## INFRARED SPECTRAL CLASSIFICATION OF HETZLER STARS

Using the 10 -inch Bruce telescope of the Yerkes Observatory with infrared-sensitive emulsions, Hetzler (1937) made a survey for stars that were very red. He published identifications for 168 stars for which the infrared index, defined as the magnitude at $5600 \AA$ minus the magnitude at $8500 \AA$, was greater than 2 mag. The largest infrared index he found was 10 mag . More recently an infrared survey has been in progress at the California Institute of Technology using an array of PbS detectors with a 62 -inch reflector. This survey resulted in the discovery of several red objects (Neugebauer, Martz, and Leighton 1965; Ulrich, Neugebauer, McCammon, Leighton, Hughes, and Becklin 1966) at least one of which had already been noted by Hetzler. Multicolor data for a number of infrared stars discovered by various investigators have been reported by Johnson, Mendoza V., and Wiśniewski (1965).

The first spectroscopic study of these "infrared stars" was made by Rust (1938), who obtained spectra of the brightest Hetzler stars and showed that they were mainly of type M. The spectra of some of the stars identified by Neugebauer et al. have been observed by Münch and Scargle (1965), Ford and Rubin (1965), and Wing, Spinrad, and Kuhi (1967). Pesch (1967) has determined spectral types of thirteen of these very red stars. In view of the interest in these infrared stars and the general lack of spectral classifications available for them, some material obtained at the Warner and Swasey Observatory several years ago for the purpose of classifying the Hetzler stars has been re-examined, and the results are presented below.

A series of objective-prism plates was taken using I-N emulsion and a Wratten No. 89 filter to cover most of the regions surveyed by Hetzler. With these plates, classifications could be made to the magnitude limit of the original survey except for some variable stars that had become too faint. A total of 120 of the Hetzler stars were classified. The classifications were based on the strength of the TiO and VO bands in the near infrared as described by Nassau and van Albada (1949) and Cameron and Nassau (1955). The stars were identified by marking their position on $B D$ charts enlarged to the scale of the Schmidt plates. No difficulties were encountered in the identification of the stars.

The assigned spectral classes are given in the second column of Table 1 ; the numbers in the first column are those assigned by Hetzler. A colon indicates a doubtful classification. Variable star data are given in the third, fourth, and fifth columns, which contain the star name, the spectral class, and the type of variability taken from the General Catalogue of Variable Stars (Kukarkin, Parenago, Efremov, and Kholopov 1958) or the Catalogue of Stars Suspected of Variability (Kukarkin, Parenago, Efremov, and Kholopov 1951). The final column contains spectral classifications from other sources when they differ from those given in the GCVS. In this column a letter is used to designate the source (sources are listed after the table), and the number is the star number from the source. The agreement between the classifications is good, especially considering that a large number of the stars are variable. For completeness data are given for some of the Hetzler stars that were not classified in the writer's survey.

The majority of the stars are late M stars; a few carbon and S stars are also present. The few early M stars are mainly supergiants. At the dispersion used in this survey, about $3400 \AA / \mathrm{mm}$, the spectra of the late M stars show no peculiarities to distinguish them from other late $M$ stars found in surveys and classified at Warner and Swasey Observatory.

TABLE 1
SPECTRAL TYPES FOR HETZLER STARS

| STAR | $\begin{aligned} & \text { SPECTRAL } \\ & \text { TYPE } \end{aligned}$ | VARIABLE STAR DATA |  |  | OTHER CLASSIFICATIONS |
| :---: | :---: | :---: | :---: | :---: | :---: |
|  |  | Name | Spectrum | Class |  |
| Area I |  |  |  |  |  |
| 2 | M8 |  |  |  |  |
| 3 | M7 | 5501 |  |  | M6 (b, 348);K5 (g, 40649) |
| 4 | M8: |  |  |  |  |
| 5 | M7 |  |  |  | M7 (a);M6 (g,40723) |
| 6 | M7 |  |  |  |  |
| 7 | M7 |  |  |  | M7 (b,333) |
| 8 | M5 | CU Cep | M4 | M | $\text { M6 (b,357);M6 }(\mathrm{g}, 40921)$ |
| 9 | M5 |  |  |  | M7(a);M6(b,34I);M5 (g, 40389) |
| 10 | M6. 5 | $5454$ |  |  |  |
| 11 | M6 | AA Cep | M7 | I? |  |
| 12 | M6 |  |  |  | M6 (b,339) |
| 13 | M7 |  |  |  |  |
| 14 | M5 |  |  |  | M6 (b,346) |
| 15 | M6 |  |  |  | $\text { M7 (a) ;M4 }(g, 40171)$ |
| 16 | M7 |  |  |  |  |
| 17 | M7 |  |  |  |  |
| 18 | M4: |  |  |  |  |
| 19 | M6 |  |  |  |  |
| 20 | M5 |  |  |  |  |
| 21 | M8: | DM Lac $5486$ | M7 | M | M0 (g, 40351) |
| 23 | M7 |  |  |  |  |
| 24 4 | M5 |  |  |  | $\text { M6 }(b, 362)$ |
| $+54^{\circ} 2511$ $+54^{\circ} 2517$ | M6 |  |  |  | $\text { M6 (e) ;M6 }(g, 39414)$ |
| $+54^{\circ} 2517$ $+53^{\circ} 2684$ | M4. 5 |  |  |  | M4 (e) M6.5: (d);M6 (e) |
| $+53^{\circ} 2684$ $+54^{\circ} 2603$ | M6 | RU Cyg | M7e | SRa | M6.5: (d);M6 (e) M5 (b, 111);M6 (e) |
| $+54^{\circ} 2603$ $+58^{\circ} 2316$ | M6 | 1 Cep | M2e Ia | SRe | M5 (b,111);M6 (e) |
| $\begin{aligned} & +55^{\circ} 2737 \\ & +51^{\circ} 3117 \end{aligned}$ |  | RW Cep | MO: Ia-0 | Ic | $\begin{aligned} & \text { M1: (d);M2 (d) } \\ & \text { M5 }(b, 109) ; M 5(e) \end{aligned}$ |
| $+53^{\circ} 2736$ $+51^{\circ} 3188$ | C | V413 Cyg | N | Ib |  |
| +51*3188 | M5 | SU Lac | M5e: | M | M5 (b,112);M4 (e) |
|  | M8 | AB Cep | M8e | M |  |
|  | M6: | WZ Lac | M6 | SRa |  |
|  | M8 | BQ Cyg DQ Cyg | M7 | M M |  |

