

THE PERKINS CATALOG OF REVISED MK TYPES FOR THE COOLER STARS

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ABSTRACT

The catalog contains 1054 stars of types G, K, and M, plus a few stars of type S. Since the Revised MK types have been estimated on spectrograms taken at four observatories, extensive comparisons have been made to ensure consistency between the types for stars in the northern and southern hemispheres.

Subject headings: stars: catalogs — stars: late-type — stars: spectral classification

The catalog lists the spectral types of the G, K, M, and S stars that have been classified at the Perkins Observatory on the Revised MK system. Although the accuracy of the temperature types, luminosity classes, and abundance indices necessarily varies with the number and quality of the spectrograms available, and with the character of each spectrum, there are two reasons for making them available in one comprehensive list. First, the types have all been estimated consistently, with many cross-checks to eliminate systematic differences that might arise from the need to use several different spectrographs in the northern and southern hemispheres. Second, the resolution of about 2 Å of the slit spectrograms allows classification to be done somewhat more accurately than is possible for even the best surveys carried out with objective-prism spectrographs.

The Revised MK system has been described most recently by one of us (Keenan 1987), and a selection of the best determined types, that can serve as standards, has been published by Keenan and Yorka (1988). In the present catalog these standard stars are identified by a dagger (†) in front of their names in the first column. For stars of normal solar composition the scheme of classification is exactly that of the original MK system (Morgan, Keenan, and Kellman 1943). It is only for the spectra that show “peculiarities” due to differences in chemical composition that abundance indices of the revised system have been added. These indices run usually from 1 to 5, with decimal fractions (normally 0.5) indicating the marginal spectra in which the peculiarity is just detectable.

The types were determined visually from photographic spectrograms taken at scales of 68–78 Å mm^{−1} with the instruments listed in Table 1. The behavior of the criteria actually used in the classification is illustrated in the *Atlas of Spectra of the Cooler Stars* of Keenan and McNeil (1976). Fine illustrations of the original MK spectral types of both hot and cool stars will be found also in *An Atlas of Representative Stellar Spectra* of Yamashita, Nariai, and Norimoto

TABLE 1
TELESCOPES AND SPECTROGRAMS

Observatory	Telescope (m)	Spectrograph	Scale (Å mm ^{−1})
Perkins.....	0.8	Meinel	85
Lowell	1.8	Boller & Chivens	77
Cerro Tololo	1.5	Hiltner	78
Las Campanas ...	0.6	Garrison	68

(1977), which, however, does not include the strong-line or the metal-deficient stars of types G and K.

The accuracy of the classification is necessarily much less than the precision with which the indices can be reproduced when different spectrograms of the same star are classified. This is due in large part to the fact that most of the observable criteria that we can use are sensitive to more than one of the variables that we are trying to separate, and the classifier must carry out a sort of mental component analysis. Thus, for example, the luminosity classes of normal stars near the giant branch, for temperature types between G5 and K5, are good to about 0.4 mag. If, however, a star has appreciable excess of the heavy *s*-process elements, several of our most sensitive luminosity criteria become unreliable. The barium stars are an extreme case, and for them the luminosity classes are followed by a colon to indicate that they are uncertain estimates.

In spite of these inherent limitations, it has been possible, over the many years since the Yerkes Atlas (Morgan, Keenan, and Kellman 1943) came out, gradually to improve the quality of classification spectrograms. At the same time the increase in the number of stars for which good spectrograms were available made it easier to notice slight differences in the line strengths. Thus the “boxes” into which we classify the spectra could often be made smaller. One result has been that occasionally a standard star that fell near the edge of an

TABLE 2
 TYPE G, K, AND M STARS

Name	BS	HD	$\alpha(2000)$	$\delta(2000)$	V	Type
			^h ^m			
+ 30 Psc	9089	224935	0 02.0	- 6°01'	4.41	M3 III
+ 85 Peg	9088	224930	0 02.2	+27 05	5.75	G5 Vb Fe-2
+ 33 Psc	3	28	0 05.3	- 5 42	4.61	K0 III-IV
G24 B	-	388	0 05.6	+45 49	8.97	M0.5 V
-	-	218	0 06.8	-70 35	8.03	M0.5 III Ba2.5
-	-	688	0 11.3	- 3 19	6.8	K0- III
+ -	-	936	0 14.0	+60 00	6.88	G7.5 IIab
+ X Peg	45	1013	0 14.6	+20 12	4.80	M2+ III
+ AD Cet	46	1014	0 14.5	- 7 47	5.06 to 5.16	M3+ III
+ G215A	-	1326	0 18.3	+44 01	8.08	M2V
+ -	-	1364	0 17.8	+20 14	7.3	M3.5 IIIa
-	72	1461	0 18.7	- 8 03	6.46	G3V Fe-0.5
+ i Cet	74	1522	0 19.4	- 8 50	3.55	K1 IIIb
41 Psc	80	1635	0 20.6	+ 8 11	5.37	K3- III Cal CNO.5
+ 9 Cet	88	1835	0 22.8	-12 13	6.38	G2.5 V
-	-	1879	0 23.9	-15 57	6.45	M2 III
10 Cet	101	2273	0 26.6	- 0 03	6.18	G6 IIIb Fe-1
DL Cas	-	--	0 30.0	+60 13	8.6-9.2	G1 Ib-II H ₆ -1*
BD +59°70	-	--	0 31.4	+60 05	8.71	K5 III*
BD +58°73	-	--	0 31.6	+60 05	8.57	K2.5 II-III*
+ -	-	2901	0 32.7	+54 07	6.93	K2 III Fe-0.5*
+ -	152	3346	0 36.8	+44 29	5.13	K6 III
+ -	157	3421	0 37.4	+35 24	5.49	G2 Ib-II
+ ε And	163	3546	0 38.6	+29 19	4.37	G6 III Fe-3 CH1
+ δ And	165	3627	0 39.3	+30 52	3.28	K3 III
+ 54 Psc	166	3651	0 39.4	+21 15	5.88	K0 V
α Cas	168	3712	0 40.5	+56 32	2.23	Ko- IIIa
G228	-	3765	0 40.8	+40 10	7.35	K2 V
β Cet	188	4128	0 43.6	-17 59	2.04	G9 III CH-1 CNO.5 Cal
+ -	-	4404	0 47.2	+60 17	7.6:	G9 II-III Fe2
+ 57 Psc	211	4408	0 46.6	+15°29'	5.38	M4 IIIa
RX Cep	-	4499	0 50.0	+81 52	7.5-7.8	G6 IIa H ₆ 1
η Cas A	219A	4614A	0 49.1	+57 49	3.44	F9V
η Cas B	219B	4614B	0 49.1	+57 49	7.50	M0- V
+δ Psc	224	4656	0 48.7	+ 7 35	4.44	K4.5 IIIb
-	237	4817	0 51.3	+61 49	6.07	K2 Ib-II CN-1 CH-0.5
υ ² Cas	265	5395	0 56.7	+59 10	4.62	G8 IIIb Fe-0.5
η And	271	5516	0 57.2	+23.25	4.39	G8- IIIb*
WW Psc	284	5820	0 59.8	+ 6 28	6.0 var.	M2.5 III
+ε Psc	294	6186	1 02.9	+ 7 53	4.28	G9 III Fe-2
-	-	6474	1 06.5	+63 47	7.4-7.8	G4 Ia
+ -	316	6497	1 07.0	+56 56	6.42	K2 IIIb Fe-0.5*
+η Cet	334	6805	1 08.6	-10 11	3.44	K2- III CNO.5
μ Cas	321	6582	1 08.3	+54 55	5.16	G5 Vb
+ -	-	6833	1 09.9	+54 45	6.74	G9.5 III Fe-2*
+β And	337	6860	1 09.7	+35 37	2.05	M0+ IIIa
+ψ ³ Psc	339	6903	1 09.8	+19 40	5.55	G0 IIIa*
+ -	341	6953	1 10.3	+25 28	5.80	K7 III
+χ Psc	351	7087	1 11.5	+21 02	4.65	G8.5 III
+τ Psc	352	7106	1 11.7	+30 06	4.51	K0.5 IIIb
39 Cet	373	7672	1 16.6	- 2 30	5.46 var	G6 IIIe Fe-0.5*
V 466 Cas	-	236697	1 19.8	+58 19	8.60	M0.5 Ib*
ξ And	390	8207	1 22.3	+45 31	4.86	K0- IIIb
ψ Cas	399	8491	1 25.9	+68 08	4.72	K0 III CN 0.5
+θ Cet	402	8512	1 24.0	- 8 11	3.60	K0 IIIb
+ -	-	8701	1 27.6	+66 04	6.98	K2+ IIab
+ 46 Cet	412	8705	1 25.6	-14 36	4.89	K2.5 IIIb
+ γ Phe	429	9053	1 28.4	-43 19	3.40	M0- IIIa
+ μ Psc	434	9138	1 30.2	+ 6 09	4.83	K3 III
+η Psc	437	9270	1 31.5	+15 21	3.62	G7 IIIa

TABLE 2—Continued

Name	BS	HD	$\alpha(2000)$	$\delta(2000)$	V	Type
			h m			
B0 +59° 274	-	-	1 33.5	+60°39'	8.50	M0.5 Ib*
BD +60° 265	-	-	1 33.5	+61 33	8.7:	M1.5 Ib*
+ δ Phe	440	9362	1 31.2	-49 04	3.95	G9 III
χ Cas	442	9408	1 33.9	+59 14	4.70	G9 IIIb
-	452	9712	1 35.9	+41 05	6.38	G9.5 III-IV
40 Cas	456	9774	1 38.5	+73 03	5.28	G7 III
+	-	9852	1 34.5	+61 56	8.0:	K0.5 IIIa CN1
50 Cet	459	9856	1 36.0	-15 24	5.5	K1.5 III CNO.5
+ 51 And	464	9927	1 38.0	+48 37	3.57	K3- III
+	483	10307	1 41.8	+42 37	4.95	G1.5 V
+ ν Psc	489	10380	1 41.4	+ 5 29	4.44	K3 IIIb
+G δ 69	-	10436	1 43.8	+63 50	8.40	K5 Vbe
+	-	10465	1 43.2	+48 31	7.00:	M2 Ib
+ 107 Psc	493	10476	1 42.5	+20 16	5.24	K1 V
+	506	10647	1 42.5	-53 44	5.52	F9 V
109 Psc	508	10697	1 44.9	+20 05	6.27	G3 Va
+ τ Cet	509	10700	1 44.1	-15 56	3.50	G8 V
+ σ Psc	510	10761	1 45.4	+ 9 09	4.26	G8 III
-	-	11092	1 51.2	+64 52	6.56	K4+ Ib-IIa
ξ Psc	549	11559	1 53.6	+ 3 11	4.60	G9 IIIb Fe-0.5
+ ζ Cet	539	11353	1 51.4	-10 20	3.72	Ko III
ψ Phe	555	11695	1 53.6	-46 18	4.41	M4 III
χ Eri	566	11937	1 56.0	-51 37	3.70	G8 III-IV CN-0.5 H δ 0.5
+ η^2 Hyi	570	11977	1 54.9	-67 39	4.69	G8.5 III
-	574	12055	1 57.2	-47 23	4.83	G7 III
G δ 83.1	-	-	2 00.4	+13 04	12.27	M5- V
57 Cet	583	12255	1 59.8	-20 49	5.41	M1: IIIa
ν Cet	585	12274	2 00.0	-21 05	4.00	M0 IIIb
+ χ Phe	602	12524	2 01.7	-44 43	5.14	K5 III
+ γ And A	603	12533	2 03.9	+42 20	2.10	K3- IIb
61 Cet	610	12641	2 03.8	- 0°20'	5.93	G4 IIIb CH-0.5 +G0?
-	611	12642	2 03.7	- 4 06	5.62	M1- III Cal
+ α Ari	617	12929	2 07.2	+23 27	2.00	K2- IIIab Ca-1
KK Per	-	13136	2 10.3	+56 33	7.7-8 (7.73)	M2 Iab-Ib
-	637	13445	2 10.4	-50 49	6.12	K1 V
+ 60 And	643	13520	2 13.2	+44 14	4.84	K3.5 III
-	645	13530	2 13.6	+51 04	5.29	G8 III CN1 CH0.5 Fe-1
+ 65 ξ^1 Cet	649	13611	2 13.0	+ 8 51	4.36	G7 II-III Fe-1
-	-	13686	2 15.9	+63 14	6.98	K2.5 Ib-II
+	-	13738	2 15.5	+52 31	7.3:	K3.5 III
+ δ Tri	660	13974	2 17.0	+34 13	4.86	G0.5 V
67 Cet	666	14129	2 17.0	- 6 25	5.50	G8.5 III
π^1 Hyi	667	14141	2 14.2	-67 51	5.55	M1- III
+ PR Per	-	14404	2 21.6	+57 52	8.6:	M1- Iab-Ib
SU Per	-	14469	2 22.1	+56 36	7.6:	M3-M4 Iab
RS Per	-	14488	2 22.4	+57 06	8.0-9.4	M3+ Iab var.?
-	-	14595	2 22.1	+22 52	6.6:	G3 III: +F6:
κ For	695	14802	2 22.6	-23 49	5.19	G0 Va
+65 And	699	14872	2 25.6	+50 17	4.71	K4.5 III
-	-	16068	2 36.9	+55 55	7.38	K3.5 II-III Cal*
+	-	16139	2 36.3	+27 28	8.1:	G8.5 IIIa Fe-0.5
+ G δ 105	753	16160	2 36.1	+ 6 53	5.82	K3- V
ι Eri	794	16815	2 40.7	-39 51	4.11	K0.5 IIIb Fe-0.5
+ 14 Per	800	16901	2 44.1	+44 18	5.43	G0 Ib Cal
+ 39 Ari	824	17361	2 47.9	+29 15	4.51	K1.5 III
+ η PerA	834	17506	2 50.7	+55 54	3.76	K3-Ib-IIa
+ β For	841	17652	2 49.1	-32 24	4.46	G8.5 III Fe-0.5
+ 17 Per	843	17709	2 51.5	+35 04	4.53	K5.5 III
+ 45 RZ Ari	867	18191	2 55.8	+18 20	5.9:var	M6- III:
+ η Eri	874	18322	2 56.4	- 8 54	3.89	K1 IIIb

