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THE SPECTROSCOPIC ORBIT OF H.R. 6532 AND THE RADIAL VELOCITIES OF TEN STARS

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ABSTRACT

Thirty-three single-prism spectrograms of the star H.R. 6532 ($\alpha=17^{\rm h}\ 27^{\rm m}\cdot 6$, $\delta=+12^{\circ}\ 00'$) form the basis of the determination of its orbit. The period is 6.7984 days, the eccentricity .069 and semi-amplitude 50.18 km.

Ten stars of different types were selected from the Harvard Catalogue to fill in a program and the radial velocities of these are determined. From three to eight plates of each star were measured.

I—THE SPECTROSCOPIC ORBIT OF H.R. 6532

The star H.R. 6532 (1900 $\alpha = 17^h \ 27^m \cdot 6$, $\delta = +12^\circ \ 00'$, visual magnitude $6 \cdot 18$, type Ao) was found to be a spectroscopic binary by the first two plates secured in July 1921. Nineteen plates were secured in July and August from which a period of approximately $6 \cdot 7$ days was evident. Their distribution along the velocity curve was not such as to warrant a solution, without additional plates, and these have been obtained, fourteen in number, during the summer months of 1922. The thirty-three plates thus obtained form the basis of the solution.

In the latest Harvard Catalogue this star is classified as B9, but the line $\lambda 4026$ does not appear on any of the plates taken here. The principal lines used in measurement were $\lambda\lambda 4481$, 4549, 4233 and the silicon pair 4131 and 4128.

The following is the table of observations in which the phases are those given by the corrected period and periastron passage.

47958--1

OBSERVATIONS OF H.R. 6532

Plate Number	Date	Julian Date	Phase	Lines	Wt.	Velocity	0-C

6225	1921 July 12	2,422,883.745	5.591	6	.2	+20.9	+ 0.2
6246	" 15	2.886.790	1.836	2	.2	$-53 \cdot 3$	- 1.8
6290	" 22	2,893.798	2.047	6	.2	-48.7	- 3:5
6327	" 28	2,899.767	1.218	5	·1	-65.7	- 1.9
6345	" 30	2,901.762	$3 \cdot 213$	8	•2	- 1.3	- 2.3
6363	Aug. 1	2,903.800	$5 \cdot 251$	2	•1	+27.8	- 2.8
6369	" 2	2,904.785	$6 \cdot 236$	2	.2	- 7.6	+ 0.4
6379	" 4	2,906.744	1.396	2	.2	61 ⋅8	+ 0.
6381	" 6	2,908.751	3.403	8	.2	+9.4	- 1.4
6387	" 7	2,909.711	4.363	7	.2	+35.9	$+2\cdot$
6406	" 8	2,910.739	5.391	3	·1	$+22 \cdot 2$	- 4.0
6420	" 10	$2,912 \cdot 718$	0.572	5	•1	-63.8	- 5.
6421	" 10	$2.912 \cdot 734$	0.588	3	•1	-52.9	+5
6435	" 11	2,913.693	1.547	5	\cdot_2	-50.2	+ 9.
6474	" 25	$2,927 \cdot 692$	1.949	3	, •1	-50.5	- 2.
6475	" 25	$2,927 \cdot 709$	1.966	5	1	-61.0	-13.
6493	" 28	2,930.662	4:919	4	•1	+39.9	+ 4.
6494	" 28	2,930.683	4.939	4	1	+33.0	- 2.
6506	" 31	2,933.730	1.189	3	.1	-66.5	- 2.
7687	1922 June 17	3,223.841	5.767	6	.2	+15.2	+ 0.
7699	" 24	3,230.770	5.798	4	.2	+ 7.7	- 5
7713	" 25	3,231.768	6.795	$\overline{2}$	•15	-38.6	- 2.
7717	" 25	3,231.848	0.177	4	.15	-44.6	- 0
772 5	" 26	3,232.782	1.111	7	.2	-64.0	+ 0.
77 4 0	" 28	3,234.714	3.043	5	.2	- 1.2	+ 4.
7761	July 1	3,237.833	6.162	5	•2	$+$ $5 \cdot 2$	+ 9.
7783	" 4	3,240.764	2.295	2	.1	-27.9	+ 8.
7801	" 12	3,248.807	3.539	7	.2	+7.3	- 5.
7825	" 18	3,259.736	2.670	$\frac{1}{2}$.1	-18.1	+ 2.
7827	" 19	3,255.734	3.668	5	.2	+20.5	+ 3.
7840	" 23	$3,259\cdot790$	0.926	7	.2	-68.1	- 4.
78 5 5	" 29	3,265.743	0.080	3	.2	-43.7	- 3.
7926	Aug. 16	$2,423,283\cdot726$	4.465	5	.2	+32.6	- 2.

