 km) one finds 50 and 20 km, respectively, for g = 10 and 12.

The logarithmic increase in numbers for the three distance groups on the *bright* side of the discontinuities occurs with the coefficients 0.33, 0.56, and 0.75, approximately; and on the *faint* side with 0.60, 0.35, and 0.73. For the entire zone, 2.0–3.5 a.u., the average coefficient between 7 < g < 13 is found to be about 0.56, but this high value is largely due to the large numbers of faint asteroids derived for the 3.0–3.5 zone. For the 2.0–2.6 zone the average coefficient between 7 < g < 15 is 0.37. The significance of these coefficients becomes clear when it is remembered that the coefficient 0.60 would cause each magnitude interval to contribute the same total mass.

It is now clear that the curvature found in relation (6) may actually be real within a limited range of p and due to the blending of the discontinuities in $\log N_g$ of the three zones, all of which are centered on apparent magnitude 13.5–14.0. It is further clear that the simple relation (7) cannot be safely extrapolated beyond the region where it has been tested. Finally, the occurrence of the three bright asteroids (Ceres, Pallas, Vesta) seems to be a *separate* phenomenon, somewhat outside the smooth frequency-curve representing the normal asteroids. Whether the discontinuities in the frequency-curves of Figure 5 represent a division of the asteroids into two classes (possibly original condensations and collisional fragments) remains to be determined.

The question naturally arises whether the differences in the three frequency-curves are due to population differences between the Hirayama families. The distance group 2.0-2.6 comprises the families 6, 7, 8, 9 (2.15 < a < 2.316) and 5, 27, 25, 26, 24, 4, 23, 29, and 22 (2.316 < a < 2.6), according to the important investigation of Brouwer (1951); besides, a number of non-family asteroids are included. The percentage of non-family asteroids becomes much larger in the middle and outer zones; this fact, together with the large number of families present within each zone, introduces many complexities not properly belonging to this investigation.

Some comments must now be made on the asteroids of Table 7 outside the zone 2.0–3.5 a.u. They number 54, of which 51 fall into three groups: 27 from 3.5 to 4.3 a.u., 11 from 1.85 to 2.00 a.u., while 13 are Trojans. The remaining objects are Hidalgo (a = 5.79), Eros (a = 1.48), and Icarus (a = 1.08).

In Table 18 the frequency distributions of p_0 are given for $p_0 > 14.5$, the range where incompleteness increases rapidly. Table 18 may be examined from two points of view: if the distribution in g is exponential, with the same coefficient as for the asteroids at large, then Table 18 gives information on the incompleteness as a function of p_0 , simply from comparisons with the corresponding numbers for all asteroids combined. So interpreted, Table 18 gives no strong evidence that the three special groups were either favored or disfavored in the detection and orbit work, because, apart from statistical fluctuations, the distribution with magnitude is not dissimilar to that for asteroids at large. This is not true for Hidalgo ($p_0 = 19.16$), which obviously would not have been followed except for its exceptionally slow motion; and for Icarus, also discovered by Baade, which is very rarely bright enough for observation, although $p_0 = 12.35$. Alternatively, one may take for granted that the discovery probability of the three groups in Table 18 is essentially the same as for asteroids at large, and then look upon the table, or its equivalent in g, as defining, in conjunction with the f-values of Table 17, the true distribution of absolute magnitude. We shall take the second point of view (although this may not be quite correct for the Trojans) and thus derive the distributions of g found in Table 19.

Parenthetically, if this second point of view were grossly in error, the 54 objects under review should have been omitted in the derivation of f from Tables 14 and 17 and equations (6) and (7). A single object, like Icarus, does not, of course, affect the f-values noticeably; but one cannot apply the derived values of f to asteroids near the earth because of the long trails on the plates and the short periods of visibility.

Of the three groups shown in Tables 18 and 19, the 3.5 < a < 4.3 group is large enough to give fair results. The mean value of a is 3.867 a.u., corresponding to $p_0 - g = 5.22$ mag. Adopting the averages of $f(\min)$ and $f(\max)$, we find from the smoothed numbers in the third column of Table 19A the rectified numbers in the fourth column.

TABLE 18 DISTRIBUTION OF p_0 FOR SPECIAL GROUPS

p 0	N (Table 7)	N(3.9)	N(1.9)	N(5.2)	N(Sum)
14.75	. 234	3	3	0	6
15.25	. 244	5	3	2	10
15.75	. 242	5	1	3	9
16.25	. 159	6	0	6	12
16.75	. 100	3	2	2	7
17.25	. 33	1	0	0	1
17.75	. 9	0	0	0	0
18.25	. 3	0	0	0	0

 ${\it TABLE~19A}$ Observed and Rectified Numbers for a=3.9 Group and Trojans

g	N(3.9) (Obs.)	Smoothed	N(3.9) (Corr.)	N(5.2)	N(5.2) (Corr.)
8.25. 8.75. 9.25. 9.75. 10.25. 10.75. 11.25. 11.75.	2 1 0 6 5 7 3 3	1 2 2 4 5 6 4 3	1 2 2 5 10 19 (27) (64)	0 2 5 4 2 0 0	0 4 15 (25) (36) (0) (0) (0)

 $\label{eq:table 19B}$ Observed and Rectified Numbers for a=1.9 Group

g	N(1.9) (Obs.)	Smoothed	N(1.9) (Corr.)	g	N(1.9) (Obs.)	Smoothed	N(1.9) (Corr.)
11.75 12.25 12.75 13.25 13.75	1 0 0 0 4	1 0 0 1 3	1 0 0 1 4	14.25 14.75 15.25 15.75 16.25	3 0 1 0 1	2 1 1 1 1 1 2 1 2	4 3 (6) ((10)) ((40))

The log N_g -curve defined by these numbers is found to be nearly linear, with the coefficient 0.51, the same as for the asteroids of 2.0 < a < 3.5 in the same interval of $g(8\frac{1}{2}-12)$. The 3.5-4.3 group is thus average in composition; the abundance equals 3 per cent of the total asteroid ring.

If the reduction method applies to the Trojans, these may be fairly numerous, say, between 50 and 100 and centered on $g \cong 10$; but they have probably been looked for preferentially and will be less numerous than is estimated in Table 19A. The mean value of a is 5.180, and $p_0 - g = 6.67$ mag.

The 1.85 < a < 2.00 group is inadequate for statistical treatment. The results depend on the smoothing process adopted and are therefore largely arbitrary. The numbers around g = 14 are roughly $\frac{1}{2}$ or 1 per cent of those in each of the distance groups 2.0 < a < 2.6 and 2.6 < a < 3.0, but at g = 15 the fraction seems to be smaller. The mean value of a is 1.910 a.u., corresponding to $p_0 - g = 1.20$ mag.

The limiting g-value for the statistics in Table 15 is 13.0 mag. for 3.0 < a < 3.5, corresponding to $d \cong 12$ km; for 2.6 < a < 3.0 it is g = 15 mag. of $d \cong 5$ km; for 2.0 < a < 2.6 it is g = 15.5 mag. or $d \cong 4$ km. Bodies of $d \cong 1$ km, as have been observed occasionally near the earth, are as yet unknown statistically in the asteroid ring proper. It is further seen from Table 15 how the center of gravity of the asteroid ring shifts to the larger a-values for the fainter objects. The ring 3.0 < a < 3.5 contributes 23 per cent of the total (2.0 < a < 3.5) for 4.0 < g < 8.0; the percentage is 39 for 8.0 < g < 10.0; 70 for 10.0 < g < 11.0; 89 for 11.0 < g < 12.0; and 95 for 12.0 < g < 13.0.

It may not be amiss at this point to emphasize that the statistical analysis in this section was possible only because of the 1600 orbits computed and that an extension to smaller asteroids will require orbit computation for representative samples of still fainter asteroids. This aspect appears to be overlooked in current proposals to discontinue orbit computation of new asteroids.

This analysis of this section was carried out jointly by Mr. Van Houten and Mr. Kuiper.

XI. FUTURE SURVEYS

It may be necessary at some future time to carry out another systematic asteroid survey and obtain another general check and inventory of asteroid positions. In that case it is recommended that larger plates be used, so that the observing can be done in a shorter interval and hence under better sky conditions. If time permits, a third plate on each field should be taken, perhaps immediately following the second, which would remove all doubt about the reality of image pairs found in the blinking. A more rapid reduction, after the taking of the plates, would allow following up objects of special interest, like the Trojans, although one would not wish to endanger the continuity of the main program.

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	SURVEY NUMBER	Q-34.2 Q-44.1 L-23.2 L-33.1 Q-43.1	B-44.1 Q-33.5 Q-44.2 E-34.1	U-73.9 I-34.4 I-44.1 K-73.6 L-85.3	L-84.2 M-15.1 M-24.1 M-25.1 F-43.2	U-24.1 D-64.6 S-53.2 S-43.2 H-13.2	X-34.1 J-32.2 E-72.2 E-82.3 F-12.1	1-82.4 0-74.3 K-74.3 B-53.2 B-54.1
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	MAG	13.81 13.76 13.28 14.03	12.13 10.82 12.58 12.68 12.60C	11.95C 12.08C 13.26A 11.69	13.93 14.06C 14.35 12.76 13.13	12.19 11.38C 11.75	12.76 14.32 13.96 13.62 11.72	14.53 13.40 13.66C 13.47
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	E U.T.	5.1687 20.2229 12.2764 13.2375 26.2215	27.3521 30.3118 1.3069 7.2715 6.2576	2.2729 5.2583 27.1826 5.3007	18.2229 4.2236 4.3410 7.1604 23.3535	22.1688 22.1778 4.3139 5.2674 8.3792	27.2611 11.3590 11.3701 24.1937 10.3035	8.3681 29.3201 5.1569 5.1687
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	SURVEY NUMBER	C-1244 G-1424 G-153.5 C-163.2	P-33.2 P-53.2 O-52.1 N-744.5	K-35 0-84.2 N-644.3 L-64.5	B-54.3 Q-71.1 Q-81.1 H-13.3 V-63.5	W-12.1 W-13.1 Q-75.5 Q-85.1 K-82.1	L-12.2 I-33.1 I-34.1 X-44.3 H-54.2	L L L L L L L L L L L L L L L L L L L
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	MINOR	151 152 152 153	156 156 158 158	158 159 160 160 160	161 161 161 162 162	163 163 165 165	165 166 167 167 168	168 169 169 171

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	10 - MOTI	-8.7 -9.1 -11.3 -8.4	-9.6 -11.6 -11.1 -8.1	-7.0 -7.8 -9.2 -9.2	-8.0 -9.1 -9.3	-6.6 -9.2 -7.7 -8.0	-13.3 -10.1 -7.9 -9.1	15.6
	o	10.25C 9.31 10.08 9.57C 9.04C	9.65 9.36 9.47 9.63	9.43 9.61 10.65 11.06	9.39 11.05 11.48C 11.52	10.99 9.30C 9.61 8.65A 10.16	10.58C 10.52C 9.39 9.94 11.00C	8.57 8.81 8.50 9.97
	MAG	14.69C 12.44 13.20 12.58C	13.81 13.70 13.81 14.28 13.83	14.22 14.17 13.22 13.90 13.61	13.77 15.19 15.31C 15.37 12.97	12.92 13.37 13.40 12.91A 12.84	13.16C 13.12C 12.98 13.70 13.54C	13.03 13.30 13.80 14.01
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TABLE	DEC 1950.0	-22 28 04 -27 17 03 -27 17 12 -26 24 48 12 13 49	24 22 18 37 05 24 37 05 17 36 03 10 35 59 17	-14 10 11 -13 13 48 25 27 29 - 0 18 44 23 22 57	- 3 15 47 -21 30 50 -24 23 17 -26 23 16 -20 03 46	-23 25 53 24 45 04 5 17 59 15 59 53 -36 49 25	-37 55 44 -37 56 27 -26 55 10 -18:33 32 10 09 39	11 35 29 11 39 29 - 4 39 52 - 2 20 53 - 6 50 57
	R• A• 1950•0	19 21 52 1 14 02 23 1 14 01 24 4 13 37 45 0 9 57 25 0	1 19 32.0 6 28 03.1 6 28 02.3 6 04 14.1 6 03 27.5	14 14 11.6 14 00 57.4 5 15 53.6 0 18 51.8 8 03 34.3	10 36 47.1 15 09 17.5 17 39 10.4 17 39 09.6 20 32 22.3	23 27 32.0 5 23 05.3 11 07 44.1 13 23 47.7 16 05 47.7	15 40 33.9 15 39 11.1 21 55 08.0 12 42 00.9 1 54 50.6	8 49 50.4 8 49 12.3 13 08 41.5 12 51 25.7 17 52 11.3
	E U.T.	4.3486 9.3368 10.3222 2.2118 19.2208	10.2146 28.3090 28.3181 21.1431 22.1160	10.3438 28.2139 27.1736 8.1667 26.2229	6.2653 22.2639 8.2375 8.2472 3.2132	5.2535 13.2701 24.2493 9.1958 8.2931	28.1896 29.2312 1.2111 4.2514 1.1736	8.2667 9.2208 23.3535 22.1778 8.3340
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	SURVEY NUMBER	M-74.1 J-76.2 J-86.2 K-26.3 U-33.4	0-51.1 S-52.2 S-51.1 S-42.1 S-41.1	J-84.1 K-24.2 S-23.2 D-24.3	H-85.3 X-24.1 L-74.3X L-75.2 N-44.2	0-76.2 F-63.2 U-63.3 J-51.2 K-86.1	L-16.1 L-26.1 O-26.3 J-36.1 Q-34.1	H-25.2 H-34.1 V-63.2 W-13.2 L-82.1
	MINOR PLANET	171 172 172 172 173	174 174 174 174	175 175 177 178 178	179 179 180 180 182	183 184 185 185	186 186 187 188 189	190 190 190 190