TABLE 2—Continued

Name	BS	HD	$\alpha(2000)$	$\delta(2000)$	V	Type
			h m			
-	-	18391	2 59.8	+57°40'	6.89	G5 Ia-Iab
τ^1	885	18474	2 59.8	+47 13	5.47	G5: III: CN-4 CH-1 Cal H γ -1
-	-	18636	2 58.6	-37 58	7.6:	G8III: CN-1 CH-4
ρ^1 Eri	907	18784	3 01.2	- 7 40	5.74	G9 III H γ -0.5
α Cet	911	18884	3 02.3	+ 4 06	2.53	M1.5 IIIa var?
τ	918	18970	3 05.5	+56 43	4.76	G9.5 III
$\tau\rho$ Per	921	19058	3 05.2	+38 50	3.2-4.1	M4 II
V Hor	-	19285	3 03.5	-58 56	8.7-9.8pg	M5- III
G& 123	-	19305	3 06.4	+ 1 56	9.10	M0 V
τ_1 Per	937	19373	3 09.1	+49 37	4.05	G0 V
$\tau \kappa$ Per	941	19476	3 09.5	+44 52	3.80	K0 III
$\tau\delta$ Ari	951	19787	3 11.6	+19 44	4.35	K0 III
-	953	19826	3 10.6	-23 44	6.37	K0 IIIb
-	966	20063	3 14.9	+42 30	6.06	K1+ III-IV CN1
-	965	20084	3 32.3	+84 55	5.61	G5 IIa CH1 Fe-1*
-	969	20123	3 16.2	+50 56	5.03	G6 Ib-IIa
-	-	20394	3 16.8	+ 2 22	8.6	G9- IIIa: Ba2.5
τ	991	20468	3 18.7	+34 13	4.82	K2 IIb Fe0.5
τ 59 Ari	995	20618	3 19.9	+27 04	5.90	G6 IV
τ	-	20619	3 19.1	- 2 51	7.05	G1.5 V
$\tau\kappa$ Cet	996	20630	3 19.4	+ 3 22	4.84	G5 V
-	999	20649	3 20.3	+29 03	4.47	K3 IIIa Ba0.5
16 τ^b Eri	1003	20720	3 19.5	-21 45	3.70	M3+ IIIa Ca-1
$\tau\zeta^1$ Ret	1006	20766	3 17.8	-62 35	5.53	G2.5 V H γ 1
τ 97 Cet	1007	20791	3 21.1	+ 3 41	5.68	G8.5 III $^{\circ}$
τ	1008	20794	3 19.9	-43 04	4.27	G8 V
-	1009	20797	3 24.7	+64 35	5.22	M0 II
$\tau\zeta^2$ Ret	1010	20807	3 18.2	-62 30	5.24	G1 V
-	1016	20894	3 21.4	-23 38	5.50	G6.5 IIb Cn-1 H δ -1
-	1023	21018	3 23.6	+ 4 53	6.38	G1 III: CH-1:
τ o Tau	1030	21120	3 24.8	+ 9°02'	3.60	G6 IIIa Fe-1
σ Per	1052	21552	3 30.6	+48 00	4.36	K3 III
5 Tau	1066	21754	3 30.8	+12 57	4.11	K0- II-III Fe-0.5
τ e Eri	1084	22049	3 32.9	- 9 28	3.72	K2 V
-	1085	22072	3 34.1	+17 50	6.17	K1 IV Fe-1
τ 10 Tau	1101	22484	3 36.9	+00 24	4.29	F9 IV-V
-	1112	22764	3 42.7	+59 58	5.73	K3.5 III+G*
τ	-	23082	3 47.6	+45 03	7.81	K2.5 II CNO.5
$\tau\delta$ Eri	1136	23249	3 43.2	- 9 46	3.53	K0+ IV
τ	1155	23475	3 49.5	+65 32	4.47	M2+ IIab
τ	1168	23697	3 44.6	-54 16	6.30	K2 IIIb
$\tau\beta$ Ret	1175	23817	3 44.2	-64 48	3.85	K2 III
$\tau\gamma$ Eri	1231	25025	3 58.0	-13 31	2.95	M0.5 IIb Ca-1
-	-	25150	4 04.6	+56 27	8.29	K0.5 III CNO.5
G& 158	-	25329	4 04.5	+35 22	8.51	K1 Vb Fe-2
τ 37 Tau	1256	25604	4 04.7	+22 05	4.37	K0 III
τ -	1270	25877	4 09.5	+59 54	6.28	G8 IIa
-	1286	26311	4 11.0	+33 35	5.72	K1 II-III
τ	1299	26575	4 10.8	-35 16	6.44	G9.5 III CH-2
τ_1 Per	1303	26630	4 14.9	+48 24	4.12	G0 Ib
47 Tau	1311	26722	4 13.9	+ 9 16	4.83	G3 III Fe-0.5
$\tau\omega^2$ Eri	1325	26965	4 15.3	- 7 40	4.43	K0.5 V
τ	1327	27022	4 20.7	+65 09	5.28	G5 IIb
$\tau\alpha$ Ret	1336	27256	4 14.4	-62 28	3.34	G8 II-III
-	-	27277	4 20.9	+50 16	8.09	G6 III*
-	-	27292	4 20.9	+50 15	7.13	K2.5 III
τ 54 Per	1343	27348	4 20.4	+34 34	4.92	G8 III Fe0.5
$\tau\gamma$ Tau	1346	27371	4 19.8	+15 37	3.65	G9.5 IIIab CNO.5
DG Eri	-	27598	4 20.7	-16 50	6.9-7.1	M4- III
$\tau\delta$ Tau	1373	27697	4 22.9	+17 32	3.76	G9.5 III CNO.5

TABLE 2—Continued

Name	BS	HD	$\alpha(2000)$	$\delta(2000)$	V	Type
			^h ^m			
+ VB 64	-	28099	4 26.7	+16°45'	8.12	G2+ V*
π Tau	1396	28100	4 26.6	+14 43	4.69	G7 IIIa Fe-1
75 Tau	1407	28292	4 28.4	+16 22	4.96	K1 III
+ ϵ Tau	1409	28305	4 28.6	+19 11	3.54	G9.5 III CNO.5
+ θ^1 Tau	1411	28307	4 28.6	+15 57	3.85	G9 III Fe-0.5
+	-	28487	4 29.7	+ 5 10	7.25:	M3.5 III
50 ν^1 Eri	1453	29085	4 33.5	-29 46	4.50	K0+ III Fe-0.5
+G λ 172	-	232979	4 37.6	+52.55	8.61	M0.5 V
-	-	29122	4 36.6	+36 57	6.62	K2 IIIb CN1
+ τ Tau	1457	29139	4 35.9	+16 30	0.86	K5+ III
BD+26°730	-	283750	4 36.8	+27 08	8.42	K2.5 Ve*
+ 52 ν^2 Eri	1464	29291	4 35.6	-30 34	3.81	G8.5 IIIa
53 Eri	1481	29503	4 38.2	-14 18	3.87	K1.5 IIIb
-	1487	29613	4 39.3	-14 21	5.45	K1 IVa
+ R Dor	1492	29712	4 36.8	-62 05	5.0-6.0	M8e III:*
+58 Eri	1532	30495	4 47.6	-16 56	5.51	G2.5 IV-V
-	1533	30504	4 49.9	+37 30	4.88	K3.5 III
+	-	30793	4 52.5	+39 17	7.91	K1.5 IIIb CN1
+ 2 Aur	1551	30834	4 52.6	+36 42	4.77	K2.5 III Ba 0.5
+	-	31274	4 52.0	-46 51	7.13	G9 III CH-3.5
+ ν Aur	1577	31398	4 57.0	+33 09	2.68	K3 II
+ θ^2 Ori	1580	31421	4 56.4	+13 31	4.07	K2- III Fe-1
+ 10 π^6 Ori	1601	31767	4 58.6	+ 1 43	4.46	K2- II
β Cam	1603	31910	5 03.4	+60 27	4.04	G1 Ib-IIa
+ R59	-	268757	4 54.3	-69 12	10.2	G8 0 LMC
G λ 182.1	-	31966	5 00.6	+14 23	6.65:	G2 IV-V Fe-0.5
-	-	32712	5 01.6	-58 31	8.53	K1 III: Ba3-
-	-	33299	5 10.6	+30 48	6.68	K0-Ib
-	1685	33555	5 11.0	- 2 15	6.24	K1- IVa
+ G λ 191	-	33793	5 11.2	-44 56	8.81	M1p VI*
-	1697	33833	5 12.8	- 6°03'	5.91	G7 III
+ ρ Ori	1698	33856	5 13.3	+ 2 52	4.45	K1 III CNO.5
α Aur	1708	34029	5 16.7	+46 00	0.70:	G4:iii:Comp.*
+ 16 Aur	1726	34334	5 18.2	+33 22	4.54	K2.5 III Fe-1
+ λ Aur	1729	34411	5 19.1	+40 06	4.71	G1.5 IV-V Fe-1
-	-	34450	5 25.0	+73 42	7.0	M0.5 III
+ θ Dor	1744	34649	5 13.8	-67 11	4.82	K2.5 IIIa
+ 29 Ori	1784	35369	5 24.0	- 7 48	4.14	G8 III Fe-0.5
27 Ori	1787	35410	5 24.5	- 0 53	5.06	G9 III-IV Fe-0.5
-	-	35601	5 27.1	+29 56	7.35	M1.5 Iab-Ib
-	-	271182	5 21.0	-65 51	9.71	F8 0 *
+ ϕ Aur	1805	35620	5 27.6	+34 28	5.07	K3.5 III CN2
G λ 204	-	36003	5 28.4	- 3 31	7.65	K5 V
-	1824	36040	5 30.8	+41 28	6.00	K0.5 IIIb CN1
β Lep	1829	36079	5 28.2	-20 45	2.84	G5 II
+	1830	36134	5 29.4	- 3 27	5.80	K1-III
+ 119 Tau	1845	36389	5 32.2	+ 5 52	4.73:	M2 Iab-Ib
+G λ 205	-	36395	5 31.4	- 3 38	7.97	M1.5 V
-	-	36552	5 30.3	-43 42	8.0:	G7III: CN-1 CH-2.5
-	-	36598	5 26.6	-70 04	8.05	K0.5 III: Ba3
+ 40 ϕ^2 Ori	1907	37160	5 36.9	+ 9 17	4.09	K0IIIb Fe-2
-	1909	37192	5 35.3	-33 05	5.74	K2 IIIa
+ LMC G367	-	269723	5 32.8	-67 42	9.91	G4 0
51 Ori	1963	37984	5 42.5	+ 1 29	4.91	K1 III
-	1970	38099	5 43.2	- 1 37	6.30	K4+ III
BD-16° 1217	-	-	5 43.1	-16 46	10+:	K0.5 IIIa CN2
+ LMC	-	269953	5 40.4	-69 41	9.94	G0 0
+ τ Aur	1995	38656	5 49.2	+39 11	4.51	G8 III Fe-1
132 Tau	2002	38751	5 49.0	+24 34	4.86	G9 IIIb
+ ν Aur	2011	38944	5 51.0	+37 18	4.73	M0 III

TABLE 2—Continued

Name	BS	HD	$\alpha(2000)$	$\delta(2000)$	V	Type
			h m			
+	2018	39045	5 51.4	+32°08'	6.25	M3 var: III*
+	2028	39225	5 52.7	+33 55	5.98	M1.5 II-III*
+ δ Lep	2035	39364	5 51.3	-20 52	3.28	K0 III Fe-1.5 CH0.5
+ 56 Ori	2037	39400	5 52.4	+ 1 51	4.75	K1.5 IIb
+ χ^1 Ori	2047	39587	5 54.4	+20 16	4.39	G0-V Ca0.5
-	-	39628	5 57.1	+55 57	6.97	K2 IIIb
α Ori	2061	39801	5 55.2	+ 7 24	0.1-1.2	M1-M2 Ia-Iab
π Aur	2091	40239	5 59.9	+45 57	4.25	M3 II
+ δ Aur	2077	40035	5 59.5	+54 17	3.71	K0-III
+	2102	40409	5 54.1	-63 05	4.65	K2 III
+	2113	40657	6 00.0	- 3 05	4.52	K1.5 III Fe-1.5
-	2121	40827	6 05.1	+59 24	6.34:	K1 III CN1-
δ Lep	2156	41698	6 05.8	-24 12	6.5-7.2	M6+ IIIa
TV Gem	-	42475	6 11.9	+21 52	7.0-7.8	M0-M1 Iab
+6 BU Gem	2197	42543	6 12.3	+22 55	6.1-7.5	M1-M2 Ia-Iab
η Gem	2216	42995	6 14.9	+22 30	3.28:	M2.5 III
+ κ Aur	2219	43039	6 15.4	+29 30	4.33	G9 IIIb
+ γ Mon	2227	43232	6 14.9	- 6 16	3.97	K1 III Ba0.5
+ η^2 Dor	2245	43455	6 11.2	-65 35	5.00	M2.5 III
+ κ Col	2256	43785	6 16.6	-35 08	4.37	K0.5 IIIa
-	2267	43993	6 18.8	- 9 23	5.35	K1.5 III
+	-	44362	6 18.8	-50 22	7.03	G2 Ib
μ Gem	2286	44478	6 23.0	+22 31	2.97	M3 IIIab
46 ψ^1 Aur	2289	44537	6 24.9	+49 17	4.95 var.	K5-M0 Iab-Ib
-	2290	44594	6 20.1	-48 44	6.60	G2 Va
-	-	44896	6 22.9	-33 37	7.2	K4.5 IIIa: Ba4
G δ 233	-	45088	6 26.2	+18 46	6.76	K2 V H,Kem1
+	2318	45184	6 24.7	-28 47	6.38	G2 Va
6 Lyn	2331	45410	6 30.8	+58 10	5.86	K0 III-IV
+	2392	46407	6 32.8	-11 10	6.27	G9.5 III: Ba3
50 ψ^2 Aur	2427	47174	6 39.3	+42°30'	4.79	K2- III
+ ν^2 CMa	2429	47205	6 36.7	-19 15	3.95	K1.5 III-IV Fe 1
+ ν^3 CMa	2443	47442	6 37.9	-18 14	4.43	K0.5 III
-	2450	47667	6 39.3	-14 09	4.81	K2 IIIa CN1 Ca1
+ ϵ Gem	2473	48329	6 43.9	+25 08	2.99	G8Ib
30 Gem	2478	48433	6 44.0	+13 14	4.49	K0.5 III CNO.5
57 ψ^6 Aur	2487	48781	6 47.6	+48 48	5.21	K0 III
-	-	49068	6 45.7	-20 51	7.06	K1.5 II-III*
-	-	49091	6 45.9	-20 47	6.93	K3 IIIa*
-	-	49105	6 46.0	-20 36	7.77	K0 II-III*
BD-20°1568	-	49212	6 46.6	-20 49	7.76	K0 II-III*
+ 18 Mon	2506	49293	6 47.9	+ 2 24	4.46	K0+ IIIa
+	2508	49331	6 47.6	- 9 00	5.06	M1+ Ib-IIa
-	2513	49396	6 45.4	-52 12	6.3	G5 Iab
-	-	49500	6 49.7	+25 30	6.92	K2- 000 Fe-2 CH-1 H δ 1
-	-	50264	6 51.2	-29 35	8.7	G2III: Ba2 CH0.5
+ ϕ^1 Cma	2580	50877	6 54.1	-24 12	3.87:	K2 Iab*
41 Gem	2615	52005	7 00.3	+16 05	5.68	K3 Ib
+	-	52220	6 59.0	-32 43	6.9:	G1 Ib
σ Cma	2646	52877	7 01.7	-27 56	3.46	K7 Ib
-	-	52938	7 02.8	- 8 27	7.8:	K3.5 II-III Ba0.5*
+	2667	53705	7 04.0	-43 36	5.54	G1.5 V
-	2685	54153	7 06.0	-38 23	6.11	G0 IIa CH1
+ G δ 268	-	--	7 10.0	+38 32	11.48	M4.5e V
63 Aur	2696	54716	7 11.7	+39 19	4.91	K3.5 III
+ τ Gem	2697	54719	7 11.2	+30 15	4.40	K2 III
AC Car	-	54795	7 06.8	-58 23	8.11-9.4:	M6 to M6+ III:
-	2703	54895	7 13.4	+51 26	5.47	M3 III
-	-	55690	7 13.3	-14 36	8.3:	M5 III
BD-9° 1935	-	--	7 14.6	-10 10	9.46	G7 IIIa*