By the additional plates of 1922 the period was slightly modified and the value of 6.800 days satisfied all observations and was accepted as definite. Eleven pairs of the observations were grouped to form the normal places. Of these four pairs were taken on the same night, six pairs differed by one cycle each and one pair by five cycles. In the case of the pair differing by five cycles the observations occurred at minimum and it was quite evident that their grouping would not interfere with including the period in the least-squares solution.

NORMAL PLACES

	Mear	Phase	Mean	3374	0-C	
	Prel.	Final	Velocity	Wt.	Prel.	Final
1	0·293 0·358 0·876	0·078 0·136 0·580	-43·7 -41·6 -58·4	·2 ·3 ·3	$ \begin{array}{r} -3.5 \\ +1.4 \\ +0.8 \end{array} $	$ \begin{array}{r} - 3.5 \\ + 1.2 \\ - 0.1 \end{array} $
4	$1 \cdot 140$ $1 \cdot 332$ $1 \cdot 499$	0.924 1.110 1.200	$ \begin{array}{r} -68 \cdot 1 \\ -64 \cdot 0 \\ -66 \cdot 1 \end{array} $	$\begin{array}{c} \cdot 2 \\ \cdot 2 \\ \cdot 2 \end{array}$	$ \begin{array}{rrrr} & -5 \cdot 2 \\ & -0 \cdot 4 \\ & -3 \cdot 1 \end{array} $	$ \begin{array}{r} -4 \cdot 2 \\ +0 \cdot 4 \\ -2 \cdot 0 \end{array} $
7	1.769 $2 \cdot 244$ $2 \cdot 251$	1.471 1.942 1.958	-56.0 -51.0 -55.8	$\begin{array}{c} \cdot 4 \\ \cdot 4 \\ \cdot 2 \end{array}$	$\begin{array}{ c c c c } + & 3 \cdot 2 \\ - & 4 \cdot 4 \\ - & 9 \cdot 4 \end{array}$	$ \begin{array}{r} + 4.8 \\ - 2.2 \\ - 7.6 \end{array} $
0	2.514 2.886 3.264 3.606	$2 \cdot 293$ $2 \cdot 668$ $3 \cdot 042$ $3 \cdot 307$	$ \begin{array}{c c} -27 \cdot 9 \\ -18 \cdot 1 \\ -1 \cdot 2 \\ +4 \cdot 1 \end{array} $	·1 ·1 ·2	$ \begin{array}{c c} + 9.1 \\ + 4.0 \\ + 5.0 \\ - 3.4 \end{array} $	$ \begin{array}{r} + 8.3 \\ + 3.1 \\ + 4.7 \\ - 0.4 \end{array} $
3	3·821 4·661 4·676	3·602 4·363 4·465	$\begin{array}{c c} + & 4 \cdot 1 \\ + & 13 \cdot 9 \\ + & 35 \cdot 9 \\ + & 32 \cdot 6 \end{array}$	$egin{array}{c} \cdot 4 \\ \cdot 4 \\ \cdot 2 \\ \cdot 2 \end{array}$	$ \begin{array}{c c} -3.4 \\ -1.4 \\ +0.9 \\ -3.4 \end{array} $	$ \begin{array}{c c} -0.4 \\ -1.0 \\ +2.6 \\ -1.9 \end{array} $
78	5\223 5\620 5\895	4.930 5.321 5.591	$+36.5 \\ +25.0 \\ +19.4$	$\begin{array}{c} \cdot 2 \\ \cdot 2 \\ \cdot 2 \\ \cdot 2 \end{array}$	$ \begin{array}{c c} -3.4 \\ +2.0 \\ -1.2 \\ +2.6 \end{array} $	$\begin{array}{c c} -1.9 \\ +1.8 \\ -4.1 \\ -1.9 \end{array}$
0	6·056 6·383 6·535	5.830 6.161 6.236	$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$	·4 ·2 ·1	$\begin{array}{c c} + 2.0 \\ + 1.3 \\ +10.6 \\ + 5.5 \end{array}$	-0.6 $+8.9$ -0.2

