

RADIAL VELOCITIES OF 204 STARS IN THE REGION OF THE HYADES*

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ABSTRACT

This investigation completes the determination of the radial velocities of all suggested members of the Taurus cluster brighter than 10.0 visual magnitude, with seven fainter stars. Of the 204 stars recently observed at Mount Wilson for radial velocities, 93, all located within 15° of the cluster center, are classified as definitive members.

The motions of the members of the nearer star clusters, such as the Hyades, Pleiades, and Praesepe, constitute one of the best sets of data available for the determination of stellar distances. If the population is large, the cluster parallaxes in combination with reliable magnitudes and spectral classifications become very useful in studies of the relation between absolute magnitude and spectral type.

The early Groningen photographic surveys of the Hyades,¹ though subject to considerable systematic and accidental errors, suggested a fairly large cluster population. Under the stimulus of L. Boss's classical study² of the motions of the brighter stars, a number of investigators have devoted considerable time during the last thirty years to a search for additional members.³ As a result, by 1941 some three hundred stars were suggested as "certain," "probable," or "doubtful" members of the Taurus cluster. The principal criteria in the selection of these stars were proper motions of about $0''.1$ per year in the general direction of the Boss convergent. Rather large departures were allowed, especially among the stars listed as "probable" or "doubtful," and among these appear a number of faint early-type stars for which actual cluster membership would entail luminosities not at all in accord with those generally associated with their spectra. The question of cluster membership is further complicated by the proximity of the Taurus convergent to that of Kapteyn's stream I, which is especially pronounced among stars of types A and F, and by the possibility that the cluster itself may share in the motion of the larger stream. It had been clear for some time that the proper motions alone do not suffice to establish cluster membership, especially for stars at considerable distances from the center, and that additional information concerning the radial velocities was highly desirable.

J. H. Moore's catalogue⁴ gives radial velocities for 68 cluster stars, a few of them, however, being based on one plate only. In the decade following its publication only ten more velocities were published. In 1941, however, I found in the Mount Wilson files spectrograms of 89 additional stars and decided to complete the determination of the velocities of all the stars in the Hyades list brighter than visual magnitude 10.0. Two plates were deemed sufficient if the measures agreed reasonably well with each other and with the cluster velocity; in all other cases three or more plates were secured. All the plates were

* *Contributions from the Mount Wilson Observatory, Carnegie Institution of Washington*, No. 741.

¹ J. C. Kapteyn and W. de Sitter, *Groningen Pub.*, No. 14, 1904.

² *A. J.*, **26**, 31, 1908.

³ For a history of that period and a complete bibliography see J. M. Ramberg, *Stockholms Obs. Ann.*, Vol. **13**, No. 9, 1941.

⁴ *Pub. Lick Obs.*, Vol. **18**, 1932.

taken with a one-prism spectrograph attached to the 60-inch reflector. Two cameras were used, the brighter stars being observed with an 18-inch camera giving a dispersion of 38 Å/mm at $H\gamma$, the fainter with a 9-inch camera, dispersion 75 Å/mm.

The radial velocities of the 204 stars in Table 1 are based upon 605 plates, 75 per cent of which were secured in the period 1942–1946. All the plates of this epoch were measured by the writer. The headings of the columns of the table are largely self-explanatory. The magnitudes are mainly (*a*) photovisual magnitudes on the International System by Eric Holmberg⁵ for stars within 4° of the center, and (*b*) Harvard Draper magnitudes, further from the center. Among the stars observed are seven fainter than the limit set when the program was begun. The classifications of spectra were made by the writer on the Mount Wilson system. The proper motions are compilations by the writer, based upon a number of sources, all of which were roughly reduced to the system of the *Albany General Catalogue*.⁶ The probable errors of the radial velocities were computed from the relation

$$r = 0.67 \sqrt{\frac{v v}{m(m-1)}}.$$

Of these stars, 93 are considered to be cluster members, 8 others are probable, and 32 must be considered, mainly on evidence other than the radial velocities, as doubtful. On the basis of the radial velocities alone, 71 of the stars may be eliminated from further consideration as cluster members. Among these are all the faint stars with early-type spectra and nearly all the stars situated more than 15° from the center. The probable spectroscopic binaries number 23, of which only 6, all located near the center, appear to be members; 7 seem to show two spectra. The most interesting of the binaries is BD+16°516 (9.2, G9), which from measures on five plates shows a velocity range of 242 km/sec. The velocities of both components of 5 visual binaries were measured. Only 1 of these, HD 26015, is considered a definitive member, although 3 others—HD 20430–20439, HD 21663, and HD 33204—are probable outliers. Eight of the stars have large velocities (Table 2). Two of these, though widely separated, may have common motion.