TABLE 2—Continued

Name	BS	HD	$\alpha(2000)$	$\delta(2000)$	V	Type
			h m			
-	-	56438	7 14.6	-47°08'	8.5:	K0.5 III-IIIb CH-3.5
+ 145 CMa	2764	56577	7 16.6	-23 19	4.82	K3 Ib-II*
-	2786	57146	7 18.8	-26 35	5.27	G2 Ib*
†	2802	57615	7 21.1	-25 53	5.87	M3 III
66 Aur	2805	57669	7 24.1	+40 40	5.17	K1 IIIa Fe 1
-	-	58134	7 23.0	-29 45	7.7:	G5 Ib H&K shallow.
† ₁ Gem	2821	59207	7 25.7	+27 48	3.79	G9 IIIb
† ₂ CMi	2828	58367	7 25.6	+ 9 16	5.00	G6 II
-	2834	58535	7 24.7	-31 49	5.34	K0.5 II-III*
† ₃ CMi	2854	58972	7 28.2	+ 8 56	4.31	K3 III Fe-1
+ 6 CMi	2864	59294	7 29.8	+12 01	4.53	K1 III
YY Gem	-	60179C	7 34.6	+31 51	9.07	M0.5 Ve, Fe-2*
-	-	60197	7 32.1	-29 38	7.73	K3.5 III: Ba 3.5
+ _U Gem	2905	60522	7 35.9	+26 54	4.06	M0 III-IIIb
-	-	60898	7 36.2	-14 18	7.83	K0 II-III*
-	-	60899	7 36.2	-14 36	7.93	G8.5 III*
74 Gem	2938	61388	7 39.5	+17 40	5.05	K5 III
-	2967	61913	7 42.0	+14 13	5.81	M3 II-III
† ₄ Mon	2970	61985	7 41.2	- 9 33	3.93	G9 III Fe-1
+ R Pup	2974	62058	7 40.9	-31 40	7.5:	G2 0-Ia*
+ CPD-31° 1790	-	--	7 41.0	-31 41	8.3:	M3 Iab-Ib*
76 Gem	2983	62285	7 44.1	+25 47	5.30	K4.5 III
+ _K Gem	2985	62345	7 44.4	+24 24	3.56	G8 III
† ₅ Gem	2990	62509	7 45.3	+28 01	1.14	K0 IIIb
+ 81 Gem	3003	62721	7 46.1	+18 31	4.87	K4 III
-	3026	63302	7 47.6	-15 59	6.34	K1 Ia-Iab
COD-32°4417	-	--	7 47.9	-33 17	9.1:	M3 III
-	3043	63660	7 49.0	-24 55	5.33	G1.5 IIIb Fe-1*
† ₆ Pup	3045	63700	7 49.3	-24 52	3.43	G6 Iab-Ib
-	-	64616	7 53.8	-26 16	6.92	K0 III
-	-	64858	7 55.0	-26°28'	9.2:	G4 Iab-Ib
1 Cnc	3095	64960	7 57.0	+15 47	5.78	K3 III
-	-	65412	7 57.9	-20 25	7.8:	M2.5 Iab
-	3120	65662	7 56.3	-60 32	5.77	K3.5 II-III*
+ 12 Pup	3123	65699	7 59.1	-23 19	5.10	K0 Ib-II
V 341 Car	3126	67750	7 56.8	-59 08	6.03	M1 II*
-	3145	66141	8 02.3	+ 2 20	4.38	K2 IIIb Fe-0.5
X Gem	3149	66216	8 03.5	+27 48	4.94	K1 III
†	3153	66342	7 59.6	-60 36	5.16	M1.5 II*
†	-	66478	8 01.8	-44 36	6.55	K2 IIIa Fe1
CoD-29° 5430	-	--	8 03.5	-29 59	8.9:	K3+ Iab-Ib
_U Cnc	3176	67228	8 07.8	+21 35	5.30	G1 IVb
† 55 Cam	3182	67447	8 12.8	+68 28	5.32	G7 II
† ₇ Mon	3188	67594	8 08.6	- 2 59	4.34	G2 Ib
+ 19 Pup	3211	68290	8 11.3	-12 56	4.71	G9 III
†	3212	68312	8 11.5	- 7 46	5.34	G7 III
NS Pup	3225	68553	8 11.3	-39 37	4.45	K4.5 Ib
-	-	68879	8 14.2	- 5 54	8.23	G9 III K-2 CH-0.5*
+ _B Cnc	3249	69267	8 16.5	+ 9 12	3.52	K4 III Ba0.5
-	3259	69830	8 18.4	-12 38	6.00	G7.5 V
† 31 Lyn	3275	70272	8 22.8	+43 12	4.25	K4.5 III
-	3282	70555	8 21.4	-33 03	4.83	K2.5 II-III
-	3296	70946	8 23.3	-38 17	6.32	M3- III
†	3306	71115	8 25.9	+ 7 34	5.12	G8 II-III Fe-1
-	3315	71176	8 25.1	-24 03	5.28	K4.5 III CN1
† 27 Cnc	3319	71250	8 26.7	+12 39	5.50	M3 III Ca-1
† ₈ Cha	3340	71701	8 20.6	-77 29	4.35	K2 III CN0.5
-	3360	72184	8 32.9	+38 01	5.90	K1.5 IIIb CN0.5
-	-	72268	8 30.5	-36 43	6.70	M3+ Iab-Ib
-	-	72696	8 31.7	-55 49	8.0:	K1+ III Fe0.5

TABLE 2—Continued

Name	BS	HD	$\alpha(2000)$	$\delta(2000)$	V	Type
			h m			
π^2 UMa	3403	73108	8 40.2	+64°20'	4.60	K1+ IIIb
c Vel	3407	73155	8 34.7	-49 57	5.00	K1.5 Ib-II
-	-	73273	8 35.4	-51 49	7.6:	K1 III: CN0.5
σ Hya	3418	73471	8 38.8	+ 3 21	4.43	K1 III
-	-	73598	8 39.8	+19 33	6.59	G8 III
+ 39 Cnc	3427	73665	8 40.1	+20 00	6.39	G8 III*
+	3428	73710	8 40.3	+19 41	6.44	G9 III*
-	-	73884	8 38.7	-47 47	7.85	K2.5 Iab-Ib Cn1
-	-	73974	8 41.8	+19 53	6.90	G8 III*
+ 9 Hya	3441	74137	8 41.7	-15 57	4.86	G9.5 III
(31 Mon)	3459	74395	8 43.7	- 7 14	4.62	G1 Ib
+ δ Cnc	3461	74442	8 44.7	+18 09	3.94	K0 IIIb
+ ι Cnc	3475	74739	8 46.7	+28 46	4.03	G8 II-III
d Vel	3477	74772	8 44.4	-42 39	4.05	G6 II-III
-	-	74868	8 44.8	-44 33	6.56	F9 III-IV
+ 12 Hya	3484	74918	8 46.4	-13 33	4.32	G8 III Fe-1
-	-	75022	8 46.6	-29 45	7.6	K2+ IIb
-	-	75156	8 48.7	+12 33	6.61	M3 II-III
γ Pyx	3518	75691	8 50.5	-27 42	4.01	K2.5 III
+ 6 UMa	3531	75958	8 56.6	+64 36	5.58	G6 III
+	3538	76151	8 54.3	- 5 27	5.99	G2 V
+ ζ Hya	3547	76294	8 55.4	+ 5 57	3.11	G9 IIIa
BD+6° 2063	-	--	8 54.6	+ 6 09	9.0:	M2 IIIb:
τ p UMa	3576	76827	9 02.6	+67 38	4.76	M3 IIIb Cal
+	3577	76830	8 59.2	+18 08	6.38	M4+III
-	3585	77087	8 59.3	-28 48	6.25	G9 IIIa
-	-	77247	9 03.6	+53 06	7.1:	G7 III: Ba0.5
-	3612	77912	9 06.5	+38 27	4.56	G7 Ib-II
ξ Cnc	3627	78515	9 09.4	+22 03	5.14	G9 IIIa Fe-0.5 CH-1
+ κ Pyx	3628	78541	9 08.0	-25 51	4.58	K4 III
+ λ Vel	3634	78647	9 08.0	-43°26	2.21	K4.5 Ib
17 UMa	3662	79354	9 15.8	+56 44	5.27	K4 III
-	3692	80108	9 16.4	-44 15	5.08-5.14	K3 Ib-II
-	3696	80230	9 16.2	-57 32	4.34	M0.5 III Ba0.3
+	-	80431	9 18.7	-34 06	7.46:	M4 III
+ α Lyn	3705	80493	9 21.0	+34 24	3.13	K7 IIiab
-	-	80721	9 20.2	-36 57	7.8:	K2.5 III:
+ θ Pyx	3718	80874	9 21.5	-25 58	4.72	M0.5 III
+ κ Leo	3731	81146	9 24.7	+26 11	4.46	K2 III
λ Pyx	3733	81169	9 23.2	-28 50	4.69	G8.5 III Fe-1.5
-	3741	81567	9 26.4	- 1 28	6.01	K2+ III
+ α Hya	3748	81797	9 27.6	- 8 40	1.98	K3 II-III
+ δ Hya	3749	81799	9 27.3	-22 20	4.69	K2.5 III
+	3751	81817	9 37.1	+81 20	4.28	K3 IIIa
-	-	81893	9 26.1	-59 46	7.4	K1.5 IIIa
+ 24 UMa	3771	82210	9 34.5	+69 50	4.56	G5 III-IV
+ 6 Leo	3779	82381	9 32.0	+ 9 43	5.08	K2.5 IIIb Fe-0.5
+ λ Leo	3773	82308	9 31.7	+22 58	4.31	K4.5 IIIb
+ ξ Leo	3782	82395	9 31.9	+11 18	4.96	G9.5 III
+ 10 LMi	3800	82635	9 34.2	+36 23	4.55	G7.5 III Fe-0.5
+ N Vel	3803	82668	9 31.2	-57 02	3.13	K5 III
+	3809	82741	9 35.1	+39 37	4.80	G9.5 III Fe-1
+ 11 LMi	3815	82885	9 35.7	+35 49	5.41	G8 IV-V
+ RR Ant	-	83199	9 35.6	-39 54	9.1v	M5+ to M5.5 II
-	3842	83548	9 38.0	-43 11	5.49	G8 IIIa Ba1
-	-	83564	9 41.3	+55 52	6.50	K1 IIIb Fe0.5
+ ι Hya	3845	83618	9 39.8	- 1 08	3.89	K2.5 III
G2 364	3862	84117	9 42.2	-23 55	4.93	F9 V
ϵ Leo	3873	84441	9 45.8	+23 46	2.97	G1 II
-	-	84678	9 41.8	-75 41	9.00	K2.5 Ba3 CH1 C ₂ -1- Ca-1