The following preliminary elements formed the basis of the least-squares solution.

P = 6.800 days

e = 0.05

 $\omega = 105^{\circ}$

T = J.D. 2,422,877.850

K = 50 km.

 $\gamma = -13 \text{ km}.$

OBSERVATION EQUATIONS

					*		
1	1·000x	- ·545y	$-\cdot 771z$	+ ⋅895u	+ 361v	+ ⋅930w	+ 3.500 = 0
2	1.000	600	- ⋅853	+ .858	$+ \cdot 314$	+ .888	-1.400
3	1.000	925	907	$+ \cdot 459$	$+ \cdot 015$	$+ \cdot 438$	— 0⋅800
4	1.000	- ·999	•564	$+ \cdot 213$	$+ \cdot 066$	$+\cdot 172$	+ 5.200
5	1.000	-1.013	$- \cdot 259$	+ .048	+ .000	+ .000	+~0.400
6	. 1 • 000	-1.001	$+ \cdot 055$	→ ·109	- ⋅003	- ·159	+ 3.100
7	1.000	924	+ .553	- ⋅364	- ·012	- ·40 6	-3.200
8	1.000	− ·673	$+ \cdot 977$	- ·703	006	710	+ 4.400
9	1.000	669	+ .978	− ·707	- ⋅036	- ⋅713	+ 9.400
10	1.000	- ·481	$+ \cdot 925$	- ⋅83 6	- ·297	- ⋅819	-9.100
11	1.000	- ·182	$+ \cdot 549$	- ·938	- ⋅339	899	4.000
12	1.000	$+ \cdot 135$	040	- ·941	- ⋅318	- ⋅892	-5.000
13	1.000	- - ·4 08	- ·564	- ·859	- ⋅020	- ·8 2 0	+ 3.400
14	1.000	+ .564	- 812	769	— ·275	$-\cdot 742$	+ 1.400
15	1.000	$+ \cdot 959$	662	− ·183	- ⋅007	- 220	- 0.900
16	1.000	+ .968	685	— ·170	085	- ·208	+ 2.600
17	1.000	$+ \cdot 949$	+ .287	+ .320	$+ \cdot 014$	$+ \cdot 273$	- 2.000

OBSERVATION EQUATIONS-Concluded

1 8	1.000	$+\cdot 782$	+ ⋅876	$+ \cdot 654$	$+ \cdot 016$	+ .630	+ 1.200
19	1.000	$+ \cdot 594$	+1.022	+ .843	$+ \cdot 005$	+ .845	-2.600
20	1.000	+ .460	+ .973	+ .929	$+ \cdot 327$	+ .946	-1.300
21	1.000	$+ \cdot 150$	+ .576	+1.035	+ .389	+1.079	-10.600
22	1.000	- ⋅004	+ .290	+1.048	+ .030	+1.099	- 5.500
where	x	$=\delta\gamma$					
	v	$= \delta K$					
	v						
	\mathbf{z}	$= \mathbf{K} \cdot \delta e$					

$$u = -K.\delta\omega$$

$$v = -\frac{1000 \text{ K}}{(1-e^2)^{\frac{3}{2}}} \cdot \delta\mu$$

$$w = \frac{K\mu}{(1-e^2)^{\frac{3}{2}}} \cdot \delta T$$

NORMAL EQUATIONS

The solution of the normal equations resulted in the corrections

$$\delta\gamma = +0.21 \text{ km./sec.}$$

$$\delta K = +0.18 \text{ km.}$$

$$\delta e = +0.019$$

$$\delta\omega = +11^{\circ} 26$$

$$\delta P = -0.00157 \text{ days}$$

$$\delta T = +0.304 \text{ days}$$

and the corrected elements with their probable errors are

The probable error of a normal place of weight unity is +1.40 km. sec. and the probable error of a plate ± 2.94 km. sec. The least-squares solution reduced Σpv^2 from 90.51 to 68.85.