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	MAG	3.88 1.62C 1.39C	5 31 2 22 3 88C 3 72 3 87	3.25 3.51 1.54 2.01C 4.71	39 34 39 31 30 31 65	2.41C 2.42 2.42 3.70 2.88C	2.89 2.74 3.02 3.76	3.64C 3.77 4.19 3.73C
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	SURVEY NUMBER	M 12.2 L-76.5 L-86.1	146 142 152 193	V-13.5 V-14.2 K-63.4 K-73.4	U-41.1 H-66.1 W-76.3 H-61.1 X-32.1	M-35.1 F-746.1 C-63.5 F-45.1	F-46.1 U-83.9 U-92.1 V-12.1 N-74.3	N-84.1 0-14.1 G-65.1 W-74.3 I-65.3
	MINOR PLANET	191 192 192 192	, 00000	195 195 196 196	197 198 198 199 199	200 200 201 201 202	202 202 202 203 203	2003 2003 204 205 5

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	SURVEY NUMBER	U-54.4 U-53.3 B-45.2 H-74.6 H-83.1	G-63.4 W-74.5 X-13.1 D-44.4 D-45.1	S-54.2 Q-45.1 Q-44.3 H-74.6 X-14.7	X C - 1 - 2 - 3 - 3 - 3 - 3 - 3 - 3 - 3 - 3 - 3	R-44.16X I-74.2 I-73.3 H-52.3 H-62.1	M-72.2X B-62.2 B-72.3 C-12.3X	S-32.3 0-21.1 U-24.2 L-71.1 M-11.1
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	E U.T.	30.3271 2.1924 1.2014 1.2687 13.2299	4.1757 4.1868 27.3528 3.2021 2.2056	9.2160 9.1785 19.2299 31.1778 24.3681	26.2708 23.1903 1.2736 1.3486 26.2056	12.2583 20.2319 3.3500 9.2889 27.2375	8.2569 27.2278 22.1688 5.1451 26.3701	23.1993 23.2174 24.2222 3.2229 14.3111
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	E U.T.	28.2868 20.2458 7.2944 18.1944 19.2028	7.19 6.17 8.23 5.24 8.20	11.2875 26.2236 13.3812 4.2118 7.2354	27.3799 28.2868 20.2458 9.3187 26.2486	23.1993 4.2201 4.2306 18.3111 28.2819	22.1250 23.1347 8.2625 8.2729 4.3174	26.3389 13.2375 23.3535 22.1688 12.2868
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TABLE	DE(195(90000	24.00.	121 121 125 125 125 125 125 125 125 125	113 113 2 174 4 24 3	-16 4 9 4 8 0 8 1 -27 1	111 223 223 200 200 200 200	-33 -16 -16 -16 -16
	A. 50.0	24.4 14.3 45.6 22.7 21.8	4 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	02.1 12.1 52.5 44.9	29.6 41.8 06.7 52.7 03.6	50.3 17.7 27.6 16.5 56.4	36.7 24.3 23.5 07.3	28.2 00.1 13.3 17.0 50.1
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	⊢	FEB OCT FEB FEB	DEC SEP OCT DEC	AUG APR APR APR	AUG AUG NOV MAY DEC	MAR PEB MAY	A P R NOV NOV DEC	FEB SEP NOV
	٥	51 52 51 51	50 50 50 51 51	52 52 52 53 54 54 54 54 54 54 54 54 54 54 54 54 54	50 50 50 50 50	52 52 52 52 51	51 50 50 51	52 51 52 51
	SURVEY NUMBER	H-55.1 D-33.1 U-55.3 G-76.3 H-16.1	S-15.5 C-53.3 C-54.2 R-56.3 S-45.2	N-75.2 N-85.2 W-36.3 W-46.1 W-45.1	B-13.1X B-23.2 Q-73.4X L-34.2 F-53.4	V-36.1 G-45.3 U-83.1 U-73.10 L-35.2	1-75.3 J-35.2 E-62.5X E-63.1 S-51.3	U-64.7X W-67.3 O-73.6 W-64.1 Q-52.2
	MINOR	665 666 667 667	668 669 669 670	671 671 672 672 672	673 673 673 674 675	675 676 676 676 676	678 678 680 680 680	681 683 685 685

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	ဖ	13.02 112.45 111.63	8 4 8 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	10.00 9.87 9.99A 9.81	10.04A 10.31 10.87C 10.32	10.34 8.10 9.63 9.49 13.94C	11.78 9.88C 12.46 11.98 13.08C	9.20 9.76 9.83 11.00
	MAG	5 5 5 6 5 6 5 6 6 6 6 6 6 6 6 6 6 6 6 6	3.44C 4.23 4.00 4.14	3.90 4.67 4.84A 3.46	4.89A 5.28 6.25C 4.26	6 4 4 6 6 7 6 9 6 9 6 9 6 9 6 9 6 9 9 9 9 9 9	4.53 4.03C 6.19 5.74 6.28C	24.00 5.00 5.46 1.56
	VAR	2005	, , , , , , ,	4 0 0 0 1 1	-41 1 -39 1 -8 1 -14 1	-14 1 -61 1 -61 1 -61 1	39 11 0 1 0 1 0 1	00676
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	E U.T.	8.1583 12.1528 6.2028 23.2083	3.33 3.33 7.370 9.211 5.198 9.211	20.2229 14.3347 31.3083 26.2556 7.2243	13.3465 31.3083 27.2535 27.2896 10.3319	2.2729 11.2083 12.1618 12.1750 30.2139	4.2090 11.2285 6.2757 29.2701 1.2479	27.3243 27.3146 27.3521 5.2674 8.1583
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	SURVEY NUMBER	G-31.2 G-41.3 G-36.3X V-23.4	35.	U-42.2 I-56.2 I-46.3 X-82.2 G-75.1	I-46.1 I-46.1 W-77.1 W-67.4 J-85.4	K-35.1 B-21.1 D-61.1 M-45.5X	K-55.5 H-64.2 M-84.6X N-13.4X H-63.3X	P-41.3 P-32.2 P-33.3 G-21.1X G-31.1
	MINOR PLANET	687 688 688 688	0 00000	693 694 694 695	696 696 696 696 701	701 702 705 705	708 709 710 710 711	712 713 713 715

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	MAG	14.78 15.86 16.05 14.87 14.07A	14.73 16.36 14.91 13.71	14.47 14.62 14.39 14.45 15.12C	15.77 14.10 13.89 14.20C 13.98	14.86 15.00 14.75 14.90	14.78C 14.97C 15.00 14.55	15.75 15.24C 13.73 14.00
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	E U.T.	22.3021 11.2722 11.2847 26.2618 14.2910	6.1958 25.2889 5.2785 24.2042 14.2576	23.2521 27.2875 25.2278 25.2361 27.2090	2.2729 7.2819 31.1826 31.2215	22.1431 22.1521 23.1528 23.1618 29.2132	29.2229 2.2507 2.2597 24.1750 24.1847	8.3889 9.2062 1.2250 26.2708 7.2611
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	SURVEY NUMBER	V-42.2 F-35.2 F-36.1 U-73.7 C-53.2	C-53.2 T-23.3X G-22.2X W-34.1 C-54.4	S-13.8 P-35.1 X-64.7 X-65.1 M-23.4X	K-35.3 N-84.2 0-15.1 0-14.3 D-45.3	S-51.2X S-52.3X S-61.1 S-62.2 M-41.2	M-42.1X P-73.1 P-74.1X Q-12.1 Q-13.1	0-52.2 0-52.2 0-55.2 0-84.5 0-84.5
•	MINOR PLANET	715 716 716 716	717 717 718 718 720	720 721 722 722 722	726 727 727 727 731	731 731 731 731	732 734 734 734	735 735 736 737

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	o	11.08 10.21C 9.90 10.17 10.08	11.47 10.60 10.72 11.22	F1.13 10.60 10.74 10.78	9.85C 9.65 9.87 13.01	11.64 11.57 10.60	10.82C 10.97 10.07C 11.23C	9.26 9.63C 9.63C
	MAG	14.61 14.37C 14.06 13.82	15.21 14.73 14.85 15.14	15.46 15.84 15.04 14.24	14.08C 14.03 14.67 15.10	15.18 13.53 15.46	15.69C 15.89 14.64C 14.48C	14.82C 13.77 14.14C
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	R• A• 1950•0	13 03 31.7 23 26 18.2 23 26 18.1 14 57 58.0 14 57 57.1	0 41 03•1 15 29 25•8 15 29 18•4 14 15 50•5 14 01 38•5	5 23 32•3 9 48 01•7 12 45 50•8 12 45 49•9 5 54 46•8	5 54 41.5 5 41 39.4 10 59 09.7 12 28 25.4 8 26 45.7	17 07 18•2 17 07 18•4 5 40 49•7 13 11 20•1 2 06 27•8	2 05 04.6 2 05 03.5 5 59 15.4 17 32 53.3 17 32 52.8	17 31 41.5 11 45 39.8 11 34 28.5 11 34 23.8 15 34 33.0
	E U.T.	9.2160 5.2437 5.2535 4.3000	30.3028 5.2208 5.3757 10.3438 28.2139	27.1917 10.2187 23.3083 23.3264	14.3111 31.1562 24.2583 2.3271 7.1604	26.3389 26.3472 14.1833 23.1722 5.1687	7.2632 7.2736 28.3271 7.2389	8.2569 15.3243 31.2236 31.3410 22.2819
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	SURVEY NUMBER	J-53.2 0-75.2 0-76.1 K-52.3 K-61.1	P-466 K-63.5 L-84.5 K-24.3	S-25.2 H-51.3X V-44.2 V-54.2 F-73.2	F-83.5 F-73.2 U-64.4 I-73.5 H-13.5	X-73.3X X-83.1 F-72.3 W-22.1 E-24.2X	E-34.8 E-35.2 S-44.3 L-65.4 L-66.2	L-76.1 1-53.3 1-43.4 1-52.1X X-33.1
	MINOR PLANET	738 739 740 740	741 742 742 743	744 746 746 746 748	748 748 748 750	752 752 753 753	755 755 755 757	757 758 758 758 758

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	ს	11.64 9.59 9.67 9.22C 9.72	12.19 11.69 9.46 9.65 13.45	10.45 10.69 11.12 10.95 11.45C	9.88 11.87 10.38 10.76C	10.24C 9.62 11.41 11.76	9.06C 9.07C 9.32C 11.09	10.22 10.03 10.53 12.55
	MAG	13.57 14.29 14.34 13.89C	15.40 14.89 14.06 14.25 14.72	15.04 14.15 14.85 14.69 14.75C	13.69 14.14 14.09 14.44C	14.45 14.52 15.10 15.43 15.02	13.06 13.02 13.27C 15.94 15.57	14.05 14.59 14.56 14.73 14.79
	VAR	14 -17 -17 -17	70077	04920	6044	1 2 4 1 2 4 1 3	1 1 1 1 0 w w 4 w	777
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TABL	DE(195(04446 0004	26 36 32 03 32 00 19 46	22 23 24 33 22 22 24 05	23 37 23 25 25 25 25 25 25 25 25 25 25 25 25 25	20 23 17 59 0 55	21 22 21 22 21 21 24 11 24 11 24	18 18 22 22 22 55 55
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	+	7 7 7 7 7 7 7 7 7 7 7 7 7 7 7 7 7 7 7 7	2 Z Z Z Z Z Z Z Z Z Z Z Z Z Z Z Z Z Z Z	MAY NOV FEB JUL	A A A B B B B B B B B B B B B B B B B B	M M M M M M M M M M M M M M M M M M M	N N N D D O O O O O O O O O O O O O O O	OCT JAN JUL DEC
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	SURVEY NUMBER	N-35-1 M-75-3 M-76-1 N-15-1	M-35.3 M-45.2 M-55.1 M-56.4 E-26.4	L-24-2 E-42-1 U-32-8 U-42-1 M-64-2	J-53.7 I-43.3 X-85.2 P-41.1 P-31.1	E-53.1 T-14.4 H-75.1 H-85.1 X-55.3	L-54.2 L-63.1 L-64.1 N-72.3 N-74.2	D-27.1 S-46.1 M-61.1 S-14.3
	MINOR PLANET	759 760 760 760 760	761 761 762 762 763	764 766 766 766	769 770 772 773	774 774 775 775	776 776 776 777	780 780 781 782 782

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	MAG	15.07C 15.25 15.33C 15.45 15.37	14.29 15.03 14.60C 14.75	14.31C 12.81 15.40 16.69 16.53C	13.54C 13.24 15.45C 14.75	14.57 15.26 15.24C 15.31 14.55C	14.90 14.90 14.46 14.74C 15.15	14.91C 14.91 14.65 14.48
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	SURVEY NUMBER	E-44.2	-54	-24.	-35.	45.	54.11	F-52.3	54.	45.	D-63.2	G-34.4X	G-45.4X	6-44.3	Xۥ96-U	7	-26	C-32.2X	-64	-83	12.	22.	98	M-17.2 M-26.2X	0-84.9	0-73.7	V-34.1X	0-44.3	B-64.4X	-	7	7	R-36.6X I-52.1	
	MINOR PLANET	1142	14	7	14	14	14	1145	14	14	14	14	1149	14	14	14	14	1153	15	15	15	15	15	1157	15	15	16	1162	16	16	16	16	1163 1164	