TABLE 2—Continued

Name	BS	HD	$\alpha(2000)$	$\delta(2000)$	V	Type
			h m			
+ 3881	3881	84737	9 48.6	+46°01'	5.10	G0.5 Va
+ ν^1 Hya	3903	85444	9 51.5	-14 51	4.10	G6.5 IIIa
μ^1 Leo	3905	85503	9 52.8	+26 01	3.88	K2 III CN1 Cal
+ 3919	3919	85859	9 54.2	-25 56	4.88	K2.5 III
- 3912	3912	85622	9 51.7	-46 33	4.57	G5 Ib-II
- 3922	3922	85945	9 57.2	+57 25	5.93	G6 III Fe-0.5
+ RR Car	-	86655	9 58.1	-58 52	7.6-8.5	M6.5S= S6.5/1-
π Leo	3950	86663	10 00.2	+ 8 02	4.70	M2- IIIab
+ 20 LM1	3951	86728	10 01.0	+31 55	5.36	G3 Va H ₂ 1
BD+14° 2203	-	--	10 07.4	+13 56	9.0:	K3.5 IIb Fe1 CH0.5
- 3977	3977	87808	10 07.2	-17 08	5.60:	K5 IIb
31 Leo	3980	87837	10 07.9	+10 00	4.37	K3.5 IIb Fe-1:
- 88009	-	88009	10 09.2	+18 32	7.1:	G8+ IIIa CN1
+ Gl 380	-	88230	10 11.6	+49 27	6.60	K6e V Fe-1
λ Hya	3994	88284	10 10.6	-12 22	3.61	K0 III CN0.5
+ 4006	4006	88639	10 13.8	+27 08	6.04	G3 III Fe-1*
+ 35 Leo	4030	89010	10 16.5	+23 30	5.97	G1.5 IV-V
+ Q Car	4050	89388	10 17.1	-61 20	3.38	K2.5 II
γ Leo C	-	--	10 19.6	+19 52	9.43	M4 Ve*
γ Leo A	4057	89484	10 20.0	+19 51	1.99:	K1- IIIb Fe-0.5
γ Leo B	4058	89485	10 20.0	+19 51	3.2:	G7 III Fe-1.5
π UMa	4069	89758	10 22.3	+41 30	3.03	M0 III
+ EV Car	-	89845	10 20.4	-60 28	7.7-7.9	M4.5 Ia
- 89848	-	89848	10 22.4	-29 32	7.7:	F8 II: CH2
- 4092	4092	90362	10 25.7	- 7 03	5.57	K6: IIIb Fe-0.5 *
+ CK Car	-	90382	10 24.4	-60 11	7.2-7.6	M3.5 Iab
+ μ Hya	4094	90432	10 26.1	-16 51	3.82	K4+ III
Gl 392	4098	90508	10 28.1	+48 47	6.45	F9- V
+ β LM1	4100	90537	10 27.9	+36 42	4.20	G9 IIab
- 90586	-	90586	10 26.4	-53 53	7.02	M2- Iab
+ 91093	-	91093	10 29.2	-57 59	8.31	M3 Iab
- 4126	4126	91190	10 35.1	+75°43'	4.84	G8 III
+ 46 Leo	4127	91232	10 32.2	+14 08	5.46	M1 IIIb
+ 48 Leo	4146	91612	10 34.8	+ 6 57	5.07	G8.5 III Fe-1
- 91629	-	91629	10 33.5	-59 25	7.9	G0 Iab
- 4154	4154	91805	10 35.2	-43 40	6.08	G7 IIIb: CH-3
+ 91810	-	91810	10 37.3	+56 26	6.55	K1 IIIb CN1.5 Cal
CPD-57° 3502	-	--	10 35.7	-58 15	7.3-7.4	M1.5 Iab-Ib
+ 37 LM1	4166	92125	10 38.7	+31 59	4.67	G2.5 IIa
ϕ Hya	4171	92214	10 38.6	-16 52	4.91	G8 III Fe-0.5
- 92501	-	92501	10 40.0	-49 16	7.0:	K5 II-III
+ 4181	4181	92523	10 43.1	+69 05	5.00	K3 III
+ 4184	4184	92620	10 42.2	+31 49	6.02	M3.5 III
RT Car	-	--	10 44.8	-59 17	8.36	M2+ Iab
B0 Car	-	93420	10 45.8	-59 29	7.2-8.5	M4 Ib
ν Hya	4232	93813	10 49.6	-16 12	3.11	K1.5 IIIb H ₂ -0.5
+46 LM1	4247	94264	10 53.3	+34 13	3.81	K0+ III-IV
+ 4255	4255	94481	10 54.3	-13 46	5.65	G4 III
- 94599	-	94599	10 54.1	-61 06	8.15	M5 Iab
+ BZ Car	-	94613	10 54.1	-62 03	7.71	M3+ Ib
+ 56 VY Leo	4267	94705	10 56.0	+ 6 11	6.0var.	M5.5 III
+ 47 UMa	4277	95128	10 59.5	+40 26	5.05	G1- V Fe-0.5
τ Crt	4287	95272	10 59.8	-18 18	4.06	K0+ III
58 Leo	4291	95345	11 00.6	+ 3 37	4.89	K0.5 III Fe-0.5
+ 61 Leo	4299	95578	11 01.8	- 2 29	4.74	M0 III
+ α UMa	4301	95689	11 03.7	+61 54	1.79	K0- IIIa
+Gl 411	-	95735	11 03.5	+36 02	7.51	M2+ V
- 4305	4305	95808	11 03.2	-11 18	5.6	G7-III
- 4328	4328	96700	11 07.9	-30 11	6.54	G1 V Fe-0.5
- 96746	-	96746	11 08.2	-32 05	9.1:	G2 Iab-Ib
- 96765	-	96765	11 08.6	-15 34	8.8:	G3: III: CH2 Fe -1.5 H ₂ -1*

TABLE 2—Continued

Name	BS	HD	$\alpha(2000)$	$\delta(2000)$	V	Type
			h m			
+ψ UMa	4335	96833	11 09.7	+44 29	3.01	K1 III
+x Car	4337	96918	11 08.6	-58 59	3.91	G4 O-Ia
-	-	97671	11 13.5	-60°06'	8.39	M3 Ia-Iab
+72 Leo	4362	97778	11 15.2	+23 05	4.62	M3 IIb
+75 Leo	4371	98118	11 17.3	+2 01	5.17	M0 III
+ξ UMa B	4374	98230	11 18.2	+31 32	4.87:	G2 V
+ξ UMa A	4375	98230	11 18.2	+31 32	4.41:	F8.5 V
+ν UMa	4377	98262	11 18.5	+33 05	3.49	K3- III
-	4379	98282	11 17.3	-67 49	6.05	M2+ S IIIa:
δ Crt	4382	98430	11 19.3	-14 47	3.56	G9 IIb CH+0.2
+	-	98817	11 19.4	-60 42	8.30	M1 Iab-Ib
+56 UMa	4392	98839	11 22.8	+43 29	4.98	G8 II-III
79 Leo	4400	99055	11 24.0	+1 24	5.38	G8 III Fe-0.5
-	-	99154	11 24.1	-57 04	8.6:	G2: II: Fe-0.5:
+τ Leo	4418	99648	11 27.9	+2 51	4.96	G7.5 IIIa
+λ Dra	4434	100029	11 31.4	+69 20	3.83	M0 III Ca-1
+87 Leo	4432	99998	11 30.3	-3 00	4.76	K3.5 III Fe-1
+o ¹ Cen	4441	100261	11 31.8	-59 26	5.10	G3 O-Ia
ξ Hya	4450	100407	11 33.0	-31 51	3.54	G7 III
+ν Leo	4471	100920	11 37.0	-0 49	4.30	G8+ IIb
-	-	100930	11 36.6	-61 20	8.1	M2.5 Iab-Ib
-	-	101007	11 37.0	-61 10	7.9:	M3- Ib + B*
+	4474	101013	11 37.9	+50 37	6.14	G9 III: Ba2.5
+	4491	101370	11 39.8	-16 37	6.5	M3.5 III Cal
+61 UMa	4496	101501	11 41.0	+34 12	5.33	G8 V
+V 810 Cen	4511	101947	11 43.5	-62 29	5.05	G0 O-Ia Fe1
ζ Crt	4514	102070	11 44.8	-18 21	4.72	G8 IIIa
-	-	102155	11 44.9	-73 09	6.93:	K0.5 IIb
+ν Vir	4517	102212	11 45.8	+6 31	4.05	M1 III
+χ UMa	4518	102224	11 46.0	+47 47	3.70	K0.5 IIb
+	4521	102328	11 46.9	+55°38'	5.27	K2.5 IIb CN1
G& 442A	4523	102365	11 46.5	-40 30	4.91	G3 V
+	4532	102620	11 48.8	-26 45	5.11	M4+ III
+	4544	102928	11 51.0	-5 20	5.63	K0+ III
1 Com	-	104452	12 01.7	+22 06	6.6:	G0 II: Fe-1
+o Vir	4608	104979	12 05.2	+8 44	4.14	G8 IIIa CN-1 Ba1 CH1
-	4609	104985	12 05.3	+76 74	5.80	G8.5 IIb
ε CrV	4630	105707	12 10.1	-22 37	2.99	K2.5 IIIa
+5 Com	4643	106057	12 12.2	+20 33	5.57	G8 III CN-0.5 CH0.5
-	4665	106677	12 15.7	+72 34	6.29	K0 IIb*
+2 CVm	4666	106690	12 16.1	+40 40	5.66	M0.5 III
7 Com	4667	106714	12 16.3	+23 57	4.94	G8 III Fe-0.5
-	4668	106760	12 16.5	+33 03	4.99	K0.5 IIb
-	-	107003	12 18.4	-20 49	7.7	M2 III
-	4693	107325	12 20.3	+26 37	5.56	K3 III Ca 0.5
+16 Vir	4695	107328	12 20.4	+3 19	4.96	K0.5 IIb Fe-0.5
+11 Com	4697	107383	12 20.7	+17 48	4.74	G8+ III
-	-	107541	12 21.8	-34 40	9.3	K1: Ba2+ CN1 CH2 Ca-2 Fe-1=C3,1-
+5 CVn	4716	107950	12 24.0	+51 34	4.77	G6 III
-	-	108164	12 25.7	-28 24	8.4:	K2.5 IIb CN0.5
+6 CVn	4728	108225	12 25.8	+39 01	5.01	G9 III
γ Com	4737	108381	12 26.9	+28 16	4.35	K1 III Fe0.5
+	4742	108477	12 27.8	-16 38	6.35	G3 III
-	4755	108759	12 30.0	-41 44	6.0	M3- IIIa
+BK Vir	-	108849	12 30.4	+4 25	7.9-8.7	M7- III:
+γ Crv	4763	108903	12 31.2	-57 07	1.66	M3.5 III
+	-	109011	12 31.2	+55 07	8.11	K2 V
+β CVn	4785	109358	12 33.8	+41 21	4.27	G0 V
+β Crv	4786	109379	12 34.4	-23 24	2.66	G5 IIb
-	-	109534	12 35.6	-29 52	9.3	K1.5 III CN1

TABLE 2—Continued

Name	BS	HD	$\alpha(2000)$	$\delta(2000)$	V	Type
			h m			
+ 6 Dra	4795	109551	12 34.7	+70°01'	4.94	K2.5 III CN-2 H ₈ 1 Fe-1
†	4807	109896	12 38.4	+ 1 51	5.71	M3+ IIIb Ca0.5
+ χ Vir	4813	110014	12 39.2	- 8 00	4.64	K2 III CN1.5
26 Com	4815	110024	12 39.1	+21 04	5.46	G8.5 III
† SV Crv	-	111499	12 49.8	-15 05	7.0	M4.5 II
†	-	111680	12 51.2	-30 28	9.1:	K1-III Fe1
+ 31 Com	4883	111812	12 51.7	+27 32	4.94	G0 IIIp*
† CPD-59°4547	-	--	12 53.7	-60 21	7.6-7.7	M2- Iab
35 Com	4894	112033	12 53.3	+21 14	4.93	G7 III + F:
†	-	112127	12 54.0	+26 46	6.92	K2.5 III: CN3 Ba-0.5 C ₂ 0.5
† ψ Vir	4902	112142	12 54.4	- 9 33	4.80	M3- III Ca-1
TU CVn	4909	112264	12 55.0	+47 12	6.1var.	M5- III
† δ Vir	4910	112300	12 55.6	+ 3 23	3.38	M3+ III
+ 36 Com	4920	112769	12 58.9	+17 25	4.78	M1- IIIb
+ 37 Com	4924	112989	13 00.3	+30 47	4.89	G9 III CH-2
† ϵ Vir	4932	113226	13 02.2	+10 58	2.83	G8 IIIab
-	-	113801	13 06.4	-20 04	8.7:	C1, 0.5=G9 III CN4 C ₂ 0.5
-	-	113842	13 07.4	-60 16	7.3:	M1+ Ib-II
41 Com	4954	113996	13 07.2	+27 38	4.84	K5- III
49 Vir	4955	114038	13 07.9	-10 44	5.19	K0 III Cal
† β Com	4983	114710	13 11.9	+27 53	4.26	F9.5 V
†	4991	114873	13 14.0	-43 09	6.15	K4 III
55 Vir	4995	114946	13 14.2	-19 56	5.63:	G7 IV-V
-	-	114960	13 14.0	+ 1 27	6.59	K4 III CN0.5
SW Vir	-	114961	13 14.1	- 2 48	7.4-8.5	M7 III:
-	-	114988	13 13.7	+32 32	6.67	G1 II Fe-1
-	4997	115004	13 13.7	+40 09	4.92	G8.5 IIIa CN0.5
†	-	115043	13 13.7	+56 40	6.85	G1 Va
† 57 Vir	5001	115202	13 16.0	-19 57	5.21	K1+ III-IV
-	-	115506	13 18.1	-34 32	8.9:	G7.5 IV
† σ Vir	5015	115521	13 17.6	+ 5°28'	4.79	M1 III
61 Vir	5019	115617	13 18.4	-18 19	4.74	G6.5 V
† γ Hya	5020	115659	13 18.9	-23 11	3.00	G8 IIIa
†	5058	116713	13 26.1	-39 45	5.09	K0.5 III: Ba3
-	-	116931	13 27.6	-39 22	9.6:	G7 IIIa CH1
† 69 Vir	5068	116976	13 27.5	-15 58	4.75	K0- IIIb CN1 CH-0.5
† 70 Vir	5072	117176	13 28.4	+13 47	4.97	G4 V
†	5091	117566	13 27.0	+78 39	5.77	G2.5 IIIb CH1
† 74 Vir	5095	117675	13 32.0	- 6 15	4.68	M2.5 III
75 Vir	5099	117789	13 32.8	-15 22	5.55	K1.5 IIIb
†	-	117877	13 33.2	+ 5 51	6.8	G8 IIIa Fe-1.5
EQ Vir	-	118100	13 34.6	- 8 20	9.36	K5 Ve H,Kem 2
† 82 Vir	5150	119149	13 41.6	- 8 42	5.00	M1.5 III
† 83 UMa	5154	119228	13 40.8	+54 41	4.66	M2 IIIab Ba0.5
-	5161	119458	13 42.7	+34 59	5.98	G6 IIIb Fe-0.5
† 83 Vir	5165	119605	13 44.5	-16 11	5.59	G0 Ib-IIa
†	5171A	119796	13 47.2	-62 35	6.80	K0 0-Ia
† G λ 526	-	119850	13 45.5	+14 57	8.48	M2 V
† 87 Vir	5181	120052	13 47.4	-17 52	5.44	M2 III
-	5188	120213	13 55.7	-82 40	5.95	K2.5 IIIa: CH-3 Ba0.5
† 2 Cen	5192	120323	13 49.4	-34 27	4.19	M4.5 III
89 Vir	5196	120452	13 49.9	-18 08	4.97	K0.5 III
† G λ 529	-	120467	13 49.9	-22 06	8.15	K6 Va
† υ Boo	5200	120477	13 49.5	+15 48	4.06	K5.5 III
-	5209	120690	13 51.3	-24 23	6.45	G5 Va
-	5219	120933	13 51.8	+34 27	4.74	M3- III
† 10 Dra	5226	121130	13 51.4	+64 43	4.60	M3.5 III
† η Boo	5235	121370	13 54.7	+18 24	2.68	G0 IV
†	-	121447	13 55.9	-18 18	7.81	K4 III: Ba4.5 Ca0.5*
†	5241	121474	13 57.6	-63 41	4.71	K1.5 III

