SPECTRAL TYPES OF SOME ECLIPSING BINARIES

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ABSTRACT

Spectral types are given for one or both components of sixty-three eclipsing binary systems.

For the solution of the light-curves of an eclipsing binary, it is useful to know the spectral types of the component stars. Therefore, a program was begun several years ago to obtain spectral types for all eclipsing binary systems easily observed with the Yerkes 40-inch refractor. The plan was to take two or three plates of each star, distributed in phase to allow some separation of the spectra of the two components. The program is still incomplete, but the results obtained to date are collected in Table 1.

This table lists the Universal Times of the plates obtained, the phases at mid-exposure counted in fractions of a period from the preceding primary minima and computed with the elements in Kukarkin and Parenago's catalogue (1948), the spectral types, and additional remarks. Where only one type is given for several plates, either no change in type could be detected between them, or each is a composite of the two types given. The A-type stars without luminosity classes are probably dwarfs but could be as bright as class III. Remarks about broad and sharp lines refer to the low dispersion used (125 A/mm at $H\gamma$); the broad lines near quadrature probably result from double lines.

REFERENCES

Kukarkin, B. V., and Parenago, P. P. 1948, General Catalogue of Variable Stars (Moscow).

TABLE 1
OBSERVATIONAL DATA

Star	Year	Date	Begin	End	Phase	Sp. Type	Remarks
RT And	{1954 \1954	July 15 Aug. 12	6:05 4:58	7:37 5:44	0.03 .45	F8 V F7 V	
ST Aqr	{1954 1954	Aug. 29 Sept. 22	5:03 4:50	5:57 6:00	.02\ .75}	G8 IV::+A7:	The A star is brighter than the G8 star
σ Aql .	{1954 1954 1954	July 15 Aug. 1 Oct. 7	6:	53 42 44	.70 .50 .75	B3 V	Lines broad
KO Aql	{1954 1954	Aug. 12 Sept. 11	3:10 4:50	4:18 5:56	.54 .04	A2 A0	
OO Aql	1953 1954 1954 1954 1954	June 23 July 15 Sept. 8 Sept. 11	5:50 8:10 1:18 2:17	7:02 8:47 2:24 2:57	.02 .81 .79 .77	G5 V: G2 V G5 V:	Lines are very broad on all plates
V 337 Aql. WW Aur	{1953 {1954 1954	June 23 Aug. 12 Jan. 15	4:48 2:04 5:	5:38 3:05 43	.75\ .51} .76	B0 5 V A7	Broad lines. The K line is rather broad but quite shallow; this could be explained by a composite spectrum such as A0+F2, but the hydrogen lines are probably not strong enough for this
BF Aur.	1954	Nov. 17	4:17	5:29	.22	B5 V	Very broad lines, possibly double
RS CVn RZ Cas TW Cas YZ Cas AO Cas AR Cas BM Cas CC Cas U Cep VW Cep	1953 1954 1954 1953 {1954 1954 {1954 1953 1954 1953 (1954 {1954 1954	June 9 Nov. 17 Nov. 17 Sept. 5 Sept. 16 Nov. 17 Sept. 22 Aug. 1 Nov. 3 Sept. 16 Nov. 17 Sept. 6 July 2 July 15 Aug. 29	3:32 8:55 8: 0: 4: 4:	5:59 10 4:02 10:05 39 32 30 34 44 8:28 0:22 6:44 7:00 4:17 2:51	.01 .49 .91 .75 .50 .01 .43 .04 .54 .92 .20 .98 .31 .58 .03	K2 III A2 A1 B1 V A2 A3 O9 V B3 V F0 Ia O9 V G8 III	Broad lines
WX Cep XZ Cep AH Cep U CrB RV Crv Y Cyg.	{1954 1954 1954 {1954 1953 1955 {1954 1954 (1954 1954 1954	Aug. 1 Aug. 26 Sept. 8 Sept. 22 Sept. 26 May 31 Feb. 12 July 2 Aug. 29 July 15 Aug. 1 Aug. 12	4:43 2:28 4:03 2:10 1:43 4:27 3:12 6:22 3:37 8:51 3:59 4:25	5:29 3:43 4:44 2:17 1:50 4:44 4:37 6:32 3:55 9:11 4:25 4:50	62\ 00\} 85 49\ 73\} 56 61 17\ 49\} 01\ 41\ 0 76	A7 O9 5 V B0 5 III B7 V F0+G0: B0 V	Broad lines