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	DAY	256 160 82 -33	10 10 10 10 10	14477	1 5 3 3 4 5 5 5 5 5 5 5 5 5 5 5 5 5 5 5 5	-17 26 21 21 -35	-41 66 -8 -11	23 19 19 19
	10 - DA) MOTION	1 1 2 2 4 5 5 5 5 5 5 5 5 5 5 5 5 5 5 5 5 5	0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	17. 17.8 1.8.1 1.8.1	-111.3 -111.3 -10.4	-111.0 -10.6 -9.2	-10•1 -10•1 -9•2 -9•8 -9•8	-10.7 -12.8 -12.4 -11.9
	ဖ	14.27C 11.69 10.91 10.97	10.89 11.00 10.54 9.03C	11.60 11.33C 10.09	9.62C 12.48 12.90C 13.24C 12.87	12.98C 12.34 12.34 12.27C 13.25	13.51C 13.08 10.90 10.47C	11.96C 13.57C 13.33 13.08 12.67
	MAG	16.02 15.26 15.53 15.62	15.59 16.08 15.37 15.26C 15.68	15.83 15.53C 14.52 15.44	15.35 15.04 15.98 16.41C 15.06	16.43C 15.61 15.61 15.61 15.54C	15.97C 15.29 14.51 14.07C 15.00	15.00 16.10C 15.86 15.68
	VAR	66611	0 0 1 1 7 1	00000	34 0 72	46661	1152	98888
	- C DEC	112	1 0 16 16	0000	30 16 15 68	1 1 1 1 1 2 2 2 2 2 1 1 1 1	132	44421
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TABLE	DEC 1950.0	17 18 58 22 49 51 -10 29 32 14 24 28 14 14 01	17 12 05 17 15 35 -20 18 29 - 6 35 09 -12 44 52	-21 49 09 -21 41 57 -23 19 46 -23 14 19 11 50 03	24 04 11 - 0 16 06 18 46 16 18 29 48 7 40 13	-25 37 23 -25 30 41 -25 30 42 -25 29 22 -20 45 58	-20 50 01 16 37 35 18 19 28 18 19 25 - 0 22 40	10 20 46 23 43 17 23 43 14 -26 58 45 -26 54 14
	R• A• 1950•0	11 43 22.5 11 34 09.5 14 33 44.9 2 09 09.7 2 07 42.9	6 30 16.3 6 12 44.1 17 23 21.5 19 34 13.4 20 57 21.5	15 04 07.5 15 03 22.4 16 44 50.4 16 43 58.8 3 15 01.0	6 06 52.8 11 20 23.9 2 47 55.8 2 43 30.9 11 16 43.5	14 22 48.0 14 00 56.1 14 00 55.5 13 59 50.1 22 25 29.7	22 24 31•1 11 09 14•6 3 06 45•8 3 05 51•1 22 53 17•3	10 11 10.9 9 00 21.9 9 00 21.2 15 47 04.6 15 44 39.9
	E U.T.	15.3833 31.2014 26.3049 5.1569 7.2632	28.3451 21.1340 7.2174 1.2556 3.3410	4.2090 5.2312 25.2278 26.3299 12.2687	27.2910 13.3722 8.1840 12.1500 26.2618	10.3222 2.2007 2.2118 3.2847 10.1819	11.2181 24.2403 3.3167 4.2958 11.2861	19.3210 9.1979 9.2097 22.2007 24.2028
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	_	W W W W W	տտտտտ		տտտոտ	വസസസസ	տատատ	വവവവവ
	SURVEY NUMBER	I-51.1 I-41.3X W+63.5 E-23.2 E-34.3X	S-54.4 S-44.6X L-63.6 M-71.2X N-53.4X	K-55.7 K-64.2 X-64.6 X-74.1 E-64.3	S-43.7X I-44.4 E-43.2X E-53.2 U-73.4	J-86.3X K-25.3 K-26.1 K-36.1 C-15.2	C-25.3X U-62.9 Q-63.2X Q-73.1 C-32.3	U-44.2 H-32.1X H-33.2X X-55.3X X-45.6X
	MINOR PLANET	1164 1164 1165 1167	1167 1167 1171 1172	1175 1175 1177 1177	1180 1182 1183 1183	1184 1184 1184 1184 1185	1185 1185 1186 1186 1187	1187 1188 1188 1158

	U Z	41 130		1 1 1 1	0 0	94440	0 7 7 4 7	4 4 4 4 4 4
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	10 - MOTI	-10.2 -7.0 -9.0	0 2 2 2 2	7 7 1 1 1 6 8 8 8 8 8 8 8 8 8 8 8 8 8 8 8 8	-5.7 -111.7 -9.2	11111	-11.0 -12.2 -12.1 -10.9	4
	g	11.29 10.79C 11.13 11.85	40040	11.73A 11.74A 12.41 12.72 13.36	13.34 12.75 13.30 10.54	11.12 11.30 11.87 12.24 11.73C	13.07 14.10 14.13 13.42 12.90	11.73C 11.25 11.63 11.08
	MAG	14.56 14.20 15.10 15.39	4 W W W W W	15.03A 14.90A 15.85 16.09 15.81	16.19 15.63 15.02 14.21	15.33 15.51 15.58 16.14 15.48C	15.62 15.94 15.97 15.14	15.24C 14.49 16.39 15.63
	VAR	41 11	44444	22792	5 - 74	-16 -16 -2	1441 80038	10 -8 -4 21 22
	DEC	41 11		1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	-104	110	1441	1 6 6 9 9 9 9 9 9 9 9 9 9 9 9 9 9 9 9 9
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TABLE	DEC 1950.0	7 55 38 22 54 26 21 55 57 -33 17 56	27 19 1 29 53 5 9 58 3 15 47 5 15 46 0	-15 13 05 -14 58 23 4 22 36 -13 26 14 15 26 30	15 25 17 15 26 10 -18 37 39 4 06 55 0 0	18 17 08 18 17 10 4 44 59 5 23 26 -21 01 52	18 21 08 20 46 00 20 45 58 - 1 34 36 - 4 20 10	- 2 15 09 -19 22 48 21 36 15 4 23 06 5 10 16
	R• A• 1950•0	22 41 55.8 6 33 42.9 6 12 30.4 16 06 54.0	7 52 04. 7 30 21. 9 25 16. 8 03 03.	15 14 16•4 15 11 56•0 9 36 16•3 15 52 55•6 6 26 51•4	6 07 28.7 6 06 55.9 14 50 45.9 23 14 26.0 0 0	11 46 13.3 11 46 12.4 4 35 02.5 4 20 17.9 12 42 52.0	9 59 21.33 4 02 33.44 4 02 32.7 0 31 09.0 13 19 14.3	12 49 17 9 21 24 09 1 3 50 19 7 12 01 26 6 11 40 16 7
	Б С. Т.	10.1951 28.3451 21.1340 29.2312	13.237 4.105 29.191 27.209 28.222	5.2208 8.2507 10.3139 24.2889 28.2819	21.1340 22.1250 4.2090 12.1958 15.3833	31.3410 31.3514 4.2285 22.1611 4.2514	19.22 29.17 29.18 9.18 23.35	22 • 1778 13 • 26 94 29 • 1771 26 • 3792 23 • 1903
	D A T	DEC DEC DEC DEC DEC DEC DEC		MAY 1 MAY 1 FEB 2 MAY 1 DEC	2 JAN 2 JAN 1 MAY 0 SEP 1 MAR	1 MAR 1 DEC 1 DEC 1 APR	2 FEB 1 NOV 1 NOV 0 OCT 2 MAR	2 APR 0 AUG 1 NOV 2 FEB 2 MAR
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	SURVEY NUMBER	C-21.1 S-54.5 S-44.3 L-26.2X	33523	K-63.6 K-54.2 H-55.3X X-43.4X S-55.7	S-44.5X S-45.6X K-55.1 C-42.2 I-51.6	1-52.3 1-51.4 R-56.1 R-46.3 J-36.2	U-33.10 R-33.11 R-34.5X D-35.3 V-63.3X	W-13.8X B-34.2 R-33.2 U-93.5X V-13.7X
	MINOR PLANET	1189 1189 1189 1194	100110	1200 1200 1201 1201 1203	1203 1203 1204 1206 1210	1210 1210 1211 1211 1211	1216 1218 1218 1219 1219	1219 1223 1223 1227 1227

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	10 - MOTI	1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	-111.7 -12.0 -10.7 -10.8	-10.3 -7.7 -7.7 -9.2 -8.5	1.00.00	-8 - 6 - 7 - 8 - 10 - 6 - 11 - 4 - 11 - 4	100.8	1 1 1 1 1 8 9 9 9 9 9 9 9 9 9 9 9 9 9 9
	9	11.35 11.25 11.46 11.05	12.67 12.87 12.04 11.95	10.92 11.39 11.42 10.52	10.70 12.46C 10.72C 11.10	11.17 11.02C 10.66 12.82 12.99	13.09 11.86 11.38 11.83C	11.76 10.82 12.80 11.65
	MAG	15.40 15.30 14.74 14.31	14.64 14.84 15.08 14.99 14.79A	15.57 14.86 15.50 14.72	15.30 15.68C 14.59 15.00 16.46	14.73 14.59C 14.22 15.51 15.68	15.79 15.72 15.54 15.97C 14.88	15.70 16.47 15.09 15.88 15.74
	VAR	21 15 21 21 21 21 21 21 21 21 21 21 21 21 21	20 20 0 0	-6 2 11 11 11	1102	44000	0 1 1 1 1 0 0 1 4 0 0	011000
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4	м О А	00-1 00-0 00-0 00-0 00-0 00-0 00-0 00-0	1 3 3 4 4 5 5 5 5 5 5 5 5 5 5 5 5 5 5 5 5	11000	0.0000000000000000000000000000000000000	100.7 100.3 100.3	10.7 10.8 0.8 -14.0	-13 0 0 0 4 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0
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TABI	DE(0 4 0 4 1 26 1 26 1 24	22 1 1 2 2 2 2 4 2 4 2 4 2 4 2 4 4 4 4 4	131 5	2 - 1 - 2 - 2 - 2 - 2 - 2 - 2 - 2 - 2 -	11200	115 4 18 2 19 3 18 2	17 31 - 3 11 30 2 29 22
	• o •	19.5 19.0 15.3 25.5	000 000 000 000 000 000 000 000 000 00	30 • 9 07 • 6 37 • 3 44 • 1 42 • 5	2000 2000 2000 2000 2000 2000	56.8 112.0 449.0 53.6 48.1	51.0 02.4 32.7 099.7 55.9	2355 235 235 235 235 235 235 235 235 235
	R• 195	9 31 6 94 9 33 9 53	3 32 3 32 9 00 7 25	1 40 0 20 0 49 6 15 6 15	50 46 93 004 93 034	00 00 00 00 00 4 4 00 4 4 4	9 43 7 37 3 00 1 33 5 27	5 13 1 39 7 10 4 00 3 45
	•	6 8 5 0 9	6	99 08 11	00408	108871	60460 11.11.01	96476
	E U.1	10.313 10.325 24.229 25.227 16.232	3.156 3.168 4.245 4.255 12.197	26.279 31.211 9.266 28.254 28.254	22.125 30.334 7.261 9.216 18.221	7.236 8.191 23.161 29.191 29.270	30.208 4.222 23.335 14.233 13.281	30.132 26.279 7.217 16.222 3.156
	⊢	MAAK NOK	DEC JUL JAN	FEB JUL OCT DEC	S S S S S S S S S S S S S S S S S S S	OCT COCT JUL JUL	JUL FEB MAR DEC	D C C C C C C C C C C C C C C C C C C C
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	SURVEY NUMBER	H-55.2 X-56.2 X-544.4 X-644.3 F-844.3	F-12.7 F-13.5X M-65.1X M-56.3X G-52.1	N N N N N N N N N N N N N N N N N N N	S-45.2 P-51.2X U-43.2 U-53.1 U-24.4X	D-15.1 D-26.4 S-62.3 N-12.3X	N-23.2X G-64.1 V-53.3X B-43.1 F-64.3	FG-64.3 L-63.6X L-63.3 E-82.5 F-12.3
	MINOR PLANET	1232 1232 1232 1232 1236	22333	1240 1242 1243 1243 1243	1243 1244 1245 1245 1247	1248 1248 1248 1249 1249	2225	1255 1256 1257 1258 1258