TABLE 2—*Continued*

Name	BS	HD	$\alpha(2000)$	$\delta(2000)$	V	Type
			h m			
+ 9 Boo	5247	121710	13 56.6	+27°30'	5.01	K3 IIIb
+ θ Aps	5261	122250	14 05.3	-76 48	5.8-6.7	M6.5 III:
-	5270	122563	14 02.5	+ 9 41	6.18	G8: II: Fe-5
+ π Hya	5287	123123	14 06.4	-26 41	3.27	K2- III Fe-0.5
θ Cen	5288	123139	14 06.7	-36 23	2.06	K0- IIIb
-	5297	123569	14 09.9	-53 27	4.75	G9- III
+	5299	123657	14 07.9	+43 51	5.25	M4.5 III*
+ 13 Boo	5300	123782	14 08.3	+49 28	5.24	M1.5 III
-	-	123821	14 08.5	+51 36	8.6:	C1,0.5=G7.5 III: CN3
						C ₂ 0.5
+	5301	123934	14 10.8	-16 18	4.91	M2 If1a
κ Vir	5315	124294	14 12.9	-10 17	4.18	K2.5 III Fe-0.5
+ EV Vir	-	124304	14 13.1	-13 52	6.7to7.1	M4 to M4.5 II-III
4 UMi	5321	124547	14 08.8	+77 33	4.82	K3- IIIb Fe-0.5
-	-	124752	14 12.7	+67 35	8.54	K0.5 V
δ Oct	5339	124882	14 26.9	-83 40	4.32	K2 III
+ α Boo	5340	124897	14 15.7	+19 11	-0.04	K1.5 III Fe-0.5*
V418 Cen	-	125332	14 21.3	-64 15	8.0 var.	K4 II
+	5361	125351	14 18.0	+35 30	4.80	K0 III
ν Vir	5366	126454	14 19.5	- 2 16	5.13	G9 III CN-0.5
51 Hya	5381	125932	14 23.1	-27 46	4.77	K4 III
+ RX Boo	-	126327	14 24.2	+25 42	7.0-9.1	M7.5-M8*
+ ϕ Vir	5409	126868	14 28.2	- 2 14	4.80	G2 IV
24 Boo	5420	127243	14 28.6	+49 51	5.59	G4 III-IV Fe-1
+ ρ Boo	5429	127665	14 31.8	+30 23	3.57	K3- III
+ 5 UMi	5430	127700	14 27.5	+75 41	4.26	K4- III
-	-	127760	14 31.0	+58 36	8.29	K2 IIIb CNO.5:
-	-	128598	14 38.2	-14 57	8.79	K1.5 III Fe-1
+ α Aps	5470	129078	14 47.8	-79 02	3.83	K3 III CNO.5
+ 31 Boo	5480	129312	14 41.6	+ 8 09	4.86	G7 IIIa
+ c ¹ Cen	5485	129456	14 43.7	-35 10	4.05	K3 IIIb
+ 34 Boo	5490	129712	14 43.4	+26°31'	4.82	M3- III
+ $\circ$ Boo	5502	129972	14 45.2	+16 58	4.60	G8.5 III
+ ϵ Boo A	5506	129989	14 45.0	+27 05	2.37	K0- II-III
-	-	130255	14 47.0	+ 1 24	8.4	G9 III CN-3 CH1 Ba0.5 or G9 V Ba2
-	5518	130325	14 47.9	-12 50	6.35	K1 III CNO.5
58 Hya	5526	130694	14 50.3	-27 58	4.41	K2.5 IIIb Fe-1:
-	-	130705	14 49.5	+10 03	6.7:	K3 IIIb: Fe2
R Aps	5540	131109	14 57.9	-76 39	5.5-6.1	K4 III:
G λ 567	5553	131511	14 53.4	+19 09	6.04	K0.5 V
+ β UMi	5563	131873	14 50.7	+74 09	2.07	K4- III
G λ 570A	5568	131977	14 57.5	-21 25	5.74	K4 V
+ G λ 570B	-	131976	14 57.3	-21 21	8.10	M1.5 V
+ 18 Lib	5582	132345	14 58.9	-11 09	5.91	K3- III CN2
+	5584	132525	14 59.4	+ 4 34	5.92	M2.5 III
RR UMi	5589	132813	14 57.6	+65 56	4.58:	M4.5 III
+	5594	132933	15 01.8	-00 09	5.70	M0.5 IIb
-	-	132935	15 02.1	- 8 21	6.85:	K2 III
ω Boo	5600	133124	15 02.1	+25 00	4.80	K4 III
+ 110 Vir	5601	133165	15 02.9	+ 2 05	4.40	K0+ IIIb Fe-0.5
β Boo	5602	133208	15 02.0	+40 23	3.49	G8 IIIa Fe-0.5
+ σ Lib	5603	133216	15 04.1	-25 17	3.29	M2.5 III*
-	-	133507	15 07.2	-57 27	7.20	K1.5 III CNO.5
+ ψ Boo	5616	133582	15 04.4	+26 57	4.52	K2 III
G λ 579.2 A	-	134439	15 10.3	-16 21	9.08	K1 VI: Fe-1*
+ ν Lib	5622	133774	15 06.6	-16 15	5.20:	K5- III
-	5631	134047	15 07.7	+ 5 30	6.16	G7 IIIa: Fe-0.5:
+	5635	134190	15 06.3	+54 33	5.24	G8 III Fe-1
+ ξ Lup	5649	134505	15 12.3	-52 06	3.41	G8 III
+ δ Boo	5681	135722	15 15.5	+33 19	3.49	G8 III Fe-1
2 Lup	5686	135758	15 17.8	-30 09	4.33	K0- IIIa CH-1

TABLE 2—Continued

Name	BS	HD	$\alpha(2000)$	$\delta(2000)$	V	Type
			h m			
-	-	135902	15 20.2	-57°24'	6.9	M3.5 IIIa
-	5692	136138	15 18.4	+20 34	5.70	G8 IIIa
+ ϕ^1 Lup	5705	136422	15 21.8	-36 16	3.56	K4 III
32 ζ^1 Lib	5743	137744	15 28.3	-16 43	5.64:	K4.5 III
+ ι Dra	5744	137759	15 24.9	+58 58	3.29	K2 III
+	5757	138289	15 39.3	-77 55	6.2	K2.5 IIIb CN1.5 Ba-0.5
+ ν^1 Boo	5763	138481	15 30.9	+40 50	5.03	K4.5 IIIb Ba0.5
+ 37 Lib	5777	138716	15 34.2	-10 04	4.61	K1 III-IV
γ Lib	5787	138905	15 35.5	-14 47	3.90	G8.5 III
+ ν Lib	5794	139063	15 37.0	-28 08	3.58	K3.5 III
+ μ CrB	5800	139153	15 35.2	+39 01	5.11	M1.5 IIIb
16 Ser	5802	139195	15 36.5	+10 01	5.26	K0 III: CN1 Ba0.7 Sr2
17 τ^b Ser	-	139216	15 36.5	+15 06	6.2-6.9	M5 var. IIIa
+ ϕ Boo	5823	139641	15 37.8	+40 21	5.24	G7 III-IV Fe-2
+ 42 Lib	5824	139663	15 40.3	-23 50	4.95	K3- III CN2
θ UMi	5826	139669	15 31.4	+77 21	4.96	K5- III CNO.5
+	5831	139862	15 40.2	+12 04	6.25	G7.5 IIIa Fe-0.5
+ κ Lib	5838	139997	15 42.0	-19 40	4.73	M0- IIIb
+ ψ Ser	5853	140538	15 44.0	+ 2 31	5.87	G2.5 V
+ α Ser	5854	140573	15 44.3	+ 6 25	2.64	K2 IIIb CN1
-	5862	140861	15 47.4	-40 12	6.42:	G5 III CH1
+ λ Ser	5868	141004	15 46.4	+ 7 21	4.43	G0- V
+ κ Ser	5879	141477	15 48.7	+18 08	4.10	M0.5 IIIab
+ δ CrB	5889	141714	15 49.6	+26 04	4.62	G5 III-IV Fe-1
+ κ TrA	5891	141767	15 55.5	-68 36	5.09	G5 Ib-II
+ ρ Ser	5899	141992	15 51.2	+20 59	4.78	K4.5 III
+ κ CrB	5901	142091	15 51.2	+35 39	4.80	K1 IVa
+ ST Her	-	142143	15 50.8	+48 29	6.8-8.5	M6S to M7S III:
θ Lib	5908	142198	15 53.8	-16 44	4.14	G9 IIIb
+ 39 Ser	5911	142267	15 53.2	+13 12	6.10	G1 V Fe-1
χ Her	5914	142373	15 52.7	+42°27'	4.62	F8 V Fe-2 H ₀ -1
-	5924	142574	15 54.6	+20 18	5.44	K8 IIIb
2 Her	5932	142780	15 54.6	+43 09	5.35	M3- III
-	-	142804	15 57.1	-16 02	6.56	M1.5 III
+ ϵ CrB	5947	143107	15 57.6	+26 53	4.14	K2 IIIab
+	5955	143346	16 05.9	-72 24	5.70	K1.5 III CN1
5 Her	5966	143666	16 01.2	+17 49	5.12	G8 IIIb
+ ρ CrB	5968	143761	16 01.0	+35 18	5.41	G0+ Va Fe-1
+	5995	144542	16 03.2	+59 25	6.19	M1 III
-	5996	144585	16 07.1	-14 04	6.31	G2 V
+ ω^2 Sco	5997	144608	16 07.4	-20 52	4.32	G4 II-III
G ξ 613	-	144628	16 09.7	-56 27	7.12	K2 V
-	-	144873	16 06.7	+34 06	8.6	G2.5 Vb
+	6014	145148	16 09.2	+ 6 23	5.95	K1.5 IV
+ τ CrB	6018	145328	16 09.0	+36 30	4.76	K1- III-IV
+ δ TrA	6030	145544	16 15.4	-63 41	3.85	G2 Ib-IIa
14 Her	-	145675	16 10.2	+43 49	6.66	K0 V
+ δ Oph	6056	146051	16 14.3	- 3 41	2.75	M0.5 III
-	-	146116	16 14.6	-00 24	7.8	G7 III CH-2 H ₀ -0.5
+ γ^1 Nor	6058	146143	16 17.0	-50 04	4.99	F9 Ia
+ 18 Sco	6060	146233	16 15.6	- 8 22	5.50	G2 Va
+ ϵ Oph	6075	146791	16 18.3	- 4 42	3.23	G9.5 IIIb Fe-0.5
+	6087	147266	16 20.1	+21 09	6.04	G8 III
+ G ξ 617A	-	147379	16 16.7	+67 14	8.62	M1-e V
ξ CrB	6103	147677	16 22.1	+30 54	4.85	G9 III Fe-0.5
ψ Oph	6104	147700	16 24.1	-20 02	4.49	K0- II-III
+ ν^1 CrB	6108	147767	16 22.5	+33 42	5.39	K5 III
-	6122	148247	16 28.2	-37 11	5.8	K1+ III CNO.5
-	6125	148291	16 30.8	-61 38	5.20:	K1 IIb
+	6128	148349	16 27.7	- 7 36	5.24	M3- III*