TABIE 1-Continued

| STAR | SPECTRAL TYPE | VARIABLE STAR DATA |  |  | OTHER CLASSIFICATIONS |
| :---: | :---: | :---: | :---: | :---: | :---: |
|  |  | Name | Spectrum | Class |  |
| Area II <br> 1 <br> 2 <br> 3 <br> 3 <br> 4 <br> 5 <br> 6 <br> 7 |  |  |  |  |  |
|  | M8 | TX Com |  | M |  |
|  | M6 |  |  |  | M6 (a) |
|  | M7 | TT Cam | мо | M |  |
|  | M ${ }^{\text {M } 6}$ |  |  |  |  |
|  | M6 |  |  |  |  |
|  |  | RV Cam | M4 II-III | SRb |  |
| $\begin{aligned} & +53^{\circ} 882 \\ & +55^{\circ} 934 \end{aligned}$ | M7 M6 | R Aur | M7e-M9e | M |  |
| $\frac{\text { Area III }}{1}$ |  |  |  |  |  |
|  | M6 |  |  |  |  |
| 2 | M8 |  |  |  |  |
| 3 | C: | TT Cep |  | SR | M6.5 (d);M7 (g,40727) |
| +620 2030 | M5 |  |  |  |  |
| +60 ${ }^{\circ} 2267$ | C | 102116 |  |  | N (HD 205777) |
| +6102134 | M5 | SW Cep | M6 | SRb |  |
| $+60^{\circ} 2217$ $+62^{\circ} 2028$ | M6 |  |  |  | Mb (HD 203265) |
| $+62^{\circ} 2028$ $+62^{\circ} 2007$ | M2 | 102144 WV Cep | M2ep Ia | EA | 18 Cep (gM5) |
| +63 ${ }^{\circ} 1770$ | M4 |  |  |  |  |
| $+64^{\circ} 1583$ | M5 |  |  |  |  |
| +63 ${ }^{\circ} 1740$ | M3 |  |  |  |  |
| +620 1924 | M6.5 |  |  |  |  |
| +62*2079 | M4: |  |  |  | M4 (e) |
| Area IV |  |  |  |  |  |
| 1 | S | 101863 |  |  |  |
| 2 | M7 S |  |  |  | $\text { M1 }(\mathrm{g}, 17576)$ |
| 3 4 | S ${ }_{\text {M }}$ |  |  |  | S3 (a) |
| 5 | M8 |  |  |  | M8: (a) |
| 6 | M7 |  |  |  | 53 (a) ; $5(\mathrm{~h}, 4)$ |
| 7 | M5 M8 |  |  |  | M8 (a) |
| 8 | M8. ${ }_{\text {M6. }}$ | CN Vul | M5 | M |  |
| 10 | M4 |  |  |  |  |
| $+22^{\circ} 3659$ | M4 |  |  |  | M6 (e) |

TABLE 1-Continued


TABLE 1-Continued


* Misidentified as SX Gem by Hetzler.
$\dagger$ Misidentified as AZ Gem by Hetzler.
SOURCES FOR TABLE 1
a: Rust, C. F. 1938, Ap. J., 88, 525.
b: Nassau, J. J., and Blanco, V. M. 1954, Ap. J., 120, 118.
c: Nassau, J. J., Blanco, V. M., and Morgan, W. W. 1954, Ap. J., 120, 478.
d: Cameron, D., and Nassau, J. J. 1956, Ap. J., 124, 346.
e: Neckel, H. 1958, A p. J., 128, 510.
f: Nassau, J. J., and Seyfert, C. K. 1946, Ap. J., 103, 117.
g: Lee, O. J., Baldwin, R. J., Hamblin, D. W., Bartlett, J. J., and Gore, G. D. 1947, Ann. Dearborn Obs., Vol. 5.
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