	UZ	91077	78776	21 821	41400	1 1 1 1	-10 -11 -11	10004
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	DAY	10 15 15 -28 -2	-19 30 23 31	36 41 -44 -60	70 74 57 48 15	6 15 31 41 41	- 49 - 33 - 66	5.2 5.5 1.9 66
	10 - MOTI	8 8 8 8 8 8 8 8 8 8 8 8 8 8 8 8 8 8 8	-7.7 -8.0 -7.8 -6.9	-11.1 -7.0 -11.4 -8.9	-8.5 -7.6 -18.5 -10.1	-8.3 -8.3 -9.5 -10.5	-100.5 -100.6 -100.5 -70.5	-7.8 -7.8 -7.3 -8.4
	ဖ	11.65 12.00A 11.88 10.38	13.56 10.01 10.28C 9.67 13.59	13.97 11.59 13.96 14.26C 13.14C	11.72 11.80 12.16 12.10 13.67	11.00 11.46 12.74 12.45 11.09	11.42 11.08 11.08 11.45	11.42 11.46C 12.13C 12.18 11.93
	MAG	6.00 6.35A 5.72 4.86 4.73	5.58 4.78 4.76 5.73	6.12 6.33 6.27 6.93C 5.40	0.000 0.000 0.000 0.000 0.000	5 6 6 6 6 6 6 6 6 6 6 6 6 6 6 6 6 6 6 6	5.37 5.01 5.01 4.79 3.89A	0.00 0.00 0.00 0.00 0.00 0.00 0.00 0.0
	VAR	-7 1 -8 1 -13 1 -31 1	0	-20 4 4 1	1 - 2 - 1 - 1 - 2 - 1 - 1 - 1 - 1 - 1 -	4 6 0 0 4	1111	W 4 4 H H
	- C DEC	-11 -30 -30	1122	123 133 134	4 5 1 1 1 1 1	7 7 7 7 7 7	10757	04410
∢	0 4	1110 120 130 130 130 130 130 130 130 130 130 13	0 1 1 1 1 1 5	1 4 8 8 6 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	1009 1009 1009	-1.9 -2.0 0.0 0.0	77 P P P P P P P P P P P P P P P P P P	1000
LL1	0.0	5 34 6 01 6 45 6 42	9 56 3 48 5 39 1 27	4 56 7 55 7 07 1 42 0 45	0 42 8 50 4 29 7 26	1 08 4 23 9 57 4 45	4 36 4 14 0 47 7 02 2 09	6 33 7 41 1 44 5 32
TABL	DEC 1950	13 35 20 36 20 26	24 4 2 2 3 3 4 4 5 4 5 5 6 4 5 6 5 6 6 6 6 6 6 6 6 6	23 44 28 47 27 21 20 10	2 56 - 3 16 -21 24 -15 06	-27 51 -27 24 5 19 5 20 35 14	35 14 35 16 29 02	1 1 1 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0
	• o	01110 66.00 66.000	80846 74701	4	200.3	004m0	00416 00000	ων. ωω 4 90
	R• 1950	11 2 11 2 11 2 49 2 12 0	28 4 22 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2	29 2 07 5 37 1 19 4 50 5	59 0 29 2 47 4 26 5 33 1	59 44 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	39 3 39 2 38 3 18 1 18 5	112 123 133 134 134 135 135 135 135 135 135 135 135 135 135
		9 19 9 9	22 15 15 15	9 4 4 4 6 20 4 4 4 6	7 13 16 12 13	17 17 9 9	7 7 20 6	6 10 10 13
	E C - 1	29.2465 29.3458 30.2181 29.2375 29.3549	11.2062 13.2597 19.2208 5.2208 5.2889	6.2139 2.3042 5.2625 22.1340 3.2910	7.2354 24.1951 2.2160 2.3479 8.2729	9.2986 28.2417 29.3188 29.3368 25.1896	25.1986 25.2979 26.1778 10.2639 22.1611	5.1785 24.2583 24.3403 13.3257 24.2042
	⊢	COCAN	SEP DEC MAY	JAN MAAY DEC AUG	A P B B A P B A P R A P R A P R	N N N N N	L A C A B B C C A B B C C A B C C A B C C C A B C C C C	L E B B B B B B B B B B B B B B B B B B
	۵	52 52 52 52	525	2000	525	51 52 52 52	50000	51 52 52 52 52
	SURVEY NUMBER	1-74.4X 1-84.5X N:24.3X 1-73.1 1-83.1	C-24.3 F-62.3 U-33.8 K-63.3	G-24.7 K-32.2 R-53.6X R-43.3 N-54.2X	G-76.2 W-33.1 L-44.1 I-71.1	L-85.5X M-25.3X T-75.3X T-85.1X T-21.1	1-22.3 1-31.1 1-32.1 B-13.2 S-53.2	G-15.6 U-64.8X U-65.3X G-66.2X W-34.7X
	MINOR PLANET	1258 1258 1259 1266 1266	1267 1268 1268 1269 1270	1270 1271 1273 1273 1274	1275 1275 1277 1278 1279	1280 1280 1281 1281 1282	1282 1282 1282 1284 1284	1286 1286 1286 1287 1287

	UZ	4 9 1 1 1	4000	1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	-1.9 -1.4 -1.5 -1.5	12851	80104	0767
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	DAY	50 1 0 41 -11	4 W W W 4 0 Q W Q R	187 145 164 154	-61 -2 -2 -44 -49	4 9 8 8 8 9 9 9 9 9 9 9 9 9 9 9 9 9 9 9	136 136 132	144 129 116 15
	10 - MOTIO	-8.1 -8.2 -8.6 -10.0	10.2 17.5 17.5	11 8 1 1 1 8 1 1 1 1 1 1 1 1 1 1 1 1 1	-10.2 -9.7 -8.1	0.000	-6.6 -8.7 -7.0 -7.0 -10.0	-10.6 -10.1 -10.8 -10.7
	g	11.46C 11.44 11.20C 11.57 12.49	11.63A 11.59 11.96	12.99 11.97 12.42 10.43	11.64C 11.43 11.48 10.67	10.50 13.45C 12.09C 11.51 10.95	11.61 12.08 11.03 11.49	11.29 11.23 11.48 13.23 11.57
	MAG	14.84 15.70 15.46C 16.10 15.37	14.53A 14.50 16.58	15.92 15.56 15.90 14.17 13.84C	15.69 15.45 15.50 15.39	15.21 15.66C 15.61 15.07 15.14	15.75 16.36 15.11 14.54 15.65	15.08 14.85 15.14 16.29 13.87
	VAR	12 2 9 9	16221	0 in 0 m m	120	-19 -07 -11	-11 -1 104 62 -2	-29 -29 -28 10
	- C DEC	1100011	1 1 1 1	7 4 7 0 1	-4 0 0 -17	111111111111111111111111111111111111111	-15 113 62 1	1 1 1 4 4 4 0 6 4 4 4 0 6
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TABLE	DEC 1950.0	- 6 48 59 10 43 11 10 43 10 - 7 12 43 23 11 22	-19 40 58 29 56 16 29 56 39 8 44 41 -16 40 32	- 4 37 18 -20 15 44 -21 20 16 -22 55 29	4 52 04 25 02 54 25 02 58 -32 32 27 -32 32 22	-32 28 03 - 0 30 43 - 8 03 02 - 7 52 48 - 7 46 48	- 7 32 01 21 38 53 - 1 14 56 3 33 46 34 08 51	-27 10 15 -28 17 56 -28 23 00 -23 22 15 -20 45 30
	R• A• 1950•0	13 05 55.2 6 52 41.5 6 52 41.0 11 35 03.9 6 03 36.7	14 56 38.6 7 20 54.9 7 20 49.1 9 59 03.3 13 19 16.7	23 44 38 1 23 20 36 4 5 25 04 1 17 49 19 9 17 33 56 9	1 10 25.9 5 58 26.5 5 58 24.8 14 48 38.5 14 48 38.7	14 47 52.9 10 55 31.8 12 46 34.5 12 44 04.9 12 31 16.7	12 29 53.3 9 43 06.0 21 57 43.5 23 12 23.8 5 07 55.7	16 05 18.1 15 45 51.7 15 43 49.2 16 31 38.9 22 15 13.4
	Е U.1.	9.2757 8.2028 8.2833 26.2799 5.1562	27.3708 25.1986 25.2889 11.2285 23.2625	27.1882 14.2799 27.1736 9.3576 26.2118	10.2917 27.2639 27.2910 3.2958 3.3062	4.2306 24.2583 4.3174 7.2715 2.2736	4.3174 18.2035 15.2368 12.1958 23.2611	8.2826 27.2312 29.2208 24.2292 1.2687
	A T	APP CAPN CABN	A P P P P P P P P P P P P P P P P P P P	SEP SEP CUN	OCT DEC MAY	MAPR APR APR APR	APR FEB AUG SEP DEC	MAY MAY MAY SEP
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TABLE	DEC 1950.0	15 52 35 24 28 36 - 7 10 08 - 6 11 34 - 1 15 17	0 50 23 0 46 08 0 45 59 21 55 39	-15 34 17 -15 34 53 24 00 47 7 55 37 -27 42 53	-27 42 06 -15 03 49 20 42 39 8 44 00 8 43 04	30 19 02 30 02 03 7 23 48 7 24 01 7 29 43	-17 33 43 -25 18 29 -25 18 18 29 24 09 21 24 34	21 23 24 36 36 38 16 00 04 -30 14 59 8 50 13
	R• A• 1950•0	9 54 40.4 1 01 00.4 15 47 34.0 15 27 14.1 11 57 39.2	11 34 46.7 0 40 32.3 23 26 45.3 23 26 44.7 0 55 21.7	22 53 47.7 22 52 57.1 0 49 31.0 9 55 56.9 14 42 54.7	14 42 50.9 11 32 45.7 0 55 49.2 10 30 36.1 10 29 37.3	4 58 16.4 4 45 55.5 13 27 19.1 13 27 17.7 13 26 40.6	22 52 30.5 13 27 00.5 13 26 59.6 3 48 43.1 1 26 25.5	1 24 37.8 7 32 44.6 9 43 05.6 14 56 53.9 2 42 32.3
	В U•T•	1.2479 30.3340 5.3861 27.2611 26.3701	23.1993 30.2215 3.3132 3.3319 30.3340	11.3076 12.1701 9.2660 19.2299 27.2347	27.3167 1.2299 9.2660 19.3210 20.2410	5.2896 22.2424 8.3556 8.3889	2.2340 23.2535 23.2806 29.2764 10.2146	12.1750 25.1896 10.2931 20.1854 3.2083
	D A T	51 MAR 51 SEP 51 MAY 51 MAY 52 FEB	52 MAR 51 SEP 51 SEP 51 SEP 51 SEP	50 SEP 50 SEP 50 OCT 52 FEB 52 APR	52 APR 51 APR 50 OCT 52 FEB 52 FEB	51 DEC 51 DEC 51 APR 51 APR 51 APR	51 SEP 52 APR 52 APR 51 NOV 50 OCT	50 OCT 52 JAN 51 FEB 52 MAY 51 NOV
	SURVEY NUMBER	H-63.3 P-51.1 K-72.1 L-12.1 U-94.9X	V-14.4X P-44.3 O-62.4X O-73.3	C-34.4 C-35.1X D-41.5 U-34.2	W-76.1 I-56.1 D-41.6 U-44.9 U-54.9X	R-62.6X C-52.8 C-61.1 C-62.7	0-55.4X W-26.3X W-36.5X R-32.1 D-51.6X	D-62.5 T-21.2X H-53.5 X-16.1 Q-54.5X
	MINOR PLANET	1487 1490 1492 1492 1493	1493 1494 1496 1496	1501 1501 1503 1503	1505 1506 1510 1510	1512 1512 1516 1516 1516	1518 1520 1520 1523 1524	1524 1524 1532 1532 1534

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	DAY ON	-13 -46 -40 -27	122 100 4	1117777777	-111 6 -71 -59	138 165 147 145	-19 -33 19 14	25 1 4 2 1 6 2 1 5 5
	10 - [MOT-10	-10.9 -10.8 -6.9 -7.6	-10.1 -7.3 -8.4	-11.2 -11.2 -10.6 -8.6	1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	-111-7 -7.2 -7.9 -6.9	-11.3 -7.4 -10.2 -6.2 -8.8	17.6
	g	12.81 12.82 12.20 12.21C 11.99C	12.56 12.60 11.51C 11.56	12.68 12.80C 13.00C 12.65	13.56 13.65 12.63 12.48 12.80	12.17C 12.42 11.49C 11.70 10.95	12.82C 12.72 10.65C 13.22 12.50	11.13 11.35 11.55 9.80 13.11
	MAG	15.06 15.07 15.33 15.12C 15.50	16.29 16.33 15.72C 15.76	15.73 15.81C 16.01C 14.53	16.05 16.60 15.85 14.24 14.55	15.65 16.41 15.67C 15.90 15.17	15.31 15.78 15.00C 16.89 16.32	16.14 15.14 16.34 16.29 15.28
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TABLE	DEC 1950.0	8 50 08 8 50 09 8 05 57 6 22 02 - 5 24 17	29 29 46 29 29 46 -20 15 33 -20 15 12 -20 03 23	-25 09 58 -25 10 36 -25 10 37 13 09 20 13 09 24	17 36 55 17 24 47 -12 20 13 25 12 02 25 01 50	- 5 38 28 38 35 26 -14 53 51 -16 57 39 -17 02 33	25 48 56 24 45 10 32 35 25 10 39 21 -12 25 39	10 35 47 23 28 50 7 43 06 20 50 13 -13 48 57
	R• A• 1950•0	2 42 32.0 2 42 27.3 1 41 52.1 1 26 05.8 23 25 57.3	4 41 38.4 4 41 38.4 17 59.29.0 17 58 40.4 17 44 40.7	17 29 31.9 17 28 27.1 17 28 26.5 10 50 54.9 10 50 54.4	4 28 25.2 4 06 02.2 16 17 11.7 2 06 05.7 2 04 20.9	23 23 17•3 6 03 43•8 23 55 07•3 23 40 40•9 23 39 57•0	6 29 00.9 6 06 38.7 4 43 05.4 13 11 04.3 14 18 12.6	11 28 33.3 5 31 52.2 10 59 04.0 0 57 04.7 23 26 51.9
	E U.1.	3.2174 3.2986 2.2597 24.1937 14.2576	22.2333 22.2424 8.3243 9.3576 26.2118	7.2278 8.2375 8.2472 24.2312 24.2403	29.2854 22.1431 24.3146 5.1451 7.1986	14.2576 27.3090 5.3201 26.1917 27.2604	28.3000 21.1250 12.1687 23.1632 2.2833	26.2618 13.2701 24.2493 9.2660 14.2687
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	SURVEY NUMBER	Q-55.5X Q-65.1 P-74.2X Q-14.1X C-54.2X	R-53.3 L-84.3 M-14.3	L-64.5 L-74.1 L-75.3X U-53.5 U-62.1	R-44.9 R-44.2 X-52.5X E-22.1 E-32.1	C-54.3 S-41.2 O-85.2 P-15.3X P-26.1X	S-53.8 S-43.3 F-42.3 W-21.1X K-34.3	U-73.8 F-63.5 U-63.9X D-41.4 C-55.3
	MINOR PLANET	1534 1534 1539 1539 1541	1541 1541 1542 1542 1542	1544 1544 1544 1545 1545	1551 1551 1554 1555 1555	1557 1557 1558 1558 1558	1560 1560 1567 1569 1570	1572 1578 1578 1583