TABLE 2—Continued

Name	BS	HD	$\alpha(2000)$	$\delta(2000)$	V	Type
			h m			
+ η Dra	6132	148387	16 24.0	+61°30'	2.73	G8- IIIIab
+ α Sco	6134	148478	16 29.4	-26 26	1.1:	M1.5 Iab-Ib
+	6135	148488	16 34.3	-70 59	5.50	K0.5 IIb CN1
+ 30 g Her	6146	148783	16 28.6	+41 53	5.0var.	M6- III
ϕ Oph	6147	148786	16 31.1	-16 37	4.28	G8+ IIIa
+ β Her	6148	148856	16 30.2	+21 29	2.78	G7 IIIa Fe-0.5
+	6152	148897	16 30.6	+20 29	5.25	G8.5 III CN-2 CH-1 Fe-1
29 Her	6159	149161	16 32.6	+11 29	4.84	K4.5 III
+	6196	150416	16 41.6	-17 45	4.96	G7.5 II-III CN1 Ba0.5
+ ζ Her	6212	150680	16 41.3	+31 36	2.80	G0 IV
η Her	6220	150997	16 42.9	+38 56	3.50	G7 III Fe-1
+	-	151061	16 45.2	- 3 05	7.25:	M5 to M5.5 IIb:
+ 18 Dra	6223	151101	16 40.9	+64 36	4.83	K0III CN-0.5 CH-2 Ca1
+ G λ 638	-	151288	16 45.1	+33 30	8.10	K7.5 Ve
+	6242	151732	16 47.3	+42 14	5.98	M4.5 III
51 Her	6270	152326	16 51.8	+24 39	5.04	K0.5 IIIa Ca0.5
+ κ Oph	6299	153210	16 57.7	+ 9 23	3.20	K2 III
+ ϵ UMi	6322	153751	16 46.0	+82 02	4.28	G5 III
(32 Oph)	6337	154143	17 03.1	+14 05	4.98	M3- III
+	6349	154417	17 05.3	+00 42	6.00	F8.5 IV-V
+	-	154430	17 10.9	-59 17	8.74	K3- IIIa Ba3-
+	6364	154733	17 06.3	+22 05	5.58	K3 III
+	6382	155276	17 12.3	-38 49	6.30	K0.5 III CN0.5
+	6392	155603	17 14.5	-39 46	6.56	K0 0-Ia
+ α^1 Her	6406	156014	17 14.6	+14 23	3.1-3.9	M5 Ib-II
-	6408	156091	17 19.2	-59 42	5.91:	K2 IIIa Ca1 CN1 Ba0.5
+ ζ Aps	6417	156277	17 22.0	-67 46	4.78	K2- III
π Her	6418	156283	17 15.0	+36 48	3.16	K3 II
-	6428	156462	17 18.3	-16 18	6.45	M2- III Ca1
VW Dra	6448	156947	17 16.5	+60 41	6.3-7.0	K1.5 IIb
-	-	157072	17 21.8	-20°34'	7.12	K5 III
-	6479	157681	17 21.8	+53 25	5.75	K4 III
-	6491	157910	17 24.4	+36 57	6.28	G5: III:
+ σ Oph	6498	157999	17 26.5	+ 4 09	4.34	K2 II
-	6516	158614	17 30.4	- 1 04	5.31	G9- IV-V H δ 1
+ λ Her	6526	158899	17 30.7	+26 06	4.40	K3.5 III
+ β Dra	6536	159181	17 30.4	+52 19	2.80	G2 Ib-IIa
-	-	159378	17 35.9	-33 26	8.5:	G3 Ia-Iab CN-1
27 Dra	6566	159966	17 32.0	+68 08	5.06	G9 IIb
+ 26 Dra	6573	160269	17 35.0	+61 52	5.24	G0 IV-V
-	6576	160342	17 42.1	-50 31	6.2:	M3+ II-III
G λ 688	-	160346	17 39.3	+ 3 33	6.52	K3- V
+ BM Sco	-	160371	17 42.0	-32 13	5.8 to 6.2:	K2 Ib H δ 1
-	-	160823	17 41.9	+ 4 22	6.91:	G0: IIIa Fe1
+ G λ 689	-	160964	17 36.1	+71 53	8.62	K4 V
+ β Oph	6603	161096	17 43.5	+ 4 34	2.77	K2 III CN0.5
+ 84 Her	6608	161239	17 43.4	+24 19	5.71	G2 IIb
-	6617	161664	17 47.8	-22 29	6.2:	G6 Ib H δ 1
+ μ Her	6623	161797	17 46.5	+27 44	3.41	G5 IV
87 Her	6644	162211	17 48.8	+25 37	5.12	K0 IIb
+ 90 Her	6677	163217	17 53.3	+40 01	5.15	K1+ IIb CN-1
+ ξ Dra	6688	163588	17 53.5	+56 52	3.74	K2 III
+	6693	163755	17 59.1	-30 16	4.99	M1 Ib
+ θ Her	6695	163770	17 56.2	+37 15	3.84	K1 IIa CN2
ν Oph	6698	163917	17 59.0	- 9 47	3.34	G9 IIIa
+ ξ Her	6703	163993	17 57.8	+29 15	3.70	G8.5 III
+ γ Dra	6705	164058	17 56.6	+51 29	2.23	K5 III
+ 93 Her	6713	164349	18 00.0	+16 45	4.66	K0.5 IIb
+ χ Oct	6721	164461	17 54.8	-87 36	5.28	K3 III
+ γ^2 Sgr	6746	165135	18 05.8	-30 26	2.98	K0+ III

TABLE 2—*Continued*

Name	BS	HD	$\alpha(2000)$	$\delta(2000)$	V	Type
			h m			
70 OphA	6752	165341	18 05.5	+2°30'	4.22	K0- V
+	6756	165438	18 06.2	-4 46	5.76	K1 IV
-	6757	165462	18 06.1	-00 27	6.34	G9 III CH-3
-	-	165553	18 07.3	-21 14	7.58	G0 Ib-II
+ 98 Her	6765	165625	18 06.0	+22 14	5.1	M3- S III
.						
+	6766	165634	18 08.1	-28 27	4.56	G7: IIIb CN-1 CH-3.5 HK+1
-	6769	165687	18 07.8	-17 09	5.52	K1 III
+ 71 Oph	6770	165760	18 07.3	+ 8 44	4.64	G8 III
AX Sgr	-	165782	18 08.4	-18 34	7.4var.	K0 Ia
+	6791	166208	18 07.5	+43 28	4.99	G8 III CN-1 CH-3
.						
+	6793	166229	18 08.8	+36 24	5.48	K2 III
+ 104 Her	6815	167006	18 11.9	+31 24	4.97	M3 III
-	6842	167818	18 18.1	-27 03	4.65	K3 II
-	6853	168322	18 17.1	+40 56	6.12	G8.5 IIIb CH0.5 Fe-1
+ δ Sgr	6859	168454	18 21.0	-29 49	2.70	K2.5 IIIa CNO.5
.						
105 Her	6860	168532	18 19.2	+24 26	5.27	K3 III: Ba0.5
-	6861	168574	18 21.5	-24 55	6.25	M5- IIIa:
106 Her	6868	168720	18 20.3	+21 58	4.96	M1 III
+ η Ser	6869	168723	18 21.3	- 2 53	3.26	K0 III-IV
+ κ Lyr	6872	168775	18 19.8	+36 04	4.32	K2- IIIab CNO.5
.						
-	-	168815	18 22.2	-15 06	7.3:	K5+ II
ζ Sct	6884	169156	18 23.7	- 8 56	4.68	G9- IIIb Fe-0.5
+ 109 Her	6895	169414	18 23.7	+21 46	3.84	K2 IIIab
λ Sgr	6913	169916	18 28.0	-25 26	2.84	K1 IIIb
-	-	170174	18 27.9	+ 6 35	8.33:	G9 III Fe 0.5:*
.						
BD +6° 3796	-	--	18 28.5	+6 45	8.96	G6 IIIb Fe-1*
BD +6° 3798	-	--	18 28.5	+6 43	8.79	G6.5 III Fe-0.5 CH0.5*
+ 42 Dra	6945	170693	18 26.0	+65 34	4.82	K1.5 III Fe-1
-	-	170820	18 32.2	-19 08	7.37	G9 II CN1 H ₂ 1*
-	6959	170975	18 32.7	-14 52	5.9	K3- Ib-II CN1
.						
+ BD+ 36° 3157	-	170970	18 30.6	+36 15	7.62:	M3+=S3+/1- III:
+ α Sct	6973	171443	18 35.2	- 8°15'	3.83	K3 III
BY Dra	-	234677	18 34.8	+51 43	8.4-8.5	K4 Ve + K7.5 Ve
-	-	171911	18 34.7	+51 48	6.67	M4 III-IIIb
+XY Lyr	7009	172380	18 38.1	+39 40	5.8-6.8	M4.5-M5+ II
.						
+ ϵ Sct	7032	173009	18 43.5	- 8 17	4.89	G8 IIb
+ β Sct	7063	173764	18 47.2	- 4 45	4.22	G4 IIa
+	7064	173780	18 46.5	+26 39	4.81	K2 III
35 ν^2 Sgr	7120	175190	18 55.1	-22 41	4.99	K3- II-III: Ba1 CN1
-	7123	175225	18 51.6	+52 59	5.51	G9 IVa
.						
+ σ Dra A	7125	175306	18 51.2	+59 23	4.64	G9 III Fe-0.5
V913 Aql	-	175309	18 54.5	+10 38	7.9var.	M5- to M5.5
-	7137	175535	18 53.2	+50 42	4.92	G7 IIIa Fe-1
+ δ^2 Lyr	7139	175588	18 54.5	+36 54	4.30	M4 II
η Sct	7149	175751	18 57.1	- 5 51	4.83	K1- III
.						
-	7162	176051	18 57.0	+32 54	5.23	F9 V
-	7164	176123	18 59.4	-18 34	6.34:	G3 II
+ ϵ Aql	7176	176411	18 59.6	+15 04	4.02	K1-III CNO.5
ν Dra	7180	176524	18 54.4	+71 18	4.82	K0 III CNO.5
-	7187	176598	18 56.4	-65 15	5.62	G8 III
.						
-	-	176617	19 02.7	-36 40	8.58:	M3 III
+ λ Lyr	7192	176670	19 00.0	+32 08	4.93	K2.5 III: Ba0.5
+ 12 Aql	7193	176678	19 01.7	- 5 44	4.02	K1 III
+ σ Sgr	7217	177241	19 04.7	-21 44	3.77	G9 IIIb
+ 49 Dra	7218	177249	19 00.7	+55 40	5.48	G6 IIb Fe-0.5

TABLE 2—Continued

Name	BS	HD	$\alpha(2000)$	$\delta(2000)$	V	Type
			h m			
+ τ Sgr	7234	177716	19 06.9	-27° 40'	3.32	K1.5 IIb
+ "	7252	178208	19 05.2	+49 55	6.4:	K3 III
-	7260	178428	19 08.0	+16 51	6.09	G4.5 V
+ "	-	178717	19 09.3	+10 13	7.5:	K3.5 III: Ba 4+
V 842 Aql	-	179820	19 13.9	+ 2 36	9.2var.	M6+ III:
-	-	179821	19 13.8	+00 07	8.4:	G4 O-Ia
+ "	-	179870	19 13.9	+ 9°01	7.35	K0 II
43 Sgr	7304	180540	19 17.6	-18 57	5.0	G8 II-III
+ G λ 752A	-	180617	19 16.9	+ 5 09	9.12	M3- V
+ δ Dra	7310	189711	19 12.6	+67 40	3.07	G9 III
+ θ Lyr	7314	180809	19 16.4	+38 08	4.35	K0 II
+ 24 Aql	7321	181053	19 18.8	+00 20	6.40	K0- IIIa: Ba0.5 CH1
+ κ Cyg	7328	181276	19 17.1	+53 22	3.77	G9 III
+ "	-	181475	19 20.8	- 4 31	7.4	K7 IIa
-	-	181596	19 18.5	+50 13	7.62	K5 III
+ τ Dra	7352	181984	19 15.6	+73 21	4.45	K2+ IIIb CN1
-	-	182296	19 23.6	+ 8 40	7.07	G2 Ib-II CN1
G λ 758	7368	182488	19 23.6	+33 13	6.36	K0 V
+ 31 Aql	7373	182572	19 25.0	+11 57	5.19	G7 IV H δ 1
+ "	7382	182694	19 24.0	+43 24	5.84	G7 IIIa
6 α Vul	7405	183439	19 28.7	+24 40	4.45	M0.5 IIb
V 450 Aql	-	184313	19 33.7	+ 5 02	6.3-6.9	M5 to M5.5 III
μ Aql	7429	184406	19 34.1	+ 7 23	4.44	K3- IIIb Fe 0.5
+ "	-	184467	19 31.2	+58 35	6.59	K2 V
+ 37 Aql	7430	184492	19 35.1	-10 34	5.12	G8.5 IIIa
+ "	7456	185018	19 36.9	+11 16	5.98	G0 Ib-II
+ σ Dra	7462	185144	19 32.4	+69 39	4.69	K0 V
+ "	7468	185351	19 36.6	+44 41	5.17	G8.5 IIIb Fe-0.5
-	7475	185622A	19 39.4	+16 35	6.40	K4 Ib
+ "	-	185663	19 39.5	+19 10	7.7	K2 IIIa
+ α Sge	7479	185758	19 40.1	+18 01	4.36	G1 II
+ β Sge	7488	185958	19 41.0	+17 29	4.37	G8 IIIa CNO.5
+ 16 Cyg prec.	7503	186408	19 41.8	+50 32	5.96	G1.5 Vb
+ 16 Cyg fol.	7504	186427	19 41.8	+50 31	6.20	G3 V
56 Sgr	7515	186648	19 46.4	-19 46	4.86	K0+ III
15 Cyg	7517	186675	19 44.3	+37 21	4.90	G8 III
+ V 973 Cyg	7523	186776	19 44.8	+40 42	6.4var.	M3.5 III*
+ γ Aql	7525	186791	19 46.2	+10°37	2.72	K3 II
BD+ 24° 3902	-	--	19 50.2	+24 55	9.3:	M1 Ia:
+ 20 Cyg	7576	188056	19 50.6	+52 59	5.03	K3 III CN2
G λ 770	7578	188088	19 54.3	-23 56	6.20	K3 Va CN1
ϵ Dra	7582	188119	19 48.2	+70 16	3.82	G7 IIIb Fe-1
-	-	350801	19 53.2	+18 21	8.90	G4 Ib*
+ ξ Aql	7595	188310	19 54.2	+ 8 28	4.70	G9.5 IIIb
+ ω Sgr	7597	188376	19 55.8	-26 18	4.70	G5 IV
+ β Aql	7602	188512	19 55.3	+ 6 24	3.71	G8 IV
+ G 208-44	-	--	19 53.8	+44 25	15.4:	M5.5 Ve
+ 59 Sgr	7604	188603	19 57.0	-27 10	4.52	K2.5 IIb
+ "	7606	188650	19 54.8	+37 00	5.76	G1 Ib-II CH2 Fe-1
+ η Cyg	7615	188947	19 59.0	+35 05	3.89	K0 III
60 Sgr	7618	189005	19 59.0	-26 12	4.83	G6 III
-	7633	189276	19 55.9	+58 51	4.96	K4.5 IIIa
+ γ Sge	7635	189319	19 58.8	+19 29	3.48	M0- III
-	-	189337	19 59.2	+11 18	6.5-	G7 Ib-II
62 Sgr	7650	189763	20 02.7	-27 43	4.46	M4.5 III
-	-	190113	20 02.0	+35 38	7.95	G5 Ib
+ BD+29° 3865	-	--	20 02.5	+30 04	9.09	G7 Ia
+ 64 Dra	7676	190544	20 01.5	+64 49	5.23	M1 III
-	-	190788	20 05.8	+25 36	8.2:	M3- Ib-II
+ ρ Dra	7685	190940	20 02.8	+67 52	4.51	K2+ III CN1 Cal