A discussion of the motions and absolute magnitudes of the members of the Taurus cluster will appear soon.

⁵ *Lund Medd.*, Ser. II, No. 113, 1944.

⁶ *Carnegie Inst. Washington Pub.* No. 468, 1938.

TABLE 1
RADIAL VELOCITIES OF 204 STARS IN THE HYADES REGION

BD	HD	(1950)		m_v	Sp.	μ	θ	ρ (KM/SEC)	PL.	NOTES
		R.A.	Dec.							
10° 374..	17663	2 ^h 47 ^m 7	+10° 25'	9.3	G8	0".144	92°	+ 54.9±1.0	3	
29 503..		54.8	+29 28	9.2	G8	.206	106	+ 25.8±0.5	3	D
11 456..	20278	3 13.2	+11 26	7.9	G1	.193	91	+ 43.6±1.5	4	D
7 493..	20430	14.8	+ 7 29	7.4	G0	.187	92	+ 31.1±0.8	3	P
7 494..	20439	14.9	+ 7 30	7.7	G1	.194	92	+ 31.2±2.0	4	P
19 505..	20600	3 16.6	+19 32	8.4	F9	.102	97	+ 33.0±1.8	3	D
11 462..	20717	17.9	+12 10	7.3	F6	.149	83	+ 44.8±1.8	4	D
8 496..	20727	17.9	+ 8 51	8.3	G3	.301	103	+ 11.1±0.6	3	
25 545..	21168	22.8	+25 41	8.6	F8	.087	116	+ 56: var.	4	1
19 547..	21663A	27.6	+19 56	7.9	G5	.170	109	+ 25.2±1.5	4	P
	21663B	3 27.6	+19 56	9.5	K2			+ 28.3±0.5	3	P
23 465..		29.9	+23 31	9.0	K0	.168	102	+ 32.7±0.5	2	P
10 461..	22254	32.6	+11 13	8.3	F9	.180	95	+ 63.4±1.9	3	
19 562..	22328	33.5	+19 54	7.6	F6	.101	114	+ 34.1±0.4	3	D
20 598..		34.6	+21 11	9.0	K0	.159	103	+ 35.4±1.1	3	H
18 514..		3 34.6	+18 27	8.6	F8	.147	96	+ 41.8±1.9	4	H
26 595..		37.1	+26 48	9.1	G8	.156	131	-198.4±0.4	3	
3 512..	22917	38.4	+ 3 25	9.3	F8	.138	86	+ 18.0±1.9	4	
26 601..	23007	39.5	+26 26	7.8	K0	.100	112	+ 33.0±1.9	4	D
27 555..		41.4	+28 03	9.0	G2	.086	134	+ 10.4±1.7	3	
24 550..	23375	3 42.6	+24 19	8.6	A8	.078	132	+ 19.8±1.6	4*	2
25 615..	23488	43.4	+25 43	8.7	A7	.078	152	+ 3.6±2.7	4	
22 550..	23514	43.7	+22 47	9.2	F8	.064	119	+ 12.0±0.4	3	
9 494..	23841	45.8	+ 9 30	7.0	G8	.072	80	- 79.8±2.2	3	
23 556..	23822	46.0	+23 42	6.6	F0	.077	133	+ 18.6±2.1	4	
22 575..	23965	3 47.1	+22 21	7.9	F7	.179	111	+ 11.7±0.8	3	
16 516..		47.4	+17 06	9.2	G9	.104	73	+ 21: var.	5	3
23 571..		48.0	+23 46	9.5	K5	.154	108	+ 38.8±1.7	5	H
21 544..		49.8	+21 50	9.5	G0	.120	126	+ 58.2±2.7	3	4
26 633..	24301	49.9	+26 31	8.0	G1	.163	132	+ 26.6±1.0	3	
27 589..	24365	3 50.7	+27 59	7.8	G5	.094	95	+ 21.7±0.8	3	
16 529..		52.3	+16 51	9.9	K2	.172	102	+ 39.1±1.8	2	H
22 596..	24570	52.4	+23 13	8.6	K0	.103	93	+ 42.9±0.6	3	D
26 645..	24690	53.4	+26 36	9.2	A8	.082	116	+ 8.1±1.0	3	
26 646..		53.9	+27 10	9.1	G9	.090	96	+ 34.2±0.4	2	D
22 608..	24844	3 54.8	+22 47	9.1	K0	.153	78	+ 27.1±0.6	3	
24 603..	24997	56.2	+24 55	9.2	F8	.106	123	- 23.6±0.2	3	
24 605..	25065	56.9	+24 33	8.7	G0	.095	116	+ 26.2±0.3	3	D
13 625..	25153	57.5	+14 10	7.7	F5	.065	101	+ 39.0±1.0	3	D
19 641..		57.7	+20 14	8.6	G2	.157	108	+ 37.6±2.0	2	H
22 626..	25532	4 01.2	+23 16	8.3	F6	.134	128	-113.0±0.3	3	
25 674..		02.9	+25 40	9.0	G8	.137	103	+ 24.0±1.6	4	D
15 582..	25825	03.3	+15 34	7.9	G2	.114	95	+ 35.6±1.5	3	H
14 657..	26015A	04.9	+15 02	5.9	F2	.136	100	+ 34.9±1.7	2*	H
	26015B	04.9	+15 02	8.7	G7			+ 40.8±0.3	2	H
13 647..	26091	4 05.5	+13 25	8.8	K1	.074	88	+ 20.3±0.9	3	
28 624..	26090	05.8	+29 03	8.5	G2	.101	99	+ 41.3±1.8	3	D
18 594..	26345	07.8	+18 18	6.6	F6	.120	110	+ 33.6±2.1	2*	H
0 711..	26623	10.0	+ 0 38	9.2	F7	.077	85	+ 25.5±1.5	3	
16 570..		10.3	+16 40	9.1	G2	0.153	109	- 7.3±1.8	3	