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	ŋ	13.33 13.19 13.10 13.22 12.62	12.98C 12.72 14.57 12.13	13.67 13.84A 12.10 11.99C	1.67 1.67 1.61 1.61 1.02	12.70A 12.90 13.47C 12.91 12.82 12.63 12.647 13.66	12.06 11.62 12.94 12.74 11.33
	MAG	16.24 16.22 16.13 15.36 15.62	14.68C 14.47 15.53 15.45 16.16	15.23 15.46A 15.76 15.83C 15.78	0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	13.82A 13.97 14.26C 13.70 15.63 16.63 14.63	13.90 13.56 15.73 15.59
	VAR	4 5 2 2 7	1	00 00	20 21 21 1	77	
	- C DEC	6 4 11 7	7	11 66	12 15 15	00	
4	0 K	11.2	-0.2	000	1122	0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	
TABLE	DEC 1950.0	21 01 04 22 40 54 22 41 08 5 32 00 13 34 51	- 2 56 24 - 5 44 42 - 3 48 31 12 44 52 25 03 13	-14 14 21 -14 14 19 14 27 05 - 8 49 18 - 8 49 17	24 34 34 10 42 25 10 42 2 2 4 4 4 4 4 4 4 4 4 4 4 4 4 4 4 4	- 0 12 29 - 0 26 44 -17 48 54 -17 49 05 -22 08 52 17 32 42 19 16 59 - 8 20 39	-23 45 07 -29 56 39 8 22 09 7 23 28 4 23 48
	R. A. 1950.0	9 33 04.2 9 16 10.2 9 16 09.8 0 00 10.7 8 00 07.7	16 37 06.8 16 38 58.2 16 32 03.2 2 15 52.5 9 17 56.1	15 11 32•1 15 09 28•9 10 50 50•9 16 10 02•5 16 10 02•1	53 15- 10 39- 11 50 34- 9 53 23- 9 53 23- 5 20 15-	21 46 46.9 22 41 11.2 22 41 10.8 22 30 12.9 9 37 52.0 9 22 54.1 23 04 41.5	19 52 45.1 19 32 25.3 23 56 39.5 23 38 52.5 6 44 30.3
	E U.T.	31.3493 17.1979 17.2069 7.1604 28.1903	2.3056 2.3056 2.3056 7.2632 31.3306	20.2458 22.2729 6.1937 24.2979 24.3146	3.237 6.361 4.196 4.205 1.228 1.290 0.281	15.2819 15.2368 19.3403 19.3514 11.2181 17.2069 3.2285	6.2757 29.2201 17.2076 6.1847 6.2028
	D A T	SZ FEB SZ FEB SZ FEB SO OCT	COUNTY OF THE CO	MAY MAY MAY MAY MAY	10001 100	AUG O AUG O AUG O SEP 12 FEB 13 SEP	JUL JUL SEP O OCT
	SURVEY	T-92.1 5 U-12.2 5 U-13.7X 5 D-12.5 5 T-45.5 5	L-41.3 L-41.4 E-41.2 E-34.4 T-82.3	X X -114.6 X -123.2 X -1483.5 X -152.0 X -52.0 S -52.0 S -50.0	1	B-75.1 B-75.1 B-75.1 C-25.1 C-25.1 C-25.1 C-25.1 C-25.1 C-25.1 C-64.1	M-84.1 N-15.2 C-62.3 C-52.3 G-36.1
	MINOR	1589 1589 1589 1590 1590	1591 1592 1593 1599 1599	1601 1601 1603 1603	0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	1606 1607 1607 1607 1607 1607	1609 1609 1613 1613 1614

	O - C MOTION					
	DAY ON	61 -39 -42 -29 -97	-83 -15 -15 -37	130 135 135 122	11 - 67 - 78 - 78 - 7 - 7 - 7 - 7 - 7 - 7 - 7 - 7 - 7 - 7	1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1
	10 - 1 MOT I	-6.8 -7.7 -7.8 -10.0	1 0 1 1 1 1 0 4 4 4 0	-5.7 -5.7 -9.9 -10.0	1100 1100 1100 1000 1000 1000 1000 100	-6.5 -7.0 -7.0 -7.0 -6.9 -7.4 -10.0 -10.0
	ŋ	11.76			14.02	12.720
	MAG	15.57 15.65A 15.50A 15.36 15.36	15.46 16.06 15.59¢ 15.50C	15.18C 15.19 15.29A 15.42 15.50C	15.26 15.84 15.84 15.64 14.91 15.94 15.29 15.23	15.37 15.29 15.83 15.70 15.29 15.534 15.66C
	VAR					
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TABLE	DEC 1950.0	5 46 38 -14 54 35 -14 54 42 -16 42 00 - 7 59 16	-11 05 32 -11 05 56 -11 14 10 -11 14 19 -11 16 51	-11 17 00 -12 13 12 - 0 30 31 0 25 27 -14 25 46	-14 44 21 - 1 14 59 - 4 16 03 - 6 16 12 - 0 25 39 0 11 53 1 40 42 -12 19 23 - 6 10 39 - 10 50 47	-10 50 52 - 4 01 12 2 01 18 2 00 23 1 41 34 -18 01 14 -22 31 44 22 29 26 2 48 07
	R• A• 1950•0	13 32 12.3 21 15 37.5 21 15 36.9 21 52 57.8 22 20 29.3	22 06 21•8 22 06 20•3 22 26 00•9 22 26 00•3 22 25 25•9	22 25 25 55 5 25 5 2 2 2 3 2 2 3 2 0 0 7 2 2 3 1 5 6 6 6	22 11 49.8 22 19 31.5 22 08 19.9 22 08 19.6 22 32 01.5 22 10 37.7 22 56 27.7 23 27 06.2 22 35 40.1 22 38 57.5	22 38 57.3 22 49 10.3 23 23 34.1 23 23 29.9 23 22 10.7 23 12 33.1 23 39 19.0 0 46 04.6 0 45 04.7 23 47 20.3
	Б О•Т•	24.1861 13.2556 13.2694 18.2229	9.2514 9.2965 18.3243 18.3389	19.3403 9.2965 19.2722 9.2396	9.2965 18.3111 9.2396 9.2514 19.2722 9.2396 11.2861 11.2861 11.944	11.2062 11.2965 12.1958 12.2986 12.2986 14.2910 14.2910 8.2819 9.2660
	D A T	2 APR 0 AUG 0 AUG 0 AUG 0 AUG	O SEP O SEP O AUG O AUG	O AUG O SEP O AUG O SEP O AUG	SEP O SEP O SEP O SEP O SEP O SEP O SEP	o sep o sep o sep o sep o sep o oct
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	SURVEY NUMBER	W-32.3X B-33.3 B-34.1X B-54.5 B-63.1	C-13.4 C-14.4X B-63.2 B-64.1 B-73.4	B-74.1 C-14.6 B-72.1 C-12.5	C-14.5 B-62.3 C-12.4 C-13.5X B-72.4 C-12.1 C-55.2X C-23.1	C-24.4X C-33.1X C-52.5X C-52.2 C-53.1 C-53.1 C-53.1 C-53.1 C-63.1
	MINOR PLANET	1614 1950PZ 1950PZ 1950QA1 1950QB1	1950QB1 1950QB1 1950QC1 1950QC1 1950QC1	1950QC1 1950QC1 1950QD1 1950QD1 1950QE1	1950QE1 1950QF1 1950QF1 1950QG1 1950QG1 1950RF 1950RZ 1950RZ	1950RA1 1950RB1 1950RC1 1950RC1 1950RC1 1950RE1 1950SL 1950SL

TABLE A	R• A• DEC O - C MAG G 10 - DAY 1950•0 1950•0 R A DEC VAR MOTION	34 29.5 - 0 00 16 15.35A -6.2 -74 39 27.7 0 34 09 15.19A -9.0 -29 23 53.3 - 0 00 13 15.43 -8.5 -10 55 57.7 0 12 26 16.29A -8.9 -5 44 57.8 - 8 22 43 15.97 -9.3 -57	46 59.8 14 41 19 15.09 -9.1 -16 55 52.1 13 51 22 15.26 -9.8 -27 42 44.9 13 21 56 15.34 -10.5 -27 36 59.8 1 36 48 15.33 -7.0 -89 54 44.3 0 47 17 14.19A -7.7 -9.1 -16	39 32.0 - 1 48 48 15.07 -7.2 -53 25 24.1 -14 43 19 14.68 -7.3 -26 35 56.2 19 19 12 15.52 -10.0 -24 35 55.5 19 19 11 15.75 -10.8 -20 59 05.7 26 11 43 16.05 -7.8 -83	57 49.2 6 42 47 15.69 -8.1 -45 57 48.5 6 42 37 15.47 -7.8 -44 40 20.1 5 08 04 15.64 -7.9 -42 40 14.3 16 29 40 16.40 -9.3 -33 43 20.1 16 55 09 15.66 -7.4 -7.4 -78	27 53.1 14 01 40 15.76 -7.5 -89 44 29.5 18 00 32 16.03 -10.6 -6 00 05.3 9 11 52 15.52 -9.3 -81 40 48.7 6 33 05 15.99 -8.8 -68 02 03.1 -18 17 30 15.53 -5.8 -5.8	01 22.7 -18 23 55 15.72 -6.8 -67 15 14.80A -7.5 -11 15.67 6 50 12 15.34 -8.5 -97 41 21.5 6 38 54 14.0 22.0 6 30 39 14.0 23	39 41.5 6 22 56 14.62 -8.2 -89 07 17.3 24 55 51 15.88 -8.6 -64 06 31.4 24 48 01 15.54 -8.3 -80 07 11.7 4 35 00 15.94 -7.3 -33
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ď	0 - C A DE							
	DEC 950.	0 00 1 0 34 0 0 00 1 0 12 2 8 22 4	4 41 1 2 3 51 2 4 4 1 1 4 4 4 4 4 1 1 4 4 4 4 1 1 1 1	1 48 4 14 43 1 19 19 1 19 19 1 26 11 4	6 4 4 4 6 6 6 6 6 6 6 6 6 6 6 6 6 6 6 6	14 01 4 18 00 3 9 11 5 6 33 0 18 17 3	18 23 5 14 50 1 6 50 1 6 38 5 6 30 3	6 22 5 4 55 5 4 48 0 4 35 0
	• A	4 29 27 3 5 5 7 5 5 7 5 7 5 7 5 7 5 7 5 7 5 7	6256 6256 6366 6366 6366	0 11 11 0 0 12 12 0 0 4 0 12 0 0 4 0 12 0	7 49. 7 48. 0 20. 0 14. 3 20.	7 53 • 6 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	1 22. 5 14. 1 56. 1 21. 0 22.	9 41. 7 17. 6 31. 7 11.
	E U.T.	6.1958 17.2187 6.1958 17.2187 17.2312	9.2764 9.2764 8.2931 12.2667 13.2299	4.1868 9.2111 2.1660 2.1764 13.1972	13.2194 13.2299 4.1757 12.1972 12.1972	2.1764 12.1972 13.2194 2.1882 7.2472	8.2028 8.1917 8.3042 9.1785	9.1785 9.2660 10.2146 9.2951
	DAT	50 OCT 50 SEP 50 OCT 50 SEP 50 SEP	50 OCT 50 OCT 50 OCT 50 OCT 50 OCT	50 NOV 50 OCT 50 NOV 50 NOV	50 OCT 50 OCT 50 NOV 50 OCT 50 OCT	50 NOV 50 OCT 50 OCT 50 NOV 50 OCT	50 OCT 50 OCT 50 OCT 50 OCT 50 OCT	50 0CT 50 0CT 50 0CT 50 0CT
	SURVEY NUMBER	C-53.4X C-63.3 C-53.6 C-63.4 C-64.5	D-42.3 D-42.4 D-32.2 D-65.3 D-75.4X	E-16.3 D-37.3 E-12.1 E-13.1 D-72.1	D-74.1X D-75.1 E-15.4X D-63.4	E-13.2 D-63.6 D-74.3 E-14.1 D-16.2x	D-27.2 D-26.3 D-33.4 D-34.1	D-34.5X D-41.1 D-51.5X D-44.2
	MINOR PLANET	1950SS 1950ST 1950ST 1950SU 1950SV	1950TC 1950TD 1950TE 1950TF2 1950TH2	1950TH2 1950TP2 1950TS3 1950TS3 1950TT3	19501V3 19501V3 19501V3 19501Y3 19501Z3	1950TZ3 1950TA4 1950TB4 1950TB4 1950TC4	1950TC4 1950TD4 1950TE4 1950TE4 1950TF4	1950TF4 1950TG4 1950TG4 1950TH4

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	ŋ	3.47	13.20C 12.98
	MAG	0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	16.227 15.659 15.659 15.659 15.559 15.500 15.75A 15.75A 15.75A 16.25 16.25 16.25 16.25 16.25 16.25 16.25 16.25 16.25 16.25
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	R• 195	MMH44 44404 00 MWW 0	2000 0000 00000 00000 00000 00000 00000 0000
	E U.1.	00.210 00.280 00.302 20.152 20.152 17.263 17	12.2042 16.2222 3.1569 16.2222 3.1569 3.1569 3.1569 16.2222 3.1681 3.1569 16.2660 10.1597 14.1937 14.1937
	⊢		ECC CCC CCCC CCCC CCCC CCCCC CCCCC CCCCC CCCC
	۵	00000 00000 00000 0	
	SURVEY NUMBER	5 7 7 7 7 7 7 7 7 7 7 7 7 7 7 7 7 7 7 7	E - 54 E - 64 E - 64 E - 64 E - 18 E - 18
	MINOR PLANET	100000 000000	1950VP 1950VP 1950VE 1950WC 1950WC 1950WC 1950WE 1950WE 1950WF 1950XM 1950XM