TABLE 2—Continued

Name	BS	HD	$\alpha(2000)$	$\delta(2000)$	V	Type
			h m			
27 Cyg	7689	191026	20 06.4	+35°58'	5.36	G8.5 IVa
†	-	191046	20 06.5	+36 14	7.03	K0 III Fe-1.5
-	-	191511	20 09.1	+33 40	8.16	G5 III Fe-1.5*
19 Vul	7718	192004	20 11.8	+26 49	5.51	K2 IIIa Ca0.5
22 Vul	7741	192713	20 15.5	+23 30	5.17	G3 Ib-II Wk H&K comp.?
† 23 Vul	7744	192806	20 15.8	+27 48	4.52	K3- III Fe-1
6 α^2 CapA	7754	192947	20 18.0	-12 32	3.57	G9 III
-	7759	193092	20 16.9	+40°22'	5.23	K3 IIIa Fe-1
BD+37°3865	-	--	20 16.9	+37 31	8.60	G9.5 IIIb*
-	7762	193217	20 17.5	+42 43	6.28	K4+ II-III
†	7788	193896	20 23.0	- 9 39	6.30	G5 IIIa
†	7800	194193	20 22.8	+41 01	5.93	K7 III
-	-	194241	20 23.1	+40 47	7.36	K1- III-IIIb Ba0.3:*
† AC Dra	7804	194258	20 20.1	+68 53	6.0var.	M4.5 to M5 III
† 39 Cyg	7806	194317	20 23.9	+32 11	4.43	K2.5 III Fe-0.5
† RW Cyg	-	--	20 28.8	+39 59	7.8-8.8	M3 to M4 Ia-Iab
-	-	195338	20 28.7	+47 37	7.48	G8 II Fe0.5
-	7845	195564	20 32.4	- 9 52	5.65	G2.5 IV
†	-	196197	20 34.7	+32 31	6.84	K1+ IIIa
70 Aql	7873	196321	20 36.7	- 2 33	4.89	K4+ III Bal
† 71 Ag1	7884	196574	20 38.3	- 1 06	4.31	G7.5 IIIa
† EU Del	7886	196610	20 37.9	+18 16	6.3var.	M6 III
† κ Del	7896	196755	20 39.1	+10 05	5.05	G2 IV
† G λ 796	7898	196761	20 40.2	-23 46	6.37	G7.5 IV-V
† G λ 795	-	196795	20 39.5	+ 4 58	8.26	K5 V
†	-	196819	20 38.3	+42 05	7.50	K2.5 IIb
† 49 Cyg	7921	197177	20 41.0	+32 18	5.51	G8 IIb
AU Mic	-	197481	20 45.2	-31 20	8.59-8.96	M1 Ve Bal
30 Vul	7939	197752	20 44.9	+25 16	4.91	K1 III*
† 52 Cyg	7942	197912	20 45.7	+30 43	4.20	K0 IIIa
† ϵ Cyg	7949	197989	20 46.2	+33 58	2.46	K0 III
† 3 Aqr	7951	198026	20 47.7	- 5 02	4.42	M3 III
† η Cep	7957	198149	20 45.3	+61 50	3.41	K0 IV
† ω Cap	7980	198542	20 51.8	-26 55	4.11	M0 III Ba0.5
† 31 Vul	7995	198809	20 52.1	+27 05	4.57	G7 III Fe-1
† 33 Vul	8032	199697	20 58.3	+22 20	5.31	K3.5 III
-	8035	199870	20 58.3	+44 28	5.53	K0 III: Fe-0.5
† AZ Cyg	-	--	20 58.0	+46°28'	8.1-9.4	M2 to M4 Iab
†	-	199939	20 59.0	+44 27	8.0:	G9 III: Ba4
†	8062	200527	21 02.4	+44 48	6.19	M4.5 III-S4/1 III
† ξ Cyg	8079	200905	21 04.9	+43 56	3.71	K4.5 Ib-II
† 24 Cap	8080	200914	21 07.1	-25 00	4.50	M1- III
-	-	201065	21 05.6	+46 58	7.56	K4 Ib-II
† 61 CygA	8085	201091	21 06.9	+38 44	5.23	K5 V
† 61 CygB	8086	201092	21 06.9	+38 44	6.02	K7 V
† 63 Cyg	8089	201251	21 06.6	+47 39	4.54	K4 Ib-IIa
† ζ Cyg	8115	202109	21 12.9	+30 14	3.19	G8+ III-IIIa Ba0.5
-	8121	202259	21 14.6	+00 06	6.38	M1-III*
-	8126	202314	21 14.2	+29 54	6.18	G6 Ib-IIa Cal Ba0.5
-	-	202380	21 12.8	+60 07	7.0:	M2 Ib
G λ 825	-	202560	21 17.3	-38 53	6.68	M1 V
†	-	203137	21 18.8	+50 04	6.96	K4+ III Ba0.5
† ι Cap	8167	203387	21 22.2	-16 50	4.28	G7 III Fe-1.5
1 Peg	8173	203504	21 22.1	+19 49	4.08	K1 III
ζ Cap	8204	204075	21 26.7	-22 25	3.74	G4 Ib: Ba2
36 Cap	8213	204381	21 28.7	-21 48	4.50	G7 IIIb Fe-1
†	8223	204585	21 29.0	+22 11	5.93	M4.5 IIIa
†	8224	204599	21 27.4	+59 45	6.2	M3+ IIIa
†	-	204613	21 27.6	+57 18	8.24	G1 IIIa: CH1.5
† 2 Peg	8225	204724	21 30.0	+23 38	4.54	M1+ III

TABLE 2—Continued

Name	BS	HD	$\alpha(2000)$	$\delta(2000)$	V	Type
			h m			
71 Cyg	8228	204771	21 29.4	+46° 32'	5.23	K0-III
+ s Agr	8232	204867	21 31.6	- 5 35	2.88	G0 Ib
-	-	205011	21 31.8	+23 51	6.5:	G9+ IIIa Ba1+
+ o Cyg	8252	205435	21 34.0	+45 36	4.02	G8 III Fe-0.5
72 Cyg	8255	205512	21 34.8	+38 32	4.89	K0.5 III Fe0.5
-	-	205905	21 39.2	-27 18	6.74	G1.5 IV-V
+ 42 Cap.	8283	206301	21 41.5	-14 03	5.13	G2 IV
+ 75 Cyg	8284	206330	21 40.2	+43 10	5.10	M1 IIIab
-	8306	206749	21 43.1	+41 09	5.49	M2 IIIb Cal
+ e Peg	8308	206778	21 44.2	+ 9 53	2.40	K2 Ib-II
+ 46 Cap	8311	206834	21 45.0	- 9 05	5.08	G7.5 II-III CNO.5
+ 9 Peg	8313	206859	21 44.5	+17 21	4.32	G5 Ib
+ u Cep	8316	206936	21 43.5	+58 47	4.0var.	M2- Ia
+ 11 Cep	8317	206952	21 41.9	+71 19	4.54	K0.5 III
EP Aqr	-	207076	21 46.6	- 2 12	6.4-6.8	M7- III:
12 Peg	8321	207089	21 46.1	+22 57	5.29	K0 Ib H ₆ 0.5
+ -	-	207119	21 45.2	+52 16	6.42	K3 II Ba0.5
+ -	8324	207130	21 43.1	+72 19	5.16	K0- III CN1
-	-	207328	21 46.3	+58 13	7.36:	M3 IIIa Ba0.5*
-	-	207585	21 50.5	-24 12	9.80	G2 II: Ba1.5 CH1 Fe-1
-	-	207991	21 52.1	+48 28	6.87	K4- III
Gz 846	-	209290	22 02.2	+ 1 23	9.17	M0.5 V
+ a Aqr	8414	209750	22 05.3	-00 19	2.93	G2 Ib
+ z Cep	8465	210745	22 10.8	+58 12	3.36	K1.5 Ib
+ 24 Cep	8468	210807	22 09.8	+72 21	4.78	G7 II-III
-	8475	210889	22 12.8	+34 36	5.33	K2- IIIb
BD+ 54° 2704	-	--	22 12.6	+55 37	9.1:	G9 III
+ -	8485	211073	22 13.9	+39 43	4.49	K2.5 III
+ 1 Lac	8498	211388	22 16.0	+37 45	4.13	K3- II-III
+ 9 Aqr	8499	211391	22 16.8	- 7 47	4.16	G9 III
+ a Tuc	8502	211416	22 18.5	-60° 15'	2.86	K3 III
DZ Aqr	-	211062	22 21.7	- 7 36	8.6var.	M7- var.
+ 50 Aqr	8534	212430	22 24.4	-13 32	5.76	G7.5 III
+ RW Cep	-	212466	22 23.1	+55 58	6.8-7.5	K2 0-Ia
+ B Lac	8538	212496	22 23.6	+52 14	4.42	G9 IIIb Cal
35 Peg	8551	212943	22 27.9	+ 4 42	4.79	K1- III-IV
+ ST Cep	-	239978	22 30.2	+57 00	8.0-8.9	M2 Ia-Iab
+ -	-	213893	22 34.6	+00 37	6.71	M0 IIIb
+ Aqr	8610	214376	22 37.8	- 4 14	5.02	K1.5 IIIb CNO:5
-	-	214471	22 36.7	+65 16	7.70	K2- IIIb
+ -	8621	214665	22 38.6	+56 48	5.21	M4+ III
+ -	8626	214714	22 39.6	+37 35	6.14	G3 Ib-II: CN-1 CH2 Fe-1
+ -	8631	214850	22 40.8	+14 33	5.72	G4 V
+ 11 Lac	8632	214868	22 40.5	+44 17	4.48	K2.5 III
+ B Gru	8636	214952	22 42.7	-46 53	2.11	M4.5 III
+ λ Peg	8667	215665	22 46.5	+23 34	3.96	G8 IIIa CNO.5
+ Aqr	8679	216032	22 49.6	-13 35	4.01	M0 III
+ μ Peg	8684	216131	22 50.0	+24 36	3.48	G8+ III
-	-	216219	22 50.9	+18 00	7.46	G1 II-III: Fe-1 CH0.5
+ λ Cep	8694	216228	22 49.7	+66 12	3.52	K0- III
λ Aqr	8698	216386	22 52.6	- 7 35	3.76	M2.5 III Fe-0.5
77 Aqr	8711	216640	22 54.8	-16 16	5.56	K2.5 III
TW PsA	8721	216803	22 56.4	-31 34	6.49	K4 Ve
-	8726	216946	22 56.4	+49 44	5.0:	K5 Ib
+ 51 Peg	8729	217014	22 57.5	+20 46	5.44	G2 IV
+ -	-	217143	22 58.6	+ 7 38	8.3	G9.5 III Ba2+
-	8752	217476	23 00.1	+56 57	5.1:	G4var. 0*
-	8761	217673	23 01.5	+57 06	6.20	K1.5 II
+ 8 Peg	8775	217906	23 03.8	+28 05	2.4:	M2.5 II-III*
+ 3 And	8780	218031	23 04.2	+50 03	4.64	K0 IIIb

TABLE 2—*Continued*

Name	BS	HD	$\alpha(2000)$	$\delta(2000)$	V	Type
			h m			
+ 86 Aqr	8789	218240	23 06.7	-23°45'	4.47	G6 IIb
+ 55 Peg	8795	218329	23 07.0	+ 9 24	4.51	M1 IIIab
56 Peg	8796	218356	23 07.1	+25 28	4.77	K0.5 II: Bal CN-2 HKem1 CH-0.5*
5 Psc	8807	218527	23 08.7	+ 2 08	5.39	G6 IIb
+ 88 Aqr	8812	218594	23 09.4	-21 10	3.67	K1.5 III
-	-	218738	23 09.8	+47 57	7.9	K2-Ve*
SS And	-	218942	23 11.5	+52 53	8.9-9.9	M7- II:
+ G λ 892	8832	219134	23 13.3	+57 10	5.56	K3 V
ϕ Aqr	8834	219215	23 14.3	- 6 03	4.22	M1.5 III
ψ Aqr	8841	219449	23 15.9	- 9 05	4.22	K1- III Fe-0.5
61 Peg	8842	219477	23 15.8	+28 15	6.5:	G2 II-III
+ γ Psc	8852	219615	23 17.2	+ 3 17	3.68	G9 III: Fe-2*
+ 8 And	8860	219734	23 17.7	+49 01	4.86	M2.5 III Ba0.5
o Cep	8872	219916	23 18.6	+68 07	4.75	G8.5 III-IIIb Fe-0.5
-	-	219978	23 19.4	+62 44	6.71	K4.5 Ib
98 Aqr	8892	220321	23 23.0	-20 06	3.95	K1 III
99 Aqr	8906	220704	23 26.0	-20 38	4.40	K4.5 III
θ Psc	8916	220954	23 28.0	+ 6 23	4.27	K0.5 III
70 Peg	8923	221115	23 29.2	+12 46	4.53	G8 IIIa
+	8924	221148	23 29.5	- 4 32	6.25	K3- IIb Fe2*
BD+ 48°4065	-	--	23 30.8	+49 15	8.95	K1 III-IV Fe-1*
-	-	221203	23 29.7	+49 07	7.74	G8 III Fe-1.5*
+	8925	221246	23 30.1	+49 08	6.17	K3 III*
+	8952	221861	23 35.0	+71 38	5.71	G9 Ib
+	8958	222093	23 37.7	-13 04	5.65	G9 III
+ γ Cep	8974	222404	23 39.3	+77 37	3.22	K1 III-IV CN1
+	8989	222670	23 42.3	+64 31	6.53	M2 III
PZ Cas	-	--	23 44.0	+61 47	8.6-9.5:	M3var. Ia
ψ And	9003	223047	23 46.0	+46 25	4.94	G3 Ib-II
τ Cas	9008	223165	23 47.1	+58 39	4.87	K1- IIIa
-	9010	223173	23 47.0	+57°27'	5.51	K3- IIb
G λ 908	-	--	23 49.1	+ 2 24	8.98	M1 V Fe-1
+	9021	223428	23 49.5	-15 52	6.24	K2 IIb Fe-0.5
-	-	223541	23 50.5	-13 06	7.1	K0+ III-IV Fe -1
-	9029	223559	23 50.6	-14 24	5.72	K4.5 IIb
+ 22 Psc	9033	223719	23 52.0	+ 2 55	5.55	K4 III Bal
-	9038	223778	23 52.4	+75 32	6.41	K3 V
TZ Cas	-	--	23 53.0	+61 00	9.0-10.5	M3 Iab
ρ Cas	9045	224014	23 54.4	+57 30	4.4-5.1	G2 0 var.*
XZ Psc	9047	224062	23 54.8	+00 06	5.5-6.0	M5 III
II Peg	-	224085	23 55.0	+28 39	7.4-7.6	K2+ IVe Fe-1*
-	-	224364	23 57.3	+61 02	7.0	M2.5 III: Cal
+ ψ Peg	9064	224427	23 57.8	+25 08	4.66	M3 III
+	-	224618	23 59.5	-16 56	8.72	K0 V
+ Sun					-26.8	G2 V

NOTES TO TABLE 2

DL Cas = BD + 59°65.—Cepheid variable ($P = 8^d0$) generally accepted as a member of the cluster NGC 129.
 BD + 59°70.—The star is Kraft AB (Kraft 1958) in the field of the galactic cluster NGC 129. It is a foreground star.