TABLE 1—Continued

BD	HD	(1950)		m_v	Sp.	μ	θ	ρ (KM/SEC)	r	PL.	NOTES
		R.A.	Dec.								
23° 649 ..	26736	4 ^h 11 ^m 5	+23° 29'	8.0	G4	0 ^h 129	106°	+ 42.2±2.2	4	H	
22 657 ..	26737	11.5	+22 20	7.0	F4	.101	108	+ 37.9±1.7	4	H	
14 673 ..	26756	11.6	+14 30	8.6	G5	.129	101	+ 37.9±1.3	3	H	
12 566 ..	26767	11.7	+12 20	8.3	G3	.115	92	+ 38.6±1.8	3	H	
10 551 ..	26784	11.8	+10 35	7.1	F7	.125	96	+ 36.6±1.7	3	H	
21 608 ..		4 12.3	+22 11	9.1	K0			+ 22.0±1.1	3		
20 721 ..	26874	12.8	+20 42	8.1	G6	.121	113	+ 27.7±0.8	3	H:	
15 603 ..	26911	12.9	+15 17	6.4	F5	.123	106	+ 37.0±1.8	4*	H	
15 604 ..		13.1	+15 42	9.5	G9	.078	134	+ 56.4±0.2	3		
21 612 ..		13.6	+21 47	8.9	G7	.119	109	+ 40.7±0.3	2	H	
4 666 ..	27089	4 14.3	+ 4 25	8.6	F8	.172	87	+ 59.8±0.3	3		
18 613 ..		14.4	+19 13	9.0	G7	.102	125	+ 4.4±2.7	3		5
17 703 ..	27149	15.1	+18 07	7.7	G4	.116	112	+ 44.0±2.7	4		6, H
19 694 ..	27250	16.0	+19 47	8.6	G7	.101	104	+ 40.6±1.6	2	H	
17 707 ..	27282	16.3	+17 24	8.5	G5	.098	98	+ 40.0±0.2	2	H	
15 612 ..	27371	4 16.9	+15 31	3.9	G8	.121	102	+ 38.5±1.4	4*	H	
13 662 ..	27372	16.9	+14 03	7.8	G8	.212	160	+ 16.6±0.8	3		
23 675 ..	27370	17.1	+23 28	7.2	G6	.148	108	+ 8.8±0.8	3		
13 663 ..	27397	17.1	+13 55	5.6	F1 _n	.118	101	+ 40: var.	3*	7, H	
18 623 ..	27406	17.3	+19 07	7.7	F9	.125	106	+ 38.7±1.5	3	H	
18 624 ..	27429	4 17.5	+18 37	6.2	F3	.118	114	+ 37.6±0.5	6	8, H	
14 682 ..	27459	17.8	+14 59	5.3	A9 _n	.113	102	+ 39.7±2.8	3*	H	
13 665 ..	27483	18.1	+13 45	6.3	F4	.118	102	+ 37.0±0.9	6	9, H	
18 629 ..	27534	18.6	+18 18	7.0	F5	.111	108	+ 36.5±2.0	4	H	