TABLE

O - C MOTION		ر د د	•		
10 - DAY MOTION	9.8 4.4 - 20 4.4 - 18 9.4 1	1006 - 22 1006 - 22 1900 - 28 -900 - 28 1203 - 11	8.2 7 7 8.2 7 7 8.5 8.5 8.5 8.5 8.5 8.5 8.5 8.5 8.5 8.5	0.2 -1 7.8 1 8.1 2 0.1 -4	12.6 15.7 17.9 15.9 16.0 17.9 17.9 17.9 17.9 17.9 17.9 17.9 17.9
ŋ	1 1	11 11	1 11	1 1	1 1 1
MAG	5 5 5 8 8 5 8 8 8 8 8 8 8 8 8 8 8 8 8 8	15.07 15.07 15.07 15.07 15.05	0000 000 000 000 000 000 000 000 000 0	7.0000	15.14 15.16 15.76 15.23 15.12 14.98C 15.48
VAR		7			
- C DEC		7	o †		
м О А					
DEC 1950.0	12939 12930 13931 13934 14934	28 17 40 27 32 96 27 14 37 23 45 00 5 36 17 20 35 29 20 35 29	2 2 3 3 4 5 6 3 5 6 3 5 6 5 6 5 6 6 6 6 6 6 6 6 6	25 1 2 2 2 2 2 2 3 4 4 5 3 4 4 5 3 4 4 5 3 4 4 5 3 4 4 5 3 4 4 5 3 4 4 5 3 4 4 5 3 4 4 5 3 4 4 5 3 4 4 5 5 3 4 4 5 5 3 4 4 5 5 3 4 5 5 5 6 6 6 6 6 6 6 6 6 6 6 6 6 6 6 6	20 35 13 15 29 44 15 29 46 30 43 00 23 33 32 23 33 34 16 36 52
R. A. 1950.0	08 17. 07 29. 01 49. 00 38. 12 20.	5 5 5 5 5 5 5 5 5 5 5 5 5 5 5 5 5 5 5	40 15 44 55 44 55 44 21 59 20	57 03. 40 50. 51 14. 51 10.	6 05 07 00 10 00 00 00 00 00 00 00 00 00 00 00
E U.T.	2.310 2.154 2.310 3.154 3.270	14.3111 31.1562 5.1562 31.1562 31.1562 8.1917 8.2951	200 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	.121 .237 .202 .283 .133	5.1674 5.1674 5.1674 5.2785 5.2889 6.2139 6.1806
⊢	DEC DEC DEC DEC	LUD DUDUE AAR EARCON	HOLOU CO	L C A B B B B B B B B B B B B B B B B B B	Z Z Z Z Z Z Z Z Z Z Z Z Z Z Z Z Z Z Z
۵	00000				
SURVEY NUMBER	0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	F-183.2 G-173.7 F-78.3 F-78.3 G-18.0 G-44.5 G-44.5	G-64. G-64. G-64. G-64.	45834	6-11-5-3 6-11-6-4 6-11-6-4 6-12-3-3 6-23-3-4 6-23-4 6-25-22 6-25-22
MINOR	99999 000000	1950XS 1950XS 1950XS 1950XA 1950YA 1951AE 1951AE	9518 9518 9518 9518	951AM 951AP 951AB 951AB	1951AF1 1951AF1 1951AF1 1951AG1 1951AG1 1951AH1 1951AH1 1951AH1

Identical with 1618

	O - C MOTION							
	DAY ON	144 145 113		11 11 36 36	-13 -15 77 75	12001	25 1 2 2 4 1 2 2 4 2 2 2 2 2 2 2 2 2 2 2 2	10 10 10 10 10 10 10 10 10 10 10 10 10 1
	10 - DA) MOTION	-10.3 -10.6 -11.0 -10.9	9 0 0 0	1000	110.0	-10.9 -10.9 -10.4 -10.6	-10.6 -9.0 -111.7	-9.7 -8.9 -10.0 -9.1
	ŋ							
	MAG	15.68 15.33C 16.27 15.07	0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	16.33 16.07 16.46 15.29	16.00C 15.99 15.97 15.88C 15.65A	15.69C 15.77 15.70 14.79	15.01 15.64 15.40C 15.09 15.10	15.70 14.84 15.18 15.70 15.80
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	MINOR PLANET	1951RJ1 1951RJ1 1951RM1 1951RQ1 1951RQ1	1951RW1 1951RW1 1951RX1 1951RX1 1951RX1	1951RY1 1951RZ1 1951SB 1951SD 1951SW	1951VA 1951VE 1951VF 1951VG 1951WG	1951WL 1951WM 1951WQ 1951WT 1951WU	1951WV 1951WW 1951WX 1951WX 1951WX	1951WY 1951WZ 1951WA1 1951WA1 1951WB1

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	MINOR PLANET	1951WB1 1951WC1 1951WC1 1951WD1 1951WE1	1951WE1 1951WF1 1951WF1 1951WG1 1951WH1	1951WJI 1951WJI 1951WKI 1951WKI 1951WMI	1951WN1 1951W01 1951WP1 1951WQ1 1951WQ1	1951WR1 1951WR1 1951WS1 1951WT1 1951WU1	1951WU1 1951WV1 1951WW1 1951WW1 1951WX1	1951WX1 1951WY1 1951WZ1 1951WZ1 1951WA2

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	MINOR	1951WB2 1951WC2 1951WD2 1951WE2 1951WF2	1951WG2 1951WH2 1951WC2 1951WC2 1951WC2 1951WM2	951X 951X	1951XL 1951XM 1951XN 1951XN 1951XO	951X 951X 951X 951X 951X	1951XR 1951XS 1951XT 1951XU 1951XV	1951XV 1951XW 1951XW 1951XX 1951XX

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	10 - MOTI	-11.6 -7.2 -9.5 -9.5	-8.9 -13.0 -12.1	-7.4 -11.7 -11.2 -9.2	8 8 6 6 9 8 8 6 6 9 1 1 1 1 1	1100 1100 1100 100 100 100 100	110000	19.6 19.8 10.8 10.8
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	M A G	15.91 16.60 16.72 17.09	15.21 16.64 15.58 16.28 16.19	16.42 16.43 15.59 16.53	16.56 16.76 16.75 16.69 16.10C	16.88 16.79 17.09 17.00	16.64C 16.74C 16.93C 16.60 16.63	16.53 16.54C 15.39C 16.43 17.54
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TABLE A	DEC 1950.0	20 13 20 18 47 59 18 40 04 19 12 03 22 19 29	23 10 05 12 26 13 35 26 18 36 13 42 36 13 44	34 48 45 35 18 23 33 42 27 34 39 28 29 12 26	29 11 41 29 11 39 33 52 57 24 25 37	24 28 20 24 16 45 24 17 28 19 08 52 19 05 39	11 35 09 11 03 59 31 04 38 28 23 30 21 29 36	16 37 13 16 37 10 20 46 00 18 47 19 21 55 56
	R• A• 1950•0	4 46 45.3 4 26 10.2 4 47 51.7 4 48 15.1 4 38 16.1	4 22 10•1 4 51 13•0 5 06 37•7 4 45 42•1 4 45 42•1	5 11 56.4 4 58 31.5 5 01 18.1 4 41 50.7 4 26 40.8	4 26 36.5 4 26 36.1 4 26 22.0 4 26 20.3 4 10 27.5	4 28 19.3 4 28 19.3 4 28 13.3 4 30 06.1 4 29 19.1	4 13 16.5 4 14 22.1 4 38 49.2 4 33 42.7 4 32 37.3	4 42 01.5 4 42 01.1 4 43 52.7 4 49 21.7 4 51 51.7
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	SURVEY NUMBER	R-54.10 R-44.11 R-54.11 R-54.12 R-54.14	R-43.4 R-55.5 R-62.3X R-52.7 R-51.1	R-62.5X S-12.2 R-62.7X R-52.6	R-42.2 R-53.5X R-42.3X R-52.1	R-43.5 R-53.67 R-53.67 R-54.15X R-54.1	RR-1-4 RR-1-4 RR-1-45 F-1-7 F-	R - 5 - 5 - 5 - 5 - 5 - 5 - 5 - 5 - 5 -
	MINOR PLANET	1951XY 1951XY 1951XZ 1951XA1 1951XA1	1951XB1 1951XC1 1951XD1 1951XD1 1951XD1	1951XE1 1951XE1 1951XF1 1951XF1 1951YB	1951YB 1951YB 1951YC 1951YC 1951YD	1951YE 1951YF 1951YF 1951YG 1951YG	1951YH 1951YJ 1951YK 1951YL 1951YM	1951YN 1951YN 1951YO 1951YP 1951YP

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¥	O - C R A DEC VA							
TABLE	DEC 1950.0	21 55 49 9 57 51 9 57 48 8 33 41 26 18 57	24 11 42 22 33 39 22 33 41 22 40 14 22 40 18	20 05 59 14 54 34 7 29 04 7 38 05 5 29 33	42)8 12 38 32 42 33 47 41 33 47 25 24 19 13	22 45 07 27 34 40 25 20 20 25 20 22 20 40 07	15 23 42 19 59 26 15 27 47 6 23 25 2 28 25	36 43 38 32 20 06 28 46 55 21 15 21 21 15 16
	R. A. 1950.0	4 51 45.1 4 36 12.6 4 36 12.0 4 39 06.4 4 50 41.9	5 07 29•7 5 10 33•6 5 10 33•1 4 54 10•6 4 54 10•3	5 02 08 0 5 01 18 5 4 52 04 3 4 52 07 6 11 09 5	5 16 07 3 5 30 45 6 5 33 59 0 5 33 51 8 5 9 0	5 19 14.5 5 24 06.5 5 33 05.3 5 33 00.1 14 29.3	5 24 09.3 5 28 02.2 5 32 04.2 5 10 26.7 5 11.9	5 39 22.3 5 45 30.3 5 49 40.0 5 37 40.1 5 36 53.9
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	SURVEY	S-14.1 R-55.5X R-56.1 R-56.2 S-13.2	S-13.5 S-14.10X S-13.7 S-14.2 S+13.11X	S-14.6 S-15.3 S-16.1 S-16.2 S-16.5	S-21.1 S-21.3 S-22.4 S-32.6X S-23.1	S-23.4 S-23.6 S-23.8X S-33.1 S-24.1	\$\\ \chi_{\chi\chi_{\chi_{\chi_{\chi_{\chi_{\chi_{\chi_{\chi_{\chi_{\chi\tiny{\chi_{\chi\tiny{\chi_{\chi_{\chi_{\chi_{\chi_{\chi_{\chi_{\chi_{\chi_{\chi\tiny{\chi_{\chi_{\chi_{\chi_{\chi_{\chi_{\chi_{\chi_{\chi_{\chi\tiny{\chi_{\chi_{\chi_{\chi_{\chi\tiny{\chi\tin\tiny{\chi\tiny{\chi_{\chi_{\chi\tiny{\chi_{\chi\tiny{\chi\tiny{\chi\tiny{\chi_{\chi\tiny{\chi\tiny{\chi\tiny{\chi\tiny{\chi\tiny{\chi\tiny{\chi\tiny{\chi\tiny{\chi\tiny{\chi\tiny{\chi\tiny{\chi\tiny{\chi\tin\tiny{\chi\tin\tiny{\chi\tiny{\chi\tiny{\chi\tiny{\chi\tiny{\chi\tiny{\chi\tin\tiny{\chi\tiny{\chi\tiny{\chi\tiny{\chi\tiny{\chi\tiny{\chi\tin\tiny{\chi\tiny{\chi\tiny{\chi\tiny{\chi\tiny{\chi\tiny{\chi\tin\tiny{\chi\tiny{\chi\tiny{\chi\tin\tii\tin\tin\tii\tin\tin\tii\tiny\tin\tin\tii\tin\tin\tii\tin\tin\tin\tin	S-31 S-32.4 S-32.5 S-33.2 S-34.2
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	10 - (1 MOT IC	-10.9 -11.5 -10.7	-9.3 -19.2 -15.0	-111.0 -10.7 -10.8 -9.6	-10.7 -11.6 -10.7 -10.4	-111.4 -111.8 -111.0 -110.5	-9.6 -9.6 -12.2 -11.7	1 1 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0
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	MAG	16.20C 16.49 15.35 16.00	16.93 16.80 16.10 16.43 16.66	15.91 16.86 17.05 16.72 15.41	16.02 16.49C 16.13 16.19	16.39 16.98 16.14 15.81 15.99A	16.21C 16.68 16.57 16.69 16.25	16.82C 16.46 16.90 17.15 17.01
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TABL	DE(195(20 4 20 4 22 5 21 5 21 5	2223	1 4 4 6 6 1 1 1 3 3 4 4 6 6 6 1 1 1 3 9 6 6 6 6 6 6 6 6 6 6 6 6 6 6 6 6 6 6	8 8 8 8 8 8 8 8 8 8 8 8 8 8 8 8 8 8 8	228 286 286 288 288 546 5	22 4 4 4 2 2 0 0 0 0 0 0 0 0 0 0 0 0 0 0	19 1 5 2 12 3 12 1 12 1
	R. A. 1950.0	5 38 14.9 5 37 17.5 5 41 02.5 5 47 11.8 5 52 04.3	5 57 33.1 5 57 31.9 5 55 53.8 5 54 32.9 5 54 18.7	5 39 29 8 5 53 49 4 5 53 39 5 5 39 57 2 5 41 48 9	5 51 43.3 5 51 38.3 6 13 45.1 6 18 40.5 6 17 38.1	6 08 15.9 6 10 32.7 6 11 06.7 6 17 24.5 6 16 20.9	6 08 22.7 6 14 16.3 6 14 15.0 6 35 42.9 6 29 34.3	6 33 42.5 6 21 10.3 6 02 01.9 6 24 28.3 6 26 27.9
	в U•T•	27.2639 28.1736 27.2639 27.2639 27.2639	27.2639 27.2910 27.2639 28.1736 28.3271	28 • 1736 28 • 1736 28 • 3271 28 • 1826 28 • 1826	28.1826 28.2549 27.3000 27.3000 28.3090	27.2910 27.2910 27.2910 27.2910 28.3000	28.3271 28.2549 28.2819 28.3000 28.3451	28.3451 28.2819 22.1250 28.2619 28.2619
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	MINOR	1951YU1 1951YU1 1951YV1 1951YW1 1951YX1	1951YY1 1951YY1 1951YZ1 1951YZ1 1951YZ1	1951YA2 1951YB2 1951YB2 1951YC2 1951YD2	1951YE2 1951YE2 1951YF2 1951YG2 1951YG2	1951YH2 1951YJ2 1951YK2 1951YL2 1951YL2	1951YM2 1951YN2 1951YN2 1951YO2 1951YP2	1951YG2 1951YR2 1951YR2 1951YS2 1951YT2