TABLE 1—Continued

Star	Year	Date	Begin	End	Phase	Sp. Type	Remarks
V 367 Cyg V 382 Cyg V 453 Cyg DM Del TW Dra S Eql WX Eri .	1954 1954 {1953 1954 1954 {1953 1953 1953 {1953 {1953	Sept. 11 Sept. 26 Sept. 26 Sept. 25 Sept. 26 May 27 Sept. 6 Sept. 6 Sept. 5 Sept. 21	1:28 1:59 1:18 4:29 4:30 3:05 2:32 1:55 7:21 9:26	1:40 3:05 1:51 5:05 4:58 3:48 2:41 2:16 8:42 10:45	0 25 .75 24\ 00\ 99 76 09 11 02\ 56\	A7 Ia O9 V B1 III A3 A5 K0 III B8 V A7+F6 V:	The A5 star is much brighter
u Her .	1953 1954 1954	June 9 Sept 11 Sept. 12	6:17 4: 4:19	6:24 16 4:24	76 .51 00	B3 III	Broad hydrogen lines
Z Her	{1953 1953	May 28 May 31	4:17 4:50	4:36 5:01	50\ 26}	F6 V	
RX Her	{1954 1954 1954	July 19 Aug. 1 Aug. 12	5:49 3:45 5:55	6:13 3:55 6:27	49 75 .99	A0	K line broad
TX Her	{1953 1954	May 31 Oct. 7	5:45 1:52	6:11 2:40	74) 49)	A7	Lines broad Lines sharp
AK Her	{1953 1954 1954	June 9 Aug. 1 Sept. 12	6:33 3:11 1:23	7:01 3:41 2:08	76 08 55	F2+F6	
DI Her	{1953 1953	Sept. 3 Sept. 6	2:02 4:26	3:02 5:25	.71 01	B5 III	
HS Her . AI Hya	{1954 {1954 1954	Sept. 12 Sept. 26 Jan. 15	2:42 0:53 5:55	3:09 1:31 6:51	48\ .98} ?	B5 III F0+F5	Period too inaccurate to compute phase; Sr II, A 4077, unusually
SW Lac AR Lac CM Lac CS Lac TX Leo δ Lib U Oph	1954 {1954 1954 1954 1954 1953 1953 {1953 1953	July 18 Aug. 12 Aug. 29 Sept. 22 Nov. 3 Sept. 26 May 31 May 31 June 9 June 11	6:22 6:39 2:59 3:47 5:54 5:04 2: 4: 6: 8:44	09 .	59 50 .00 12 .99 .8 00 00 56\ 82}	K0 V K2 III+F8:: K2 III A7 B5 V A2 A0 B5 V	strong H and K emission Sharp lines Broad lines
V 451 Oph	1953 1954 1954	June 23 July 15 Aug. 1	4:05 4:38 5:39	4:25 4:51 6:01	$egin{array}{c} .64 \\ .02 \\ 54 \\ \end{array}$	A0	
V 456 Oph.	{1954 {1954	Sept. 8 Sept. 11	5:06 3:06	6:05 4:07	5 }	A5:	Plates somewhat under- exposed; period too un- certain to predict phases
V 502 Oph	\begin{cases} 1953 \\ 1953 \\ 1954 \end{cases}	May 31 June 9 July 19	5:07 4:19 2:45	5:39 4:41 3:45	$ \begin{array}{c} 25 \\ 02 \\ 0 \ 12 \end{array} \}$	K0 IV:+F5:	Plate at primary mini- mum somewhat under- exposed

TABLE 1—Continued

Star	Year	Date	Begin	End	Phase	Sp. Type	Remarks
	(1953	June 11	8:56	9:15	0 22)		
V 566 Oph	{1954	July 15	5:02	5:30	40}	F4 V	
	[1954	Sept. 26	3:28	3:58	97		
AW Peg	1954	Sept 8	3:44	3:56	15	A5	
EE Peg	1954	July 19	5:06	5:20	00	A3	
AY Per	1953	Nov. 14	6:08	7:36	54	A0	
SZ Psc	∫1953	Sept. 21	7:25	7:54	48)	K1 III+A	H and K emission; the K
02150	1954	Sept. 25	5:09	6:05	49 }	111 111 11	star is much brighter
TT C	∫1953	May 31	6:32	6:37	50	B7 III	
U Sge	1954	Sept. 8	2:33	3:34	01	K1 III	
V 505 Sgr	1954	Sept. 26	4:04	4:22	01	F6:+A3	
RZ Sct	1953	Sept. 3	1:29	1:50	34	B0 V	
W UMi.	∫1953	May 31	3:11	3:55	50	A2	
W UMII.	1953	June 23	2:48	3:56	01	A3	
UY Vir .	1953	June 9	2:48	3:11	.73	A7 V	
	[1953]	May 31	6:42	6:51	70)		
Z Vul.	{1953	June 23	4:33	4:42	03}	B5 V	
	1954	Aug. 29	4:07	4:54	00)		,
RS Vul .	1953	May 31	6:19	6:29	0 24	B5 V	