14 685 ..		18.7	+14 44	10.1	G3	.135	120	+ 47.5±2.1	4	D	
14 687 ..	27561	4 18.8	+14 18	6.7	F6	.120	103	+ 38.4±0.5	2*	H	
5 636 ..	27610	19.0	+ 5 15	9.0	F3	.072	79	+ 22.4±1.4	4		
13 668 ..	27628	19.2	+13 58	5.7	A4	.117	102	+ 37.7 var.	4*	H	
31 769 ..		19.9	+32 06	8.8	F8	.104	133	+ 79.4±1.0	3		
16 585 ..	27685	19.9	+16 40	7.9	G5	.145	106	+ 33.2±1.0	2	H	
13 671 ..		4 19.9	+14 08	10.1	G3	.107	107	+ 39.0±0.4	2	H	
21 635 ..	27732	20.4	+21 16	9.1	G7	.105	115	+ 38.8±1.4	2	H	
19 708 ..		20.4	+19 32	9.1	G7	.094	109	+ 36.4±1.2	2	H	
24 654 ..	27731	20.5	+24 17	7.2	F6	.106	114	+ 33.6±1.2	2*	H	
15 616 ..		20.5	+15 39	10.7	K5	.130	104	+ 48.1±0.3	2	H	
14 691 ..	27771	4 20.7	+14 33	9.3	G8	.107	98	+ 44.2±1.5	2	H	
21 641 ..	27808	21.3	+21 37	8.0	F8	.127	117	+ 33.4±1.0	2	H	
14 693 ..	27836	21.3	+14 39	7.7	G1	.106	102	+ 37.9±0.9	3	H	
17 715 ..		21.4	+17 53	10.0	K4	.111	97	+ 46.9±1.0	3	H	
16 589 ..	27835	21.4	+16 16	8.3	G0	.083	100	+ 38.4±0.6	3	H	
17 716 ..		4 21.5	+17 20	10.0	K0	.028	78	+ 2.1±1.3	3		
16 593 ..		22.1	+16 52	10.4	K2	.087	101	+ 43.1±1.9	3	H	
23 692 ..	27972	22.8	+23 26	8.9	F8	.082	121	+ 27.7±2.2	3	D	
17 721 ..	27990	22.9	+17 55	9.1	K0	.098	102	+ 38.0±0.5	2	H	
17 722 ..	28007	23.1	+17 20	7.6	F2	.050	107	+ 30: var.	4	10	
21 644 ..	28033	4 23.3	+21 22	7.5	F9	.111	108	+ 42: var.	4	11, H	
4 690 ..	28069	23.3	+ 5 01	7.2	F6	.097	85	+ 31.1±0.3	3	H	
16 598 ..	28068	23.5	+16 44	8.2	G1	.099	99	+ 40.4±2.7	4	H	
13 684 ..		24.6	+14 09	10.6	K5	.118	112	+ 38.7±0.1	2	H	
11 614 ..	28237	25.0	+11 39	7.4	G0	0.138	92	+ 42.2±1.1	2	H	