II—RADIAL VELOCITIES OF TEN STARS

The stars for which observations are here given were placed on the observing programme of the writer for the past two summers as a matter of filling in. Eight stars were so taken on the first summer and four were added this year. Two of the first eight turned out to be binaries. The orbit of one binary, H.R. 5992, was completed last year and appeared as Vol. I, No. 28 of the Publications of this observatory. The orbit of the other binary, H.R. 6532, is given in the first part of this number. As the observatory does not contemplate an early publication of another list of radial velocities it has been thought advisable to include in this number the observations of the other ten stars.

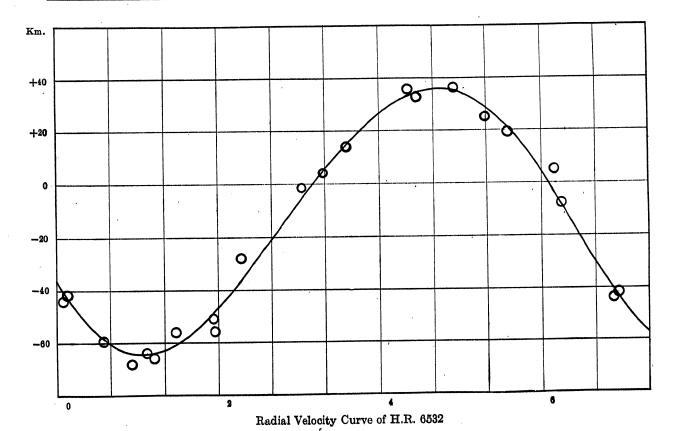
Star H.R.	Type Mag.	Date G.M	4.T.	Rad. Vel.	Regions Lines	Quality	Remarks
6057	G5	1922 June 2		-20.1	13 = 23	Good	A good spectrum.
16h 09·3m	6.05	July " 1	1·755 12·770	$-19.0 \\ -23.1$	13 = 23 $13 = 23$	"	
+06° 09′	0.00		6.700	-23.6	13 = 23	"	
,				-21·5 ±·8	13 = 23		
6176	Aop	1921 June 1		+ 2.7	5	Weak	Spectrum rather
			15.763	+2.0	8	Fair	diffuse.
16 ^h 21·1 ^m	6.29		$22 \cdot 772$	- 7·3	8	Good "	
+15° 42′	1		28.740	- 4.6	6	"	
		Aug.	1.734	- 3.9	8		
	ł	ł	8.707	+3.4	6	Fair Good	
	1		10·697 25·708	$\begin{array}{c c} + 3.2 \\ - 6.5 \end{array}$	6 5	Good	
	1	1	20.108	- 6.5	ð		
		1922 June 2	25.741	+ 1.6	5	"	
				-1·0 ±1·0			
6467	F2	1922 June 2	25 · 786	+31.4	1 = 23	Good	A good spectrum.
			28.799	+26.8	1 = 23	"	
17 ^h 17 ^m ·8	6.32		1.783	+31.1	1 = 23	"	
+48 17'		"	23 · 765	+30.4	1 = 23	"	
				+29·9 ±0·7			
6803	F2	1921 July	12.774	- 9.4	4	Good	A fair spectrum, but too diffuse
			15.820	$-17 \cdot 2$	4	Weak	for measurement on the com
18h 05m·7	6.14		$22 \cdot 824$	- 7.3	6	Good	parator. The extreme range
+16° 27′		1	28 · 815	-18.6	5 .	"	seems large, but even if vari
	ŀ		8.765	-15.5	7	"	able the mean velocity will b
			10.753	-15.6	9		very approximately the velo
	1		25.742	-12.5	6	"	city of the system.
			28.710	-15.7	8	i.	
		1922 June		- 4.3	11	"	
			29·786 24·811	-22.2	7 5	Fair Weak	
				-14.7			