BD + 59°73.—The star is Kraft AA in field of NGC 129. It is a foreground star.

n Cas B = HD 4614B.—All the normally strong lines in the blue region are shallow, but the lines in the red region appear normal.

n And = HD 5516.—This is a two-lined spectroscopic binary, but the spectrum looks normal at low dispersion. The secondary is only slightly fainter than the primary.

BS 316.— $V_r = -94.5$.

HD 6833.—A high-velocity star with $V_r = -248 \text{ km s}^{-1}$.

Ψ^3 Psc = BS 339 = HD 6903.— $V \sin i = 91$. Active chromosphere (Stickland and Williams 1983).

39 Cet = AY Cet = BS 373.—Variable by about 0.2 mag (Olsen 1974). H, K emission.

V466 Cas = BD + 57°258 = HD 236697.—The star is a well-established member of NGC 457.

BD + 59°274.—Probable member of NGC 581.

BD + 60°265.—Possible member of association Cas OB8.

HD 15068.—Possible member of cluster Tr 2.

NOTES TO TABLE 2—*Continued*

BS 965 = HD 20084.—This CH star is more luminous than most of those observed by Bond, and most closely resembles BS 8626 (see Pl. 9 of Keenan and McNeil 1976). It appears to be a halo star near the top of the Population II giant branch.

BS 1112 = HD 22764.—The weakness of the metal lines, plus slight strengthening of H δ and slight filling in of H and K, suggest that the spectrum may be composite.

HD 27277.—The star is Hoag No. 4 in NGC 1545. HD 27292 is No. 3.

HD 28099.—The star is Van Bueren No. 64 in Hyades. H and K are slightly shallow.

BD +26°730 = HD 283750.—Gl 172A, a BY Dra variable. Fairly strong H, K emission.

R Dor = BS 1492.—The star is an SRb variable.

Kapteyn's star = HD 33793.—See Yerkes Atlas, Pl. 19. The peculiarity lies in the enhancement of the Ca I and Cr I lines relative to the strength of the TiO bands. From the strength of the latter the type is assigned, placing the star more than 1 mag below the main sequence.

α Aur = Capella = HD 34029.—The type corresponds to the phase when the primary dominates the spectrum. The influence of the secondary (of slightly earlier type) is not entirely absent, but Capella is included in the catalog because of its importance as an early G giant binary with one of the best-known orbits.

HD 271182 = CPD -65°457.—LMC.

BS 2018 = HD 39045.— $V_r = +105$ km s⁻¹

BS 2028 = HD 39225.— $V_r = +101$ km s⁻¹ BS 2018 and 2028, about 2° apart, not only have similar radial velocities but both have small proper motions.

HD 49068.—Probably a member of cluster NGC 2287.

HD 49091.—Although in the field of NCC 2287, the star is probably not a member of the cluster.

HD 49105.—Probably a member of NGC 2287.

BD -20°1568 = HD 49212.—Probably a member of NGC 2287.

δ^1 CMA = BS 2580 = HD 50877.—Possible member of cluster Cr 121.

HD 52938.—In cluster M50 = NGC 2323.

BD -9°1935.—Hoag No. 3 in cluster NCC 2353. Probably a member.

145 Cma = BS 2764 = HD 56577.—Possible member of δ^1 CMA group.

BS 2786 = HD 57146.—VV Cep system with companion of type B8 III or A0 III (Parsons 1982).

BS 2834 = HD 58535.—No. 22 in Cr 140. The star is a possible but doubtful member of this doubtful small cluster. Our luminosity class gives $d = 240p$ in contrast to $d = 360$ of Williams 1967. Hagen assigned it luminosity class II, which would make it a more likely member.

YY Gem = HD 60179C = α Gem C.—BY Dra eclipsing binary. The metallic lines in the blue are all shallow, though CaI λ 4226 is wide.

HD 60898, HD 60899.—These giants are in the field of NGC 2422 but have not been considered as members because of their proper motions. The luminosity of HD 60898 ($M_v = -1.25$) places it at nearly the cluster. HD 60899 has a lower luminosity ($M_v = 0.30$) and is thus about 1 mag too faint to be a cluster member.

R Pup = BS 2974 = HD 62058.—Possible member of NGC 2439.

CPD -31°1790.—In NGC 2439.

BS 3043 = HD 63660.—The apparent weakness of the metal lines may be due to the presence of components from the companion in this spectroscopic binary.

BS 3120 = HD 65662.—Close to open cluster NGC 2516. If a member, $M_v = -2.32$.

V341 Car = BS 3126 = HD 65750.—Probable member of NGC 2516.

BS 3153 = HD 66342.—Cox's star No. 110 in NGC 2516. Member of cluster.

HD 68879.—No. 8 in NGC 2548. The weakness of the K line and the 3883 Å CN band suggest a companion of type F, but this is not certain. Wallerstein and Conti 1964 found a metal deficiency of [Fe/H] = -0.51.

39 Cnc = B53427 = HD 73665.—Probable member of Praesepe.

BS 3428 = HD 73710.—Probable member of Praesepe.

HD 73974.—Probable member of Praesepe.

BS 4006 = HD 88639.—Member of 61 Cyg group?

γ Leo C = BD +20°2465 = G1388 = ADS 7724C.—Flare star.

BS 4092 = HD 90362.—The atomic lines suggest a type of K4.5, but TiO gives M0. Probably has high velocity, since μ is large.

HD 96765.—The strong G band and weak atomic lines were noticed first by Bidelman.

HD 101007.—VV Cep type binary.

BS 4605 = HD 106677.—A double-lined RS CVn binary.

31 Com = HD 111812.—On classification spectrograms the chief peculiarity noticeable is the weakness of H and K and some other metallic lines (Morgan, Abt, and Tapscott 1978). On plates of higher dispersion the lines can be seen to be broader than in most G0 stars. The star 31 Com lies in the Hertzsprung gap, where it is difficult to find a star with a spectrum that can be termed "normal."

CPD -59°4547.—This is star D in NGC 4755.

HD 121447.—This is a barium-carbon star with the Ca λ 4737 band slightly weaker than in the two similar stars, HD 24035 and HD 107541, described by MacConnell, Frye, and Upgren 1972.

BS 5299.—There is some indication that the spectral type may vary between M4+ and M5.

α Boo = HD 124897.—Spectroscopically the star appears about 1 mag too faint compared with the value of $M_v = -0.2$ corresponding to the large trigonometric parallax. With a space velocity of 115 km s⁻¹ Arcturus seems to belong to the old disk population.

RX Boo = HD 126327.—Usually g M8, but sometimes as early as M7.5.

HD 130255.—This peculiar spectrum has a ratio λ 4376/ λ 4383 suggesting a giant that is a marginal barium star with CN unusually weak. The ratio CN/CH would be consistent with classification as a dwarf barium star.

σ Lib = BS 5603.—Possibly slightly variable in type.

GI 579.2A = HD 134439.—Has $V_r = +294$ and appears to be definitely a halo star.

BS 6128 = HD 148349.—The star has a high space velocity.

BM Sco = HD 160371.—In open cluster M6 = NGC 6405.

HD 170174.—Hiltner No. 116. In NGC 6633.

BD +6°3796.—Hiltner No. 134 in NGC 6633. The star is a member.

BD +6°3798.—Hiltner No. 140. In NGC 6633.

HD 1700820.—This is star No. 150 in M25 (IC 4725).

BY Dra = HDE 234677.—Prototype of the BY Dra variables. The spectrum is definitely composite. Strong emission in H and K.

NOTES TO TABLE 2—*Continued*

- V973 Cyg = BS 7523 = HD 186776.—Probably has a high space velocity. $V_r = -102 \text{ km s}^{-1}$. The spectral type may be variable.
- HDE 350801 = BD +17°4132.—The star is No. 8 in the open cluster Harvard 20, and is probably a member according to Turner 1980.
- HD 191511.—The star is in the field of the possible open cluster Roslund 5. Lee and Perry 1971 doubt the reality of the cluster. From their distance (330 pc) M_v of HD 191511 would be +0.37, while if it had the normal luminosity of a G7 III star, $M_v = +0.6$, which is reasonably consistent with membership. The weak-line character seems too marked for a cluster member, however.
- BD +37°3865.—Hoag No. 2 in IC 4996. Foreground star in cluster field.
- EU Del = BS 7886 = HD 196610.—Type slightly variable: M6– to M6+. The space motion exceeds 140 km s^{-1} .
- HD 194241.—Foreground star in field of NGC 6910.
- 30 Vul = BS 7939 = HD 197752.—This spectroscopic binary has a rather large proper motion, but the spectrum is that of a normal Population I star.
- BS 8121 = HD 202259.—The radial velocity of -123.5 measured by R. Griffin places the star in the high-velocity population.
- HD 207328.—In Tr 37 = IC 1391.
- BS 8752 = HD 217476.—This supergiant was given a type of G0 Ia in Johnson and Morgan 1953. Perkins spectrograms in 1950 and 1957 give types ranging from G0 to G1, but by 1970 the type appeared to be considerably later and the luminosity greater. The type given in the catalog is based on a series of matched spectrograms extending from 1970 to 1987. On these small-scale plates only the Balmer lines and blend surrounding Sr II $\lambda 4078$ showed appreciable variations, and the detailed changes in the line depths and shapes described by Harmer, Lawson, and Stickland 1978 did not affect the type appreciably.
- β Peg = BS 8775 = HD 217906.—The spectral type is probably slightly variable.
- 56 Peg = BS 8796 = HD 218356.—The type given assumes that the observed lines arise in the atmosphere of the bright giant primary of this system, in which a white dwarf secondary has been reported (Schindler *et al.* 1981). The continuing presence of strong H and K emission, plus the weakness of CN and other features, suggests the possibility that the hypothesized accretion disk affects the spectrum even in the photographic region. In that case the estimation of luminosity class is not valid.
- HD 218738 = AD 5165578.—A BY Dra variable?
- γ Psc = BS 8652 = HD 219615.—Although the radial velocity is not large, the proper motion indicates a space velocity exceeding 100 km s^{-1} .
- BD +48°4065.—The star is Hoag No. 3 in the field of NGC 7686, but must be considered a doubtful member of a doubtful cluster.
- BS 8924 = HD 221148.—The star has a space velocity in excess of 100 km s^{-1} .
- HD 221203.—Hoag No. 2. Foreground star in field of NGC 7686.
- HD 221246 = BS 8925.—The star is Hoag No. 1 in the field of NGC 7686, but must be considered a probable field star.
- ρ Cas = HD 224014.—During its minimum of 1946–1947, when about 1.5 mag fainter than normal, the spectrum changed markedly from its 1940 type of F8, developing TiO bands (Beardsley 1961). Detailed changes in the spectrum through 1960 have been described by Sargent 1961. Although the star has generally remained bright since then, the spectrum seems never to have returned to a type as early as F8. Estimates between 1971 and 1978 range between G0 and G5 (Morgan *et al.* 1981).
- II Peg 224085.—This transition star between the RS CVn and the BY Dra variables is assigned luminosity class IV because the spectrum suggests a position above the main sequence.

original spectral box now belongs in an adjacent smaller box. For example, one of the original Ib supergiant standards, ϵ Peg, is slightly less luminous than most of the other stars in that “luminosity box,” and can now be more consistently classified as K1.5 Ib–IIa.

In the catalog (Table 2) the temperature types are generally given to a quarter of a subclass (e.g., K0, K0+, K0.5, K1–, K1), but when not all the decimal divisions of a class are counted as a subclass (as in type G), the plus and minus signs are not needed. The full subclasses that we use are the following: G0, G5, G8, K0, K1, K2, K3, K4, K5, M0, M1, M2, M3, M4, M5, M6, M7, M8. The irregularities are the price that we pay for keeping the general scheme of the original Henry Draper classification.

The columns of the catalog are arranged as follows:

Column (1).—The name or constellation letter of the star (preceded by a dagger for standard-type stars). For fainter stars the variable star designation, and for dwarfs the number in Gliese’s catalog (Gliese 1969), is given when available, preceded by G1. The symbol G refers to Giclas’s lists of faint proper-motion stars (Giclas, Burnham, and Thomas 1955–1980).

Column (2).—The number in *The Bright Star Catalogue* (Hoffleit and Jaschek 1982) (= the HR number).

Column (3).—The number in the *Henry Draper Catalogue*.

Column (4).—The right ascension for the year 2000.

Column (5).—The declination for the year 2000.

Column (6).—The V -magnitude, taken whenever possible from the catalog of Nicolet (1978).

Column (7).—The spectral type. An asterisk indicates a note at the end of the catalog.

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