TABLE 1—*Continued*

BD	HD	(1950)		m_v	Sp.	μ	θ	ρ (KM/SEC)	r	Pl.	NOTES
		R.A.	Dec.								
13° 685...	28258	4 ^h 25 ^m 2	+13° 46'	9.2	K0	0".115	93°	+ 42.0±0.6		2	H
19 727...	28291	25.6	+19 38	8.4	G7	.123	114	+ 36.4±0.6		2	H
18 640...	28305	25.7	+19 04	3.8	G9	.118	109	+ 38.1±0.6		3*	H
16' 606...	28344	26.0	+17 10	7.8	G2	.105	106	+ 40.5±1.4		3	H
15 633...	28363	26.1	+16 03	6.7	F8	.111	106	+ 45: var.		4	12, H
26 722...		4 26.4	+26 34	9.0	G7	.096	117	+ 37.3±1.9		3	H
17 731...	28394	26.5	+17 26	7.0	F8	.094	107	+ 33.2±2.4		2*	H
17 732...	28406	26.6	+17 45	7.0	F7	.109	104	+ 34.4±0.1		2*	H
17 734...		26.6	+17 47	9.0	G7	.116	107	+ 44.8±0.6		2	H
15 634...		26.7	+16 08	10.5	K5	.097	99	+ 45.3±0.5		2	H
13 688...	28424	4 26.7	+13 48	7.8	G9	.138	99	+ 96.8±1.5		6	
16 609...	28462	27.1	+16 33	9.1	K0	.109	102	+ 41.7±0.1		2	H
15 636...	28485	27.3	+15 32	5.8	A8n	.108	104	+ 27: var.		5*	H
19 731...	28483	27.4	+19 44	7.2	F5	.101	113	+ 37.2±0.8		2	H
15 638...	28545	27.8	+15 38	9.0	G8	.098	104	+ 36.5±0.3		2	H
19 733...	28593	4 28.2	+20 01	8.4	G5	.090	108	+ 40.0±1.3		2	H
10 588...	28608	28.2	+10 38	7.1	F7	.120	91	+ 36.9±1.8		3	H
13 691...	28635	28.6	+13 47	7.9	F8	.093	102	+ 42.4±0.8		2	H
15 645...	28677	29.0	+15 45	6.1	F0n	.110	107	+ 39.0±1.9		4*	H
5 674...	28736	29.4	+ 5 18	6.4	F5	.113	84	+ 40.7±1.5		4	H
15 646...	28783	4 30.0	+15 54	9.1	G9	.111	106	+ 43.4±0.5		2	H
16 620...	28878	30.8	+16 38	9.4	K0	.105	111	+ 43.4±1.3		2	H
15 649...	28888	30.8	+15 52	8.5	G2	.095	131	+ 56.1±0.6		3	
12 608...	28911	31.0	+13 09	6.7	F4	.115	100	+ 34.4±2.3		3	H
15 650...	28977	31.6	+15 43	9.7	K0	.122	103	+ 40.5±1.0		2	H
15 651...	28992	4 31.7	+15 23	8.0	G0	.093	112	+ 41.2±0.6		2	H
27 667...		31.8	+27 56	9.0	F5	.108	122	+ 32.3±1.3		3	D
15 654...	29159	33.2	+15 35	9.4	K0	.104	106	+ 43.5±1.9		2	H
23 715...	29169	33.5	+23 14	6.0	F4	.131	117	+ 41.7±2.0		3*	H
15 656...	29225	33.8	+15 46	6.7	F6	.111	107	+ 33.2±0.6		3	H
14 728...	29310	4 34.7	+15 03	7.7	G1	.106	107	+ 39.7±1.9		3	H
15 662...	29387	35.4	+15 21	9.8	G7	.080	94	+ 43.5±0.9		2	D
12 618...	29388	35.4	+12 25	4.4	A5n	.099	97	+ 45: var.		2*	H
22 721...	29419	35.9	+23 04	8.4	F9	.147	108	+ 39.6±1.1		2	D
13 702...	29461	36.1	+14 00	8.0	G3	.089	100	+ 39.7±1.7		2	H
15 666...	29488	4 36.4	+15 49	4.7	A5n	.089	103	+ 43.8±2.5		2*	H
7 681...	29499	36.4	+ 7 46	5.6	F0	.089	89	+ 42.0±1.1		3*	H
26 732...		36.8	+27 06	9.2	A0p			+ 6: var.		3	13
23 722...		37.1	+23 13	9.0	G9	.093	111	+ 40.0±2.0		3	H
16 640...	29608	37.5	+16 25	9.5	K2	.095	108	+ 41.4±2.5		3	H
23 723...	29621	4 37.8	+23 43	8.8	G5	.095	126	+ 32.3±1.9		3	H
19 754...		38.5	+20 10	9.7	G8	.083	127	+ 31.1±0.6		3	H
28 681...		38.6	+29 07	9.2	K2	.084	126	+ 3.0±2.0		3	
- 1 697...	29789	39.0	- 0 51	8.4	F5	.089	68	+ 35.2±1.9		3	D
19 762...		39.8	+20 07	9.6	A0p	.092	124	- 10: ±4:		3	
18 684...	29836	4 39.9	+18 38	7.1	G3	.142	128	+ 15.0±0.7		3	
21 694...	30169	43.0	+21 12	9.1	G9	.104	105	+ 26.5±1.3		3	
3 664...	30286	43.7	+ 3 13	7.9	G4	.057	61	+ 19.6±1.5		3	
1 819...	30299	43.8	+ 1 15	8.5	F6	.053	80	+ 24.0±0.7		3	
8 759...	30311	44.0	+ 8 55	7.2	G0	0.112	90	+ 39.8±0.4		2	H