II. RADIAL VELOCITIES OF TEN STARS—Continued

Star H.R.	Type Mag.	Date G.M.T.	Rad. Vel.	Regions Lines	Quality	Remarks
					`	
6820	Ko	1922 June 25 874	-65.7	11=23	Good	A good spectrum.
10h 00m 0	0.05	" 28·829	-70.9	15 = 23 $13 = 23$	"	
18 ^h 09 ^m ·0 +21° 51′	6.25	July 23.823 Aug. 26.767	$ \begin{array}{c c} -65 \cdot 2 \\ -61 \cdot 3 \\ -66 \cdot 0 \pm 1 \cdot 3 \end{array} $	15 = 25 17 = 23	Weak	
6971	В3	1921 July 3·894	-11.1	3	Fair	Most of the plates were
,		" 15·874 " 22.857	-14.7	6	"	rather under-exposed, and
18h 29m·6 +30° 49′	6.43	" 22·857 " 28·841	$-6.5 \\ -11.7$	6	"	showed broad hydrogen and helium lines only. The K line
+90 49		Aug. 8.828	+13.6	5	"	of calcium was present on a
		" 10.797	+ 7.0	5	"	couple of plates and on one
		" 25·795 " 28.701	- 0.2	4	"	showed a suggestion of dupli-
		" 28·791	$ \begin{array}{c c} -5.3 \\ -3.6 \pm 2.3 \end{array} $	4		city.
7008	Gop	1921 July 22·899	-19.3	7	Fair	A fairly good spectrum but lines
•	_	Aug. 8.798	-23.7	14	"	too diffuse for measurement
18h 34m·7	6.30	" 10.775 " 25.761	-14.0	14	"	on the comparator.
+05° 10′		" 25·761 " 28·748	$-21.0 \\ -17.9$	5 4	"	
		1922 June 28.857	-21.0	5	Weak	
,		July 1.807	-19.9 -19.5 ± 0.8	9	Good	
7396	A	1922 July 1.890	+21.8	4	Fair	This star is listed as A in H.A.
	,	" 12.927	+ 1.8	5	"	50 but is really B-type.
19 ^h 23 ^m ⋅0	6.26	" 23.903	- 7.0	4	"	
+14° 05′		" 29.855 Sept. 1.760	+6.5	3 4	"	
		Sept. 1.700	$+3.5 \\ +5.3 \pm 3.2$	4	'	
			10020			
7482	F	1921 July 28 912	$+12 \cdot 4$	9	Fair	This star, listed as an F-type in
		Aug. 1.852	+12.9	5	Weak	H.A., is really of type B.
19h 36m·1	6.44		+ 3.6	5	17 to	The spectrum is very good.
+20° 15′		" 10·818 " 25·827	$\begin{array}{c c} + 7 \cdot 1 \\ - 1 \cdot 7 \end{array}$	5 10	Fair	
•		" 28·863	+ 1.6	2	Weak	
		1922 June 25·899	+ 3.8	3	Fair	
		July 12.906	+13.0	11	Good	
		" 23.874		8	" Ta.:	
•		Aug. 6.856	$+4.9 \\ +5.9 \pm 1.1$	5	Fair	

II. RADIAL VELOCITIES OF TEN STARS—Concluded

Star H.R.	Type Mag.	Date G.M.T.	Rad. Vel.	Regions Lines	Quality	Remarks
7839 20 ^h 26 ^m ·5 +20° 16′	A2 6·00	1921 July 22·924 " 28·934 Aug. 1·914 " 8·884 " 10·839 " 25·850	-36·1 -40·9 -43·8 -38·2 -39·8 -44·3	16 15 18 18 16 15	Good " " " "	A good spectrum.
		1922 July 1 966 " 12 952 " 23 927	$-40 \cdot 8 \\ -45 \cdot 2 \\ -42 \cdot 1 \\ -41 \cdot 2 \pm 0 \cdot 6$	18 18 23	66 66	



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