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	MAG	15.61 16.53 17.13 16.34 16.20	16.05C 13:72 13.73 16.88C 15.77C	15.97 16.10C 16.28 15.07 14.87	16.75C 16.47 15.35 16.19C 16.01	15.56C 16.23 14.44 16.19 15.52	16.52 15.98A 16.25 15.57 15.67	16.58 16.25 16.14 15.42 15.39
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TABLE	DEC 1950.0	24 14 15 20 52 38 18 10 13 20 21 06 1 59 26	1 56 37 26 04 21 26 24 44 23 13 20 31 16 27	35 39 28 29 08 06 29 08 10 16 13 20 16 13 27	23 15 34 19 26 04 18 11 56 8 45 42 8 54 59	42 24 11 34 22 46 31 25 35 31 14 30 4 34 01	14 51 43 15 13 29 19 00 35 12 21 19 11 42 09	8 57 13 8 19 44 8 20 10 1 23 38 28 25 45
	R. A. 1950.0	9 10 04.5 8 45 02.7 8 59 48.8 8 48 25.3 5 53 45.1	5 53 15.9 6 38 25.2 6 37 28.2 6 30 31.3 6 41 49.9	6 46 13.7 6 46 25.9 6 46 25.5 6 47 23.5 6 47 23.2	6 59 47•4 7 14 28•6 7 14 49•9 7 52 39•5 7 50 38•1	8 19 59.9 8 03 56.9 8 03 45.1 8 08 09.0 7 57 46.8	8 13 10.9 8 14 50.9 8 22 46.0 8 18 00.5 8 22 52.5	8 27 26.9 8 28 55.3 8 28 49.7 8 21 15.9 8 43 30.9
	E U•T•	29.2285 28.3257 29.2375 28.3257 21.1160	22.1340 22.1611 23.1708 23.1257 23.1618	23.1618 23.1618 23.1708 23.2431 23.2521	25.1535 25.2799 25.2799 26.2049 28.1903	26.2951 26.2861 26.2861 26.2861 28.1993	28.2083 28.2083 28.2083 28.2174 28.2174	28.2174 28.2174 28.3076 28.2264 28.3347
	D A T	2 C AN 2 C AN 2 C AN 2 C AN	2 2 A A A A A A A A A A A A A A A A A A	2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2	2 C C C C A N C C A N C C C A N C C C A N C C C C	2 CAN CAN CAN CAN CAN	2 C AN C C AN C AN C AN C AN	2 JAN 2 JAN 2 JAN 2 JAN 2 JAN
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	SURVEY NUMBER	1-72.4X 1-63.6X 1-73.3X 1-63.7X S-36.1	S-46.3X S-53.3 S-63.1 S-54.7X S-62.4	S-62.5 S-62.7X S-63.2 S-65.2 S-65.2	T-13.1 T-24.3 T-24.4 T-35.2X	7-41.1 7-42.1 7-42.2 7-42.3	1 - 5 4 4 1 1 - 5 5 4 4 1 1 - 5 5 4 9 1 1 - 5 5 5 1 1 1 - 5 5 5 1 2 1 2 1 2 1 2 1 2 1 2 1 2 1 2 1	1-55.3 1-55.4 1-65.3X 1-56.1
	MINOR PLANET	19528G 19528H 19528J 19528K 19528K	1952BR 1952BS 1952BS 1952BT 1952BU	1952BV 1952BW 1952BW 1952BX 1952BX	19528Y 19528Z 19528AI 19528BI 19528BI	1952BC1 1952BD1 1952BE1 1952BF1 1952B61	1952BH1 1952BJ1 1952BK1 1952BK1 1952BK1	1952BN1 1952B01 1952B01 1952BP1 1952BP1

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	DAY	25 10 16 35	11 42 56 53 80	82 50 72 76	44 -3 -77 37 -	108 99 18 -155 -48	31 33 13 8	55 45 49 32
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	g					12.47	12.57 12.77C	
	MAG	16.12 15.92 15.59 15.89 16.30	15.48C 16.63 16.36 16.40 16.36	16.32 16.44 16.79C 16.79 16.61	16.43 16.66 14.41 16.01A 15.64C	16.07 15.69 16.63C 14.42 15.51	15.65 15.85C 16.13 15.65 15.78	16.71 16.01 16.64 15.84 16.10C
	VAR							
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TABLE	DEC 1950.0	20 22 41 22 30 26 17 14 27 2 03 37 26 30 29	25 20 57 23 33 46 23 34 24 23 45 00 26 04 06	26 21 06 17 33 12 17 33 12 32 04 04 32 04 08	24 27 07 9 51 42 10 51 51 15 04 56 13 53 55	16 45 54 16 45 58 6 29 57 - 5 31 22 22 22 03	23 40 02 23 40 02 7 35 39 20 31 38 20 32 46	9 04 29 6 01 41 10 21 40 10 21 43 12 13 16
	R• A• 1950•0	8 38 24.2 8 39 26.4 8 37 03.7 8 37 59.7 8 56 12.5	9 10 43.1 9 15 11.3 9 15 04.7 9 13 25.3 9 12 13.5	9 10 13.3 8 57 22.1 8 57 21.3 9 33 32.3 9 33 31.1	9 16 55.1 9 14 55.9 9 16 17.8 9 16 07.0 9 28 55.9	9 14 58 2 9 14 57 5 9 31 22 1 9 33 22 9 9 30 44 4	9 14 35.4 9 14 34.7 10 43 35.7 10 06 29.3 10 05 40.7	11 12 52.9 11 10 15.1 10 52 37.3 10 52 36.8 11 05 05.7
	DATE U.T.	52 JAN 28.3257 52 JAN 28.3257 52 JAN 28.3167 52 JAN 28.2986 52 JAN 29.2285	52 JAN 29.2285 52 JAN 29.2285 52 JAN 29.3549 52 JAN 31.3306 52 JAN 29.2285	52 JAN 31.3306 52 JAN 29.2375 52 JAN 29.2465 52 JAN 31.3403 52 JAN 31.3583	52 JAN 31.3306 52 JAN 29.3458 52 JAN 29.3458 52 JAN 29.3458 52 FEB 1.2583	52 FEB 17.2069 52 FEB 17.2160 52 FEB 1.2764 52 FEB 1.2854 52 JAN 29.3549	52 FEB 17.1979 52 FEB 17.2069 52 FEB 20.2410 52 FEB 19.2118 52 FEB 20.2229	52 FEB 26.2618 52 FEB 24.2493 52 FEB 24.2312 52 FEB 24.2493 52 FEB 24.2403
	SURVEY NUMBER	7-63.1 7-63.2 7-64.1 7-66.2 7-72.1	1-72.2 1-72.3 1-83.2X 1-82.6X	T-82.1 T-73.2X T-74.2 T-81.2 T-91.3X	1-82.2 1-84.2 1-84.3 1-84.6X	U-13.1 U-14.3X T-95.3X T-96.2 T-83.4X	U-12•1 U-13•9X U-54•2 U-32•7 U-42•8X	U-733.1 U-633.4 U-53.8X U-63.8X U-62.8
	MINOR PLANET	1952BR1 1952BS1 1952BT1 1952BU1 1952BV1	19528W1 19528X1 19528X1 19528X1 19528X1	1952BY1 1952BZ1 1952BZ1 1952BAZ 1952BAZ	19528B2 1952BC2 1952BD2 1952BE2 1952CB	1952CB 1952CB 1952CC 1952CD 1952DA	1952DA 1952DA 1952DF 1952DJ 1952DJ	1952DR 1952DS 1952DT 1952DT 1952DW

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	DAY ON	40 94 99 199	32 31 48 100 78	22 68 98 98	110 109 80 33 37	78 14 9 -19	2 4 4 5 6 4 4 4 4 4 4 4 4 4 4 4 4 4 4 4 4	72 77 22 64 85
	10 - DA) MOTION	-11.5 -7.6 -8.2 -9.1	1 1 8 8 5 1 1 1 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	- 10 - 1 - 1 - 1 - 1 - 1 - 1 - 1 - 1 - 1	-11.0 -10.2 -9.8 -12.0	-1110 -1110 -180 -600	-7.2 -7.9 -8.7 -9.1
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	MAG	16.39 16.14 16.54 16.34 16.66	16.18 16.40 16.64 15.94 15.21	16.69C 16.68 16.84 16.64C 16.76	16.51 16.19C 16.07 16.36 16.33	15.29 15.98 15.86 15.65C 15.61	16.64 16.14 16.06 16.18 16.29	15.85 15.84 16.20C 16.10 15.84
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TABLE A	DEC 1950.0	17 08 24 13 51 42 14 01 13 19 54 49 20 51 45	16 51 47 16 51 49 17 15 50 18 22 15 10 33 18	19 45 21 17 56 27 9 24 52 9 31 24 4 54 36	5 05 00 5 04 59 23 51 20 15 55 13 16 54 19	10 45 13 5 22 21 5 22 17 8 02 29 8 02 30	23 28 29 11 16 49 11 07 24 5 41 12 5 41 10	5 47 26 5 47 30 - 0 16 41 15 21 53 18 18 07
	R. A. 1950.0	9 41 25.7 9 28 40.6 9 27 55.8 9 15 54.6 9 18 00.9	9 21 58•7 9 21 57•7 9 25 48•5 9 20 09•1 9 10 20•8	9 39 37 0 9 41 28 9 9 48 12 1 9 47 17 0	9 46 14.0 9 46 13.3 9 58 48.5 9 53 17.1 10 03 09.5	9 57 46.1 10 00 49.1 10 00 48.7 10 08 09.2 10 08 02.3	10 12 41.1 10 17 54.8 10 18 25.5 10 21 15.3 10 21 14.8	10 28 46.2 10 28 45.5 10 15 50.8 10 36 35.2 10 39 20.9
	T E U.T.	EB 18.2125 EB 17.2160 EB 18.2125 EB 17.2069	EB 17.2069 EB 17.2160 EB 17.2069 EB 17.2069 EB 17.2160	EB 18.2125 EB 18.2125 EB 18.2215 EB 19.2299 EB 18.2306	EB 19.2299 EB 19.2389 EB 19.2118 EB 19.2208 EB 19.2208	EB 19.2299 EB 19.2299 EB 19.2389 EB 19.2299 EB 19.3210	EB 20.2229 EB 19.3210 EB 19.3210 EB 19.3111 EB 19.3210	EB 19.3111 EB 19.3210 EB 19.3111 EB 24.2222 EB 24.2222
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	SURVEY NUMBER	U-23.5 U-14.5X U-23.7X U-13.2 U-13.3	U-13.4 U-14.4X U-13.5 U-13.11X U-14.1	U-23.2 U-23.4 U-24.3.4 U-34.3X U-25.1	U-35.2X U-32.0 U-32.4 U-33.2 U-33.7	U-34.3 U-34.4 U-35.1 U-34.6 U-44.1	U-42.3 U-44.3 U-44.3 U-44.3 U-44.3	U-45.5X U-44.8 U-45.1 U-52.2 U-52.3
	MINOR	1952DX 1952DZ 1952DZ 1952DA1 1952DA1	1952DC1 1952DC1 1952DD1 1952DE1 1952DE1	1952DG1 1952DH1 1952DJ1 1952DJ1 1952DJ1	1952DK1 1952DK1 1952DK1 1952DK1 1952DM1	1952D01 1952DP1 1952DP1 1952DQ1 1952DQ1	19520R1 1952DS1 1952DT1 1952DU1 1952DU1	1952DV1 1952DV1 1952DW1 1952DW1 1952DX1