TABLE 1—Continued

BD	HD	(1950)		m_v	Sp.	μ	θ	ρ (KM/SEC)	PL.	NOTES
		R.A.	Dec.							
17° 786..	30355	4 ^b 44 ^m 7	+18° 10'	8.2	G4	0.116	119°	+ 42.0±0.7	2	H
24 689..	30418	45.4	+24 40	8.0	F5	.035	147	+ 42.4±0.8	2	
20 823..		45.8	+21 02	9.0	K0	.092	104	+ 42.4±1.6	3	D
24 692..		46.1	+24 44	9.1	K3	.123	119	+ 44: var.	4	14, H
18 736..	30505	46.1	+18 33	8.8	K0	.072	142	+ 42.6±0.8	2	D
3 679..	30544	4 46.1	+ 3 34	7.1	B9	.064	66	+ 33.2±2.3	4	
15 686..	30589	46.7	+15 48	7.9	G0	.086	109	+ 39.7±2.2	4	H
23 747..	30572	46.8	+23 19	8.6	G4	.073	120	+ 32.8±0.8	2	P
16 657..	30676	47.5	+17 07	7.2	F8	.104	114	+ 41.6±0.8	2*	H
14 770..	30712	47.8	+15 00	8.2	G5	.086	102	+ 43.7±0.2	2	H
13 725..	30726	4 47.9	+14 09	8.9	G1	.084	94	+ 31.0±0.7	3	D
15 692..	30738	48.0	+16 08	7.3	F8	.098	101	+ 41: var.	5	15, H
0 873..		48.3	+ 0 30	9.2	F8	.088	79	+102.7±1.7	3	
28 706..	30754	48.4	+28 33	9.0	K5	.124	115	+ 15.8±0.9	3	
- 0 789..		49.1	- 0 10	8.9	G8	.055	46	+ 30: var.	4	16
4 769..	31003	4 49.8	+ 4 15	8.5	F8	.085	107	+ 41.2±2.0	4	D
22 769..		50.6	+22 56	8.8	F8	.083	103	+ 45.2±1.9	2	D
10 668..		51.1	+10 16	8.9	G8	.070	108	- 29.3±1.8	3	
17 807..	31181	51.2	+17 33	9.8	F7	.072	115	+ 40: var.	4	17, D
16 664..	31153	51.2	+16 57	7.1	F8	.080	93	+ 54: var.	4	18
- 1 747..		4 51.2	- 1 15	9.3	G7	.123	61	+ 7.3±0.6	3	
19 811..	31236	52.0	+19 24	6.2	F1	.072	122	+ 41.2±0.8	2	H
5 765..	31354	52.7	+ 5 33	8.2	G0	.112	72	+ 17.5±1.8	4	
13 749..	31609	55.0	+13 56	8.5	G5	.093	113	+ 46.4±1.1	2	H
30 752..	31706	56.1	+30 59	8.0	F6	.055	138	+ 14.2±2.4	3	19
- 0 815..		4 56.1	- 0 47	9.1	G7	.098	67	+ 26.9±1.1	3	
26 771..	31781	56.6	+26 10	8.6	G0	.071	145	+ 16.2±0.7	3	
0 916..	32023	57.8	+ 0 57	9.1	F8	.072	76	+105.3±2.5	3	
- 0 823..	32114	58.2	- 0 34	8.9	A0	.087	64	+ 2: ±6:	3	20
31 846..		58.7	+31 34	8.9	G0	.091	146	+ 74.9±1.8	3	
13 783..	32347	5 00.3	+13 39	9.3	K0	.064	100	+ 43.6±1.8	2	H
27 732..	33204A	06.6	+27 58	6.0	A3	.088	138	+ 41.3±0.9	2*	P
.....	33204B	06.6	+27 58	8.5	G7	.092		+ 46.4±0.4	2	P
20 897..	33400	07.8	+20 31	7.8	F5	.051	106	+ 44.8±1.5	3	D
6 865..	33662	09.4	+ 6 47	7.9	K5	.044	85	+ 22.4±1.3	3	
19 876..	34031	5 12.2	+20 00	7.7	G1	.116	109	+ 22.2±1.4	3	
12 756..		13.5	+12 10	8.9	F6	.044	104	- 23.8±0.7	3	
26 806..		16.1	+27 02	9.0	G4	.067	163	+ 10.2±1.2	3	
23 902..	34772pr	17.9	+23 59	8.9	F6	.067	149	+ 18.5±1.5	3	
.....	34772fol	17.9	+23 59	8.9	F6			+ 20.6±1.0	3	
28 783..	34987	5 19.6	+28 42	8.7	F7	.075	147	+ 37.4±0.8	2	D
24 846..	244516	28.9	+24 57	9.2	F7	.090	146	+ 41.6±0.6	2	D
20 978..		29.2	+21 00	9.3	F0	.036	118	+ 37: var.	4	21
29 936..		32.1	+29 14	9.5	G3	.074	148	+ 46.1±2.1	3	D
23 981..	37388A	36.1	+23 16	8.6	F8	.070	141	+ 40.8±1.3	4	D
.....	37388B	5 36.1	+23 16	9.0	F7			+ 44.6.....	1	D
26 899..	246128..	37.7	+26 58	9.0	G1	.087	160	+ 58.4±1.0	3	
26 907..		39.3	+26 54	8.9	F8	.097	155	+ 57: var.	3	22
13 964..	37982	40.2	+13 07	8.9	F5	.061	127	+ 49.8±0.1	2	D
9 970..		46.3	+ 9 52	9.0	K0	0.050	89	+ 51.2±1.0	3	D

TABLE 1—Continued

BD	HD	(1950)		m_v	Sp.	μ	θ	ρ (KM/SEC)	PL.	NOTES
		R.A.	Dec.							
17° 1031..	39117	5 ^h 48 ^m 2	+17° 50'	8.2	F8	0.077	137°	− 26.0 ± 1.5	3	23
24 1036..	249499	54.2	+25 00	9.3	K4	.063	151	+ 5: var.	4	
25 1089..	40895	59.7	+25 53	8.0	F8	.082	150	− 21.4 ± 1.5	3	
25 1105..	41221	6 01.6	+25 11	8.6	F8	0.073	155	− 7.6 ± 1.8	3	

NOTES TO TABLE 1

* Other measures available.

H. Member of cluster.

P. Membership probable.

D. Membership doubtful.

1. +56, +59, +37, +70.

2. Two spectra.

3. +143, +99, +38, −76, −99.

4. +51, +60, +64.

5. −4, +9, +8.

6. +49, +46, +30, +51.

7. +51, +41, +30.

8. Two spectra.

9. Two spectra.

10. +43, +11, +28, +40.

11. +51, +23, +58, +38.

12. +49, +37, +52, +40.

13. +14, −28, +32.

14. Two spectra.

15. +65, +26, +38, +41, +38.

16. −13, +42, +50, +42.

17. +30, +28, +53, +48.

18. +46, +50, +58, +60.

19. Two spectra.

20. Two spectra.

21. Two spectra.

22. +44, +60, +66.

23. 0, +21, +4, −3.

TABLE 2

LARGE VELOCITIES IN THE HYADES REGION

Star	m_v	Sp.	ρ	Star	m_v	Sp.	ρ
BD+26°595.....	9.1	G8	−198.4	HD28424.....	7.8	G9	+ 96.8
HD23841.....	7.0	G8	− 79.8	BD+0°873.....	9.2	F8	+102.7
HD25532.....	8.3	F6	−113.0	HD32023.....	9.1	F8	+105.3
BD+31°769.....	8.8	F8	+ 79.4*	BD+31°846.....	8.9	G0	+ 74.9*

* Possible common motion. For BD+31°769, $\mu = 0.104$, $\theta = 133^\circ$; for BD+31°846, $\mu = 0.091$, $\theta = 146^\circ$.