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	DAY ON	71 -10 -11 112 79	4 5 7 5 7 5 6 7 5 7 5	109 96 83 83 83	4400W W 00 4 00 4	41 91 34 132	59 61 47 75	106 99 9 14 77
	10 - DA) MOTION	-100.6 -100.6 -100.6	1	1 1 1 1 1 8 4 4 4 8 8 6 6 6 6 6 6 6 6 6 6 6 6 6 6	-8.7 -6.1 -7.2	1	-7.8 -7.3 -7.1 -6.6	-7.3 -7.1 -9.2 -8.2
	נה							
	MAG	15.69C 16.34 16.57 16.43	16.49 15.47 16.29 16.44C	15.68 15.07C 15.94 15.13C	16.03 15.54 16.41C 16.66	16.55 16.58 15.17 16.40 15.28	16.64 16.45 16.46 16.83	15.28 15.32 15.94 15.85 16.95
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TABLE	DEC 1950.0	19 04 00 - 5 12 35 - 5 12 45 19 14 00 12 26 04	12 42 16 - 0 59 26 21 40 12 14 29 47 10 02 04	8 21 43 7 45 01 6 07 23 18 20 13 4 40 21	4 09 04 5 09 11 0 35 58 1 50 35 0 31 01	2 48 20 - 7 27 08 - 4 16 00 - 4 26 41 16 32 25	0 06 26 0 06 22 - 3 04 05 - 2 40 31 - 2 43 20	1 19 49 1 19 51 - 6 25 08 - 6 25 05 - 1 26 19
	R• A• 1950•0	10 42 38.2 10 46 53.9 10 46 49.2 11 12 11.6 10 56 25.9	11 02 20.0 11 05 32.6 11 17 00.5 11 31 53.5 11 14 08.0	11 15 33.6 11 20 21.7 11 22 11.2 11 46 09.5 11 35 05.1	11 38 50.6 11 48 38.7 11 32 01.3 11 33 40.6 11 34 43.5	11 37 38.5 11 37 11.8 11 45 07.5 11 46 08.1 12 11 35.8	12 05 37.8 12 05 38.2 11 56 22.0 12 07 02.2 12 07 14.9	11 51 25.5 11 51 25.2 11 54 55.2 11 54 55.1 12 06 05.5
	DATE U.T.	52 FEB 24.2222 52 FEB 20.2500 52 FEB 20.3132 52 FEB 20.3222 52 FEB 24.2403	52 FEB 24.2403 52 FEB 24.2583 52 FEB 26.2417 52 FEB 26.2528 52 FEB 26.2518	52 FEB 26.2618 52 FEB 26.2618 52 FEB 26.2618 52 FEB 20.3590 52 FEB 24.3681	52 FEB 24.3681 52 FEB 24.3681 52 FEB 26.2708 52 FEB 26.2708 52 FEB 26.2708	52 FEB 26.2708 52 FEB 26.2799 52 FEB 26.2799 52 FEB 26.2799 52 FEB 24.3771	52 FEB 26.3792 52 FEB 26.3701 52 FEB 26.3701 52 FEB 26.3701 52 FEB 26.3701	52 MAR 23.1993 52 MAR 23.2083 52 FEB 26.3611 52 FEB 26.3701 52 FEB 26.3701
	MINOR SURVEY PLANET NUMBER	1952DZ1 U-52.5 1952DA2 U-55.2 1952DA2 U-56.2X 1952DB2 U-61.3 1952DC2 U-62.5	1952DD2 U-62.6 1952DE2 U-64.5 1952DF2 U-71.1 1952DG2 U-72.2 1952DH2 U-73.2	1952DJ2 U-73.3 1952DK2 U-73.5 1952DL2 U-73.6 1952DM2 U-81.1 1952DN2 U-83.4	1952DD2 U-83.5 1952DP2*U-83.7 1952DQ2 U-84.2 1952DR2 U-84.3 1952DS2 U-84.4	1952D12 U-84.6 1952DU2 U-85.2 1952DV2 U-85.4 1952DW2 U-85.5 1952DX2 U-85.5	1952DY2 U-93.4X 1952DY2 U-94.3 1952DZ2 U-94.1 1952DA3 U-94.4 1952DB3 U-94.5	1952DB3 V-14.5 1952DB3 V-23.2 1952DC3 U-95.1 1952DC3 U-94.7X 1952DD3 U-94.12X

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TABLE	DEC 1950.0	-12 43 38 -12 43 27 - 6 57 31 19 18 11 21 30 22	21 30 28 3 10 41 3 10 45 - 1 45 35 - 1 45 30	- 5 09 55 - 5 35 07 - 7 11 02 - 4 39 56 - 4 36 57	0 0 1 1 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	4 4 4 5 5 0 5 0 5 0 5 0 5 0 5 0 5 0 5 0	- 4 52 00 - 5 15 31 - 3 13 52 - 2 31 20	-19 07 44 - 9 51 27 - 7 52 15 - 8 30 54 - 8 33 28
	R• A• 1950•0	12 02 06•7 12 02 06•1 12 02 18•6 11 38 15•1 11 53 43•5	11 53 42.9 11 31 50.3 11 31 50.2 11 51 23.2 11 51 22.9	11 29 59•3 11 31 26•9 11 34 44•0 11 51 14•8 11 50 28•4	2 23 04. 2 15 19. 2 27 20. 2 27 20. 2 17 12.	2 25 19. 2 16 56. 2 45 00. 2 44 59. 3 13 14.	13 13 14•3 13 23 33•1 13 39 15•1 13 39 40•5	14 18 08 07 13 48 13 03 14 01 14 08 13 51 11 03 13 44 15 09
	T E U.T.	EB 26.3521 EB 26.3611 EB 26.3611 AR 22.1833 AR 22.1833	MAR 22.2014 MAR 23.1903 MAR 23.1993 MAR 23.1993 MAR 23.2174	MAR 24.1875 MAR 24.1875 MAR 24.1875 MAR 23.2174 MAR 24.1875	RR 233.2228	R 24 - 20 8 8 8 8 8 8 8 8 8 8 8 8 8 8 8 8 8 8	MAR 23.3535 APR 23.1903 APR 24.1951 APR 24.1951	PR 26.2146 PR 24.2854 PR 24.2854 PR 24.2854 PR 24.2854
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	SURVEY NUMBER	U-96.2X U-95.3 U-95.4 V-11.1	V-21.1 V-13.1 V-14.7X V-14.6 V-24.2	V-15.1 V-15.2 V-15.3 V-24.4X V-15.5	0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	V-63.4 W-24.3X W-33.8X E-33.4	W-55.3 W-43.13X W-43.11 W-43.8 W-34.8X
	MINOR PLANET	1952DE3 1952DE3 1952DF3 1952FU	1952FV 1952FW 1952FW 1952FX 1952FX	1952FY 1952FZ 1952FA1 1952FB1 1952FB1	9952F 9952F 9952F 9952F	952FG 952FU 952FU 952FU 952FK	1952FK1 1952HB 1952HH 1952HJ	1952HL 1952HO 1952HS 1952HT 1952HT

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	SURVEY NUMBER	2 4	-44-	-62.2	-61.	-13.	-22.	W-32.1	-33.2	-36	35.	45.	W-44.1	26.	36.	-40	-35.	W-42.2	-43.	-43.	-43	W-43.12	-54.1	77-	-45	-45.	-52.	W-53.3	-53.	-53•	3.1	-53	W-62.1	-54.	-54•
	MINOR PLANET	1952HV	ν υ υ	952HG	952H	952HY	952HZ	1952HA2	952HB	952HD	952HD	952HD	1952HD2	952HE	952HE	952HF	952HG	952HG	952HH	1952HJ2	952HK	1952HL2	952HL	952HM	952HN	952H0	952HP	1952HQ2	952HR	952HS	952HT	952HU	1952HU2	952HV	952HW

	O - C MOTION							
	DAY ON	65 88 12 13	6 76 67 164 62	51 57 78 -19	8 6 6 6 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	71 6 9 13 13	-112 -12 -13 13	139 139 139
	10 - DA) MOTION	0000	1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	- 9 - 0 - 9 - 5 - 9 - 9 - 11 - 4	8 - 1 - 9 - 0 - 9 - 0 - 9 - 0 - 9 - 0 - 9 - 0 - 9 - 0 - 9 - 0 - 9 - 0 - 9 - 0 - 9 - 9	-111-7 -111-7 -111-7 -11-0	110000000000000000000000000000000000000	-8.6 -8.8 -7.7 -7.1
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	MAG R	16.50A 16.62 16.21 16.88 16.88	17.04 16.11 16.62C 16.21 15.91C	16.54 16.07 16.83C 16.71 17.28	16.27 15.74 15.31	16.33 15.63	16.140	15.61 11.76C 16.07 16.06 16.52
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TABL	DE(195(14 0 11 1 15 0 17 4	-15 00 -11 59 -11 19	13 17 13 15 17 17	-1-7 -1-7 -1-4 -1-4 -1-6 -1-6 -1-6 -1-6 -1-6 -1-6	21 1 20 2 20 2 15 4 18 2	118 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2	-123 -16 -16 -16 -16 -16 -16
	R. A. 1950.0	07 38.0 -0 59.6 -0 47.7 -1 16 34.7 -1 16 33.8 -1	17 23.5 - 17 39.7 - 32 52.1 30 53.5 - 31 12.9 -	29 45.6 - 32 34.5 - 33 54.1 - 35 31.1 - 35 59.1 - 1	43 18.9 43 10.3 47 59.5 49 55.5 49 46.5	34 08.3 46 02.6 45 53.1 56 42.5	02 09.6 01 14.7 51 13.3 04 31.9	43 14.3 17 24.4 08 22.1 51 40.3
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	E U.T.	26.2056 26.2056 26.2056 26.2056 26.2056	26.2056 26.2056 26.3229 26.3049 26.3049	27.2167 27.2167 27.2167 27.2167 27.2167	27.2167 27.3708 27.2167 27.2167	27.2257 27.2257 27.3708 27.3799	27.3708 28.2778 27.3708 27.3708 28.2778	20.1764 28.2507 28.2868 20.2458 28.2868
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	SURVEY NUMBER	W-54.4 W-54.5 W-54.6 W-54.6	W-54.12 W-54.13 W-61.3 W-63.3	X - 64.2 X - 64.3 X - 64.3 X - 64.5 X - 64.5	W-64.9 W-75.1 W-64.10 W-74.1	W-65.1 W-65.2 W-75.9 W-74.4	W-75.10X W-84.1 W-75.3 W-75.8	X-15.1 W-83.2 X-14.2 X-14.2
	MINOR PLANET	1952HX2 1952HY2 1952HZ2 1952HA3 1952HA3	1952HB3 1952HC3 1952HD3 1952HE3 1952HF3	1952HG3 1952HH3 1952HJ3 1952HK3 1952HK3	1952HM3 1952HM3 1952HN3 1952HO3 1952HO3	1952HP3 1952HQ3 1952HQ3 1952HR3 1952HR3	1952HS3 1952HS3 1952HT3 1952HU3 1952HU3	1952HV3 1952HV3 1952HW3 1952HW3 1952HW3

TABLE A	•T• R• A• DEC 0 - C MAG G 10 - DAY 0 - C 1950•O R A DEC VAR MOTION MOTION	868 15 28 46.5 -15 01 08 15.23 -7.6 27 458 15 11 34.7 -14 02 51 15 12 22 -7.3 19 729 15 10 02.8 -13 58 17 15.25 -7.2 20 74 14 6 34 15.91 -7.8 52 74 15 15 -7.8 52	97 15 16 90 15 16 90 -8.8 -1 97 15 16 16 90 -8.8 -1 10 15 15 15 15 15 15 10 16 16 16 16 16 16 16 10 16 <	729 15 12 01:3 - 7 39 03 15.28C -8.8 -35 458 14 58 50.5 -16 47 11 16.19 -10.0 26 458 15 10 59.5 -11 54 52 16.39C -8.8 69 729 15 09 18.2 -11 41 19 16.65C -8.5 56	458 15 10 07.3 -17 19 44 15.32 -10.9 -28 729 15 08 00.0 -17 25 48 15.66 -10.4 -33 819 15 01 21.1 -2 52 27 16.20 -8.0 -8 819 15 16 45.9 -3 17 56 15.44 -9.2 -37 78 729 15 09 42.0 -6 17 02 16.12 -7.7 78	16.57 -10.1 42 15.32 19.0 -24 46.38 16.45 -10.2 -24 15.31 -29.30 19 16.31 -8.7 44 15.32 -29.30 19 16.13 -10.0 67 292 16.12 53.2 -19 43 05 15.97C -11.8 1	208 16 12 53.8 -19 43 02 208 16 23 48.2 -13 30 24 16.29 -9.8 52 208 16 24 48.8 -14 07 32 16.56 -8.7 53 097 16 23 00.6 -29 27 42 15.13 -12.5 1 097 16 25 45.8 -27 19 23 14.88 -9.1 51	278 16 25 44.8 -27 19 14 84 -9.5 52 215 16 24 39.0 -4 39 99 15.38 -6.7 18 132 16 41 48.6 -9 36 46 15.95 16.54 215 16 41 48.6 -9 36 33 16.54 15.56 -11.9 -40 278 16 28 11.1 -24 37 59 15.56 -11.9 -40
	C EC							
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TAB	DE 95	15 0 14 0 13 5 21 4	13 2 3 2 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	16 4 3 11 5 5 11 5 5 11 5 5 11 5 5 11 5 5 11 5 5 11 5 5 11 5 5 1 5 1 5 5 1 5 5 1 5 5 1 5 5 5 5 1 5	17 1 17 2 2 2 3 1 6 1	22 22 22 22 26 26 26 26 26 26 26 26 26 2	19 4 13 3 14 0 29 2 27 1	24 3 3 4 3 4 3 4 3 4 4 3 4 4 3 4 4 4 4 4
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	95	5 2 2 8 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	6 2 3 6 7 6 7 6 7 7 7 7 7 7 7 7 7 7 7 7 7 7	12 58 10 10 09	10 01 01 16 09	665 665 125 125 125 125 125 125 125 125 125 12	6 12 6 23 6 23 6 23 6 25	25 24 41 41 28
	-	868 458 729 778	5 9 7 2 0 8 0 4 9 2 0 8 0 4 9	729 458 729 729	458 729 819 819 729	097 937 028 028 292	208 208 208 097 097	278 215 132 215 278
	E U	28 - 2 20 - 2 22 - 2 28 - 2	0 4 4 4 4	2000	2002	22 - 2 24 - 1 24 - 2 24 - 2 24 - 2	24.2 24.2 24.2 25.2 25.2	25.2
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	SURVEY NUMBER	W-83.6 X-14.11X X-23.6X W-84.5		X-122 X-14.5 X-14.9 X-22.3 X-22.3	X-14.10X X-23.1 X-21.1 X-21.3 X-22.2	XX - 36 X - 44 X - 45 X - 45 5 - 2 X - 5 5 - 5	X - 53 • 1 × × - 53 • 1 × × - 53 • 2 × × - 55 • 3 × × - 55 • 1 × - 55 • 2 × -	X-64.1 X-61.1 X-62.2 X-61.2 X-64.2
	MINOR PLANET	1952HY3 1952HY3 1952HY3 1952HZ3	952KA 952KA 952KA 952KA 952KB	1952KJ 1952KM 1952KM 1952KN 1952KN	1952KO 1952KO 1952KP 1952KP 1952KQ	1952KS 1952KT 1952KU 1952KV 1952KV	1952KW 1952KX 1952KY 1952KZ 1952KZ	1952KA1 1952KB1 1952KC1 1952KC1 1952KC1